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UNITED STATES OF AMERICA  
NUCLEAR REGULATORY COMMISSION

CRBR SUBCOMMITTEE and  
STRUCTURES AND MATERIALS WORKING GROUP

Room 1046  
1717 H Street, N.W.  
Washington, D.C.

Thursday, ~~May~~<sup>March</sup> 17, 1983

The Subcommittee on CRBR and Structures and Materials Working Group met, pursuant to notice and recess, at 8:30 a.m., Dr. Max W. Carbon, Chairman of the Subcommittee on CRBR, presiding.

ACRS MEMBERS PRESENT:

- M. Carbon, Chairman
- J. Mark, Member

ACRS CONSULTANTS PRESENT:

- Mr. Abdel-Khalik
- W. Kastenber
- W. Lipinski
- Z. Zudans

DESIGNATED FEDERAL EMPLOYEE:

- P. Boehnert

NRC STAFF PRESENT:

- R. Stark
- W. Morris 8303210088 830317
- C. Bell PDR ACRS
- C. Allen T-1190 PDR

TR 4  
delete B. White

P R O C E E D I N G S

(8:35 a.m.)

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2  
3 MR. CARBON: The meeting will now to order. This is  
4 a continuation of the Advisory Committee on Reactor Safeguards  
5 Subcommittee on CRBR and Structures and Materials Working Group.  
6 My name is Carbon, the subcommittee chairman. The purpose of  
7 this second-day meeting will be devoted to the discussion of  
8 the HCDA issue for CRBR. We will proceed with the meeting,  
9 and I'll call upon Mr. Curtis Allen of the NRC Staff.

10 MR. ALLEN: Good morning. As Dr. Carbon indicated,  
11 this subject today is the HCDA energetics. The staff's presen-  
12 tation will be given essentially by Dr. Theofanous and Dr. Bell.  
13 They'll discuss the results of their work on developing -- that  
14 they've done for the staff in developing an assessment of the  
15 energetics in CDAs and the CRBRP.

16 As you can see, the report is extensive, it's  
17 sitting on the table in front of you, and it will be a long  
18 presentation, and we urge patience on your part in hearing the  
19 presentation. It's a complex story and they have a lot of  
20 things to say.

21 I have a few introductory comments I'd like to make  
22 before I summarize the status of the review and introduce them.  
23 These are largely a few comments about the staff's approach to  
24 CDA evaluations in general, and the role of CDA energetics in  
25 evaluating CDAs in particular.

1           As we discussed yesterday, the staff is also of the  
2 opinion that CDAs should be classed as beyond the design basis  
3 events. However, we do feel, as was also mentioned yesterday,  
4 that because they are considered to be potentially significant  
5 risk contributors, and until we understand that potential better,  
6 we feel it's prudent to be able to accommodate such events.  
7 And that's why we evaluate CDAs. We evaluate them so we can  
8 determine reasonable accommodation requirements, and the  
9 capabilities of the system to accommodate those events. Another  
10 reason is to develop information for use in making judgments  
11 about the risk of CDAs, and that's essentially the staff's  
12 attitude in that regard.

13           Before I turn to the role of energetics in CDA  
14 evaluation in particular, we had a discussion yesterday about  
15 important differences between the TMBDB and SMBDB scenarios,  
16 and I thought I'd just try to illustrate that point a little  
17 bit this morning. To do that, I've taken a figure from the  
18 applicant's CRBRP-3 TMBDB scenario. It's a sketch that illus-  
19 trates the reactor cavity domain and the reactor containment  
20 building environment domain. The reactor vessel and the core  
21 sits down inside the reactor cavity, and these are concrete  
22 walls, they're steel-lined, et cetera. This is the operating  
23 floor (indicating), this is the head axis area and here is the  
24 reactor vessel head.

25           The reactor containment building environment is

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isolated from the below operating floor structures; it is sealed off from the containment primary heat transport system.

Given a CDA, when a CDA starts developing in the core, as long as the head remains intact, the CDA is constrained within the cavity region, and if the CDA sequence is non-energetic or is of a small energetics that it doesn't challenge the integrity of the head, the progression of the CDA remains inside the cavity region. The core debris winds up penetrating the vessel, and it flushes the sodium and the debris down under the cavity, and then it progressing goes along the TMEDB long-term scenario that you heard about last week.

If the energetics are large enough to fail the head, then that opens the direct communication between the disrupted or disrupting core and the reactor containment building environment early in the transient. And you heard yesterday these things develop in the order of 15 to 20 seconds; the challenge is developed in that range, to the head. So that should that happen, if a sodium spray fire results, you might over-pressurize and challenge the integrity of the containment.

If the energetics were very large -- and this is very unlikely, -- missiles associated with a head failure like that could also present a challenge.

This is the -- the potential for head failure, therefore, is the reason for focusing on the capability -- the energy absorption capability of the head. That's the only point I

1 wanted to make about that. And that's all this slide says.  
2 That this is -- really, the role of energetics in CDAs is to  
3 determine if CDAs can lead to an early containment challenge  
4 or early containment failure.

5 And the reactor head failure provides a threshold  
6 test to examine that question. And we heard yesterday and we  
7 believe that the head can accomodate the impact of a sodium  
8 slug having kinetic energy of 75 megajoules. I emphasize it  
9 can because that's contingent on the applicant resolution of  
10 the design -- the proposed design capability deficiency that  
11 we heard discussed yesterday.

12 Therefore, that provides a good way to test the  
13 energetic potential of CDAs; against the energy absorption  
14 capability of the reactor vessel head.

15 Finally, turning to a summary of where we are in  
16 this review, the applicant's analyses were given in CRBRP's 3  
17 Volume 1 and GPR 523. We had a number of meetings with the  
18 applicant. They culminated in questions which were submitted  
19 to the applicant and they provided answers to us. At that  
20 time, we initiated a special task to develop an independent  
21 assessment of CDA energetics. This was directed by Dr. Theo-  
22 fanous and Dr. Bell. You're going to hear a lot more about  
23 that shortly.

24 They gave a preliminary progress report to the  
25

1 subcommittee back in November of 1982; the final report issued  
2 March 11, 1983, and it provides the technical basis of that  
3 report; it provides the technical basis for the staff's conclu-  
4 sions on the CDA energetics. One of the conclusions is that we  
5 recommend that the plenum fission gas compaction potential be  
6 eliminated by design, and the applicants agreed to address this  
7 concern. You'll hear a good deal about that in the course of  
8 the presentation. And assuming the elimination of that concern,  
9 we believe that the proposed structural design capability,  
10 the 661 megajoules and 75 megajoules slug kinetic impact  
11 energy is adequate, and given this capability we believe that  
12 the vessel head failure through the CDA energetics is physically  
13 unreasonable.

14 I know there was a lot of confusion about the 661  
15 75 number, yesterday, a lot of discussion about that, and I  
16 think we have the right people here today to discuss that.  
17 Theo and Charlie have done an awful lot of work in this area,  
18 and I think at the end of the day everybody in the room should  
19 have a very clear picture of how these numbers are generated  
20 and how they're applied.

21 And I think with that, I would turn the meeting over  
22 to Theo.

End tp.  
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1 MR. THEOFANOUS: First of all, my apologies for not  
2 being able to give you this document 1 or 2 days in advance  
3 so you would have time to read it before this presentation.  
4 In view of this fact, this morning I would like to give you a  
5 kind of an overview, walk you through this document in what is  
6 an overview way and, hopefully, this will help you as you try  
7 to go through it. It will help you to find things where they  
8 are and so on, and especially try to identify for yourselves  
9 the ideas that you are more interested in to look at.

10 The document appears to be very lengthy one, and it  
11 is for this reason that we separated the figures from the text.  
12 And the text is only 250 pages, and the rest is figures. And  
13 by the time the final figures are drawn and incorporated in  
14 the text, the actual size of it will be quite a bit reduced,  
15 the visual size will be reduced.

16 As you realize, we are concerned about the initial  
17 impact of people being afraid even to look at it, by it being  
18 so big. But we feel that we did not go to any unnecessarily  
19 lengthy discussions there. In fact, in some areas maybe you  
20 might find it somewhat skimpy.

21 The point is that the whole area to be assessed  
22 correctly is rather broad, has a lot of facets to it. It is  
23 very complicated. And in order for one to be able to convey  
24 exactly what one has done, it cannot be done in just a few words.  
25 It is rather complicated.

1           The structure of the report, I would say a couple of  
2 words as far as how the report can be read or how we would  
3 recommend that you read it. By looking at the table of contents  
4 you will find out that we have broken it down, the whole thing,  
5 into just a few important steps. And the report is organized  
6 so that each one of the steps is more or less like a unit, so  
7 that you, by looking at the chapter section, you can go back  
8 just going like this through the figures and the text, side by  
9 side. You can identify the section that you are interested in.  
10 And then within that, the whole thing is a complete unit.  
11 Hopefully, that will help you to read it. And so even the  
12 page numbers were according to units.

13           So to start with then, this is our goal today is to  
14 give you a summary of this independent assessment which we have  
15 just completed and is put into this NUREG document of 5224.

16           And to start off, we want to take a look at the  
17 different kinds of core-disruptive accidents. As you realize,  
18 there are many, many different ways by which one can enter into  
19 a coremelt situation. And in an attempt to assess energetics,  
20 one needs to somehow abstract the complexity and be able to  
21 come up with some generic way of looking at things. And one  
22 major classification among the different core-disruptive  
23 accidents is between protected and unprotected ones. This is  
24 important because if the reactor is scrammed, the heating rates,  
25 of course, are very low. And one needs to sort of different



1 kinds of sorts of assumptions for getting this core to melt.  
2 And normally, this involves waiting for a very, very long time  
3 without cooling.

4           The heatup rates are on the order of a degree per  
5 second. On the other hand, in the unprotected situations, the  
6 power continues to be normal and the heating rates are higher,  
7 and suddenly this implies a set of phenomena that is very  
8 different between those two cases.

9           Now, furthermore, the unprotected accidents can be  
10 further classified into two major categories, depending as to  
11 whether the sodium is in the core when the disruption takes  
12 place versus when the sodium is outside of the core.

13           And this is -- now, don't laugh at what I am telling  
14 you here -- but we have the genetic CDAs associated with lesser  
15 flow accident, while the other ones involve the sodium in,  
16 coming under the name of transient overpower.

17           Basically, here we are losing pumping capability while  
18 the core continues to produce power at normal levels. That  
19 leads to core disruption that is preceded by sodium voiding.  
20 Now, it so happens that in some of those reactors, the  
21 reactivity effects associated with the loss of sodium are such  
22 that the loss of flow eventually winds up as a transient  
23 overpower. But at least the beginning phenomenology is  
24 specific to sodium out of the core as the core begins to  
25 disrupt.

1           On the other extent here, the pumps continue to pump  
2 normal flow through the core while the power for some reason  
3 increases. And if that power increases rapidly enough, it  
4 leads to fuel melting while the sodium is still running by the  
5 pins, and therefore we begin to have a disruption with the  
6 sodium in.

7           In this case here, referred to as the loss of heat  
8 sink accident, the power is decay heat levels, and typically  
9 about 1 percent by the time that one is concerned about core  
10 disruption. And in fact, under these conditions, as you will  
11 see later, the sodium must have gone out of the core; otherwise,  
12 the core would not have disrupted. And what is more, because  
13 of the very low heating rates, even the steel has had time to  
14 melt and get out of the core. So that is still our situation.  
15 It still is out.

16           Yes?

17           MR. MARK: Theo, that 1 degree Kelvin per second  
18 of the LOHS is applicable approximately what time after --

19           MR. THEOFANOUS: This is several hours, like 10  
20 after.

21           MR. MARK: Well, then for several hours it goes  
22 through a factor of 10 in the decay heat.

23           MR. THEOFANOUS: Yes.

24           MR. MARK: So if --

25           MR. THEOFANOUS: This number here is applicable to

1 something like 10 hours, to answer your question.

2 MR. MARK: Okay. So it's much higher at time zero?

3 MR. THEOFANOUS: Well, yes, but it will drop so  
4 quickly that within just a few seconds, it is to a few degrees,  
5 maybe not out, but a few degrees.

6 MR. MARK: Yes. Okay.

7 MR. THEOFANOUS: This is just to show the order of  
8 magnitude, it is not really to look at the detail of the  
9 numbers.

10 Now, in addition, we have, of course, other  
11 possibilities. And one that is quite prominent is the  
12 very, very severe earthquakes, seismic, severe seismic we also  
13 set out. Here they have combinations of this as caused by the  
14 seismic event, and we might also have other situations such as,  
15 for example, fracture of the core support. This is an accident  
16 postulated and studied to quite a great extent by the British,  
17 in particular, for the last few years. We think that the  
18 failure of the core support is in the category of very, very  
19 low probability events, that does not deserve very detailed  
20 evaluation.

21 Fuel failure propagation also is a very -- has been  
22 a very favorite kind of scenario. From the very early days  
23 people have studied that for quite a long time. And the  
24 general conclusion there is that this is not really a problem  
25 from the point of view of achieving core disruption from fuel

1 failure propagation.

2 As a possibility of transient overpower followed by  
3 a loss-of-flow accident, this would be coming about because when  
4 one attempts to scram the reactor, the scram signal is sent,  
5 the trip signal is sent both to the pumps and to the scram  
6 system. And one might conceive of a situation where the scram  
7 system fails to work but the trip to the pump works. We examined  
8 that, and we think that it is really in the noise level of  
9 probabilities. It is such a low probability that it does not  
10 need to be considered.

11 The point here is that if one is left completely free  
12 to think of everything that comes to mind, one can always, I  
13 guess, construct situations that may be more severe than the  
14 ones that we are going to talk to you about today. We have  
15 looked at the whole spectrum of things, and we have tried to  
16 discriminate between things that we consider to be worthwhile  
17 from the point of view of maybe in the way, way out low  
18 probability range, contribute something to the risk, versus  
19 other events that maybe are in the hypothetical sense can cause  
20 higher consequences, but if one looks at the probabilities, one  
21 comes to the conclusion that this is so unlikely that really  
22 fall in the category of events of the earth opening up or  
23 things of this type.

24 MR. LIPINSKI: Will you consider, on the last item  
25 there, the DOS, what do you consider low probability? What

1 is the number? What number do you have in mind?

2 MR. THEOFANOUS: This one was examined from the  
3 point of view of the instrumentation in the reactor, and the  
4 people that looked at that at SAI came to the conclusion that  
5 it is several orders of magnitude probability than TOP. The  
6 number was not determined absolutely for this event, but it  
7 was determined in conjunction with how much lower probability  
8 versus having a straight TOP. In other words, you try to find  
9 out how many additional failures have to take place in order  
10 to cause this event.

11 MR. LIPINSKI: Okay.

12 MR. THEOFANOUS: You find out that you need all kinds  
13 of common-mode failures between electrical and mechanical  
14 systems, and it is this nature of things that make that very  
15 low.

16 MR. LIPINSKI: That's right. And to answer that  
17 question again, what is the probability of the TOP?

18 MR. THEOFANOUS: Well, I think that for this question  
19 you might get different answers depending upon whom you ask,  
20 and it is not part of our charter to look at the probabilities  
21 of initiators.

22 MR. LIPINSKI: Then how can you answer my question,  
23 because the TOP event, the last one, is the probability of  
24 failure to scram.

25 MR. THEOFANOUS: Yes.

1 MR. LIPINSKI: And given the fact that we have a  
2 number and a probability of failure to scram, then that sets  
3 this extremely low probability numerically.

4 MR. THEOFANOUS: Yes.

5 MR. LIPINSKI: And now if you are saying that that is  
6 several orders of magnitude lower --

7 MR. THEOFANOUS: Yes.

8 MR. LIPINSKI: -- then I can spot your TOP event.

9 MR. THEOFANOUS: You can what?

10 MR. LIPINSKI: The transient overpower event  
11 probability itself without the failure to scram.

12 MR. THEOFANOUS: Well, typically, this unprotected  
13 events have probabilities of the order of 10-5. That is -- and  
14 somebody might say 10-6. There are a few studies that have been  
15 prepared by Sandia and the other by SAI, and they are in the  
16 category of 10-5, 10-4, 10-6. It is, maybe all told, in  
17 magnitudes around 10-5.

18 This level of probability coupled with the potential  
19 consequences, at least as these people see those possibilities  
20 of CDAs, to assess that, a judgment was made many years ago  
21 and people have followed that through, that they have to  
22 believe that.

23 Now, if you couple on top of that additional events  
24 that make these events even lower probability and bring them  
25 down to 10-6, 10-7, 10-8, that is where you begin to lose a lot

1 of -- and that is how we discriminate between the things that  
2 we do want to consider and things that we don't want to  
3 consider.

4 Now, I don't want to get stuck in the details or  
5 of actually is it the TOP, is it 10-5 or 10-6, because even  
6 the people actually working on that have disagreements.  
7 What I want to say is that we consider the TOP is an event that  
8 has to be looked at. However, by considering this combination,  
9 it is still so much lower probability than this first kind  
10 of an event that we decided not to look at it.

11 MR. LIPINSKI: That depends on what kinds of numbers  
12 you have been given for that TOP LOF event and as to whether  
13 you believe them.

14 MR. THEOFANOUS: Well --

15 MR. LIPINSKI: Because that's based on the analysis  
16 of the hardware and the system as it stands.

17 MR. THEOFANOUS: Right. Well, I think that it is  
18 referenced in the report; in fact, the analysis that was done  
19 with respect to this and the considerations, all the detailed  
20 arguments with respect to the CRBR system in particular.

21 Now, based on those considerations, this combination  
22 is, in our opinion, much lower probability, several orders of  
23 magnitude lower than the straight TOP or the straight LOF.  
24 All right. So less than that. This is the basis for excluding  
25 that, independent of what the TOP LOF probabilities are. But

1 nevertheless, we know that these straight events are already  
2 very low probability. So I think your argument would be very  
3 appropriate if we had -- if those were probable events or  
4 not very probable, maybe in the order of  $10^{-2}$ ,  $10^{-3}$ , then  
5 maybe by losing another one or two orders of magnitude, we are  
6 not totally in the incredible domain.

7 But by those being already very extremely improbable,  
8 coupled on the top of that, another two or three orders of  
9 magnitude less in probability, that's what makes it kind of go  
10 over the hill. So that's the logic of it.

11 Another category that falls in the same situation  
12 is in the TOP itself. We are going to come to that later on  
13 in the discussion. You -- the question is what kind of TOPs  
14 do you want to consider. Obviously, you know that if you are  
15 going to pull out a bank of control rods, it's going to be much  
16 more probable than pulling out all the controls at the same  
17 time, all the rods.

18 So some discrimination has to be made because the  
19 difference of probability between all those different events  
20 are significant -- are not insignificant.

21 So we looked into that aspect of it, and in fact  
22 we determined almost like in a crossroad of a reasonable thing  
23 to look at it, to look at versus again events that they are of  
24 such a low probability that the ramp rates are really -- can be  
25 considered hypothetical.



1 MR. LIPINSKI: Let me make one comment to the  
2 chairman. Based on this TOP LOF event, the numbers are based  
3 on the ATWS probabilities, because this is a two-part system  
4 and basically, we have ATWS with this TOP event failure  
5 to respond, we also have the pump tripping for the LOF. And  
6 assuming that one part of the circ acts and the other does not,  
7 the probability of the failure to scram is now directly related  
8 to the ATWS issue. And early arguments, we had a set of  
9 data, and the industry was saying we have rectification.  
10 Now with the latest information based on parts that are being  
11 examined and that cannot scram, we have effectively negative  
12 rectification, saying we haven't learned a thing in terms of  
13 maintaining this hardware, and so far the discussion has not  
14 been held as to what the new number is for the data base.

15 And to assume that we have a nice number on this  
16 plant based on the analysis that has been done, everything is  
17 contingent upon proper maintenance. And the industry's  
18 record right now is poor in that respect.

19 So I would very much question what somebody tells  
20 you for the probability of failure to scram.

21 MR. THEOFANOUS: Well, would you encourage us then  
22 to take this kind of an event within the spectrum of events  
23 that it doesn't need to be considered as far as CDA in this?

24 MR. LIPINSKI: I don't think it's any jore or less  
25 probable than the TOP itself. In this particular event, it's

1 mechanically possible based on the way the plant is built.

2 MR. THEOFANOUS: Obviously, I don't know why you  
3 say that, because you can identify additional systems it's  
4 safe to fail for this event for this combination to happen  
5 versus the initial event on the straight event. And I think  
6 it's intuitive on this event, although I am not asking you to  
7 say that, but if you required more systems to fail requiring  
8 extreme common-mode failures, there has to be a lower  
9 probability. Maybe not as much as you are saying, but it has  
10 to be lower probability. I can't see how it can be higher  
11 probability or even the same.

12 MR. LIPINSKI: Okay. But to say it's extremely low  
13 compared to the others you are throwing out?

14 MR. THEOFANOUS: I think that -- well, I was not  
15 planning to get into this discussion here today, so I really  
16 cannot give you much more. But I think what I can do is I  
17 can get you the document on which you are really basing this  
18 judgment.

19 MR. LIPINSKI: All right.

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j-1-3,1 1 MR. GROSS: If I could add a word, I'm not sure  
2 that Mr. Lipinski was at the full committee meeting last week  
3 where we discussed the scram breakers. I think part of his  
4 concern comes from the Salem event.

5 MR. LIPINSKI: That's right.

6 MR. GROSS: The ATWS event.

7 MR. CARBON: Can you speak a little louder?

8 MR. GROSS: Last week at the full committee meeting  
9 we did present to the full committee the significant differ-  
10 ences in our design versus the typical scram breakers such as  
11 those that are at the Salem plant. Our scram breakers are  
12 entirely different. They are much smaller. They are fully  
13 enclosed and sealed. They are not anywhere near as sensitive  
14 to maintenance in order to continue.

15 Only one of the two systems, and we have two, of  
16 course, independent systems, and only one of those systems is  
17 dependent upon scram breakers. So I recognize your concern,  
18 but I believe the differences in our design are significant  
19 and need to be taken into consideration.

20 MR. ZUDANS: But this meeting has nothing to do with  
21 Walt's correction. You are talking about TOP and he is talking  
22 about the combination of two.

23 MR. GROSS: I understand, but I believe part of his  
24 concern is that he doesn't believe that the TOP, that the ATWS  
25 is as low a probability as perhaps some people felt it was

j-1-3,21

previously.

2 MR. ZUDANS: Not really. The way I heard the  
3 question, he wants to know simply what is the probability of  
4 a combination of these events. That is all.

5 MR. LIPINSKI: The point is you have got some  
6 theoretical numbers you are calculating based on the design  
7 you see on paper in front of you. But given the fact that  
8 that design is completely overridden by maintenance procedures  
9 and we end up with a new set of numbers, how can I believe the  
10 set of numbers you are presenting from a nice, clean analysis  
11 assuming perfect maintenance?

12 MR. MARK: There is a difference between the system  
13 here and the light-water reactors, which have only one set of  
14 rods. This has two redundant rods. So the ATWS number as it  
15 might be modified for light-water reactors has nothing to do  
16 with this.

17 MR. LIPINSKI: No, but their number as might be  
18 modified due to poor maintenance is the number of interest.

19 MR. MARK: One has to look at the system and ask if  
20 it is subject to the same diseases. If it's a sealed unit,  
21 then maintenance may in fact be irrelevant. Of the switch  
22 bars, anyway.

23 MR. LIPINSKI: Not looking at the details of the  
24 hardware, all I am doing is making a general statement --  
25 maybe it was the specifics -- that the vulnerability is not

j-1-3,3 1 there.

2 MR. MARK: Look, you are absolutely right, it has  
3 to be relooked at with the same question as Salem, the light-  
4 water things. If they apply, then they have to be accepted,  
5 and if they don't apply, then it needs to be pointed out.

6 MR. THEOFANOUS: To continue, then, in the follow-  
7 ing we are going to take a look at the loss of flow, TOP and  
8 loss of heat sink, in that order. Those of you who were here  
9 last time in November, you remember that we put most of our  
10 emphasis in the loss of flow accident.

11 We did that because, number one, we could identify  
12 areas that we were concerned with as far as energetics, and  
13 really this case not being so for the other two situations.

14 Having now completed the analysis, we think that  
15 this is a correct judgment, and the presentation and the  
16 report is really then with this emphasis on the loss of flow.  
17 This is not to say, however, that the unique aspects of the  
18 other two initiators were not considered, and in fact we  
19 think there are some interesting aspects associated with  
20 them and we are going to talk about them later today.

21 But the emphasis is on the loss of flow, and the  
22 overall approach here is to first consider the structural  
23 capability of the vessel. Obviously, you need to have a  
24 yardstick against which to measure energetic behavior, and  
25 that yardstick is provided by the structural capability of the

1 primary system.

2 As it turns out in this case, in addition to the  
3 head just talked about, there is another like a mini-containment  
4 inside the containment, and that is this core cage, we call it,  
5 and this is made out of structures, the core support structure,  
6 the core barrel and the upper internal structure. It is a  
7 cage because it is not completely closed up. Things can leak  
8 out, but structurally, if you like to put a big force, a  
9 high-pressure source in there, that source does not manifest  
10 itself on the outside unless there is some failure on this  
11 cage, on this structure.

12 Now, here we establish this yardstick, then. We  
13 go through the disruption, through the core disruption  
14 phenomenology, and we attempt to establish a general framework.  
15 That is, in very, very rough lines, we like to know how the  
16 accident progresses, what are the different steps, the differ-  
17 ent configurations that the core might find itself as it goes  
18 from initiation to the termination, to the end of this  
19 accident. This is what we call framework.

20 Here the discussion can be viewed or the whole thing  
21 can be viewed in two major parts. One is the initiating  
22 phase. That is where the disruption begins. The geometry is  
23 only beginning to be lost and therefore is much better defined  
24 from that point of view. And then the other one we call the  
25 disruption phase, and that takes on from the point where the

j-1-3,5 1 fuel actually begins to move at some rather large distances  
2 until the action is terminated.

3 One word about termination. Our task is to look  
4 at energetics. Therefore, when we say termination, we refer  
5 to the termination of energetic potential. What it means is  
6 the achievement of permanently subcritical state and not  
7 termination of the accident. Even after the core has gone  
8 through this termination of the energetic concerns, as you  
9 already know from previous discussions, it has to go through  
10 in the containment eventually somehow by decay heat.

11 Now then, within this framework we look at this  
12 framework and try to identify if there is any potential for  
13 energetic behavior. So we search, then, for energetic events  
14 within this general phenomenology, and we can classify them  
15 in terms of two classes, again, of energetic events. One is  
16 pertinent to the initiating phase. Here you will recall from  
17 the last discussion, in order to produce energetic events, we  
18 need to have rather forced fuel motions. Plain gravity will  
19 not do it here because the fuel just begins to disrupt and  
20 this disruption is rather incoherent because of the power  
21 distribution in the core.

22 The only way to do it, then, is to have forced  
23 fuel motion. That means you require pressure somehow. On  
24 the other hand, in the disruption phase a lot of the fuel of  
25 the core moves, therefore coherence might be achieved, and

j-1-3,6 1 therefore even gravity alone can do the job of producing  
2 energetics and that is the kind of situation that we call  
3 recriticality.

4 So we identify, then, initiating phase of energetics  
5 and recriticality. But then independently of whether we  
6 identify the possibility of such energetic events, we somehow  
7 have to see these come to a termination, come to an end, and  
8 obviously there are other non-energetic or mild ways by which  
9 the fuel can get out of the core, and we try to identify and  
10 see what is the potential for that.

11 So again, we search over this whole range of  
12 phenomenology, the whole accident progression and identify  
13 paths through which the fuel actually can get out of the core.  
14 When 40 percent of the fuel gets out, you can call it  
15 energetically terminated.

16 At the beginning, again, in the initiating phase,  
17 the only available parts for the fuel to get out is the  
18 coolant passages. This is up and down through the actual  
19 blankets and up and down. Those are very tiny, small paths  
20 and the fuel as it tries to get out might freeze there and  
21 might plug them. So one of the important considerations is  
22 are those parts available.

23 Furthermore, it is possible that those might plug  
24 up before any fuel attempts, and this is when this cladding  
25 melts before the fuel, at least in the initial stages of the



j-1-3,7  
1 loss of flow accident, and this cladding can be moved  
2 upwards and downwards and can plug and freeze and completely  
3 isolate from that point of view these actual paths.

4 Later on when the fuel actually continues to  
5 move in what we call extended motions and continues to produce  
6 heat, not only through decay heat but, we believe, also  
7 through neutronic activity, through recriticalities, because  
8 of the continuous presence of gravity, the sub-assembly walls  
9 begin to become attacked and melt, and therefore the gaps  
10 between the sub-assemblies now become available, and that is  
11 what is shown here.

12 So in the disruption phase, new paths become  
13 available for the fuel to get out. Again, those paths are  
14 about 1000 degrees lower temperature than the fuel melting  
15 temperature, a very large delta t. And again, will those  
16 paths allow for the fuel to get out before freezing occurs and  
17 therefore isolation.

18 Finally, the control assemblies, upon melting of  
19 the walls, provide additional paths for the fuel to escape,  
20 and again similar considerations apply here.

21 One important thing that I will come to again  
22 and again and I want to impress upon you is that as the  
23 disruption phase increases, as we go from the initial first  
24 pins to fail and first little fuel to begin to move, as we  
25 go down to whole subassembly molten, maybe more than one

j-1-3,8 1 subassemblies melting together and molten, more and more of  
2 those paths become available and eventually you will see that  
3 the actual system becomes overwhelmed by these relief  
4 paths. It is almost like trying to keep water in a sieve,  
5 and you know that is very difficult.

6 Here I am showing the sections or units within  
7 this loss of flow, which is Chapter 2 in the report, that  
8 addresses some of those areas, and we are going to try to  
9 walk through this, then, in a relatively brief fashion, through  
10 each one of them in that order.

11 MR. MARK: Theo, a question of a general sort.  
12 You did say that if you disperse 40 percent of the fuel,  
13 then recriticality is excluded. If I take the fuel and melt  
14 it all and magically put it in the pool in the bottom of the  
15 vessel, is it or isn't it critical if there is no sodium?  
16 Sixty percent will be critical if there is a good sodium  
17 layer on top, is that correct?

18 MR. THEOFANOUS: In fact, I don't think the  
19 sodium layer makes a big difference. Maybe if there is some  
20 steel around it might make some difference, but 40 percent  
21 is -- 60 percent is critical if you have the normal mixture  
22 of fuel and steel and internal blankets all mixed up within  
23 the core configuration.

24 MR. MARK: At the bottom of the --  
25

j-1-3,9

1 MR. THEOFANOUS: Inside the core. I don't care  
2 where you have it, but inside the original core configuration.  
3 Now, if you bring in also the radial blankets, and if you are  
4 going to do that, if you are going to let everything melt  
5 and go down to the bottom, probably also by that time maybe  
6 you bring in the axial blankets, the lower blankets, maybe  
7 even some of the radial blankets, then you have enough dilution  
8 that in fact without any loss of fuel the thing will be  
9 subcritical.

10 MR. MARK: It is that kind of question I am wonder-  
11 ing about.

12 MR. THEOFANOUS: It is this kind of thing. I think  
13 you will see that better when we discuss the loss of heat sink  
14 accident, because in this case, in fact, you don't lose any  
15 fuel. All the fuel stays inside, but what happens is more and  
16 more radial blankets come in. I think I will be able to  
17 answer your questions there.

18 Now, the details of this whole thing are sometimes  
19 so overwhelming that it can cause a lot of confusion. What I  
20 am saying is one can get hung up with details in looking at  
21 these core disruption accidents and really misunderstand  
22 certain simple features of them.

23 What we are trying to do here is trying to do  
24 exactly that, trying to look at them in a simple way, to  
25 really try to comprehend what is going on. I think this is

j-1-3,10 1 done well with this kind of picture. We call this a generic  
2 core-disruptive accident progression, and it is generic to  
3 the loss of flow accident in particular.

4 Here it is shown that from initiation, what I said  
5 before, from the initiation of core disruption involving just  
6 a few pins failing, until the complete core disruption, which  
7 is the whole core all molten inside the original core confines,  
8 this we call homogeneous whole core pool. The core will go  
9 from this initial disruption to this final disruption through  
10 successive stages of melting materials and relegating  
11 materials.

12 This is like a continuum of states here, and we  
13 describe this continuum state in terms of two representative  
14 ones. One is when the subassembly walls are largely intact,  
15 so at most we have individual subassembly pools, and this is  
16 this stage over here. The next stage comes about because of  
17 the heterogeneous core design, and here I would like to put  
18 up a figure donated to me from the last meeting from the  
19 project, and that shows very well, I think, an intermediate  
20 stage between the individual subassembly pool scale and whole  
21 core pool stage because of the pressures of the internal  
22 blankets.

23 These are low-power regions and much colder, so  
24 when the drivers are quickly melting, within a fraction of a  
25 second, they will remain intact. Of course, they will be

j-1-3,11 1 attacked by melt from inside and outside, but you can see  
2 that there is at least for a short period of time this  
3 structure over an annular pool, shown here by this red.

4 That is the reason, then, we have identified an  
5 intermediate stage, the annular pool stage.

6 Now, through this continuum of advanced disruption  
7 stage, there is a potential for the system to terminate the  
8 accident energetically, that is, to obtain energetic  
9 termination. This termination can happen either energetically,  
10 either in a forceful way, and by that we mean pressures  
11 developed of significant magnitude to forcefully eject the  
12 fuel materials into the sodium pool and to actually do that  
13 with enough force that some energy is delivered to the  
14 sodium pool, or it can terminate in a mild way again by  
15 escape through all those different relief paths. This  
16 process here we call hydrodynamic disassembly or disassembly,  
17 this process over here we call dispersal.

18 Now, depending on which of those exit paths the  
19 system can take for energetic termination, we have a different  
20 probability or a different chance of failing the system, and  
21 that is why, then, we have identified here four paths that  
22 go from the disassembly process into failing the reactor  
23 vessel, and that is what we mean by ex-vessel containment, and  
24 each one of those paths will have its own number, its own  
25 probability.

j-1-3,12 1                    Now, to have a quantitative view of which way  
2 things are going here, one could put numbers in each one of  
3 those paths, probability numbers. You see there they are  
4 split point decisions. At this point what is the chance of  
5 going into energetic disassembly, what is the chance of going  
6 into subassembly pool, and what is the chance of terminating?

7                    Now, one could express those things by words, and  
8 I would say the most likely possibility here is to go into a  
9 pool, and then from here on the most likely probability is  
10 to disperse. On the other hand, as you see, there is more  
11 than one step here and very quickly one gets confused with  
12 words if one tries to do this way.

13                    So we have bitten the bullet, I guess, and tried  
14 to put numbers here, and this has caused some controversy  
15 because people in general get nervous when one talks about  
16 numbers, especially about things that are somewhat maybe  
17 not amenable to exact quantification.

18                    But we think that it is an important exercise to  
19 put this whole thing into perspective, not to put in perspec-  
20 tive the situation to a whole PRA, because that involves another  
21 step, but to put in perspective the whole sequences within  
22 a core disruptive accident.

END  
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23  
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25

1 (Slide.)

2 So I would like, then, to show you the approach that  
3 we took in doing that. The procedure is, first, to identify  
4 or define a set of probability numbers on an order of magnitude  
5 basis and give definitions to them. And we have chosen three  
6 levels; one is characterized by  $10^{-1}$  as an event that really  
7 can be obtained by age of spectrum assumptions. This is an  
8 event, for example, that one normally would not expect on a  
9 best estimate, not even on a low probability, but one has to  
10 really go to the edge of the spectrum before one can claim that  
11 this event has occurred.

12 The next level,  $10^{-2}$ , is when one would require to  
13 make out of spectrum assumptions in order for this event to be  
14 realizeable. And this implies that we understand the phenomena  
15 controlling, and therefore, we can come up to the edge of the  
16 spectrum. Here, the trend is different. Here we are trying  
17 to look at the phenomena from the other end of it and consider  
18 what is physically possible, and really, this kind of a circum-  
19 stance is so far out from what we expect that this will be  
20 outside of the spectrum.

21 So when we are at this level, I guess I am saying  
22 we know better the best estimate behavior when you come to the  
23 edge of it. When we are here (indicating) we're exploring more  
24 to the outer spectrum and coming up to the edge on that side.

25 Now, if you have two events, one like this and one

1 like that in series, one following the other, you would have  
2 a probability of  $10^{-2}$ ,  $10^{-3}$ , and physically, I think it is  
3 reasonable to claim that such a sequence with two very unlikely  
4 situations would really be classified as physically unreason-  
5 able. So that's how our definition of what is physically  
6 unreasonable.

7 Now, having done that, we need to develop the basis  
8 for judging each one of the parts that were given before, that  
9 you saw in the generic accident progression, so we can identify  
10 put numbers to them. Now again, to do this precisely, one  
11 would have to have the complete probability distributions, and  
12 actually, quantify each one of the parts not with a number but  
13 with a probability distribution.

14 We claim that really, we don't have the information  
15 to do that. It's a step that cannot really be accomplished in  
16 a reliable and meaningful way at this time. One can always  
17 put probability distributions and carry through the exercise,  
18 but nobody has done it yet and I think it will take some time  
19 before anybody actually attempts to do that in a reasonably  
20 confident way.

21 On the other hand, if we ask a different question,  
22 say what if we can assign those numbers as high confidence  
23 numbers at the end of those probability distributions. We  
24 don't have to know the details in order to know what is out at  
25 the end, and if we try to do that in a conservative fashion, we



1 can actually accomplish this step as we have done in the  
2 report.

3 Now after this, of course, things are simple, we  
4 can carry out the arithmetic, following each one of the  
5 composites we are interested in, and in particular, we can come  
6 out with the probability for vessel failure by summing up  
7 the probabilities of those four parts that were shown on the  
8 previous slide.

9 Now we have a number at the end, and that is the  
10 probability for failing the vessel, and this number was  
11 obtained in terms of these definitions. It would be very  
12 wrong if one would take that number now and go and stick it  
13 into a probability risk assessment study that is based on a  
14 different set of numbers; a lot of them actually based on  
15 experience. Therefore, we emphasize that with the start, one  
16 has to take this number and go back and convert it into the  
17 physical meaning associated with this based on those definitions  
18 here. If the number is less than  $10^{-3}$ , then we can say that  
19 the failure of the vessel is not considered physically  
20 reasonable; it is physically unreasonable. And that is really  
21 the bottom line of that.

22 And you will see, we will come to the bottom line.  
23 Now, if one wanted to see what is the impact of this kind of  
24 treatment to a total probabilistic approach, considering both  
25 the front end of the core disruption, the initiator side, as

1 well as the back end of it, considering the radiological  
2 aspects, one has to go back and look at the context of the  
3 numbers assigned to both of the sides, what would be a correct  
4 number to assign to physically unreasonable states. And  
5 that's the number that has to go in there, then, in the  
6 analysis.

7 MR. KASTENBERG: Do you also come to the converse  
8 of that? In other words, you must have assigned number 1 to  
9 your best estimate; something that is. I mean, if you heat  
10 up the core you'll boil the sodium. That has to have a 1.  
11 So then if you propagate everything and you conserve proba-  
12 bility, you have to have something that is physically reasonable  
13 at the end, I presume.

14 MR. THEOFANOUS: Right, sure.

15 MR. KASTENBERG: And you do have it.

16 MR. THEOFANOUS: Yes, and you will see that. At  
17 the end we're going to put the numbers, and you will see the  
18 many of the parts, in fact, have 1, and therefore, this way if  
19 one wanted to arrive at the best estimate scenario, or if you  
20 wanted, you could find out which way it goes.

21 So with these more or less introductory materials,  
22 I think we are going to go into some of the technical details,  
23 and to start out by considering the CRBR's actual capability.

24 This is a schematic of the system. And here are  
25 some of the acronyms we use; this is the UIS or operating

1 internal structures, and that is a rather large structure that  
2 supports from the top with four strong columns. Here you see  
3 the core barrell, and this is the radial shield; there's a  
4 number of subassemblies there that contain the steel mass,  
5 there's no core. The core is in the middle, and this is, of  
6 course, the core structure.

7 Also, there is another structure up here, the upper  
8 core structure, and this is made up of blankets and the  
9 fission gas plenum, steel.

10 So to a large extent, this is just an empty hole  
11 kind of structure, like a honeycomb.

12 Now, the failure mode through the system, then,  
13 the US -- see, there's pressure, there's an energetic event  
14 suddenly releasing high pressure into the core and can, of  
15 course, block the UIS and the cones can buckle, and that would  
16 make this whole thing be translated upwards and the high  
17 pressures, then, would manifest themselves in the bottom of  
18 the pool, the pool will accelerate and heat the head. This  
19 is the vessel head structure, and this slug impact, then, can  
20 cause damage to the head, and the appropriate way to look at  
21 that is in terms of the impact kinetic energy.

22 For the UIS, the failure mode is buckling, as I  
23 said before, and for that, one needs to know the pressure  
24 history on the UIS.

25 Now, the structural evaluations we have done show

1 that the core support structure is by far stronger than any of  
2 those 2 structures, so it doesn't even enter the picture.  
3 That is, the thing is going to fall apart in other places  
4 much before the core support structure gives way.

5 On the core bottom, however, that can fail also by  
6 radial strain under this body, and again, for one to evaluate  
7 that one needs to know what is the pressure time history of  
8 the bounding of this core bottom.

9 We have done the analysis, then, according to this  
10 kind of a logic, and we found out that the approximate approach  
11 to failing this cage, this structure here, is at an event of  
12 about 100 dollars per second in a two-phase disassembly mode.  
13 And I emphasize that because there is some very great difference  
14 between the configurations to which these assemblies take place.

15 And this is one that basically turns around and  
16 terminates by variable pressures.

17 Now, after the structure fails, as I said before,  
18 the full accelerates, this work is done, and obviously, this  
19 requires a high level of energy, and for this energy, then,  
20 to be of significance as far as the vessel head structure is  
21 concerned one must approach the 200 dollars per second, two-  
22 phase disassembly rates.

23 MR. ZUDANS: Theo, what makes the core accelerate  
24 upward? Just the pump pressure?

25 MR. THEOFANOUS: No, no, this is an exiting event

1 now, and how ones goes about getting an energetic event you  
2 will hear in the rest of it. But here we are hypothesizing  
3 that suddenly the core becomes molten and has developed high  
4 pressures.

5 MR. ZUDANS: And it can't go down.

6 MR. THEOFANOUS: These pressures, then, develop  
7 within this cage, so it's going to load all the sides of it,  
8 upward and otherwise. And if the thing fails, then, of course,  
9 we're going to load also this other side.

10 The reference yields, just for normalizing with  
11 what you know from previous discussions, the 100 dollars per  
12 second, two-phase disassembly has an 1150 megajoule ultimate  
13 work potential. This is kind of short terminology to use  
14 instead of this expansion to an atmosphere. Normally, if we  
15 want to characterize how much an energetic event is, one  
16 takes the initial state, following the neutronic termination;  
17 this has a temperature distribution, a pressure distribution,  
18 and then one lets that expand pocket by pocket adiabatically  
19 until it all comes out to one atmosphere pressure. And then  
20 one calculates the PV curve from that, the PV work, and this  
21 is what we are referring to as ultimate work potential; that  
22 is, the most possible work that anyone can ever do from this  
23 kind of an event.

24 MR. MARK: Isn't that similar to PVs throughout  
25 the vaporized region, initially?

1 MR. THEOFANOUS: Initially, yes. And here, for the  
2 200 dollars per second is 25, 50 megajoules; very significant  
3 amount of energy.

4 MR. MARK: It would seem to me more graspable if  
5 instead of the ultimate work potential you just said the  
6 product of PV throughout the core initially is, like, 1150  
7 megajoules.

8 MR. THEOFANOUS: Well, that would require many more  
9 words to say it --

10 (Laughter.)

11 Or, the usual expression is expansion to an atmosphere; it  
12 is the same thing.

13 MR. MARK: I'm familiar with that, but obviously  
14 meaningless.

15 MR. THEOFANOUS: This is really a controversial  
16 item because many people don't like to see these big numbers  
17 and they think really more meaningful things. Essentially,  
18 everything is going to be contained inside the vessel. A more  
19 meaningful thing is to do this PV expansion only to the cover  
20 gas volume, but then again, one loses a little bit from doing  
21 that because every system has its own cover gas, and maybe  
22 even the same system that might have a different cover gas.  
23 So this is really a thermodynamic definition, generic, that  
24 really conveys the magnitude of the energy possibly available,  
25 and as long as we remember that really, those levels of energy

1 never manifest themselves -- one should never make the  
2 mistake that those numbers actually show up in terms of the  
3 work .

4 MR. MARK: Well, you mean they never get converted  
5 to kinetic energy.

6 MR. THEOFANOUS: Yes.

7 Now, the applicant has used a 661 megajoule source  
8 to record. They have neglected all those structures. Basically  
9 they let this equivalent PV carb that gives this 661 megajoules  
10 by again, adiabatic fuel, and let it expand upward unrestrained,  
11 accelerate the slug. At the time of impact, there was a kinetic  
12 energy in its volume and all the materials in the core, and  
13 they counted up all that kinetic energy and they found that  
14 to be around 75 megajoules. Some of the energy, of course,  
15 if you do the PV work for this source here, it is about 100  
16 megajoules, but they did that taking into account the structural  
17 deformations and some of that went into strain, so the total  
18 kinetic energy was 75 megajoules, counting the core materials.

19 And then they made the step that said we will design  
20 the head for a 75 megajoules slug impact energy. This involves  
21 a situation now to try to account for a kinetic energy that's  
22 inside the bubble with liquid moving inside the vapor space  
23 and then going and heating the liquid.

24 So as you see, the approach that we took is somewhat  
25 idealized, in the sense that it neglects important, in our

1 opinion, very important mitigator of energetics into the  
 2 pool. And this is an important mitigator because as you will  
 3 see, this 100 dollars per second energy level really we don't  
 4 even expect to have it under the worst possible conditions in  
 5 the core.

6 So from that point of view, it is really an ideal  
 7 situation. If there was a range of expected ramp rates that  
 8 go up to 150, let's say. So maybe in that case, by neglecting  
 9 all this mitigator, may be not too bad because maybe a portion  
 10 of events only are covered by that portion, and there's  
 11 another large portion that is not covered.

12 But what you are saying here is the whole thing is  
 13 covered by this, and if you neglect that, you basically  
 14 neglect the necessity of completely mitigating the whole  
 15 energy release.

*End  
 TP 1-4  
 MS*

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j-1-5,1

1           Now, what I have to do is actually walk you through  
2 the different steps of showing how this characterization of  
3 \$100 and \$200 per second energy levels was obtained for  
4 the failure of those structures.

5           MR. ZUDANS: Theo, the 75 megajoule impact,  
6 kinetic energy impact, of course neglected all the internals,  
7 and that means they made a clear path for this. Now, if you  
8 do not neglect all the internals, you would get less kinetic  
9 energy at the impact but you would strain the vessel more.

10           Now, I know that the tests show that you can get  
11 twice the strain with the upper internals than you get without.  
12 Did you analyze this?

13           MR. THEOFANOUS: Yes. In fact, we are looking at  
14 that in a realistic way. We are looking with the internals  
15 and everything there. The reason I am bringing this up here  
16 is because if one were to look at these numbers here and  
17 these numbers, one might say, like somebody told me in a  
18 previous presentation I made, the Staff, who is more conser-  
19 vative here? You give higher numbers but you are coming at  
20 the bottom line and saying that this very, very high level  
21 of energy still, as you say, will give you the tendency the  
22 same that the applicant gets.

23           What I want to say is that the two approaches are  
24 so different, and the one is so idealized that I don't  
25 want to venture into saying who is more conservative, who is

j-1-5,2 1 less conservative. We will just give you our story here,  
2 that we think is consistent, and then you have to make your  
3 judgment as far as how the two compare.

4 MR. ZUDANS: Well, you didn't get my point. Which  
5 failure modes?

6 MR. THEOFANOUS: We are going to consider all three  
7 failure modes.

8 MR. ZUDANS: But it doesn't show vessel failure  
9 mode. You only show column buckling --

10 MR. THEOFANOUS: We want to talk about the vessel  
11 failure mode also, but the point here is that we don't really,  
12 are not really overly concerned about the vessel failure,  
13 although we are looking at it, because the primary concern is  
14 the failure of the upper head. Like Carson said, the failure  
15 of the head has the potential of releasing materials into the  
16 upper containment, and one is concerned from the point of view  
17 of violating the containment. Now, if materials were released  
18 into the cavity, you are talking about a different set of  
19 events from the point of view of failing, basically releasing  
20 any radioactivity to the outside.

21 MR. ZUDANS: I just want to know whether your  
22 analysis considered vessel failure as a failure mode.

23 MR. THEOFANOUS: As a failure mode.

24 MR. ZUDANS: Correct. Because you list failure  
25 mode- and it does not include vessel -- core barrel strength,

j-1-5,3

1 column buckling --

2 MR. THEOFANOUS: You see, I am keying the whole  
3 discussion here, and as it should be, the whole energetic  
4 discussion, keyed to generating failure of the upper -- of  
5 the vessel head structure, potential generation of missiles,  
6 basically potential threat to the containment.

7 Now, there are other ways, there are other parts  
8 of the system, and there may be other parts that can fail. We  
9 are not concerned about all of those other parts from the  
10 point of view of energetics.

11 MR. DICKSON: Excuse me. I would like to add  
12 something here because I read yesterday from a figure that  
13 contained a typographical error. The numbers I read were  
14 that SM3 had a 2.8 percent strain, and SM2 had a 1.4.

15 MR. THEOFANOUS: That is correct.

16 MR. DICKSON: The truth was -- that's a typo --  
17 that should have been 4.4. It turns out that when you include  
18 the upper internals in the test, the thing that is strained  
19 more is the core barrel and not the vessel up near the head.

20 MR. ZUDANS: What was the strain in the vessel near  
21 the head on that figure that was labeled 1.4, 2.8? I thought  
22 the labels were switched.

23 MR. DICKSON: It was 2.8 and 4.4. The graph at  
24 the bottom is correct.

25 MR. THEOFANOUS: The 1.4 and 2.8 still is essentially

j-1-5,4 1 close to 80 percent higher.

2 MR. DICKSON: Yes, but it was the SM2 that had  
3 the greater strain. SM3 had less, the reason being that more  
4 of the energy went into straining the core barrel.

5 MR. THEOFANOUS: That doesn't make physical sense.

6 MR. DICKSON: Well, it does. If the UIS is  
7 sufficiently strong, the vessel isn't strained at all. All  
8 the strain will be taken up in the core barrel. With an  
9 infinitely strong upper internal structure, the only thing  
10 that can take up energy will be straining of the core barrel.

11 MR. ZUDANS: In other words, it prevents the  
12 pressure variable penetrating upward.

13 MR. DICKSON: That is correct, and that was  
14 experimentally determined.

15 MR. THEOFANOUS: What is happening here is this  
16 slug accelerates and hits the top, and because of the impact,  
17 the pressure is what strains the vessel. Now, obviously if  
18 you keep the slug from accelerating, and you can do that by  
19 keeping the UIS in place, obviously you are going to be having  
20 less strain up here and more strain down here.

21 MR. ZUDANS: It is not you that confused me. It  
22 was yesterday's incorrect number.

23 MR. THEOFANOUS: I heard about yesterday's discus-  
24 sion.

25 So now, then, looking at what we have done in terms

j-1-5,5 1 of analysis to analyze these failure modes, we need the  
2 pressure history of the boundaries. After we get this  
3 pressure history, we can do a structural analysis to take the  
4 strain of the barrel, and the same thing for the upper  
5 internal structure.

6 For one to do this correctly, one would require to  
7 do a couple of fluid structural analyses. This capability  
8 really is within reach. Rexco can do that but it is not really  
9 all geared up to do this now. Therefore, we had to resort  
10 to an approximation technique, and the way we did it is to  
11 get the pressure histories on the boundaries, assuming these  
12 boundaries are rigid first. That maximizes the pressure  
13 history on the boundaries.

14 Now, having done that, go and do a structural  
15 analysis and find out about how much of a strain we have. Now,  
16 having gotten the strains, we can go back and do the initial  
17 expansion analysis again to find the new pressure histories on  
18 the boundaries but allowing for the displacement that we  
19 calculated before.

20 The new pressure histories now are applied to the  
21 structural, this two-step approach, applied to the structural  
22 analysis, and now we come up with new strains. We have done  
23 that, considered the two levels, the \$100 and the \$200  
24 per second level, and we obtained in the first one convergence  
25 in both steps. The strain was about 2 percent, 2.5 percent,

j-1-5,6 1 while in the other level, in the 200 level, the second time  
2 around the strain was somewhat lower than the first one, so  
3 it was not an exactly convergent situation but it was on the  
4 conservative side.

5 Now, remember that this is a case in which already  
6 we know we were going to fail that. We are primarily concerned  
7 now for failing the head, and therefore this lack of conver-  
8 gence is not really a big problem.

9 MR. ZUDANS: Theo, the first step would not be con-  
10 servative pressure distribution, if I understood you correctly.  
11 If you took rigid walls and derived the pressure history and  
12 turned around and applied that pressure history to the walls  
13 to get the deformation strength and repeated the pressure  
14 history computation but now with the displacements as a func-  
15 tion of previously computed pressure, then that expansion  
16 would be larger than it really is because your actual  
17 pressure --

18 MR. THEOFANOUS: That's right.

19 MR. ZUDANS: And you repeated those steps?

20 MR. THEOFANOUS: But if you go back and you  
21 converge your --

22 MR. ZUDANS: That is what I am asking.

23 MR. THEOFANOUS: And you converge. That's what  
24 I'm saying. That's right.

25 MR. ZUDANS: Then it would be okay.

j-1-5,7 1 MR. THEOFANOUS: So now I want to give you the  
2 picture here again of the energetic event. There is no  
3 more energy released in the core, and typically there will be  
4 a pressure distribution in the core. There is a high flux  
5 region in the center, and that's where the peak pressures,  
6 peak temperatures are going to be.

7 Now therefore, one would expect to have a temperature,  
8 radial and axial temperature distribution and pressure distri-  
9 bution. Obviously, the pressure up here will not be the same  
10 as the pressure up there.

11 Now, if this whole thing was solid or liquid and  
12 had good transmission characteristics, that would be very  
13 close to being true. However, we have a two-phased mixture  
14 here that is very compliant. It has a lot of void inside it.  
15 As a result of that, the pressures that are manifested in  
16 the center, actually they never materialize on the walls.  
17 That is a very important thing because if this mitigating  
18 aspect, this compliance of the medium inside the core, was  
19 not there, there would be no mechanism to get the energy in  
20 the first place. One really needs to understand that.

21 If this whole thing was all liquid, let's say, one  
22 can postulate all liquid, and say I'm going to put into that  
23 all liquid \$500 per second or \$1000 per second, one  
24 could generate almost no energy at all because of the thermal  
25 expansion characteristics of the liquid. You know, very high

j-1-5,8 1 pressures, very quick shutdown, and very low energy deposited  
2 in the fuel. That is well known for many, many years. So  
3 the main reason that one can ever expect -- I don't say will,  
4 but ever expect to get energy here is because of the compli-  
5 ance of the system. Therefore, it is a very gross conserva-  
6 tism in error if one does not take into account this compli-  
7 ance in calculating the loads on the system, and that is what  
8 we were trying to do here.

9 The core barrel will strain under loading, and  
10 that will provide additional expansion, additional energy  
11 absorption within that case than will ever appear up on the  
12 expansion of the pool. Therefore, that represents an impor-  
13 tant loss. The same thing with the upper core structure.  
14 The upper core structure, as I said, is very much a void  
15 structure and is going to be accelerated upwards and is going  
16 to hit the UIS. But again, it is a crushable thing and it  
17 is going to take energy to crush it, and also it will  
18 produce an expansion, which means a drop in pressure here  
19 before even any of this pressure can appear up in the pool.  
20 So this is some of the important phenomena, and you can see  
21 here the trends.

22 This is the center core pressure as a function of  
23 time from an expansion, and this is the \$200 per second,  
24 two-phase disassembly.

25 You see, we are starting out at extremely high



j-1-5,9 1 pressures, 1500 bar, but it very quickly drops down to just a  
2 few hundred, 200 to 300 bar, and then goes like this.

3 Now, this asymptotes out because remember that we  
4 did this part of the calculation for the cage for a completely  
5 closed system. We didn't allow for the fuel to get out. As a  
6 result of that, these are going down.

7 On the other hand, this expansion, that same  
8 expansion -- which, by the way, was done using the SIMMER  
9 code, and this is just a simple state adiabatic expansion,  
10 and the fact is you will see later we accounted for every bit  
11 of the energy in the thermodynamic sense at the end.

12 So in the core barrel, however, the pressure, at  
13 most it is about 700 bar for a short time and then quickly  
14 drops down again. Of course, eventually all the pressures  
15 have to come together.

16 As far as straining the core barrel, this part  
17 here is very important, this high part, and the short-livedness  
18 of it is important.

19 Now, I want to along the way contrast that with  
20 what the project has done. They have taken basically the  
21 peak pressure, not the 1500, because they didn't consider  
22 such high energy levels, but they took the peak pressure of  
23 the 661 case and applied it directly, I guess, on the pool  
24 and in accelerating from time zero, and in the study -- not  
25 the project, but in the study that was published in the

j-1-5,10 1 Chicago conference of reactor safety, in which the core  
2 barrel was examined from this set of conditions to the 661,  
3 that peak core pressure was applied directly on the core  
4 barrel.

5 As a result of that, they obtained displacements  
6 on the order of 12.5 percent or strains of 12.5 percent. For  
7 our cases we need much higher energy to obtain the same  
8 displacements, and that isn't exactly because of this  
9 attenuation of the pressure, because of the compliance of  
End 1-5 10 the medium.

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1           This is the expansion, the P delta p curve, and  
2 this 200 dollars per second. If one was to do a pocket by  
3 pocket isentropic expansion to the total free volume that is  
4 inside that enclosure, that takes into account the strain of  
5 the core barrel, as I said before, all the internal voids,  
6 the displacement of the upper core structure, adds up to  
7 about 14 cubic meters. If one was to do this PV work, one  
8 would obtain 520 megajoules; that's maximum thermodynamic work  
9 potential.

10           By doing a similar calculation in the straining,  
11 as I described before, we find that about 180 megajoules have  
12 gone into the core barrell as strain energy. And 340 megajoules  
13 have gone into the upper core structure as kinetic energy.  
14 And the sum of the two is this number up here. So you see  
15 it's a numerical calculation, and we are coming out -- we  
16 haven't dissipated energy, we haven't lost any energy; it  
17 is only that some went to strain and some went to kinetic  
18 energy.

19           Now, one might say that this upper core structure  
20 is moving upwards with this very high kinetic energy; can  
21 that add anything to the damage in the system. Because if  
22 that cannot, what this really implies, because of all this  
23 internal expansion inside the cage, really somehow were  
24 dissipated 520 megajoules, which is a really significant  
25 number.

1 Well, you can visualize this upper core structure  
2 again is a thing that is accelerated and moves very rapidly,  
3 but goes and heats up the structure or mass that is much larger  
4 in size and is strongly supported from the top. And it goes  
5 and crushes and all the damage is absorbed into ashes and  
6 dissipates. And all of that energy then will show up as  
7 kinetic energy. Up to this point.

8 So now then, having done that, we found that the  
9 100 dollars per second somehow dissipated all of it inside  
10 the bubble, gave very minimal, only 2 percent strain on the  
11 core barrel, or 2.5 percent in the core barrel and just one  
12 or two inches, very little, on the upper internal structure.  
13 So this 1150 megajoules we feel is safely and totally  
14 contained in that cage.

15 Now, we consider the second threshold which is  
16 the long-term expansion, and that only comes into the picture  
17 if this cage fails, and that requires a high level of energy  
18 than the 1150.

19 The next level we picked to analyze went all the  
20 way up to the energy level that would be required to produce  
21 enough kinetic energy in the pool to start compromising or  
22 approaching the design limit of the vessel head. And that  
23 is the 2550 megajoules at 200 dollars per second.

24 The situation here is as follows. Obviously,  
25 under this very high energy release the core barrel will strain,

1 and we estimate, following the short-term approach analysis I  
2 described before, that the core barrel will strain maybe  
3 quite a bit, maybe up to 15, 20 percent. So that is really  
4 going up there to the limit and depending on whom you talk to  
5 somebody will say maybe it will fail, or somebody may say it's  
6 really up there but if you took into account that the vessel  
7 itself is going to strain and that it's going to come up  
8 against the guard vessel and there's going to be another  
9 stiffening effect from that, so maybe this might give a little  
10 bit maybe it will not become a total catastrophic failure.  
11 But it's right up there.

12 By the way, this analysis of the strain of the core  
13 barrel was done taking into account the stiffening effect of  
14 the vessel wall and the presence of sodium also taking into  
15 account the drop in pressure from the edge of the core to the  
16 core barrel interface because of additional convergence and  
17 additional divergence, I should say.

18 MR. ZUDANS: In this calculation, was it assumed  
19 that the upper internal structure effectively close to core  
20 barrel and that reactor vessel volume above it did not  
21 participate in expansion process?

22 MR. THEOFANOUS: The first part of the calculation  
23 was done by neglecting the relief of the pressure because of  
24 the displacement of the --

25 MR. ZUDANS: So it was like a closed --

1 MR. THEOFANOUS: It was very close, it was conserva-  
2 tive. Now in reality, when that begins to move out, the  
3 pressure will be relieved and you will have less volume.

4 However, again, remember that mostly in this  
5 kind of forms -- and I must admit I don't have a very good  
6 feel for that, but just looking at the results, it's very,  
7 very clear that in a very, very short time with very, very high  
8 pressures, those are the ones that are producing most of the  
9 energy release into the structure; that has produced most of  
10 the strain. And then from then on, basically it just keeps  
11 on moving like that.

12 But I do want to make the point, however, that both  
13 this model here as well as this model, I mean the structural  
14 models, were checked very carefully because obviously, it's  
15 an important part of the story. We went into each sector,  
16 number one, benchmarking against the SRI tests, structural  
17 test, and more really because we were -- because we thought  
18 we found some interesting discrepancies between our results  
19 and this paper I mentioned to you from the Chicago conference  
20 that was giving about the same kinds of strains for what we  
21 thought were different pressure histories. We took their  
22 pressure history that they used in there to get the strains  
23 and we put it in our model and really got exactly the same  
24 deformation. So we have an additional confidence then that  
25 our structural results are consistent in terms of the methods

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1 applied and techniques to such a core disruption which has  
2 really received extensive study and verification over many,  
3 many years.

4 MR. ZUDANS: In all this process, the core support  
5 plate and the core barrel and connection to the conical skirt,  
6 were they assumed to be rigid? The bottom end was closed -- ?

7 MR. THEOFANOUS: The bottom end was closed and  
8 rigid, that's right.

9 Now, then, what happens here, if we have this  
10 2550 megajoule kind of an event, the UIS would buckle, will  
11 move up by buckling the columns, and that will then generate  
12 paths for the high pressure zone to expand. And therefore,  
13 a bubble will form here, and on the other hand, there will  
14 be another drop in pressure between the core and the bubble,  
15 and this is caused because of the additional acceleration  
16 required as the fluid flushes to fill up the bubble volume.

17 And this amounts to about a half -- the pressure  
18 here is about one-half of the pressure up here, and that is  
19 shown schematically like this (indicating). This is how the  
20 core pressure decays with time in this long-term expansion,  
21 and this is how the bubble pressure decays with time,  
22 typically maintaining this well-known critical pressure vessel  
23 about .5.

24 This we refer to as the throttle effect, and you  
25 can see that that is very important because over this part

1 of the transient the pressure really does not decay very  
2 rapidly; you are almost near the pipe, so if you're thinking  
3 in terms of the delivery to the slug in terms of P delta v,  
4 yousee here the delta v is the same in both cases, you see  
5 that in this case you're going to get a warp that's about half  
6 as much as you would have gotten if you neglected this.

7 MR. ZUDANS: I'm trying to find the slides in  
8 your handout --

9 MR. THEOFANOUS: You won't find them there because  
10 those were not made directly from the --

11 MR. ZUDANS: This is not very easy for me to follow  
12 because you have so many additional loads on the slides. It's  
13 just no way we can copy them and I think this is part of your  
14 argument.

15 MR. THEOFANOUS: Right. The idea here is that --  
16 again, I think the problem is that you do not have the document  
17 involved ahead of time. But the idea is that I want to -- I  
18 didn't want to restrain myself only to the figures that are  
19 only from the pages because then I wouldn't be able to explain  
20 to you --

21 MR. ZUDANS: Can't we get a copy of the slides made?

22 MR. THEOFANOUS: What we can do, however, is we  
23 can make copies of those on the break and you can then take  
24 notes right on it.

25 MR. CARBON: Let's just take a break at this time



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and make copies of all these slides for everyone.

Zenons, would you like to go back and start back  
someplace?

MR. ZUDANS: No, that's fine, as long as we  
get the pictures.

MR. CARBON: Let's break and start a little earlier.

(A short recess was taken.)

End Tp  
MS1-6

1 MR. CARBON: Let's resume the meeting.

2 Go ahead, Theo. Could you move your microphone up  
3 closer?

4 MR. THEOFANOUS: Okay. I was suggesting we start  
5 a couple of slides back, and now we can focus on the vuegraphs.

6 MR. CARBON: Just give the entire presentation if  
7 you will.

8 MR. THEOFANOUS: Give what?

9 MR. CARBON: Give the entire presentation.

10 MR. THEOFANOUS: Oh, the entire presentation? Yes.

11 MR. CARBON: Unless that takes you too far back.

12 MR. THEOFANOUS: No, I will do so.

13 The structure that first we see the energy in this,  
14 from the core is this enclosure, this cage made out of the UIS,  
15 the core barrel and the core support structure. The core  
16 support structure is very strong, much stronger than those two  
17 other ones. Therefore, this internal containment, or this cage,  
18 will generally fail first through here or through here. The  
19 failure modes would be radial strain of the core barrel and  
20 buckling of the coils of the UIS.

21 And what we need in order to be able to evaluate  
22 those is the pressure history of the boundaries. Assuming  
23 that -- and also we find that there is a release of \$100 per  
24 second, corresponding to 1150 megajoules ultimate worth  
25 potential, is the kind of an easy way to approach this failure

1 limit.

2 Now beyond that point, and after the UIS has failed,  
3 there is a path for this energy to show up in the sodium pool.  
4 It will accelerate the sodium pool then, and this pool is going  
5 to hit the head. That is the mode of transmission of energy  
6 from this source to the vessel head structure.

7 Here, therefore, one is interested on the slug  
8 import energy. This is quantified by the import kinetic energy,  
9 and we find about \$200 per second energy release or energetic  
10 level in a two-phase mode. And a two-phase disassembly is what  
11 is required to approach the design limit of 75 megajoules import  
12 kinetic energy.

13 This corresponds to 2550 megajoules, and this is to  
14 be contrasted to the number used as a source by the applicant  
15 as 661 megajoules. And I doubt if there is any reason to go  
16 back and repeat all this discussion about what the applicant  
17 has done.

18 Now the analysis that we carry -- and we tried to be  
19 persistent here in terms of where the source is and how does it  
20 show, how does the energy show in different points in the  
21 expansion -- can be considered in terms of two parts. One we  
22 call the short-term expansion, and that is what it takes for  
23 this to expand against these boundaries, these inner boundaries,  
24 and kind of come up to some kind of a positive charge equilibra-  
25 tion.

pv 3

1           This sort of expansion gives for us the pressure  
2 histories of the boundaries, and as a result of that, we can  
3 determine the failure limits or the failure thresholds.

4           The longer-term expansion requires that this is moved,  
5 that the sodium slug accelerate, and this is taking place on the  
6 order of magnitude of about 50 milliseconds while this first  
7 shorter expansion takes place on the order of 20 milliseconds.

8           So that's what I will show next then is these two  
9 steps. First step: The part of the core that is undergoing  
10 this disassembly is going to end up at the end of the end of the  
11 neutronic excursion with a temperature history. That implies  
12 the pressure history, and that is going to be highly nonuniform  
13 because of the peaking of the flux in the center.

14           Because of the compliance of the medium -- and again,  
15 it has to be compliant because if it is not, we wouldn't have  
16 the energetics in the first place -- this very highly  
17 centered pressures are not going to show up in the boundaries.  
18 There will be a delay in the decay and an expansion associated  
19 with this decay.

20           And that will show up here. The core pressure in  
21 the center as a function of time as determined by an adiabatic  
22 calculation, you see that within 5 milliseconds goes from a  
23 very high volume, an initial volume of 1500 bars down to just  
24 a few hundred bars.

25           The -- by comparison, the core barrel pressure

1 history, taking into account the transmission characteristics  
2 of this medium, rises to a few hundred bars and then kind of  
3 comes and merges with the center pressure as the whole thing  
4 comes together to equilibrium or requires equilibrium.

5           The upper core structure, seeing these high pressures,  
6 is going to translate upward, is going to crash. Now, that is  
7 a complicated problem that hasn't really been done before. And  
8 the way that we are doing it -- I didn't mention that before;  
9 maybe it's helpful here -- the way we are doing it is by taking  
10 the mass of this structure, just as an initial constraint; we  
11 modeled that as the mass being there, and now seeing these  
12 pressures and therefore being accelerated.

13           And we allow it, allow this mass numerically to come  
14 to a full dense condition. And until it comes to full dense  
15 condition, really the loading of the UIS is very minimal.

16           In reality, what happens between the beginning of this  
17 crossing forces and until the mass comes to full dense condition,  
18 what really happens is that the UIS is there with its strength,  
19 much, much stronger, of course, than this little structure here,  
20 so it provides the backup against which this structure can  
21 crash.

22           This takes place over a period of about 20 milli-  
23 seconds. This process of translation and crashing is only 20  
24 milliseconds then that the UIS actually begins to see the loads.  
25 At that time, therefore, the UIS is going to see the loads

1 of this magnitude instead of loads of this magnitude.

2 So the effect of the translation and crashing and the  
3 straining of the core barrel is to basically dissipate energy  
4 through our expansion and dropping pressures before the pool  
5 sees any of the pressures present in the core. We have heard  
6 about these losses.

7 So those are really, I think, the important ingredients  
8 here, the important considerations, as to why one would be far  
9 off the target here if one was to take this traditional  
10 approach of pocket-by-pocket expansion of the core in applying  
11 these resulting pressures to all the boundaries of interest.

12 Yes?

13 MR. CARBON: I would like to have a little bit more  
14 discussion about the upper end boundary conditions. The  
15 upper core structure is supported on top against this closure  
16 that really goes and supports against the head with columns.  
17 And during this calculation you are assuming that that upper  
18 -- at the top of your blue arrow -- that line does not move?

19 MR. THEOFANOUS: Yes. The shorter calculation is  
20 then with this goes and this goes and this goes. All of the  
21 first step of this calculation with the boundaries being fixed  
22 rigid. We do the expansion. And you come up with the  
23 pressure histories like is shown here in these two boundaries.

24 Now, since we know that, we go in and do a structural  
25 calculation for this and for that. This will give us some

1 strains. Depending on the pressure, the strains might be quite  
2 a bit.

3 Now, this strain then is taken back into the original  
4 expansion. We allow this boundary now to move and get us back  
5 in basically up to an equivalent amount of strain as we  
6 calculated before.

7 MR. ZUDANS: Including the top?

8 MR. THEOFANOUS: And including the top. Right. Well,  
9 in fact, the top, in the case of the -- in the top we didn't do  
10 that in fact, because, you see, the top before it even begins  
11 to work, the two come together. So the transmission  
12 characteristics and so on are not all that important.

13 So we never really -- for some of them, you  
14 never really took credit for any losses in venting or  
15 displacement of the UIS in that short of a period. However,  
16 this is in the immediate vicinity; it has to do a lot with the  
17 dissipating of this and the pressure history, and that's the  
18 reason we take into account this early straining of the core  
19 barrel.

20 So in the second time around, and allow this to  
21 expand to an equivalent amount of strain that we calculate in  
22 the previous step, now we calculate pressures in the core  
23 boundary as a result of this new expansion. Take that new  
24 pressure now, put it back into the structure, and calculate  
25 again. We find that for \$100 per second, the second time

pv 7

1 around we come up with the same strain. Therefore, we think  
2 we have conversed enough, and although the calculation is done  
3 in another or a couple of fashions, I think we are in good  
4 shape as far as predicting the ultimate strain of that.

5 We find there is only 2 percent or 3 percent. It is  
6 of no consequence therefore to the failure of the core barrel.

7 Now, for the UIS, you see the UIS loading only becomes  
8 relevant at a much later time, about 20 milliseconds. And that's  
9 after all these very early dynamic effects have really  
10 dissipated. At that time, as you see, the pressure here and  
11 the pressure on the UIS itself is very similar.

12 MR. ABDEL-KHALIK: I find this somewhat counter-  
13 intuitive. 1150 megajoules is about half a ton of TNT.

14 MR. THEOFANOUS: Yes.

15 MR. ABDEL-KHALIK: And to say that this will result  
16 in only a 2 percent strain is quite counter-intuitive. I  
17 realize that you have to take into account the fact that you  
18 have a very compressible fluid or compressible medium in there  
19 to allow for the voiding. But my concern then, should one  
20 worry about less energetic events with considerably less than  
21 1150 megajoules where you would not have such a compliant  
22 medium?

23 MR. THEOFANOUS: Well, first of all, let me make  
24 a remark about the intuition aspect of it. In this business,  
25 we cannot afford to go by intuition. If you like intuition,



1 you can come up to some very strange results. All I can say is  
2 that we have done that systematically and correctly and  
3 benchmarked every step of the way.

4 Now, what I am saying is that if this was a medium  
5 with a void fraction of something less than -- less than 10  
6 percent, maybe 5 percent -- well, in fact, even less than that  
7 -- it has to be essentially a pretty solid system in order to  
8 transmit and transmit loads directly. And by that time, you  
9 should have not enough energy.

10 You see, even at the \$200 per second level, you build  
11 up pressures of a few hundred bars for a very, very short time,  
12 and that is 2500 megajoules of energy that you put in there.  
13 If you take events less than \$100 per second, it will be of no  
14 consequence at all.

15 MR. ABDEL-KHALIK: If one were to do a thought  
16 experiment, and let's look at the different combinations that  
17 would result in 2 percent strain, and you are saying that one  
18 of these is the case that we have here, 1150 megajoules with  
19 highly voided system --

20 MR. THEOFANOUS: Well, it is not highly voided. Okay.  
21 It is not highly voided. We are talking about the void  
22 fraction of about only about 40 percent, 40-50 percent.

23 MR. ABDEL-KHALIK: Completely solid system, how much  
24 energy would I have to have in order to produce 2 percent  
25 strain? Is it 100 megajoules, 50 megajoules, 10 megajoules?

1 MR. THEOFANOUS: In terms of the equivalent amount of  
2 energy that we are talking about here, I think that to produce  
3 2 percent strain, you would have to have probably, I don't know,  
4 I am only guessing, 600.

5 MR. ABDEL-KHALIK: I was going to say I think the  
6 REXCO application gives you basically that kind of an answer,  
7 and it's about a factor of 2 lower.

8 MR. THEOFANOUS: Yes. 600 megajoules.

9 MR. BELL: So that if you have a completely  
10 incompressible fluid in there, you don't have -- you have  
11 liquid sodium in the channels, you would get 2 percent strain  
12 if your energy release is half of 1150 megajoules. Is that  
13 what you're saying?

14 MR. CARBON: Would you give us your name?

15 MR. BELL: Charles Bell, Los Alamos.

16 MR. THEOFANOUS: Really, that's what the applicant  
17 has done. They did not take into account the losses because  
18 of the compliant nature of the systems. So that whatever  
19 pressure was calculated to be right here, they're pulling it  
20 up to the boundaries. In fact, not only that, but they put it  
21 up into the pool also. I don't think it's really too much  
22 counting on intuition. I don't think it comes to mind.

23 MR. FAUSKE: Hans Fauske. I think it's also  
24 misleading to try to compare this kind of situation with TNT.  
25 It's very different. In the TNT equivalent, you're talking

1 about very high shock waves and the damage potential as a  
2 result can be very different. Your intuition, if you go into  
3 the TNT, may indeed be correct, but that's not applicable in  
4 this case.

5 MR. ABDEL-KHALIK: I don't have a pressure history  
6 here to compare with, so all I had was total energy release.  
7 The time scales are different than I understood. The  
8 response would be different.

9 MR. FAUSKE: They would be very different. You are  
10 talking about very different pressure levels.

11 MR. THEOFANOUS: All right. So I think that I was  
12 here in the process of describing or discussing how this  
13 calculation was done did not take into account the venting that  
14 would occur as soon as the UIS begins to move.

15 Here we go to the energy partition. There is a  
16 couple of 14 cubic meters of expansion taking place in the  
17 short term. And this corresponds to -- under the pressure curve  
18 -- to 520 megajoules. And that is the worth potential  
19 according to thermal dynamics, pocket-by-pocket maximum worth  
20 potential.

21 And you look at the results of the calculation now,  
22 the numerical calculation, and you see that 100 megajoules has  
23 gone into a strain in the core barrel and 340 megajoules has  
24 done into this, not by a mass -- remember now how we model that  
25 -- not by a mass of the upper core structure. It's moving

1 upward with some velocity as in a mass.

2           The -- not much difficulty, I guess, in understanding  
3 what this means -- the considerations concerning this number  
4 may be a little more entangled here. The way we approach it  
5 is that this kinetic energy is inside the bottle, so to speak.  
6 It cannot manifest itself, certainly not to the head, and even  
7 less so to the pool.

8           This kinetic energy will hit the UIS bottle. All  
9 right. Now, a lot of that energy is going to go into crossing  
10 the structure itself. Remember that this we did not take into  
11 account because we did basically remodel that as a fluid;  
12 the whole construction was a fluid and was allowed to basically  
13 squeeze out the void and come closer and closer together  
14 without any dissipating characteristics of that thing as it was  
15 doing that.

16           So some of that then is going to go into strain energy  
17 with that structure itself as it compresses and crosses into  
18 a solid mass. Some of it is going to be absorbed in the UIS  
19 cones. Some of it, in fact, might be even transmitted to the  
20 head as some load. But those loads are much lower mechanically  
21 than the capability of the head. They are coming in at a very  
22 different time frame. And again, I think, to visualize that,  
23 let me just show you the time frame here.

24           This is the time at which the UIS is impacted and  
25 begins to feel the force and therefore begins to move. Now, it

1 takes time to accelerate, of course. All the mass takes time  
2 to overcome pressure and accelerate. And during this time,  
3 where the venting process and the formation of the bubble begins,  
4 as you will see in a minute. The bubble hits. At about this  
5 time, assuming as you will see in a minute that the UIS has  
6 failed instantaneously and was displaced instantaneously to  
7 its maximum position when the bubble hits. That is for the  
8 purpose of estimating conservatively this complicated process  
9 from the point of view of getting maximum slug impact energy  
10 at that time.

11 So, in reality, in fact, the bubble will hit somewhat  
12 later, maybe 50-60 milliseconds. So the time at which the head  
13 is going to see this impact kinetic energy is around that time.  
14 The head is going to be loaded by the slug at around that time.  
15 A significant time difference. So it's just like you take that  
16 energy and give a little kick to the head before, and you wait  
17 and give it a real big burst later from the slug.

18 MR. ZUDANS: Okay, Theo, but the real question in my  
19 mind is how much this little initial hit is, because this is  
20 where the problem lies. It's 350 megajoules.

21 MR. THEOFANOUS: Are you saying the head --

22 MR. ZUDANS: Hold on. Hold on. 340 megajoules. Some  
23 of those megajoules, although you call them all losses, but  
24 they are not, some of those megajoules would be absorbed in  
25 crashing the upper core structure, some of them. Some of it

1 will be absorbed in crashing the support columns. The rest of  
2 it will be compacted mass in the form of kinetic energy flying  
3 upward. And there's nothing else to stop it but the head.

4 MR. THEOFANOUS: Well, no, really what happens then  
5 is an exchange of energy. And in fact, we did the calculation  
6 and you can see this by the problem, for example, we can  
7 completely neglect the strength of the columns, find out what  
8 would b the kinetic energy of the UIS if it was a free body hit  
9 by that 340 megajoules. They didn't talk about the very  
10 different masses involved. Okay. And we did that, and we  
11 found out -- I forget the number, Charlie, do you remember the  
12 number?

13 MR. BELL: 80.

14 MR. THEOFANOUS: It was 80. The kinetic energy of  
15 the UIS.

16 MR. ZUDANS: What was the --

17 MR. THEOFANOUS: Because of that exchange of energy.

18 MR. ZUDANS: Well, you call this 340 kinetic energy  
19 of the UCS.

20 MR. THEOFANOUS: That is the upper core structure.  
21 What's that?

22 MR. ZUDANS: All right.

23 MR. THEOFANOUS: No, what was it hits this? That's  
24 a very heavy mass.

25 MR. ZUDANS: All right.

1 MR. THEOFANOUS: You see that in the limit. I think  
2 it's a very unrealistic way of seeing it, very conservative.  
3 You can see that it's a two-body problem. One must --

4 MR. ZUDANS: That's fine but it's still an energy  
5 conservation problem.

6 MR. THEOFANOUS: Right.

7 MR. ZUDANS: 340 unless you dissipated something  
8 in terms of inelastic deformations or something in the fluid;  
9 it's still there whether it has a low velocity or not, whether  
10 it's a large mass or a small mass.

11 MR. THEOFANOUS: Well, it really depends. And I  
12 think what I am saying here -- I don't know what you're driving  
13 at -- but what I am saying is that we have this 340 megajoules.  
14 That is calculated as the maximum kinetic energy in this  
15 structure, neglecting its own dissipating characteristics, its  
16 own own class of characteristics.

17 Now, a portion of that, as I said before, is going  
18 to go into strain energy, into just what it takes to make that  
19 into a bundle, a complete mass. We don't know how much that  
20 is because it is not easy to calculate.

21 Secondly, another portion of that is going to go  
22 to the cones, and some residual portion is going to go maybe,  
23 if that is to fail -- but if that fails, remember, that fails  
24 up here and it fails gradually because of the quasi-static  
25 levelling of the pressure. That, if you took only that mass

1 hitting the head, that UIS was not going to fail. Okay.

2 So a portion of that then only is left to add to the  
3 top of this pressure level as it tries to displace the UIS  
4 and fail it by buckling the cones. All right.

5 What we are saying is really it is of no consequence  
6 at all because we are going to fail the UIS anyway. Now, I  
7 think it is also clear that you can't fail the head by pushing it  
8 from the bottom through those four cones. There's no way they  
9 can do that. So the most that you have is the cones to fail.  
10 Now, if that was to happen, another way of looking at the  
11 problem is see it as a two-body problem, forget about the  
12 crashing, which I think is extremely significant of this and  
13 the crashing of that, just take the 340 megajoules and do a  
14 two-body problem. Even if you did that you don't end up with  
15 much kinetic energy on the UIS. By the time the bubble forms,  
16 this will begin to see forces from above also that will tend  
17 to make it all disappear.

18 MR. ZUDANS: I guess I begin to see what your  
19 reasoning is. In other words, there will be a number of  
20 interactions with the UCS and UIS as the elastic plastic mass  
21 and then some of the energy through these interactions would  
22 be either transferred to through the sodium or lost otherwise,  
23 but it will never be available as a net kinetic energy to move  
24 it end to end.

25 MR. THEOFANOUS: That's right. The important point



1 here is those interactions are happening in this time frame.

2 MR. ZUDANS: Okay.

3 MR. THEOFANOUS: And the head is going to be really  
4 loaded from the sustained pressures accelerating the sodium  
5 slug is going to be all the way up here at a much later time.  
6 So they are not additive.

7 MR. ZUDANS: All right.

8 MR. THEOFANOUS: All right. So I guess we covered  
9 that.

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1 Long-term expansion will take place because of  
2 the buckling of the UIS cones; a bubble will form. The pressures  
3 in the core are going to be much higher than ever possibly  
4 you could have in the bubble because it takes pressure to  
5 accelerate the material to produce the flashing, and that  
6 flashing is what is keeping the pressure in the bubble up.

7 As a result of that, we find the typical kind of  
8 thing that you see in experiments and is common knowledge;  
9 that the pressure in the bubble will be about one-half of the  
10 pressure in the core. Just doing a simple PV expansion work,  
11 you can see then that just due to this mechanism that we call  
12 throttle effect, we have a reduction by a factor of 2 of the  
13 potential kinetic energy of the slide at time of impact.  
14 That is an effect, also, that the applicant has not taken into  
15 account.

16 The calculation was done assuming an adiabatic  
17 bubble. That is, we did not take into account any augmenta-  
18 tion of those pressures because of sodium entering the bubble,  
19 introducing additional pressures. We have looked into this  
20 problem by means of calculations, sensitivity calculations,  
21 in which we put sodium coming in from the pool in the bubble  
22 to build up pressure. But when that happens, this pressure  
23 builds up but then the expansion slows down somewhat, there's  
24 some kind of a cancelling effect and you don't really get  
25 much of an effect.

1           Furthermore, what we know from experiments --  
2           and there are a couple more experiments done on this --  
3           when you blow down a flashing fluid into a pool that is  
4           highly subcooled, as happens, for example, by flushing hot  
5           water into cold water, there's a lot of mixing going on, a  
6           lot of entrainment and a lot of heat losses, as you might  
7           expect, so that the actually water that you deliver in terms  
8           of slug energy is much, much less than even this factor of 2  
9           that you see up here.

10           On the other hand, however, we also know from  
11           experiments that if the pool is volatile, this volatility of  
12           the pool can interfere with the mixing from the process, and  
13           the results look much more like adiabatic results. In fact,  
14           we have a couple of experiments that we have run in which we  
15           used freon here and not water here. And we found out that  
16           the results are exactly adiabatic.

17           So it's on this basis, then, that we did a calcula-  
18           tion on the best estimate basis, and we think also it's  
19           conservative in terms of the adiabatic bubble.

20           Now then, let's see what happened to the rest of  
21           the energy. We are talking about a delta v expansion after  
22           the 14 cubic meters, as happened here in the short term.  
23           Now we have another 15 cubic meter expansion that is the  
24           covered gas volume. This 15 cubic meters, if it was done  
25           isentropically according to thermodynamics, should give us 160

1 megajoules. This number is 160 megajoules. Well, we got  
2 in the calculation 80 megajoules.

3 Another point that has to be made, both during  
4 the short-term as well as the long-term expansions there is  
5 other heat transfer processes in the losses that we think are  
6 present, and those result because of the interaction with hot  
7 materials with cold steel that is present both here, as well  
8 as in the upper core structure. These losses are significant  
9 and if one actually had done some similar calculations they  
10 would show that one might be talking about another reduction  
11 maybe by a factor of 3 or 4 in energy releases. But those are  
12 a little more difficult to take into account, and people tend  
13 to ask a lot of questions about them because we don't like to  
14 be questioned and we're not going to talk about that. And we  
15 are not basing our results on these mechanisms.

16 Now then, to summarize all this situation, here we  
17 have the structural capability of the vessel head structure  
18 in particular, kinetic energy versus reactivity rate. Again,  
19 this is for two-phased disassemblies. What we are showing  
20 here are the two limits I just talked; one limit is the 175  
21 to 200 megajoule impact energy in the head, and we think that  
22 is the -- well, that is the limit extended by the project to  
23 the design of the head. And this is the 100, 275 megajoule  
24 with some uncertainty for the threshold for finding the UIS.

25 Now, already we discussed that unless this fails,

1 there is no way that we can release any energy in the sodium  
2 pool. And that is shown here by this green line, that up to  
3 this kind of level of energetics we have really almost no  
4 impact energy to speak of. However, as soon as the failure  
5 point is reached, then there is a jump because now suddenly,  
6 the minute the cones buckle, you're going to get the bubble,  
7 there'll be some release, and that release may be close to  
8 50 megajoules. And from then on, gradually this increases  
9 with the ramp rate, and the other point we have made is this  
10 calculation of this point, which is a 200 dollars per second  
11 approaching the 75. In fact, for that one case that we did  
12 under adiabatic conditions, we got exactly 80 megajoules  
13 for this 200 dollars per second case.

14 Therefore, what this implies is two things. Number  
15 one, if one was to expect ramp rates releasing energies  
16 equivalent to less than this 100 dollars per second, there  
17 would be no structural damage to the head at all.

18 Furthermore, one concludes from that that in order to  
19 cause structural damage you don't need 110 or 120, but there  
20 is still another very big significant margin on top of it that  
21 the vessel can take. We think that this is completely outside  
22 the realm of possibility, but it is good to know that it has  
23 also got that additional margin on top.

24 So now, then, once we go through the rest of it,  
25 we should be keying our thoughts to, number one, to this

1 number, and number two, remembering that there is also this  
2 margin on top.

3 MR. KASTENBERG: Theo, do I interpret that graph  
4 correctly if I assume that the green bar means you're somewhat  
5 insensitive to the ramp rate between 100 and 200 dollars a  
6 second in respect to impact kinetic energy? Or is it just the  
7 way you've drawn it?

8 MR. THEOFANOUS: Well, this is a qualitative graph,  
9 but I think that this is true, there is some -- there is not  
10 a very great sensitivity, put it that way.

11 MR. KASTENBERG: Can you give us a physical feeling  
12 for why?

13 MR. THEOFANOUS: No, this has already been taken  
14 into account. This is not taken into account here. We take  
15 that in the case of when the UIS fails. After the UIS fails,  
16 we assume it to be completely displaced to its maximum position  
17 and really do the expansion process.

18 MR. ZUDANS: Theo, the transition is if the UIS fails.

19 MR. THEOFANOUS: That's why the transition. That's  
20 why you see that great sensitivity up here, so if you actually  
21 wanted to try it without the UIS, it would be more of a kind  
22 of smooth curve. Maybe that's why -- .

23 MR. KASTENBERG: I see.

24 MR. THEOFANOUS: Now we are going to go through  
25 the accident sequences. We are going to begin with the

1 initiating phase and walk through all the terminology and  
2 the various stages until we come to the termination. Each  
3 step of the way, we are going to be concerned about one thing,  
4 and that is, what is the potential for generating ramp rates  
5 in two-phase systems on the order of 100 dollars per second.  
6 That's the question we want to answer in every step of the  
7 way.

8 MR. ZUDANS: I'd still like to return back to the  
9 previous -- I'd just like to tell you what bothers me, and  
10 maybe if you think about --

11 MR. THEOFANOUS: There's still something that bothers  
12 you?

13 (Laughter.)

14 MR. ZUDANS: Oh, yes. You see, your argument about  
15 the two-body interaction seems to hold water until the UIS  
16 phase.

17 MR. THEOFANOUS: You mean, there's no interaction  
18 after it fails?

19 MR. ZUDANS: I don't think so, because there's no  
20 reason for the motion to reverse. Once it began to go up --  
21 for the both masses, UCS and URS, will follow up until they  
22 compact the columns completely.

23 MR. THEOFANOUS: Right. We aren't saying that this  
24 occurs.

25

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1 MR. THEOFANOUS: I know. Sure. We are saying inter-  
2 action is not only this way. It can be like this. The  
3 question if there is a big body and there is another smaller  
4 body, and it comes up and hits it. The question is how much  
5 energy it can transmit to that big body.

6 MR. ZUDANS: Inelastic impact.

7 MR. THEOFANOUS: Not all of it because of the  
8 process of transmission.

9 MR. ZUDANS: That's what I said.

10 MR. THEOFANOUS: It goes and hits a car. The tree  
11 doesn't go down the road. The car crashes.

12 MR. ZUDANS: All of the energy will go into the  
13 combination of the bodies.

14 MR. THEOFANOUS: That's what I'm saying. One goes  
15 into dissipation in the first body. The other part goes into  
16 buckling the columns. The residual will go into the kinetic  
17 energy of the bodies.

18 MR. ZUDANS: And that is the residual that I am  
19 concerned about. How much is that?

20 MR. THEOFANOUS: The maximum of that can be 80  
21 megajoules.

22 MR. ZUDANS: Eighty?

23 MR. THEOFANOUS: Eighty.

24 MR. ZUDANS: How did you arrive at that figure?

25 MR. THEOFANOUS: Just taking a big mass and a small



1 mass and hit it and just the two move together.

2 MR. ZUDANS: Where the energy would dissipate in the  
3 buckling of the UIS, can that be computed? You could dissi-  
4 pate some in the sodium because there is a flow, but there  
5 would be significant piece of 340 that would be of kinetic  
6 energy of two bodies, and it will continue to crash the  
7 columns against the head, and that could be substantially  
8 more significant than that you get through the buckle.

9 MR. DICKSON: Conservation of momentum. If you have  
10 two bodies in collision, a small body colliding purely  
11 elasticity with no loss of energy, they reverse themselves  
12 to conserve the total outward momentum. The kinetic energy that  
13 is distributed between them is conserved, but it's not in the  
14 same direction.

15 MR. ZUDANS: That's correct, but there is no way  
16 for this other -- the smaller body to go back, because it  
17 has a pressure load that is accelerated in the first place.

18 MR. DICKSON: But if you do the calculations con-  
19 serving momentum and then worrying about where the kinetic  
20 energy went, you find you have to worry about where a lot  
21 of kinetic energy went when the little body hits the big  
22 body. And that has to be conserved, independent of whether  
23 you are making it an idealized two-body problem or not.

24 MR. BELL: Well, perhaps. I do agree. Whether  
25 you do the inelastic two-body problem, the final kinetic

1 energy of those two bodies moving together, it is also  
2 about 80 megajoules. In other words, the 340 is reduced to  
3 80.

4 MR. ZUDANS: Okay.

5 MR. BELL: That 80 is continuing to move up.

6 MR. ZUDANS: Okay. And that would be helped by  
7 the head earlier than the pressure --

8 MR. BELL: No, no. It is removed several meters  
9 from the head.

10 MR. ZUDANS: How?

11 MR. BELL: Because of its initial location. It has  
12 to transfer other than what could be transferred directly  
13 through the columns which are buckling, and that is fairly  
14 small.

15 This kinetic energy is in the body, and it's  
16 located I think on the order of four or five meters down  
17 below the head, so by the time of a few meters -- I don't  
18 know, 20 meters per second -- by the time that moves through  
19 the distance to hit the head, the pool has already been  
20 accelerated to hit the head long, long before that, simply  
21 because the pool only has to move two-thirds of a meter,  
22 and this other object down deep in the pool has to move four  
23 or five meters.

24 MR. ZUDANS: That sounds reasonably all right.

25 MR. BELL: So you effectively have a staged impact,

1 if you will. The pool hits, and if this body were to continue  
2 to move up, it would impact the head after the pool impact  
3 has taken place.

4 Let's go back to the other article where a small  
5 body impacts a big body. The upper core structure is supported.  
6 As soon as it feels the pressure, it transfers over to the  
7 big body?

8 MR. THEOFANOUS: No, no, no.

9 MR. ZUDANS: No?

10 MR. THEOFANOUS: This upper core structure is a,  
11 you know, is the fission -- it is a very porous thing. It's  
12 supported, but you cannot -- you cannot fail the columns here  
13 by crashing the fission gas plenum.

14 MR. ZUDANS: I am not saying you can. I am only  
15 against the argument that the small body hits the big body.  
16 There is no such physical process.

17 MR. THEOFANOUS: That's why I think we got the thing  
18 with this problem in this context.

19 (Slide.)

20 This really is the 340 number. We did that, by the  
21 way, only to account for the total, and maybe it wasn't a  
22 good idea.

23 MR. ZUDANS: It certainly wasn't.

24 MR. THEOFANOUS: But I did talk about it.

25 MR. ZUDANS: Yes.

1 MR. THEOFANOUS: -- Is that this is as you say,  
2 is in contact.

3 MR. ZUDANS: All right.

4 MR. THEOFANOUS: However, this is a flimsy structure  
5 compared to this. This is a very high load. You cannot keep  
6 this on the form, and this to start filling the loads. Really  
7 what happens is that this crashes here. It crashes only to  
8 the extent that it can transmit forces, but its own very  
9 flimsy nature.

10 MR. ZUDANS: Right.

11 MR. THEOFANOUS: And that is not until all of the  
12 walls were squeezed out of it and it becomes a compact  
13 mass. Therefore, we -- and in this number there is no portrayal,  
14 of course, of energy absorbed in crashing that. That is why  
15 it is wrong to think of a body with 340 megajoules and hits  
16 another body here.

17 What happens is that during this transmission process,  
18 20 milliseconds, this is crashed continuously. That energy  
19 is the equivalent of 340 megajoules. That is being absorbed  
20 here, and it will reach the fully crashed state not with a  
21 zero velocity, and that is the kinetic energy that is  
22 concerned with from the point of view of adding up to the  
23 top of this.

24 There is a loading, and this additional kinetic  
25 energy of that structure will be negligible. We said forget

1 about all of this to satisfy your question.

2 Now, what is the most we can get? The most is 80,  
3 and that can appear -- the velocity of that, since it fails,  
4 but meanwhile remember that that is doing work. It fails  
5 because of this loading, and that hit will be moving upward;  
6 and typically it will have been displaced to something like  
7 maybe 3 feet, like 2 to 3 feet by the time we are in this  
8 time frame, so now you figure this out.

9 The upper internal structure moves with this  
10 velocity, but only 3 feet away from its original position.  
11 The head -- the sodium already has hit the head at that point  
12 and hits -- and it bounces from that. Not all of it will go  
13 in the head and comes out and hits it, displacing the UIS  
14 and comes back a wave. I don't think one should waste more  
15 time on this problem.

16 MR. ZUDANS: If that is the way things work out,  
17 that makes sense because you could observe it in the sodium  
18 and never see the head.

19 MR. THEOFANOUS: Sure.

20 MR. ZUDANS: Okay.

21 MR. THEOFANOUS: I don't see why it would work  
22 any differently.

23 MR. ZUDANS: Well, because of your statement. It  
24 says 340 megajoules in kinetic energy of that mass, and it  
25 has to go some place.

1 MR. THEOFANOUS: I am accounting for it.

2 MR. ZUDANS: Now, you are accounting for it? Let's  
3 see. The little piece that you showed there, it's connected  
4 to the top piece that you -- that is closed.

5 MR. THEOFANOUS: Yes.

6 MR. ZUDANS: There is no gap, so you begin to expand  
7 your interfering volume. There is a little bit of gap. I  
8 don't know what you consider the end of that, but, in fact,  
9 the end -- they are not smack up against the big plate, so  
10 that will translate -- that will move a little bit before it  
11 begins to stop, therefore begin to crash.

12 MR. ZUDANS: And in the process of this first phase  
13 that you described, the blue piece which is in the upper core  
14 structure collapsed. It's most of the time supported against  
15 the upper part and in the part some kinetic energy on that --  
16 some of that will show up there. But I am buying your  
17 explanation that as it goes up in the sodium, doesn't move  
18 too far before the sodium begins to come back, and it will  
19 be in that fluid. That's okay.

20 MR. THEOFANOUS: All right. This interface,  
21 initiating phase phenomenon. The important physical processes  
22 here are sodium voiding; that is followed by clad relocation  
23 and eventually fuel motion.

24 You will remember from the last presentation we  
25 made that there is no way that one can precede energetic

1 events by moving additional fluids. On the other hand, they  
2 do have important consequences. They affect the power, and  
3 they are important because they set the stage as far as the  
4 fuel.

5           There are two different aspects that have to do with  
6 those processes which are the result of radial boiling  
7 incoherence, and because of the tilt and because of the  
8 presence of the walls. That causes a known one-dimensional  
9 boiling process which translates into a known one-dimensional  
10 clad melting and clad process.

11           We can calculate those things, and I talked about  
12 that. I didn't want to take time and talk about it here except  
13 the bottom line, that is on the boiling process. We feel  
14 that the 1D compilation correctly compiles the voiding,  
15 the microscopic voiding, both in time and in rate and, number  
16 two, the cladding location problem is going to be much, much  
17 accentuated.

18           As a result of that, we do cite compilations, take  
19 into account the melting coherency, and we see the effect  
20 of this plus the plenum fission gas. We believe that in  
21 the UC-3 and UC-4 cases there is some pressure for fission  
22 gas in the plenum. It will produce no net cladding location  
23 at all, and I really agree. That means that the core is  
24 going to go into the fuel motion stage with a total incomplete  
25 cladding.

1           Now, if ever of these two aspects -- that is,  
2 the -- if the fission -- if there is no fission gas to speak  
3 of, there could be some upward location, but that would happen  
4 only in a very, very early part of it, and it would be very,  
5 very minimal. And I don't want to go through the arguments  
6 because they really don't even pertain too much to the story  
7 because we don't want to count on anything like that in this  
8 stage anyway, but that it is important to keep that in perspec-  
9 tive.

10           Now, for the fuel motion, if we are interested to  
11 look at what possible energetic events we can get from that,  
12 we have to look for forceful motion, not just normal melting and  
13 that kind of running around. And the only thing that, of  
14 course, happens in a core like this is very low sodium voiding  
15 activity and not very high power to start with, is to look  
16 at fission gas pressures.

17           These pressures can manifest themselves in two ways.  
18 One is for the fission gases that are retained and they  
19 produce pressure as the fuel melts inside the pin, and  
20 that can only happen in the early part of conditions. And  
21 in this core, again, we don't expect that to happen under  
22 loss.

23           However, the other way that they can manifest  
24 themselves is to during the early disruption stage of the  
25 fuel, let's say there's pressure on the top, it is accumulated,



1 so those are the areas in which these considerations enter,  
2 so it's because of this then that one is interested to know  
3 the fuel failure mode and how is actually the fuel going to  
4 fail in any certain phase.

5 Well, one can obtain this result by carrying out  
6 system calculations, and to cover the range of the activity  
7 effects, we have gone all the way from very, very accentuated  
8 feedbacks like saying the sodium --

9 (Slide.)

10 -- Then taking the least possible doppler and all  
11 of those things into account, and you can make it slow but --  
12 very slow activity and increase the negative ones.

13 Here is an example of a slow loss of flow accident  
14 for the end of Cycle 4 configuration. This is for the purpose  
15 of determining the potential for separating out the steel from  
16 the fuel. The slower the accident becomes, the more time  
17 there is between the clad melting and fuel disruption, and  
18 as long as this time becomes larger and larger, there is more  
19 opportunity for it to -- well, I made the statement before  
20 that because of radial incoherencies and because of the  
21 plenum fission gas blowdown that interferes, that we don't  
22 accept. Even the calculations we wanted to really -- could  
23 be conservative, and really there are several points, and  
24 they are not appropriate for a presentation, but we have  
25 them all in the report. We just thought that we would allow

1 the calculation to move the cladding, that some of the things  
2 are reasonable in the upward direction, although we don't  
3 expect that to be the case.

4 We find even if we did that, because of the activity  
5 it needs and the power, very little time for the cladding goes  
6 to move even if you allow it move, and that is the process  
7 whereby the power increases enough to disrupt the fuel, and  
8 as the cladding mixed up well enough with it to cause a core  
9 disruption.

10 We referred to that as a core disruption of fuel  
11 and cladding. That is a result of the power increases, and  
12 it is observed in phases where the activity feedbacks were  
13 chosen so as to make the overall accident progression slow.

14 This figure contains all the important aspects of  
15 the initiating phase. We have plotted here the time, and  
16 here is a group of subassemblies grouped together from the  
17 point of view of doing thermal hydraulics, and the yellow  
18 line is sodium boiling. The green is clad melting. This  
19 is clad motion, and this is fuel motion.

20 Looking at the lines straight up this way, you can  
21 find out at what time boiling started and how much later  
22 cladding melted, how much later cladding began to move, and  
23 how much later the fuel began to disrupt.

24 By looking at a line across this way, at any time  
25 you can actually have a visual effect of what the core is

1 doing. You find out that some of them haven't boiled yet.  
2 You can look at 17 seconds. This part up here just began to  
3 boil. This has boiled earlier, about a second and a half  
4 earlier, and it begins to move the cladding, and there is  
5 no fuel flowing anywhere.

6 What is the state of the core during this phase?

7 I have shown here with arrows the way by which you take these  
8 terms of blockage. There is enough time. There is enough  
9 time between the melting of the load and the disruption of  
10 the fuel so that the cladding can separate by motion.

11 To figure out at what parts of the core this is  
12 possible, basically you are looking at differences between the  
13 fuel disruption and clad melting of the order of more than  
14 a quarter of a second. If there is more than a quarter of  
15 a second time there for -- conservatively, one might claim  
16 there is blockage. Of course, it will be incomplete. They  
17 are self-limiting because it's like a self-limiting situation.

18 The stream that causes the upward location also  
19 is cut off by the blockage formation, so before you have  
20 a complete blockage, a completely zero streaming, zero loss  
21 of the vapor, the flow is going to be reduced to the point  
22 where it is insufficient to levitate the cladding, and that  
23 will happen well before the core exit of the subassembly is  
24 blocked. So at worst you have some parts of blockages, and  
25 you can visualize in terms of the timing to form that.

1           So you can see here that suddenly you have -- here  
2 is 6 -- this is the first to enter this disruption phase, so  
3 here you say well, maybe there is some blockage here. Maybe  
4 some -- a little bit more, a half a second, but there is no  
5 blockage here and not in this other -- where the cladding  
6 just melted every time that whole core disruption because  
7 of fuel motion occurs.

8           So this core disruption is promoted by increased  
9 power and by fuel motion that results from that, and it is  
10 from that point on the whole accident is controlled by the  
11 fuel motion process itself, so that from the moment the  
12 fuel begins to move, the cladding activities play a relatively  
13 minor role.

14           I want to show you a picture of what that core  
15 disruption looks like, and this is a schematic.

16           (Slide.)

17           We have a result of a computation here, and you  
18 can see that this is slanted lines. This represents liquid  
19 fuel. This is the actual condition, and here is fractions  
20 of material that are in different states, and those states  
21 are -- the key for those states is given here.

22           The important thing is that for the active core  
23 we have essentially all the fuel melting at this time. All  
24 this is melted steel also, so you can see there is a real  
25 intimate mixture of melted steel, molten steel. The boiling

1 point of steel and the determination will provide higher  
2 vapor temperatures, so the presence of the two together, the  
3 steel produces the vapor pressures, so the steel around. If  
4 no blockage is formed, then you can see here the cladding  
5 did not have time to move up there before it melted completely,  
6 before the fuel melted.

7 Now, we have that fixture really -- as is shown  
8 here. It is not operated steel. It is a mixture. Now, even  
9 if that mixture which is going to try to exit were to block,  
10 the blockages would not be in the blockages. They would be  
11 head-producing blockages, so the next time there is a power  
12 increase, the blockages will melt and open up.

13 MR. KASTENBERG: The ordinate on the bottom?

14 MR. THEOFANOUS: There is the mil fractions,  
15 fraction of material. For each of those things, if you  
16 look at a fuel element or a pool, you also might think in  
17 terms of a fuel pin or subassembly. It will be the same.

18 Here is the picture of the end of Cycle 4, increasing  
19 the limits of the activities for clad worth and so on. What  
20 you see here is only channel 6, and it allows maybe half a  
21 second. Under these conditions we didn't think so, but that  
22 is all.

23 See, in all those cases, in all those, more than  
24 half of the core, there is just a very small fraction of a  
25 second available between the melting of the cladding and the

1 fuel disruption, so that all this clearly is going to core  
2 disrupt and mix all together.

3 MR. KASTENBERG: Six is a high-powered channel?

4 MR. THEOFANOUS: That is true in all the even-cycle  
5 cores. The odd cycle, they have replaced it by blanket, so  
6 in that case that is not.

7 MR. KASTENBERG: I think you said before that in a  
8 high-powered channel you expected core disruption.

9 MR. THEOFANOUS: No, no, no. I didn't say that.  
10 I said because of the high power of the core, we bring core  
11 disruption, because as the power increases, this interval  
12 becomes smaller and smaller.

13 (Slide.)

14 Now, going back to looking for energetic events  
15 in this general framework of fuel motion, we identify two  
16 possible mechanisms. One is portrayed here mechanically. That  
17 is referred to as the LG-driven TOB. In the previous  
18 homogenous core it boosts the power early enough so that the  
19 very significant fraction of the core had not voided yet at  
20 a time in which the inner pin melting occurred.

21 Now, if the core is irradiated, this inner pin  
22 melting will produce high pressures. That is going to produce  
23 pin failure, and then that will produce an ejection of  
24 molten fuel into the sodium.

25 Now, depending where the fuel failure occurs, that

1 may be a good or a bad event. If this failure occurs outside  
2 the mid-range of the core, this ejection of the fuel and the  
3 motion of the pin will produce motion away from the high  
4 flexation and have negative effects, and therefore a shutdown  
5 situation.

6 Now, we have axial motion of the fuel inside the  
7 pin. That was the mid-plane, and that, of course, is a strong  
8 positive effect. Of course, as soon as the fuel comes out,  
9 it will have to get away from the assembly somehow; but this  
10 process has to happen first before that happens.

11 It is important also for the timing that this process  
12 happen between the different subassemblies in the compilation,  
13 but anyway, you can see that after it begins to happen, that  
14 will increase the power in the activity, and that will  
15 accelerate the time that the other unfailed subassemblies  
16 might enter this kind of picture. So you can see the  
17 tendency.

18 Again, if the failure is to happen in the core  
19 mid-plane or in that immediate vicinity there. Now, for  
20 the heterogeneous core, we don't have the sodium void activity  
21 strong enough to accomplish this kind of thing by itself.  
22 However, we have also clad motion and maybe some fuel motion.

23 You see that channel 6, and if that solid was to  
24 melt and collapse, that in itself might have given enough of  
25 an activity boost to produce this kind of situation before

1 completely voiding the core. So we assessed that situation  
2 by picking all, in every part picking the worst possible  
3 situation.

4 The sodium activity, that is the maximum. That is  
5 the normal plus the sigma. We took the fuel to be -- we  
6 took the lowest possible. All right. On the fuel we  
7 assumed that it was not very expensive, which would offer  
8 experiments when fuel is subjected to high power, which you  
9 expect if you approach this condition. It is very -- we  
10 took it to be very mild.

11 Even with all of that, one on top of the other,  
12 we find that that that situation, it is not out of the books,  
13 and you are not concerned with it. We thought that it was  
14 very important to establish that limit because that was one  
15 of the sticky points with the previous core; so we are very  
16 careful to kind of clear away from that.

17 Now, with that out of the way, we have no other  
18 mechanism for producing energetics in this case except for  
19 this mechanism of the plenum fission gases pushing on the  
20 fuel column, and therefore introducing fuel into the high  
21 flex area.

22 This happens because in an irradiated core the  
23 pressures can be very high, 30, 40 atmospheres. And as  
24 the accident progresses and the gases hit, they can become  
25 50, 60; so they are very, very high pressures. And as long



1 as the column is integral, of course, the pressures are  
2 balanced.

3 If you take away suddenly the center part, basically  
4 by melting it, and you see that this pressure will be unbal-  
5 anced, this cladding cannot go upward, and it will push it  
6 upwards. If that process was to happen unimpeded by the  
7 interaction between the pellets and the cladding, one is  
8 concerned by channel 6 doing that, and all of the rest of  
9 the channels will be joining in the process.

10 And so when we approached this problem -- and the  
11 Applicant has considered this probably -- we were very concerned  
12 about how to have catalytic behavior because of this  
13 phenomenon. As it turns out, it takes time before pellets  
14 can be accelerated, and even if we assume that there is no  
15 interference between the cladding and the pellets -- and many  
16 people really disagree with that, and they think -- I agree  
17 with them that it would be hard to push a lot of pellets, if  
18 you have seen some operation, through a very, very tight  
19 clearance cladding.

20 We do not have any actual evidence to say they  
21 will be subject there or it will not move at all, or that  
22 they will move at some small velocity. Therefore, we take  
23 the approach let us see how bad that situation can get. We  
24 carry out the compilation as before for the reference cases  
25 at the time of fuel failure. We allow the acceleration of

1 the pellets downwards, and from this it allows it to do that.  
2 Not all the while at the same time, and maybe like \$20 per  
3 second.

4 MR. KASTENBERG: Is this consistent with the picture  
5 of core disruption that you showed us that I asked the question  
6 on the previous vu-graph from the appearance that you have  
7 coherent melting in a channel over a large portion of the  
8 active core, and here you are postulating something which looks  
9 more like what you would expect in a TOT where you have cool  
10 clad because you have coolant there?

11 MR. THEOFANOUS: No, no. This doesn't -- this is  
12 not done to scale, and the time frames are different. In  
13 the case here, the initial melting of the fuel will be maybe  
14 just a few centimeters, just showing here, just compress  
15 this compressible region here and the possibility of getting  
16 pressure from above, right?

17 The other picture was shown on a different time  
18 scale which is many tenths of a second, and this is only a  
19 fraction of a second. Obviously, you don't expect to have  
20 melting of the fuel across the -- in the core pin instantaneously.  
21 There will be some small area disruption first. So the rest  
22 of it is going to be compacted, and after this process begins,  
23 then you are talking about going to millisecond time scale.  
24 And what the rest of the fuel will be doing of its own  
25 collision is a different story. From that point on the problem

1 is dominated by this process and not by anything else.

2 MR. KASTENBERG: Yes, I understand that. That is  
3 why I made the comment that if you were to believe that this  
4 were to occur.

5 MR. THEOFANOUS: Yes.

6 MR. KASTENBERG: This picture.

7 MR. THEOFANOUS: This picture?

8 MR. KASTENBERG: Then you may never get to this  
9 other state.

10 MR. THEOFANOUS: Yes, sure.

11 MR. ZUDANS: It starts to act like a TOP.

12 MR. THEOFANOUS: If it happens, you will never enter  
13 the state. In fact, you will probably have an energetic  
14 behavior. Again, if the failure -- if the pellets were  
15 allowed to move again, I think that is an important qualifica-  
16 tion of that, and we only assume that because we have other  
17 evidence.

18 MR. KASTENBERG: Do you have a picture in your  
19 packet or in the report which shows the temperature profile  
20 as you approach melting?

21 MR. THEOFANOUS: The temperature profile along  
22 the pin?

23 MR. KASTENBERG: Right.

24 MR. THEOFANOUS: That gives you a feeling that you  
25 can have such a vocalized phenomena, because I tend to think

1 of it as you have more uniform heat operation. That's true.

2 MR. KASTENBERG: And this tends to say that you  
3 don't, that you have more of a gradient along there, and that  
4 space in the middle of the core you melt and start a little  
5 disruption, and it leads you on this path.

6 MR. THEOFANOUS: Well, I am only at this point. I  
7 am looking at this picture here, Bill.

8 MR. KASTENBERG: That's what I am saying. Is this  
9 consistent with what?

10 MR. THEOFANOUS: It is consistent because it never  
11 fails the whole pin at the same time. It will be some portion  
12 of the pin that is going to fail. It is not coherent over  
13 the whole length of the pin, and I don't care how much of it  
14 fails at the same time. At a very short time it begins to  
15 become soft. There is no cladding. As soon as you have  
16 that fuel becoming soft at some spot, it will begin to fill  
17 in compaction, and it is controlled by this process only.

18 Of course, the other important aspect of it is that  
19 there is a power distribution across the floor, the inter-  
20 subassembly. We don't expect that to happen all at the same  
21 time. If it were to happen all together, it would have been,  
22 of course, a different situation, a different story. But  
23 because of this intersubassembly coherencies, because there  
24 is not enough time for different failures to come into the  
25 picture and have time to accelerate and add their activity

1 to that whole core, begins to act coherently, so only a few  
2 of the subassemblies undergoing that. And before the other  
3 end of the picture there is never vapor pressure developed  
4 that we have this assembly and the whole thing is over.  
5 However, there is one --

6 Yes?

7 MR. MARK: You said the cladding was already?

8 MR. THEOFANOUS: In this particular tunnel.

9 MR. MARK: Only in the place that red is?

10 MR. THEOFANOUS: No. Only --

11 MR. MARK: If the cladding is gone and you don't  
12 have any friction moving the pellets?

13 MR. THEOFANOUS: Well, you have friction in the  
14 blanket area. The cladding is only going in the active core  
15 region.

16 MR. MARK: Okay.

17 MR. THEOFANOUS: This gas pressure has to transmit  
18 itself -- has to be transmitted by pushing the blanket pellets.  
19 So just like a piston, the blanket pellets come in and  
20 push the rest of it down.

21 MR. CARBON: In your calculation do you only allow  
22 for the initial?

23 MR. THEOFANOUS: Right. We asked the question to  
24 the project, and they told us that in fact it is possible  
25 that some of the fission rods that have migrated might even

1 make like an easy sliding condition.

2 MR. CARBON: So you assume that the cladding is  
3 gone in the entire active region?

4 MR. THEOFANOUS: We assume that cladding is gone in  
5 the active region. We don't assume anything. We are just  
6 calculating according to the calculations, and they tell us  
7 what is gone.

8 MR. CARBON: But when it comes to calculating the  
9 rate at which the pellets are pushed into the center to give  
10 you your reactivity experience --

11 MR. THEOFANOUS: Only on the basis of the pellets  
12 that have not disrupted. In reality, there will be significant  
13 interaction here in the blanket period. So you won't be  
14 able to transmit this force in the velocity downward. We  
15 look at it in a very pessimistic, limiting kind of condition.

16 MR. CARBON: When you calculate the pellets in the  
17 upper part of the active region moving toward the center do  
18 you allow them to move out into the channel, or do you assume  
19 that they stay in the state column? And my reason for  
20 asking is it seems that it has been moved out into the  
21 channel. Then you would squeeze more of it in a hurry.

22 MR. THEOFANOUS: But now the whole calculation is  
23 in a one-dimensional sense. We can't do that in two dimensions,  
24 so whatever pellets we have, they can only move downwards,  
25 not in a radial direction. However, as you try to visualize

1 subassembly, is a very, very tight diameter, there isn't  
2 much of a space for things to move over in radial directions.  
3 I think the real -- if you ask me, the real conservatism here,  
4 if you wish, is not allowing any interaction between the  
5 blanket pellets and the cladding, which is true, and that is  
6 going to be -- many people feel it will be significant, and  
7 many people in fact don't like the idea of analyzing this,  
8 and see the pellets can never move, but we like to see somebody  
9 prove that in a reasonable way, and that we don't see because  
10 there are no experiments at all with this kind of -- for this  
11 situation.

12 MR. CARBON: You spoke of the different times at  
13 which from pin to pin or subassembly to subassembly. What  
14 kind of time differences is there? You haven't given any  
15 numbers. What do you mean? What are the magnitudes of time?

16 MR. THEOFANOUS: I think you are asking about  
17 corewide incoherencies, and that gives you an idea of the  
18 difference in times. If you look at it this way, you can see  
19 by how much different parts of the core lacking other parts.  
20 That is for the fast case. That is aggregating all positive  
21 effects and degrading all negative effects. You could find  
22 that between the channel 6 when this reaches a point of  
23 failure of the fuel, and that would be 50 percent melt fraction.

24 So the changes were time, and it becomes more and  
25 more coherent. However, the process of concern also becomes

1 shorter and shorter in time, so that coherence throughout  
2 the scenario is important.

3 MR. ZUDANS: On this section, how long is time  
4 compared to the core where you have one that is no longer  
5 existing in the core? By the time you reach that boundary,  
6 it will discharge?

7 MR. THEOFANOUS: What you are thinking here, yes,  
8 I understand the question. This blanket area is 30 centimeters,  
9 and you are talking about movement here. They are a fraction  
10 of a centimeter.

11 MR. ZUDANS: I see.

12 MR. THEOFANOUS: Before the thing disassembled. That  
13 is why it is limited.

14 MR. CARBON: Theo, let me go back to the question  
15 of forcing the pellets toward the center. Am I understanding  
16 this correctly? The pellets may be broken up into bits and  
17 pieces, is that so?

18 MR. THEOFANOUS: Not the blanket pellets. They  
19 might break. I think it is doubtful in the time frame all  
20 the way observed to the very, very end of the core they will  
21 be all broken up, because first they are going to disrupt  
22 in the high power ridges.

23 MR. CARBON: Let me ask a question. If the pellets  
24 are considerably broken up, could you get a higher reactive  
25 rate from pellets not only going straight down in a column but



1 bits and pieces of pellets going out and filling up the  
2 channel so as to give you greater compaction and more fuel  
3 in the matter of the core?

4 MR. THEOFANOUS: No, Max, I don't think so. But  
5 the time there is expansion associated with the heating that  
6 is required. There isn't much room there.

7 Now, that activity rate is really controlled by  
8 the flux of fuel coming in times the worth grading, so if  
9 you allow any movement in this direction because of continuity,  
10 and your flux is going to drop, so it's already radial motions  
11 are mitigated. We have them anyway. I don't follow your  
12 explanation. If you move material in the one direction, in  
13 one direction, whatever you push in, the top must come out  
14 on the bottom. Therefore, you have the maximum flex.

15 MR. CARBON: Wait a minute.

16 MR. THEOFANOUS: I am assuming you will not only  
17 push it downwards but that some of the bits and pieces move  
18 radially outward into the channel. I mean this way.

19 MR. CARBON: Yes. And could you get more fuel  
20 moved into the center of the core than you did get?

21 MR. THEOFANOUS: No. Any of the mass from that  
22 one additional motion by radial motion, that is not available  
23 to move downwards. If you look at the sample fuel that is  
24 disrupted, cut out the core, say where can that go? All the  
25 gradients are for it to go upward because it is controlled by

1 the fission gases that are being released in the active  
2 part of it, and the gradient is for the fuel to move out.  
3 In fact, that is exactly what happens. If this thing is  
4 going to come in, this part here actually is dispersed. It  
5 goes out in both directions. In fact, that is what we are  
6 doing in the calculations, and that is what all the experiments  
7 show, so that any disruptive fuel, the natural tendency is  
8 to move outwards from the core.

9 MR. CARBON: Well, it appears by the same token  
10 at the top of the active core region you assume that the  
11 cladding is gone, your fission gas pressure is going to  
12 apply throughout, not only going to apply at the top of  
13 the column of fuel pellets, but the fission gas pressure  
14 would go out into the channel.

15  
16  
17 HEMLOCK  
18 BRAGAZALE  
19 COTTON CONTENT  
20  
21  
22  
23  
24  
25

1 MR. THEOFANOUS: I know where you are coming from.  
2 The problem here with this logic is that the pressure -- the  
3 high pressure is up here. Now, in order for this pressure to  
4 manifest out here, it has to go through some very, very tiny  
5 clearances between the blanket pellets and the cladding, so  
6 you talk about a very effect cladding. In fact, that is the  
7 reason that we have a problem. Really, if this clearance was  
8 somewhat larger, this gas would have vented well before the  
9 fuel disrupted, so because of the tight clearances here, we  
10 retain the pressure at high.

11 This retained pressure is effective in pushing the  
12 blanket pellet, but not effective in pressurizing the vessel  
13 because of the larger volume out here. Much larger space.  
14 So all the gradients, then, all the pressure gradients in the  
15 channel itself are from the core out, from the center out.  
16 The gradients are from the core out, and only inside the pins  
17 and only because this pressure here is basically kept there  
18 by this blanket pellet.

19 MR. CARBON: So we have two reverse pressure  
20 gradients?

21 MR. THEOFANOUS: One is inside downwards, and the  
22 other is on the other side and is upwards.

23 MR. CARBON: I understand what you are saying, but  
24 I guess in my own mind I am not really convinced that there  
25 couldn't be more fuel pushed in there.

1 MR. THEOFANOUS: Well, let me give you an example  
2 that may be of help to you, because this process is only very  
3 appropriate for the cladding location. The pressurization of  
4 the tunnel by gas coming out is important in the cladding  
5 location picture. What do we have here? We have another -- we  
6 all it something like a 14, 15 psi, and that is the velocity  
7 that causes the upward cladding position, and considered one  
8 dimensional. And there is no bypass around the molten fuel.

9 Another question is, in other words, the story when  
10 nobody considers these releasable stresses. Now, we did that.  
11 You remember in the November meeting a very detailed discus-  
12 sion in the report, and we are saying this gas there, of course,  
13 is going to push the pellets, but even before that is going  
14 to start venting into the coolant tunnel, and maybe now it  
15 can push the cladding downwards. The same reasoning that you  
16 are thinking of. And, therefore, it will interfere very  
17 drastically with the cladding location process.

18 Well, we took the gas out. We compared that against  
19 the experimental data which was done with the pressurized  
20 plenum on the top, so we did a very thorough, very detailed  
21 study of that, and we found out that even in the beginning  
22 the first time the cladding fails, and that is typically about  
23 .6seconds to .2 seconds earlier from when this happens here.  
24 That means you have the highest pressures and the highest  
25 blowdowns. Even then, there is not enough flow that they come

1 out to even reverse the pressure gradient, and that is only  
2 a pressure gradient of only a few psi.

3 Do you feel better about that? So it cannot inter-  
4 fere with the cladding under much more benign conditions. In  
5 this case, we have lower pressures pushing in. We have much  
6 higher pressure here because of the distribution and the  
7 fission gases being released, and there is no way in the world  
8 that could have anything -- that the released pressure from  
9 here can interfere with the overall pressure gradient itself.

10 The two processes are so -- you know, going to the  
11 cladding that is much more. I guess one would expect, if this  
12 was the case, what you say was correct, that we should have  
13 a very clear picture of cladding moving downwards under this  
14 effect, and we can grant that is going to happen.

15 MR. CARBON: I suspect you are right. Let's move  
16 on.

17 MR. ZUDANS: Could I ask a quick question, Theo?  
18 What process creates this delta t in the core?

19 MR. THEOFANOUS: In the tunnel?

20 MR. ZUDANS: That's correct.

21 MR. THEOFANOUS: The process I was just talking  
22 about, or what generates that adverse pressure?

23 MR. ZUDANS: Just the buoyancy, because you don't  
24 have any driving power there, do you?

25 MR. THEOFANOUS: Well, you have very little. The

1 flow is about 20 percent down in the pump. There is only  
2 10 seconds after the thing starts.

3 MR. ZUDANS: The other thing is if you apply that  
4 full pressure base of  $g$  that you have on the figure, what  
5 happens to those pellets at the point where they get out of  
6 the core? They still move the same amount?

7 MR. THEOFANOUS: That's right.

8 MR. ZUDANS: So it's really immaterial.

9 MR. THEOFANOUS: Because the motions are so little  
10 that can affect that.

11 So with this, then, we satisfied ourselves that  
12 there was no auto catalytic problem even under the very, very  
13 worst kind of assumption, but another problem came up as a  
14 result, and we did really bring this up in the previous  
15 meeting in which you remember that we had the vessel failures  
16 from all events, and we had one going from -- in fact, it was  
17 stretched out and was made to be 3 over 1000 because we were  
18 concerned about that, and since that time, we have been able  
19 to put some numbers to this, and it does turn out that because  
20 of the channel fix and the next channel to fail coming in early  
21 enough, and because of the whole accident progressing slow  
22 enough, and because this is like a slow accident progression,  
23 so everything is slow.

24 Fuel motion, voiding of the core is slow, and now  
25 this brings us back to the whole problem that we have happening

1 because the primary reason being sodium voiding. Now, it  
2 happens because sodium voiding is not fast enough so to speak,  
3 and because we have the additional kick that comes in and  
4 generates an over-power condition, so I want to emphasize,  
5 so that this is not taken out of context, that this is putting  
6 too -- I really don't want to call it pessimistic, but put  
7 two bad things one on top of the other.

8 Well, the question now is can that be auto catalytic.  
9 By the way, this was driven at a much slower rate because this  
10 may be driven by as much as 20 dollars per second, and we  
11 are concerned about that because there isn't much known about  
12 this situation.

13 This is with sodium fuel tunnels, and we feel it  
14 has been irradiated. How does it fail, and how much activity  
15 can one get by this? So we are very hesitant to try to cope  
16 with that in the old ways of the homogeneous review. In  
17 fact, this was an attempt to mitigate this situation here, but  
18 if there was anything accomplished from that or learned, you  
19 can get any numbers you like by just changing things in the  
20 order if one wants new certification one way or the other.

21 We want to calculate this. We have done some calcu-  
22 lations, but we don't report it because we don't think it's  
23 meaningful. It can be a bad situation, and then, it also has  
24 the ability -- we have the ability, or the project has the  
25 ability to eliminate it altogether if this pressure vent is

1 made to act on the blanket, and that looks like it was the  
2 best solution of all, and brings us to the plenum fission gas  
3 composition resolution, which is to eliminate it by design.

4 (Slide.)

5 We have been talking to the project, to the applicant  
6 and they have agreed to consider changing their design and are  
7 writing us a letter to that effect, also.

8 MR. KASTENBERG: Can you give us a hint as to what  
9 the design solution will be?

10 MR. THEOFANOUS: One of the possible ways that one  
11 can do that --

12 (Slide.)

13 -- between the moment that the cladding fails, of course, there  
14 will be some time between the failure of the cladding and  
15 disruption of the fuel, especially in this core. It is  
16 typically half a second or more, so that the gas would have  
17 come out well before it's relevant if there were a small volume.  
18 We have a large volume and it stays there -- the pressure stays  
19 up at the time this comes. So the thought, then, was if we  
20 take a volume and break it into two volumes, one volume --  
21 that is, the two volumes separated from each by a very, very  
22 small clearance.

23 Well, if that is the case, the upper volume being the  
24 big volume, when this disruption process happens, only the  
25 lower volume will be really effective in pushing that, and that



1 lower volume can expand very quickly because of the normal  
2 blowdown while the upper volume will decay much slower, in a  
3 longer period of time, and because of that, because of the  
4 clearance, will not be able to apply any pressure, like putting  
5 another impedance up here. That means this cannot be pressurized.

6 So during normal operations and because of the  
7 small clearance -- and it will be a calculation and enough  
8 space for the gases to flow, but under the rapid conditions  
9 there will be no way by which this pressure here can manifest  
10 itself on the top of the fuel.

11 MR. CARBON: You will effectively put a little  
12 orifice there?

13 MR. THEOFANOUS: Yes.

14 MR. KASTENBERG: Do you know if there is anything  
15 in terms of other accident scenarios or normal operations?

16 MR. THEOFANOUS: I think the project might like to  
17 take this question.

18 MR. DICKSON: Maybe I could explain that a little  
19 more clearly. At the top of the axial blanket is a spring  
20 about 7 inches long. It rests at its top on what we call a  
21 plenum spacer which is -- think of it as a closed tube except  
22 that it has a tenth of an inch hole in the top and the bottom  
23 just to provide a landing at the bottom and the top, which  
24 is almost to the top of the whole pin. Four inches away, a  
25 space for the tag gas capsule. We could close that up and make

1 it hermetically sealed and make some very fine holes in it  
2 like a couple of two-mil holes in the bottom and a couple of  
3 two-mil holes in the top. It would absorb, as Theo said, all  
4 the fission gases slowly, but would let them out only slowly.

5           We have actually built such pins for experimental  
6 purposes; not for this reason whatsoever. We regarded it as  
7 undesirable to do that. Obviously, you have to tailor it so  
8 that you will get the tag gas out when you want if you have  
9 a failed pin, for normal operation. You don't want to get it  
10 out so slowly that it impacts your failed fuel monitoring system.  
11 And secondly, it is going to cost something to do that.

12           You realize that there are some 50,000 pins in every  
13 core, so it doesn't have to cost very much to have an impact.  
14 And third, the only reason all that junk is in there is for  
15 shipping. The objective of the plenum is to give you a lot of  
16 room. The advanced fuel are trying to save the cladding from  
17 15 mils down to 12, and here is a whole mil inside there that  
18 if we follow this path, we never get out.

19           There are a lot of other ways you can think of to  
20 keep those -- this pellet back from moving during shipping,  
21 besides putting in plenum spacers that take up space and use  
22 much more as a gas plenum, reducing the height of the thing,  
23 reducing the cost of the fuel and the capital cost of the plant.  
24 So we do regard it as undesirable, but can do it if necessary.

25           MR. THEOFANOUS: So with this conclusion, then, we

1 feel that the energy, particularly, the energetic behavior,  
2 is physically unreasonable.

3 MR. KASTENBERG: Just as a follow-up, when one has  
4 a bulletin that says that applicant has agreed to consider  
5 it, does that mean that this will be revisited at sometime  
6 later, or does it just mean --

*End  
J-C*

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RG  
j-2d-1

1 MR. THEOFANOUS: I think this is important for the  
2 NRC Staff to answer. With the way you are approaching it, is  
3 this going to be fixed one way or another and therefore we  
4 don't have to consider it any further from our point of view?

5 MR. DICKSON: Let me add that we haven't had a  
6 chance to evaluate these results in detail. The project has  
7 heard of it but we haven't been able to evaluate it. Since  
8 we know of the design fix, we can commit to yes, if we agree,  
9 and see that that does need to be done. We can do such a  
10 thing, but we have not yet agreed that that mechanism is  
11 really operative, considering that Theo himself says that's  
12 a very conservative basis.

13 We reserve the right to look at it first, and if  
14 we can convince the NRC that maybe we don't need to do that.

15 MR. THEOFANOUS: Yes.

16 MR. CARBON: Along the same lines, how strongly do  
17 you feel the need for this? I think we have certainly -- we  
18 obviously have concern about safety, but I don't think we are  
19 pushing for something that is grossly overconservative.

20 MR. THEOFANOUS: Well, our own feeling is that we  
21 have taken the approach throughout this study here that we  
22 don't become so entangled with the details that we start  
23 really chasing an imaginary event. I think based on what we  
24 have seen in the calculations, we think that it is a reasonable  
25 concern that we would like to see it fixed.

j-2d-2

1                   On the other hand, it is not something that we  
2 can say that you are going to get all the time, but we certainly  
3 see -- and I think this is a good picture to show you that --  
4 you see that you become susceptible to detail. If you grind  
5 it down to the point where you try to separate a few tenths  
6 of a second between the core voiding or 20 or 30 percent  
7 left over, you begin to have a problem.

8                   Like here, for example, you can see that this  
9 shows that the time that the fuel disrupts, essentially all  
10 the core is voided except for the very low-powered parts  
11 which have just begun to boil.

12                   (Slide)

13                   In fact, boiling has started everywhere. It is not  
14 the classical one that you have seen before where maybe 50  
15 percent of the core had not even come even close to such.  
16 That in itself is important because it affects the failure  
17 location. If you have time to heat up the core, the failure  
18 mechanisms become different, so there are several aspects of  
19 the problem.

20                   What we are saying here is they are of detailed  
21 nature and they have to do with processes that we don't know  
22 how to characterize yet, so taking this plus the fact that  
23 we are saying it in calculations and also taking into account  
24 that our knowledge of it, or our direct experimental evidence  
25 on the effect of this pressure on the fuel pellets is non-existent

j-2d-3 1 and not likely to be discovered in the next one or two years,  
2 it is the kind of calculation that led us to the conclusion  
3 that we would like to see the core fixed in a different way.  
4 But what I am saying is if you sharpen your pencil, you can  
5 see the way out of it, but the question is as you sharpen  
6 your pencil, you lose a lot of the confidence that we have --  
7 that doesn't allow you anymore to make the statement that I  
8 made before, that you strictly cannot see.

9 MR. DICKSON: Two more points I would like to make  
10 if I could. One is that we have a significant amount of time  
11 before we order the fuel to resolve this problem. The second  
12 is the probability exists that we would choose to make the  
13 design fix rather than attempt the analysis because it might  
14 be cheaper to fix it than to go through the analysis.

15 This pencil-sharpening costs a lot of money, so  
16 we have left that open yet, as I say. We haven't had an  
17 opportunity to relook at it, but since we could identify a  
18 fix, we could go forward with the NRC that we can fix this if  
19 necessary.

20 MR. STARK: This aspect matured very late, and the  
21 applicant found out about it quite late, so what they did,  
22 they sent a letter to us, which I have here, dated March 8th,  
23 saying that they are aware of the aspects that we are looking  
24 at but they haven't had time to analyze it. In addition, they  
25 indicate that the design change that we suggest to them is

j-2d- 4 1 feasible, and since the fuel won't be constructed for several  
2 more years, what they have done in the SER is saying if you  
3 want some more time to study and analyze it, you can always  
4 come back with another way of solving the problem.

5 We know of one fix that exists, and unless they  
6 could show us another way, which we are always open to, we  
7 are not the designers. We just kind of set the regulations and  
8 the criteria. We know of one fix, so therefore, since a known  
9 fix is technically feasible and that exists, we indicate  
10 our criteria or equivalent, and they will have to come back  
11 and convince the Staff, whatever the equivalent might be,  
12 which happens all the time, but it's possible. So I think that  
13 is why we have handled it in the way we have.

14 MR. CARBON: Theo, it's 12:00. How much time will  
15 your next topic take? Is this a good point to break for  
16 lunch?

17 MR. THEOFANOUS: I think it is a good time. I am  
18 in between topics now, so it would be a good time. None of  
19 the topics, as you see, none of the topics are very big but  
20 there are a number of them. I guess if you wanted to go over  
21 the next topic, it could take maybe an hour.

22 MR. CARBON: Let's break for lunch, then.

23 (Whereupon, at 12:00 noon, the meeting was recessed,  
24 to reconvene at 1:00 p.m. the same day.)  
25

END 2d

## AFTERNOON SESSION

(1:00 p.m.)

MR. CARBON: Let's resume our activities.

Theo, go ahead.

MR. THEOFANOUS: We will continue with Mr. Bell, taking a bit of the time to kind of break him in here.

MR. DICKSON: Mr. Chairman, could I take this opportunity to ask a question that was asked yesterday and we said we would try to get an answer?

MR. CARBON: Yes, go ahead.

MR. DICKSON: Yesterday the question was asked if we accounted for the lack of symmetry in the loads in the bolts because of the assymetric head. We have checked. The answer is yes, we do. It is not as strong an effect as you might see. The variation from the mean is only about plus or minus 10 percent while they are in the elastic mode.

MR. ZUDANS: I would expect larger radiation in the elastic mode than in a plastic mode. Integration related more to the sheer key load. The large sheer key that you had -- sheer range.

MR. DICKSON: That too has a variation around it, yes.

MR. ZUDANS: About the same?

MR. DICKSON: I didn't check that number but we did calculate it.



j-1-2

1 MR. ZUDANS: Because I think when you do go out you  
2 find there is a significant difference, maybe 50 percent  
3 minimum to maximum load. Then you would have to model the  
4 flange in the plant test scale rather than put some plant that  
5 kind of hides this effect. If it is 10 percent, I don't care.

6 MR. DICKSON: Fine, thank you.

7 MR. CARBON: Go ahead.

8 VOICE: I am going to pick up, then, where Theo  
9 left off. He just finished the discussion of the initiating  
10 phase and the energetics potential for that phase of the  
11 accident, and what we would like to do now is move on into  
12 what we have called the disruption phase, which is roughly  
13 that part of the accident that continues from the early  
14 disruption in the initiating phase on out to the termination,  
15 either by fuel removal, which we have termed dispersal, or  
16 by energetic events, which we call disassembly.

17 So a lot of the perspective that we have looked  
18 for in the initiating phase was to gain a handle on what the  
19 conditions for this disruption phase legitimately might be.  
20 To do that, we have looked at the other end of the spectrum,  
21 so to speak, of the uncertainties and initial conditions of  
22 the initiating phase to look for that behavior which would  
23 tend to make the disruption phase more prone to energetic  
24 events. Therefore, we are particularly sensitive to anything  
25 that might change the extent of blocking, for example, of the

j-1-3  
1 normal coolant channels and the extent of disruption that  
2 They talked about. These things will play a role later on  
3 in the fuel removal processes and recriticalities to come.

4 In the disruption phase, as we move into that, I  
5 will review just briefly what are the major aspects of this  
6 phase of the accident that we are concerned with.

7 (Slide)

8 As Theo was pointing out, what we find at the end  
9 of the initiating phase is that most of the instantaneous  
10 neutronic activity is controlled by the fuel motions. The  
11 voiding and the cladding tend to control the state of the  
12 reactor in terms of how far subcritical or how near critical  
13 it may be, but the instantaneous reactivity effects are  
14 primarily controlled by the fuel motion.

15 Now, what we will do is follow those fuel motions  
16 on as we progress in time along this sequence to see how they  
17 further control the overall neutronic behavior of the system.  
18 What we generally will be looking for is the neutronic activity  
19 both from the standpoint of how it manifests pressures for  
20 fuel removal and also from the standpoint of what the energetic  
21 potential is. We have laid the groundwork now in terms of  
22 what the system appears to be able to take in terms of specific  
23 ramp rate events in a two-phase disassembly. The remainder of  
24 our discussion will be to look at this scenario itself and  
25 see to what extent we develop ramp rates that even come close

j-1-4

1 to those limiting situations.

2 So neutronic activity, of course, is a fundamental  
3 part of that, and as I mentioned, we will be looking for the  
4 transient pressures in terms of fuel removal and how that  
5 influences the scenario, and we will be keeping an eye out  
6 for this progressive disruption as we go through these  
7 different identified phases of subassembly disruption, and in  
8 a large scale, annular disruption and eventually to the whole  
9 cylindrical pool.

10 We will want to watch as we go along to see what  
11 kind of fuel removal paths are becoming available as we go  
12 through that sequence. We will also want to be looking for  
13 the other major aspect that tends to control the magnitude of  
14 the recriticalities or their severity, and that is the way  
15 in which this extended fuel motion becomes more coherent as  
16 the physical structures within the core are broken down through  
17 this progressive disruption process.

18 So we will go through this phase by phase, starting  
19 with the subassembly disruption phase, and look at some of  
20 the main features of that and particularly try to finally  
21 come to grips with what the energetics potential of this phase  
22 of the accident is.

23 A typical type of activity which goes on in a  
24 subassembly phase following the initiating phase, where we may  
25 have fuel still largely distributed in the subassembly structure,

j-1-5  
1 still at a high fuel inventory state at this point, and there-  
2 fore the fuel cannot really be in a gross slumping mode because  
3 it would have been critical long before that. In general the  
4 fuel tends to be distributed. It may be bunched up more on  
5 the ends.

6           There is a lot of coherence between channels, but  
7 on the average it is still roughly a distributed core in  
8 terms of the fuel location.

9           We found in every case from the SAS analysis of  
10 the initiating phase that we began this disruption with a  
11 highly neutronicly active system. It was not a quasi-steady  
12 state but a system in which power transients are continuing.

13           In that process what that means in terms of fluid  
14 motions is the potential, at least, even in the subassembly  
15 phase, for these multiple neutronic events, which can act  
16 across the structural boundaries not impeded by them, put power  
17 into the different subassemblies in essentially the same  
18 time frame, causing a progressive increase in the coherent  
19 behavior of the fluid.

20           One typical subassembly, if we follow through a  
21 sequence, might look like this, where we have the fuel largely  
22 distributed but over a longer time frame of a second or more.  
23 This fuel will -- much of the fission gas will be de-entrained.  
24 It will begin to slump. With the high inventory, it will  
25 tend to go recritical into this active neutronic mode, and

j-1-6

1 with the more or less classic but mild disassembly -- not really  
2 disassembly. We will have to watch that terminology. We are  
3 using disassembly to be a neutronic terminating event. This  
4 would be mod recriticality, with the power and the pressure  
5 tending to be centered about the mid-plane, the result being  
6 that the materials are pushed axially towards the ends of  
7 the subassembly, creating a highly subcritical state.

8 Now, at this point it is not likely that massive  
9 fuel removal would have occurred. It is still very early in  
10 the transient. We haven't even failed subassembly walls.  
11 The thing that will happen will be that the system will try  
12 to obtain more energy. It does not have enough energy at  
13 this point to overcome all the heat sinks available. It cannot  
14 maintain a steady dispersed state, and since it can't do that,  
15 it has no choice but to go recritical and try to obtain a  
16 higher energy state.

17 It does that in one of a couple of ways. We have  
18 defined two possible modes in which this state breaks down to  
19 achieve a second recriticality state. One is what we call a  
20 drainback mode. If you imagine the configuration looking like  
21 this, that material will not be stable there at the top. If  
22 sometime over at your kitchen sink you take a glass of water  
23 and turn it upside down with your hand over it and pull your  
24 hand away real quick, you will see this mode of fluid dynamic  
25 drainback occurring, where a bubble will grow up through the

j-1-7  
1 liquid slug and the material will drain down around the outside  
2 of it. That is a well-known solution.

3 Through that mode one can get the mass reflux into  
4 the puddle at the bottom, and as the puddle grows, the  
5 criticality state will be reestablished and this drainback  
6 rate would give a ramp rate and thereby will define a second  
7 recriticality.

8 If we do that and if we arbitrarily say that every  
9 subassembly in the core is undergoing this process coherently,  
10 that is, in the same time frame, and add all these mass reflux  
11 rates up together, and if we take then what the differential  
12 reactivity worth of this puddle increasing per centimeter  
13 really is and multiply those two things together, we find that  
14 we would have a second recriticality of about \$30 or \$35 per  
15 second.

16 MR. ZUDANS: Is this physically possible? You have  
17 both walls to contain.

18 MR. BELL: At this stage, yes. In fact, any other  
19 way would be really rather difficult to explain, I think. We  
20 are still at a state where these walls have not had time to  
21 be heated to their failure point.

22 MR. LIPINSKI: These are the subassembly walls?

23 MR. BELL: Yes. Remember they are starting out  
24 roughly at the sodium boiling temperature significantly below  
25 their melting temperature.

j-1-8 1 MR. CARBON: What is the temperature of the molten  
2 fuel inside the subassembly?

3 MR. BELL: A few hundred degrees, it is, melting  
4 point.

5 MR. CARBON: Which is what?

6 MR. BELL: That would be around 3230 degrees  
7 Kelvin.

8 MR. CARBON: It seems that is so high that in  
9 heat transfer to the stainless steel subassembly walls, it  
10 doesn't seem possible that the subassembly walls would exist  
11 there.

12 MR. BELL: It wouldn't except that what happens is  
13 the material on contact with the cold wall freezes the crust  
14 of material. The uranium dioxide has a very low conductivity,  
15 one-tenth of that of the stainless steel. So that crust of  
16 material that forms on the wall is basically an insulator,  
17 and even a crust a millimeter thick will require nearly 2  
18 seconds of time in order to melt that wall.

19 MR. ZUDANS: But looking at this, your column number  
20 2, when that happens the wall would dry out on the outside.

21 MR. BELL: No, not necessarily. You won't  
22 necessarily strip this the way I have got it shown.

23 MR. ZUDANS: If it doesn't dry out, it will not  
24 happen. It will remove the heat and it will --

25 MR. BELL: No, there is a film -- even if there is

j-1-9  
1 a liquid film on here, the fuel just does not have the ability  
2 to reject heat to the wall because of its low conductivity.  
3 It just cannot reject it that fast, particularly at these  
4 low temperature states. It will always form a solid crust on  
5 the wall, with the wall temperature what it is, yes.

6 MR. THEOFANOUS: There is no disagreement here. We  
7 are not saying that this transient is going to last for a very  
8 long time. It is going to be maybe a matter of a second or a  
9 second and a half. Now, if the walls were to fail, we would  
10 go to the next stage. What we are doing here is taking  
11 snapshots, and this is the first snapshot. It is not long-  
12 lived by any stretch of the imagination. In a second it will  
13 be over,

14 MR. ZUDANS: Well you see, for the fuel to melt,  
15 you would have to dry out the outside surface.

16 MR. BELL: No, no.

17 MR. THEOFANOUS: That is the wall of the subassembly.  
18 The fuel melts inside.

19 MR. ZUDANS: This is the outside can. Then I don't  
20 disagree with you.

21 MR. BELL: I'm sorry, we didn't clarify. We are at  
22 a state where the pins are completely disrupted and we are into  
23 a subassembly scale of fluid motion.

24 MR. ZUDANS: And you say you might have a second or  
25 two of this.



j-1-10

1 MR. BELL: That's right.

2 MR. MARK: So there is a fair amount of steel in  
3 that red solid stuff.

4 MR. BELL: If the cladding still has not been  
5 relocated into blockages, then this would be roughly one-third  
6 steel; that's right.

7 MR. KASTENBERG: Charlie, you mentioned at the  
8 beginning that your power is moving dynamically during this  
9 period, or at the beginning of this period. Are you at a  
10 critical situation? Are you at low power, high power, where  
11 are you power and reactivity wise?

12 MR. BELL: All over the map, literally. I mean  
13 you are coming into here (indicating). Typically when you are  
14 finished with the fission gas control dispersal phase and so  
15 on, you may be at most a few dollars subcritical. It doesn't  
16 take very much fuel slumping to bring you back up to a critical  
17 state.

18 MR. KASTENBERG: My experience has been at these  
19 stages and calculations, any little change makes you diverge  
20 either up, down, any little motion, very sensitive reactivity  
21 changes here.

22 MR. BELL: That is exactly right. You don't  
23 necessarily have high ramp rates.

24 MR. KASTENBERG: Right.

25 MR. BELL: But when you are near that critical

j-1-11

1 point -- in fact, that's one thing that keeps the ramp rates  
2 from being very high, because you are never in a position to  
3 get momentum built up before you are recritical again.

4 MR. KASTENBERG: I guess the next thing is how  
5 do you justify these very clear pictures from one bar to the  
6 next?

7 MR. BELL: What we are trying to do is to establish,  
8 if you will, a bounding situation. This is as big as it can  
9 get even in the worst circumstances, and if that bounding si-  
10 tuation -- and we agree that it is a bound, and if it is of  
11 this order, then we are still not challenging the system.

12 MR. KASTENBERG: This is really not a snapshot --  
13 it is not a progression as you look across the picture, then.

14 MR. BELL: No. This is a representation of what  
15 could be going on in a given subassembly. I think that is real.  
16 The thing that every subassembly in the core is doing this  
17 simultaneously is simply a way to get to a bounding situation.  
18 If you wanted to be concerned in the limit of this neutronic  
19 tuning, having everything going up and down together, and we  
20 don't think that's really possible because this state breaks  
21 down before about two or three neutronic cycles and therefore  
22 that is not enough to tune the whole system and therefore  
23 clearly this is a bounding type number. It cannot get any  
24 bigger than that.

25 Now, you can get a somewhat bigger number if you

j-1-12

1 don't take what I would consider to be a more realistic type  
2 of drainback or reassembly of the puddle. If one were to  
3 postulate that this upper half of the material on its way up  
4 were to completely uniformly distribute itself and not get too  
5 much momentum such that it would free agglomerate at the top  
6 like it did here but just enough till it gravity acted on it,  
7 then it turns out that the reflux rates are a bit higher in  
8 that mode. And again on this whole core basis, it is around  
9 \$82 per second.

10 MR. KASTENBERG: Why do you rule out a slump flow?  
11 Why can't that whole --

12 MR. BELL: For the same reason that I can't get it  
13 to fall out of my glass of water when I turn it upside down.  
14 It is unphysical. You would be pulling a vacuum up here at  
15 the top in order for that to do that. It would not fall. In  
16 order for it to fall, a gas bubble has to grope through it.

17 MR. KASTENBERG: But you are not closed on top of  
18 the channel, are you?

19 MR. BELL: If we weren't closed, we would remove  
20 the fuel here and we wouldn't have the problem to begin with.  
21 This channel is what I would call a leak-tight channel where  
22 you would expect some degree of blocking.

23 MR. ZUDANS: I am just following Bill's comment.  
24 You could begin to slow down and form bubbles.

25 MR. BELL: Well, I have never seen that physically.

j-1-13

1 MR. ZUDANS: A glass never falls along the wall;  
2 it always falls in the middle.

3 MR. BELL: I think at this point these details  
4 don't matter.

5 (Laughter)

6 MR. BELL: Yes.

7 MR. THEOFANOUS: I want to clarify this point that  
8 Bill brought up. The process of having liquid fall down is  
9 one in which the acceleration record is the record from the  
10 light phase to the heavy phase. That is known as the classical  
11 stability. There is no way, though, which you can have it  
12 fall down independently, regardless of what is on the top or  
13 the bottom, just looking at the interface.

14 Now, the size of this thing is such that at the  
15 most it will generate one or two wavelengths, and that is why  
16 the picture of this first thing that you see there kind of big  
17 like a bubble. So it is inherent that the slug would break out  
18 because of the instabilities.

19 MR. BELL: Now, what this does for us in this  
20 stage of the accident where we are talking about the subassembly  
21 scale fluid flows, what we are led to is that unless some  
22 mechanism exists to induce higher ramp rates than these that  
23 we are getting from this oscillatory recriticality behavior,  
24 we will not have a challenge to the system, the structural  
25 system. So the only potential challenge will come if we get

j-1-14  
1 into phenomena that we discovered as we were looking at some  
2 postulated configurations, primarily from the standpoint of  
3 seeing how recriticality or disassembly yields would be  
4 dependent on these configurations. Obviously, you could have  
5 a whole range of configurations here, depending on what the  
6 fuel inventory was and when it would go recritical.

7 So, in that process we came across something which  
8 we have termed a disassembly or recriticality best phenomenon,  
9 and it has to do with this kind of a picture here where the  
10 material is puddling in the bottom with the flux peak because  
11 part of the reactor may not have disrupted yet and be in a  
12 configuration more like this. If you take the reactor as a  
13 whole, the flux peak might be somewhat above the top of this  
14 puddle.

15 If that's true and now you put a mild burst on that  
16 puddle, the thermal expansion of this puddle would drive mater-  
17 ial up the flux gradient and actually give you a positive  
18 reactivity effect right at the peak power. In other words, you  
19 get a disassembly or recriticality-type boost.

20 That worried us for a while, but it is only this  
21 phenomenon. If this phenomenon does not exist in any signifi-  
22 cant way, then I think we are prepared to say that this  
23 early phase of subassembly behavior cannot challenge the  
24 structural part of the system.

25 So we have investigated that to some extent. I will

j-1- 15 1 not try to go into all the details of it at this point, but  
2 basically what we find is this: that if you only had this  
3 occurring on a very local scale -- in other words, the rest  
4 of the reactor is not disrupted and not coherently doing this,  
5 then the ramp rates, because only a small fraction of the  
6 core is involved, are relatively small. The ramp rates coming  
7 into this recriticality are small, and if one does have any  
8 of this phenomenon going on, you are amplifying something that  
9 is already small and therefore you never get again to the  
10 challenging energetics level.

11 If, on the other hand, a large part of the system  
12 has been tuned or coherent so that a lot of it has begun to  
13 puddle all at once, we have to remember that the worth curve  
14 tends to follow the mass centroid of the material.

15 Now, starting out with half of it at the bottom,  
16 roughly, that means the mass centroid is very close. In this  
17 case it is up here somewhere. But as the material moves  
18 down, what will happen is that the peak of the flux will  
19 actually move into the puddle, and now the recriticality,  
20 rather than boosting, is actually mitigated.

21 So the only possible place this can happen is if  
22 you are in the very early stages of disruption in the  
23 subassembly phase, in which time the initial ramp rates are  
24 going to be very small in this mode and therefore the boost  
25 is never powerful enough to achieve a threat to the system.

j-1-10

1                   There are many, many conditions that have to be  
2 satisfied in order to get the boost in the first place, some  
3 of which I have listed down here at the bottom. Obviously,  
4 if there was radial compliance in the system, any radial  
5 expansion, which typically generates quite pressures, might  
6 just as easily move sideways. If there is any void in here  
7 to start with, that local compliance will observe the  
8 thermal expansion and therefore you won't get the boost.

9                   There are a number of things that you have to have  
10 just right in order for this boost process to come about.  
11 Obviously, in just postulating configurations we managed to  
12 find some that were very, very capable of producing these  
13 boosts, but we, I think, as our perspective has matured a  
14 little bit, we find out that those idealizations really have  
15 no place in this at all. They just cannot come about.

End  
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25

(Slide.)

If I try to put the entire sequence into perspective and in the framework of the system continually losing fuel, what we find is that the boost mode will typically have a fairly narrow range in which it can occur; either because of the high inventories and therefore low reductions of inventory, you just can't get the ramp rates. That's what I was talking about. You don't have the ramp rates. You don't get a significant boost effect.

On the other hand, if you've lost nothing but you are such that you have to puddle the reactor extensively in order to go critical again, then the flux peak moves into the puddle, and you very quickly cause disassembly or mitigation of any recriticalities from that point on. So at best it can only possibly happen in a very narrow range of inventory and for a very specific set of conditions.

And furthermore, the only way you can get that is through this ideal rainback mode of reassembly rather than in the drainback mode that we showed on those two situations. So, therefore, we are very comfortable with the conclusion here that we cannot get a significant boost, and therefore, we cannot get a significant energetics threat during this early disruption phase. We just cannot see that happening.

Let's move on then to the next stage, generic stage



1 of disruption where we have now seen the can walls between  
2 the driver subassemblies disrupted due to this rapid heatup  
3 from the fuel material. But because the internal blankets are  
4 starting from a colder state and they have no internal power  
5 generation, at least of the same magnitude they have, which  
6 is roughly one-quarter as much in terms of specific power,  
7 that their disruption will lag behind the disruption of the  
8 can walls, the subassembly walls, and the drivers themselves.

9 So we can see that the generation of the pool or  
10 the progression of the pool becoming larger as different  
11 driver subassemblies merge into this annular pool. I've got  
12 this displayed here as again we look for in a generic way the  
13 kind of bounds on ramp rates that one might expect in this  
14 kind of configuration.

15 Again we start with a gravity-controlled fuel  
16 motion as the primary motivator of the neutronic activity. At  
17 this point just about everything else is gone in terms of  
18 reactivity effect. If you visualize this as a particular  
19 starting point for one of these neutronic cycles -- and of  
20 course this is not unique; there's nothing terribly generic  
21 or unique about this, but it is simply meant to show a tendency  
22 for puddling at the bottom because typically you will have  
23 lost some fuel by this time, so you have to have some degree  
24 of slumping to achieve a criticality state.

25 But beyond that, this is a relatively chaotic

sc 3

1 distribution material. Now, it goes recritical. I have  
2 a power shape that looks something like this. Again, it's  
3 peaked down at the mass centroid, and it's also peaked at  
4 the radial inside of this annulus simply because the rest of  
5 the core is over in here tending to pull the flux up, whereas  
6 at the outside you have nothing but blanket, and you have a  
7 leak flux going on there. So the power leak radial shape  
8 will tend to be peaked in here and die off on the outside.  
9 Therefore, when I have this next recriticality, my hot point  
10 will be right in here at the inner radius at roughly the  
11 axial mass centroid.

12           The result then if you look at the flow dynamics  
13 of that kind of a bubble growth through heating a region here  
14 that's preferentially expanding just like a disassembly, that  
15 vaporization and pressurization at this point will cause  
16 these fluids to want to move in a typical type of motion that  
17 we calculate and observe experimentally, because we've actually  
18 set up some experiments nearly full-scale with water with  
19 introducing gaseous sources to represent the disassembly here  
20 to try to follow this fluid dynamics. And indeed it does  
21 appear to have these kinds of characteristics where the bubble  
22 grows and because of the inertia being less at the top, there  
23 is a bias in the growth of that bubble upward and outward.

24           Momentum in this fluid here is given in the early  
25 expansion phase, and it's given a momentum upward and outward.

1 And as that momentum continues to drive the fluid, the  
2 fluid tends to collect out of the wall, and its vertical  
3 component tends to carry it on towards the top as the bubble  
4 breaks through.

5 That momentum, depending on the strength of this  
6 recriticality, if it's very mild this momentum will not be  
7 enough to carry it all the way around the circulation. It will  
8 tend to climb the wall, turn around and fall back down. That's  
9 a very low ramp rate as a result of that.

10 On the other hand, if the momentum is high enough,  
11 it tends to have the circulation pattern. Now, what that  
12 does effectively is if you look at this picture, half of the  
13 mess is down here, the other half is up here, but this half  
14 that's up here is really distributed over two lengths of the  
15 system. So, therefore, as it circulates and comes back again,  
16 it will have roughly one-half the reflux rate as it would  
17 have if it were all just draining down together.

18 Consequently, you calculate in a circulating mode  
19 ramp rates of roughly \$35, \$36 per second -- again, no real  
20 threat to the system. Again, if on the other hand you know  
21 there may be some subassembly wall stubs sticking down here  
22 at the top or other dispersion sources here at the top such  
23 that this climbing sheet is dispersed into a rain here at  
24 the top, I think that's very ideal in the sense that it's  
25 perfectly distributed and then rained back.

1           What we would visualize as a real upper limit on  
2 this would be like \$72 per second -- again, well below the  
3 structural threshold for vessel head failure.

4           MR. MARK: Don't put the other slide back on. But  
5 in a slide or two ago you had that thing that Bill Kastenberg  
6 was asking about where half the fuel is condensed at the top  
7 and half at the bottom. And you pointed out -- and I don't  
8 argue against that -- that you can't think of that top thing  
9 coming down as a slug in real life. It will break up like  
10 water out of a bucket, agreed; but nevertheless, I think if  
11 you imagine that slug coming down under gravity, you would  
12 break it off these other, more complicated variants you have  
13 shown of rain, et cetera, and have a ramp rate that's probably  
14 higher than any of the ones you put on the slides.

15           Is it so that that ramp rate even is not so high  
16 as to cause you a problem, or is it too high?

17           MR. BELL: I don't know that magnitude. Do you, Theo?

18           MR. THEOFANOUS: That would be about three times higher,  
19 so 36 times 3 is a little bit over 100. And as you remember  
20 from the last concluding slide of my structural presentation,  
21 it removed that problem. If you applied the factor of 3 on the  
22 rainback on the \$70, then you were pressing the limits.

23           MR. MARK: I was wondering if you knew a number to  
24 go with that, because that's a geometrically simple set of  
25 assumptions.

1 MR. BELL: I guess what you're saying, this would  
2 be roughly a factor of 3 larger than the numbers we have.

3 MR. MARK: Three on top of one of the numbers you  
4 have.

5 MR. BELL: Of which one, the rainback number?

6 MR. MARK: If it's 3 times 36, then you would have  
7 said well, there's a limiting case, and everything is still  
8 fine.

9 MR. THEOFANOUS: Yes.

10 MR. MARK: If it's 3 times 70 or 80 cents, then it's  
11 not so immediately obvious.

12 MR. THEOFANOUS: You mean \$70 or \$80.

13 For the rainback we have about \$70 times  
14 3. That's about \$200. And what we're saying is in that  
15 extreme limit we are approaching the threshold.

16 MR. MARK: Is that dollars or --

17 MR. THEOFANOUS: It's dollars, dollars per second.

18 MR. MARK: Oh, you said it's \$82. You'd get \$240,  
19 and that would be a little high.

20 MR. THEOFANOUS: A little high, yes.

21 MR. BELL: But that would in fact require you to  
22 stretch your imagination even further, because now if you're  
23 bringing things back in a given subassembly more rapidly,  
24 that means I have to require a closer coherence.

25 MR. MARK: I wasn't arguing for it as probable event.

1 It seemed to me it had to be limiting, and if it were also  
2 tolerable, then, of course, you could save some of these  
3 drawings.

4 (Laughter.)

5 MR. BELL: That's right.

6 MR. LIPINSKI: I think that last slide of your  
7 annular pool -- has this phenomenon been verified through  
8 analysis with the code?

9 MR. BELL: These patterns are actually calculated  
10 by the code.

11 MR. LIPINSKI: What are you using and what are  
12 the assumptions in terms of boundary conditions?

13 MR. BELL: We did this with the SIMMER code, the  
14 same one we are using all the way through this, and  
15 seeing a lot more of it through here. These are rigid  
16 boundaries all the way around here.

17 MR. LIPINSKI: Are they six-sided?

18 MR. BELL: These are annular.

19 MR. LIPINSKI: If you don't have a six-sided can,  
20 if you assume a cylinder in order to make the analysis --

21 MR. BELL: There are no cans at this point, right.  
22 We are out in advanced disruption stage.

23 MR. LIPINSKI: Okay, I see. I'm with you now.

24 MR. BELL: This outer boundary might in fact be  
25 jagged if you chased it around.

1 MR. LIPINSKI: Now you've smoothed it out.

2 MR. KASTENBERG: The line on the left in each figure  
3 which looks like a center line --

4 MR. BELL: That's really the inner radius of the  
5 annulus.

6 MR. KASTENBERG: Okay. And that is solid still.

7 MR. BELL: It's porous.

8 (Laughter.)

9 MR. KASTENBERG: I guess that's what you want it to  
10 be.

11 MR. BELL: Well, you can imagine it as being --  
12 there are subassembly structures there, but the gaps between  
13 subassemblies are open, so there can be some small passage of  
14 material back and forth between this region. And those cans  
15 are being melted, so you know at one instance in time it may  
16 look like subassemblies, and the next instance it may look  
17 like a half a subassembly and on and on.

18 (Slide.)

19 Now, we move on to the whole core pool, and I think  
20 not surprisingly that if there's a threat to be had to the  
21 system it is in this phase that it would come from. And I  
22 think we probably all were able to forecast that months and  
23 months ago, but perhaps there are some new things here that  
24 we have not up until now really been concerned about. But they  
25 turn out to be very major aspects of this early whole core pool

1 behavior.

2 I'm going to make the distinction here in the whole  
3 core cylindrical pool between its early behavior, which I will  
4 also call its unhomogenized state, and its longer term behavior  
5 in which the internal blankets, which I have shown here in  
6 purple, it's simply a collapsed rubble form. This long-term  
7 phase would be when these blankets are homogenized into the  
8 system. It turns out that the whole-core pool has a very  
9 fundamentally different behavior in those two arrangements.

10 MR. LIPINSKI: Is that core center line on the left?

11 MR. BELL: This is true core center line, so this  
12 is roughly to scale, roughly a meter radius and a meter high.  
13 All the structure within the core boundary has been disrupted.  
14 It's all in a mobile, largely fluid state at this point. And  
15 again, as we look at some of these idealized calculations, we  
16 can think of these boundaries as being solid.

17 The other reason that the early behavior of the pool  
18 or the unhomogenized behavior of the pool is different is  
19 because all of this blanket material in the middle does one  
20 fundamental thing, and that is to shift the power peak outward.  
21 Instead of one peak at the center line in a normal fundamental  
22 mode, it is peaked out here in this annulus region.

23 Now, that means that whenever I have a recriticality  
24 event, my peak energy deposition is going to be out here, and  
25 a slushing mode, that's the one we typically worry about the



1 most in terms of getting larger ramp rate events. It's not  
2 the classical slushing mode that starts at the middle where  
3 material sloshes out and then runs back in a bowl-shaped  
4 slosh. It's going to be more of a confused slosh now where  
5 material is going to tend to split at this radial location and  
6 go both ways as well as upward.

7 Now, it turns out that is a very, very fundamental  
8 change to the way the ramp rates are developed and tends to  
9 keep the pool in a more or less confused state rather than in  
10 this highly organized state where you can slosh back in with  
11 high ramp rates.

12 So the early behavior is not a central sloshing. And  
13 furthermore, the ability to homogenize this material, if you're  
14 tending to generate the sloshes out in this region, there's  
15 not very much of a homogenization potential here. If you  
16 were centering your power in here, then obviously you'd be  
17 driving all this material out and mixing it up. But if you're  
18 centering your power out here, and it's actually more  
19 dramatic than what I've got it shown here, then you're tending  
20 to, if anything, keep the material bunched in the middle.

21 Now, what this does for the whole scenario is that  
22 it allows a time period beyond the initial formation of this  
23 pool for additional fuel removal. When you get to this state  
24 you've got enormous fuel removal paths along this outer  
25 boundary. If there's time available to move material into

1 those paths and remove it from the core, you can literally  
2 have a very high assurance of reaching permanent neutronic  
3 subcriticality before this homogenization process is completed.  
4 And that's the theme then that we will want to try to walk  
5 through as we go on down through this disruption scenario.

6 MR. MARK: In this picture you're not assuming that  
7 any fuel has already penetrated down into the lower blanket.

8 MR. BELL: It need not. In reality it really has.

9 MR. MARK: You say it might freeze, so we'll assume  
10 it does.

11 MR. BELL: When we talk about the actual analysis  
12 of the scenario, these are just kind of postulated configura-  
13 tions at this point, but the actual analysis, we are in fact  
14 having fuel move out of these regions continuously all the way  
15 up to this point in time. And what we find is that with  
16 this material in this general configuration -- in other words,  
17 nonhomogenized -- that the threshold for subcriticality has  
18 been changed; that we really only need to remove something on  
19 the order of 20, 25 percent of the inventory to reach a  
20 subcritical state, continuously subcritical state, even with  
21 the material completely slumped.

22 Now, without that or in other words if this is all  
23 homogenized, then you have to remove something like 35 percent  
24 of the inventory to achieve that permanently subcritical  
25 state, so the lack of homogenization not only keeps the

1 neutronic oscillations under control, keeps them to a lower  
2 ramp rate, but it also tends to put you to an arrest, if you  
3 will, to an arrest period in this whole scenario where you  
4 are temporarily subcritical and maybe even more than that  
5 because obviously if you're subcritical, how's this stuff  
6 ever mix.

7 Let's go on now then --

8 (Slide.)

9 -- And explore this a little bit further in kind  
10 of a bounding way again. To simply say what the situation  
11 would be if I went all the way to the limit of having things  
12 completely homogenized, we have all had some concern about  
13 these sloshing modes and what kind of ramp rates could result  
14 from them. So we decided that we would in fact actually  
15 attempt to do some sloshing-type calculations with full  
16 coupled neutronics to see just what the feedbacks really were  
17 from these things.

18 Here is an example. If we imagine again some  
19 partially slumped initial state with everything homogenized  
20 and then assume that there is a perturbation applied to this  
21 system, either mild recriticality or some sort of pressuriza-  
22 tion source that is centrally located, now a neutronic event  
23 will tend to be centrally located; that's one reason we have  
24 been worried about it in the past.

25 Pressurization events if they can occur -- there's

1 nothing that says they need to be at the center. They could  
2 be anywhere. But the most organized sloshing behavior and  
3 therefore the highest potential for a high magnitude ramp rate  
4 must come from a centrally located slosh.

5 So now we apply our perturbation here to center line  
6 and again at the axial mass centroid of this. What you find  
7 is again that the material that tends to be driven to the  
8 outside -- this is the classical disassembly, only it's going  
9 way beyond disassembly. Normal disassemble calculations move  
10 material a centimeter or two, and you have neutronic shutdown,  
11 but you put momentum in the system.

12 What we're doing is following that momentum on out  
13 in time and asking the question what happens then. If I  
14 haven't removed fuel, I temporarily disassemble. This is a  
15 disassembled configuration. It might be \$20 subcritical,  
16 but the accident is not over if there is still a high inventory  
17 of fuel within that reactive core region.

18 So now what happens at this state in time which might  
19 be several tenths of a second after the original neutronic  
20 event, which, by the way, might be a very large event; this  
21 might be just a few dollars per second type event initially --  
22 we have a system here now that's pretty much pressure  
23 equilibrated, nothing but gravity to act on it. Most of  
24 the momentum has been dissipated, and gravity simply begins  
25 now to pull this. It would go down to here. This material

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1 at the top falls down. The material at the edges will tend  
2 to drain down and then have to drain in as it wants to reach  
3 a common level here. And that's what I have depicted at this  
4 point. The top material has fallen down and formed a region  
5 in here. The other material is simply draining down the wall.  
6 Remember, there is a radial convergence here. So even though it  
7 doesn't look like there's much material here, by the time  
8 you start moving it across the bottom, it starts to look bigger  
9 and bigger.

10 MR. LIPINSKI: What provides that top boundary?

11 MR. BELL: Here?

12 MR. LIPINSKI: No, the next one.

13 MR. BELL: This top boundary here would be normally  
14 the upper axial blanket which has been plugged up with either  
15 fuel trying to escape and freeze in that location, or from  
16 prior steel blockages from the initiating phase. Typically  
17 it is not an absolute boundary.

18 MR. KASTENBERG: These are SIMMER calculations again?

19 MR. BELL: This is a pictorial representation of  
20 the actual SIMMER calculation.

21 Now, from these calculations what we see is that  
22 when the pool is reconfiguring in about this state -- in other  
23 words, the mass of the material really has not reached the  
24 center line yet; it's what we call a partial in-slosh or  
25 inward slosh. What you find if you simply do some K calculations

1 in a configuration like this and simply advance its state,  
2 pull this down and move this down a little bit, you can actually  
3 get a progression of ramp rates if you know the velocity of  
4 material here, or you can construct a radial differential  
5 worth curve.

6 What you find out is that typical velocities across  
7 here are like a meter per second, and the differential worths  
8 are on the order of a centimeter again. That's a very magical  
9 number. It seems like it crops up all the time.

10 Consequently, in this configuration one can envision  
11 for this organized axisymmetric centrally located slosh ramp  
12 rates on the order of \$100 per second. If I let that slush  
13 progress under this state where it's come on down all the way  
14 in and starts to overslosh and build up in the center, which  
15 again is not an unreasonable thing to expect since this  
16 momentum has to be dissipated somehow, and the only way to do  
17 it really is to turn the corner and let gravity work on it,  
18 so it bulges up in the middle.

19 What we find again from doing the differential K  
20 calculations is that the minute this material reaches the center  
21 line and starts moving up, we find that the differential  
22 worth changes dramatically. You're going into more of almost  
23 a hemispherical type of reactor, and that's saying the inward  
24 velocity when it's in this configuration starts to look like  
25 \$300 per second. And it's the worrisome-type numbers that we've

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been confronted with in the past.

But let me point out one very significant point here. This configuration at best must have a fairly low void, and in fact if does have a low void, that \$300 per second is nearly meaningless. It does not produce a significant energy yield. In fact, the actual calculation that we did had one that went recritical at about this configuration with \$100 per second and another one that went recritical at \$300 per second. This one produced twice the energy of this one, and it's simply because to get that high ramp rates in the filling of the central region and this turning of the momentum, it has to be single phase.

end AR 2

1 MR. MARK: This has to depend upon the power level  
2 from which you start. If you are at fairly high power you  
3 will reverse that thing on the lower right before it ever gets  
4 there.

5 MR. BELL: Well, you're coming from a highly sub-  
6 critical state, so typically you will be coming in there at  
7 relatively low powers.

8 MR. MARK: You will start to vaporize junk in the  
9 middle, and as soon as it starts that--

10 MR. BELL: That will eventually crash around, but  
11 what I'm saying is that with these high ramp rates, we're only  
12 talking, what, three more seconds of time between critical  
13 and prompt critical, and it's a very, very short time to  
14 overcome any momentum, and if the powers in that range are 10,  
15 20 times nominal, you are only heating material up, you know,  
16 a few tenths of degrees. The calculations show --and I think  
17 that kind of a fairly simplistic reasoning also confirms that  
18 you really won't get much of that happening with these high  
19 ramp rates. At low ramp rates you will tend to get that  
20 mitigating feedback coming in due to vaporization on the way  
21 to prompt critical. But here you get there so fast it really  
22 doesn't do much.

23 MR. KASTENBERG: Charlie, could you go back one  
24 viewgraph?

25 MR. BELL: Yes. To understand what differentiates



1 this case or this kind of a scenario is the heterogeneous core  
2 giving you that initial power shake.

3 MR. BELL: Exactly.

4 MR. KASTENBERG: Did you do some sensitivity calcula-  
5 tions to see what would happen if you got some redistribution  
6 of the blanket materials and you changed that?

7 MR. BELL: In the whole core calculation; these are  
8 just spatial effects, calculations I've talked about so far  
9 where we've started with idealized initial conditions.

10 Now, what I'm going to show you next are a few  
11 results of an attempt to mechanistically calculate all the way  
12 through this thing from the SAS conditions, and there, of  
13 course, you have a continually changing configuration. I  
14 don't think I have enough details here to really eliminate  
15 that too well for you, but we will see some of the effects of  
16 it, nevertheless, in some of the overall results here. But  
17 indeed, that is continually changing. And what you find is  
18 that every time you have a mild burst, this whole power shape  
19 changes dramatically.

20 MR. KASTENBERG: That's what I was wondering. Once  
21 you get into some movement it will forget what the initial  
22 condition would be.

23 MR. BELL: Exactly.

24 MR. KASTENBERG: Then I wouldn't see why the hetero-  
25 geneous core would be different than the old homogeneous core.

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1           MR. BELL: When you have a mild event and you move  
2 this material, typically you're moving it into a higher leakage  
3 condition. What that does, if you take this material, for  
4 example, here, that's on the average much higher enrichment than  
5 this material, the flux tends to be peaked out here. The  
6 power burst then is centrally located here and will tend to  
7 move this material up on a sheet on the side in some confused  
8 motion back in here. This material on the side will become  
9 a very high leakage configuration, and what you will find is  
10 that this flux will peak and go back to the fundamental mode  
11 kind of thing. The coincidence with that is the fact that  
12 your systems is 10 dollars sub-critical. Then when it  
13 reconfigures again into a critical state, you're right back to  
14 the same over-enriched region out here, and it comes right  
15 back to that same kind of general power profile.

16           And what we find is that after a number of these  
17 events, this will tend to start to be mixed, and what you  
18 find is that you'll go recritical with this thing being  
19 essentially flat, for example, and eventually as it's mixed  
20 more you go back to the fundamental mode and you'll see that  
21 whole progression changes.

22           MR. KASTENBERG: But all along you're assuming this is  
23 all bottled up, right?

24           MR. BELL: No. In the actual calculations you're  
25 not assuming that. In fact, that is the key thing you're asking;

1 this is what we would call the -- in the development of the  
2 clearance in the system, going hand in hand with that is the  
3 loss of fuel from the system which is trying to take it sub-  
4 critical, and these are two very important features that one  
5 tries to consider, and we will talk about this whole business  
6 of dispersal or removal here in a little bit, because we've  
7 tried to tag it, also, in these different stages in kind of  
8 a generic way, using this whole core calculation as sort of a  
9 background perspective to give us an orientation of what's going  
10 on.

11 In this by the way, we basically ran across this  
12 whole behavior as a result of that whole core perspective.

13 (Slide.)

14 Let me go on and just say a few words about the type  
15 of results that one gets when you try to make a stab at  
16 analyzing this thing all the way through. There haven't been  
17 very many of these kinds of calculations done, and they are  
18 literally of a project nature to turn in and do them. So  
19 you don't expect to do a whole lot of sensitivities in these  
20 kinds of things, and, therefore, we have chosen to use them  
21 not in a mode of saying this is the answer, but more in a mode  
22 of establishing perspective for us, and then we'll go off and  
23 do the separate effects and even idealize things a little bit.  
24 But at least we have this background which is a kind of bench-  
25 mark for idealizing. We just don't pick things out of the air.

1 Well, you see a very complicated picture here. What  
2 we have is the green lines are the reactivity state of the  
3 system as a function of time. Times 0 here is at 19.757  
4 on the SAS calculation that preceded it. It happened to be the  
5 SAS calculation that Theo was terming as the slow development  
6 of that transient. We chose that one because it would tend  
7 to maximize the blocking of the system prior to this whole  
8 phase, and we were trying to edge towards the conservative  
9 side of things, so we chose it that way.

10 The SAS calculation up to this point had already  
11 manifested some neutronic activity. There had already been  
12 several swings up and down in the reactivity rates. The  
13 SAS calculation was, in fact, run out to this point. It had  
14 run through this burst. We did it purposely so that there  
15 was an overlap in the transient, and we can check the SIMMER  
16 calculation of this power burst versus the SAS calculation to  
17 see if we have made the transition in a reasonable fashion.

18 Indeed, in this case, the power burst occurs at  
19 very nearly the same time and is of almost identical magnitude.  
20 We feel like we have preserved the system reasonably well in  
21 making the jump, but what we see here is the very cyclic  
22 neutronic behavior during this first 1 1/2 seconds. And it  
23 turns out to be synonomous with the subassembly phase of  
24 disruption.

25 Now, what is going on there is literally the simple up

1 and down motion that we were looking at earlier in the cartoon  
2 fashion where the neutronic event burps this material up into  
3 the air, gravity pulls it back down. You look at the timeframe  
4 and it's exactly the time you would expect for gravity to  
5 reassemble the material. It also manifests a very gassy phase  
6 of the exit of fission gas and there was some cushioning going  
7 on. In the reassembly process you have these nice single-phase  
8 puddles at the bottom.

9           Then there is a very distinct change in behavior  
10 here at this point. Also notice I have tried to plot on here --  
11 the red curve is the fuel inventory in the active core as a  
12 function of time. It starts at about 16/100 kilograms of  
13 driver material, so what's happening is that every one of  
14 these little bursts you can see there is a change of slope here  
15 on the inventory, and it's simply the material being burped  
16 up against the top. Its momentum is actually carrying some of  
17 it into the axial blanket. There's also some fuel removal  
18 going on here through the gaps between the subassemblies of  
19 the internal blankets, and it turns out in this calculation  
20 that those were not very effective; they do not add very much  
21 to the fuel removal, primarily because of the way we've modeled  
22 them, and we intentionally did that because we wanted, again,  
23 this perspective to tilt on the conservative side of things.  
24 We didn't want in fuel removal that couldn't clearly be  
25 defended.

1           So what you see here is a reduction during the  
2 subassembly phase of nearly 13 percent of the inventory.  
3 Roughly speaking, 1 percent of inventory is almost 1 dollar  
4 in reactivity, so you would expect that at that point I have  
5 not removed enough material to keep the system sub-critical if  
6 indeed it puddles at the bottom.

7           This minimal change in behavior here comes about  
8 because at this point, the internal blankets have begun to  
9 break down, become mobile. In other words, the annular pool  
10 space is deteriorated at this point and we're going into the  
11 cylindrical pool phase.

12           This last recriticality actually was able to cause  
13 radial movement of material. Therefore, it does not reassemble  
14 in the same timeframe as it was in the simple up and down motion.

15           The other thing that we see going on here is the  
16 gradual puddling of the core in the reactivity state, coming  
17 back. Now we are basically in this phase, and here to here we  
18 are in this non-homogenized whole core pool situation, and again,  
19 it comes back to critical. And then what happens here is  
20 exactly what we expected would happen. You start with a very  
21 small reactivity event and the thing grows with time. This  
22 has been seen in a number of these kinds of calculations and  
23 it's simply a manifestation of the fact that you've disturbed  
24 the puddle a little bit and it comes back slowly. You disturb  
25 it a little more and the next time it comes back a little

1 faster and then more and more. And obviously, the thing is  
2 growing. This is a log scale so this one last one doesn't  
3 show it very much, but it also would manifest a very large  
4 swing. This is 30 dollars sub-critical at this point; very  
5 large reactivity swings because these motions are very  
6 coherent and on a core-wide basis. But the interesting thing  
7 that we see here in these very mild events, this last one was  
8 sufficient again to drive a lot of material in the axial  
9 blanket. Let me point out that this calculation was done  
10 assuming that this outer circumferential boundary of the whole  
11 core pool was 100 percent solid. No fuel escaped past at all.

12 We will see here in a minute that the gaps that come  
13 in from that boundary -- in other words, the gaps between  
14 subassemblies of the radial blanket and radial reflector  
15 region constitute a very large fuel-removal area. Those gaps  
16 are guaranteed to be open; there is no mechanism to close  
17 them. They do have sodium in them. The sodium has to be  
18 ejected, but all during this period, basically from here to  
19 here (indicating) those gaps would have been opened had we  
20 allowed them to be open.

21 This calculation shows that even without those  
22 large gaps and without control rod fuel removal we have  
23 driven the system down to something like 23, 25 dollars  
24 sub-critical at this point, and note that it does not come  
25 back critical again. We have gone below the criticality

1 threshold of the system.

2           Now, that is completely in line with what we spoke  
3 about earlier. When completely slumped, the non-homogenized  
4 pool -- it has a criticality threshold which is higher than  
5 the homogenized pool. It's around that 20, 25 percent level,  
6 so what this simply shows is that we've achieved neutronic  
7 shutdown here. If there were some mechanism to rehomogenize  
8 and get about 10 or 15 dollars of homogenization reactivity  
9 coming in, then it would bring it back again. Our perspective  
10 at this point is that even this is pessimistic, and if this  
11 had actually been open it would have been clearly shut down  
12 by this point.

13           MR. MARK: In all of these considerations there  
14 is no control rod material, no absorber?

15           MR. BELL: The control rod for this particular core  
16 state, which is into cycle 4 -- six control rods are only  
17 roughly one-sixth of the way in from the top. In other  
18 words, it's basically at the end of the burnup cycle, so  
19 they're nearly all out when the accident starts.

20           MR. MARK: This is what you call an unprotected  
21 accident?

22           MR. BELL: Yes.

23           MR. MARK: But that is not an LOFS.

24           MR. BELL: This is an unprotected loss of flow  
25 accident.



1 (Slide.)

2 So even from this unhomogenized pool phase, what we  
3 tend to see is a damped type of neutronic activity. We see  
4 a large window for fuel removal and, therefore, we don't see  
5 any credible potential for threatening energetics, even during  
6 that part.

7 Here's the point of the instantaneous ramp rates  
8 through that same sequence, and as you can see, they are  
9 pretty volatile. These large negative ramp rates are charac-  
10 teristic of disassembly ramp rates. But in general, what I  
11 want to point out is that the general magnitude of these  
12 things is technically in this 40, 50 dollar range, except out  
13 here at the end, where we see this oscillation building up and  
14 then we see an upward trend in the magnitude of the ramps when  
15 it goes critical again, and that, again, is perfectly to be  
16 expected, that is what would happen.

17 This last one gives gets up to around 70 dollars,  
18 and, of course, we were subcritical by this time so it never  
19 had a chance to do anything.

20 (Slide.)

21 I'll point out briefly the significance of this  
22 flux tilt. If I look at an RC map of the specific power which  
23 is proportional to the flux, here is the center line of the  
24 core. We are going radially outward. This is a situation of  
25 three seconds in that calculation right before the centerline

1 began, and that is the actual flux peaking. A very dramatic  
2 peaking out there in the outer region. In fact, it begins  
3 to look very much like a spherical reactor -- more like a  
4 toroidal reactor, I guess. But this is the kind of power peaking  
5 that has a sort of fundamental damping built into it, because  
6 you can move materials literally in four directions away from  
7 the power peak and get very efficient neutronic feedback from  
8 the fuel motion.

9 Now, just a little bit later in time -- we just  
10 happened to pull this one off after a power burst where the  
11 material had spread out, and you see the complete shift in  
12 the flux shape. There's a very dynamic system when you're  
13 moving materials around to this degree. It's very important --

14 MR. KASTENBERG: R is in the radial direction. What  
15 is C?

16 MR. BELL: This is the bottom of the core and this  
17 is the top axial.

18 MR. KASTENBERG: Okay.

19 (Slide.)

20 MR. BELL: Let's summarize ramp rates, and then  
21 we'll go into the fuel dispersal part of it, and I want to  
22 try to walk through with you what we really think the fuel  
23 removal or fuel dispersal characteristics and possibilities  
24 at least are for this core, even though we, in this calculation,  
25 took it to be very, very conservative in terms of fuel removal

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1 by simply not allowing it to take place. So by summarizing  
2 here in the subassembly pool phase, we clearly see energetics  
3 less than 100 dollars per second. In fact, I guess it's only  
4 in the very idealized situations that we see anything even  
5 approaching 100 dollars per second.

6 In the annular pool phase this situation is essentially  
7 identical, recalling again that the 100 dollars per second  
8 threshold was the structural threshold at which you just begin  
9 to get loads in the primary system on the head structure. You  
10 just fail the UIS at about this condition, and in both these  
11 phases we see the situation has been well below that, well  
12 below that. So it's only in the whole core pool phase then  
13 that one sees any possibilities for threats at all.

14 In the non-homogenized pool, from that calculation  
15 we just showed we have several opportunities there for the  
16 ramp rates to grow, and they did, in fact, grow, but still  
17 never even came close to 100 dollars per second. It's only in  
18 the homogenized pool phase that we have even identified ramp  
19 rates that would get above the 100 dollar threshold. And in  
20 fact, in the one very idealized calculation, we actually did  
21 get to the 300 dollar threshold, but it's non-energetic. So  
22 in terms of energy threats, we see the potential being  
23 maximized out here in the homogenous phase, the pool phase,  
24 which is certainly no startling revelation. But what is  
25 interesting is that we believe from those calculations we have

1 done, it seems to be fairly clearly indicated now that the  
2 only way you take the leap, if you will, from this class of  
3 ramp rate up to this class is by that central compaction into  
4 the middle into the centerline region, and that fundamentally  
5 has to be connected with a low void fraction configuration of  
6 material which, in turn, is non-energetic and basically immune  
7 to ramp rate.

8           So if I were to put a bottom line on that, I guess  
9 what I'm saying is that we simply see no critical threat to  
10 the structure of the system, the head structure, as we have  
11 defined its capability to withstand events up to 200 dollars  
12 per second in the two-phased disassembly mode.

13           MR. KASTENBERG: Charlies, in breaking up the  
14 calculation, or your approach, by looking at these three  
15 phases -- which I think is an interesting and reasonable way  
16 to go -- did you check to see that during the transition from  
17 your subassembly phase to your annular pool phase that you  
18 couldn't introduce something which would give you a high  
19 ramp rate?

20           For example, melt-through cans? Are there any  
21 phenomena that you may have overlooked, or are you convinced  
22 these are the places where if you would have a higher ramp  
23 rate, it would be these places?

24           MR. BELL: We tried to go to the upper limit of  
25 coherence in each phase and see what kind of ramp rates could

1 possibly be manifested from that. I don't know if there's  
2 anything significant there. Nothing has come to mind that we  
3 clearly understand to be a major problem. I don't know --  
4 again, keeping it in perspective, trying to keep fairly close  
5 to reality -- I'm sure we could dream up something, you know,  
6 postulate something that could, in fact, drive us into a bad  
7 situation.

8 MR. THEOFANOUS: Bill, in the transition from the  
9 subassembly to the annular, you enter that through reduction in  
10 the walls, and that tends to be kind of like a quenching. Is  
11 that so? If anything, during that stage I don't believe you  
12 will see the potential for developing any big things but axial  
13 things. And really, you have the same problem there.

14 The only suggestion would be if you go from the  
15 annular pool stage and you introduce a pressure source, you  
16 would --

17 MR. BELL: Yes.

18 MR. CARBON: Theo told us the last time that your  
19 analyses are based on not needing to rely on large computer  
20 codes such as SIMMER. How much of this conclusion are you  
21 really basing on SIMMER calculations of which I'm inherently  
22 suspicious, but how much can you base on reproducing them, so  
23 to speak, back of the envelope or anything that you can check  
24 that's fairly simple, that you can't have much doubt about?

25 MR. BELL: I think that's precisely why we wanted

1 to investigate some of the simple gravity slump configuration  
2 for example, I think that's something we can agree on there's  
3 nothing very mysterious about.

4           The SIMMER calculations indeed have played a big  
5 role in the understanding of the expansion process, but what  
6 we have done -- the material that Theo talked about earlier --  
7 we have intentionally driven it into a mode where it is purely  
8 fluid dynamics and not taking into account any of the fine  
9 details of heat transfer and those kinds of physics; they are  
10 difficult to quantify.

11           Some areas we are using it, we're using it primarily  
12 for perspective and then we're backing it up with some actual  
13 experience in these regimes where things are sensitive, like  
14 the annular pool and the whole core slushing pool. We're  
15 actually doing experiments and comparing the code to those.

16           If we had time we could actually show you some movies  
17 of the calculation and the experiment side by side, and you  
18 could see the report itself does speak to that issue.

19           What we are trying to quantify in particular are  
20 these ramp rates from shes. We have a section in here where  
21 we tried to benchmark  analysis of the fluid dynamics directly  
22 against these experiments. Even that alone doesn't need to  
23 stand by itself because one can go through these differential  
24 K calculations to see how the reactivity of the system is  
25 changing in a static mode. We have done that, too.

1           We have come at this from several different angles,  
2 and I think are viewing it very much on a first principle kind  
3 of basis, so in that sense I feel very confident.

4           MR. MARK: I was going to say next, you should  
5 probably withhold your suspicion of the SIMMER calculation as  
6 used. When SIMMER is trying to discuss drops in fuel pumping  
7 out or not pumping out of a little hole, that is doing a  
8 thing which I think everybody deserves to be suspicious. But  
9 these are rather large-scale things. You do have to do a  
10 complicated calculation to get this two-dimensional fluid  
11 mechanism in neutronics, but they're not -- they're sort of  
12 kinetic calculations, and you use this big machine because  
13 it's the easy way to get the integral of the ramp rate while  
14 the thing is swishing back and forth. You can't do that easily,  
15 you can't do it on the back of an envelope. All you can do is  
16 say well, it will swish back and forth. I wonder how high it  
17 will get.

18           It's not involving the things that I think must  
19 involve your suspicion about the SIMMER framework.

20           MR. BELL: I think in many ways the whole core calcu-  
21 lation does involve those kinds of things that one should be  
22 suspicious about. And in fact, I think we have been duly  
23 suspicious and that's fine. We're not saying that calculation,  
24 as it stands, is the answer. It's simply a perspective on the  
25 kinds of things that --

1 MR. MARK: Now that you've assumed a starting  
2 condition and you're suspicious as to whether it applies, but  
3 once you --

4 MR. BELL: I think what we have learned, if anything,  
5 here is that without some kind of perspective of that type  
6 you can idealize. You're into a very difficult situation  
7 simply because you don't have the perspective that you need to  
8 have to simply integrate all this high, non-linear mess together.  
9 You just cannot generate that perspective in the head.

10 MR. CARBON: Well, it isn't clear to me that a  
11 computer code that I don't fully appreciate how much has been  
12 checked and benchmarked and so on, couldn't give you the  
13 incorrect perspective, so I'm going to ask the question: have  
14 your experiments of which you speak -- are there enough of  
15 those that by themselves they put you on pretty firm ground  
16 in terms of perspective, and are they meaningful experiments?

17 MR. BELL: I think the answer is yes. Theo has been  
18 itching to respond back there.

19 MR. THEOFANOUS: On this question, we take the  
20 approach that SIMMER is a very, very general tool. You can do  
21 a lot of different things with it. Therefore, for such a  
22 thermal tool, to think in terms of it as a final tool thing is--  
23 We would use it for a specific task that is very well identified,  
24 and then we try to make sure that for each particular task we  
25 check it either through other analysis or through particular



1 experiments to make sure that the calculations for this  
2 particular task is done correctly. So it's in this light, for  
3 example, that SIMMER has shown that in this annular pool, the  
4 bubble was breaking up very quickly and producing the sloshing  
5 action. While recognizing that SIMMER does not take into account  
6 all the instabilities that are present there, we were suspicious  
7 as to whether it was calculated correctly. But in fact, we  
8 found very good agreement in the whole process that gives us  
9 a very good feeling of what we can do with it.

10 We have not covered yet, but we have a calculation  
11 now that we can apply to a particular experiment in this  
12 prototypic material and actually looking at this particular  
13 problem. We calculate -- make sure the calculation is  
14 correctly reflected in that. So in that sense, we are taking  
15 a step-by-step approach and making sure that every step of  
16 the way, SIMMER is doing a good job.

17 One aspect, however, I must say everybody seems to  
18 be emphasizing, and correctly so, is the sensitivity to  
19 detail of some of the dynamics here, but one aspect I think --  
20 in fact, we rely a lot on it -- we rely on SIMMER.

21 MR. CARBON: In which aspect?

22 MR. BELL: In the neutronics aspect. If you cannot  
23 do it in your head, you certainly cannot do it with codes any  
24 better, because SIMMER is the state of the art. When you get  
25 it slushing back in, it's not only the motion that's coming

1 back into the picture; also, you need to know what is the  
2 reactivity rate, and what ramp rates are as a result of that.  
3 We think this is the best way we know how to do it. We  
4 don't know of any other method to do it better.

5 MR. LIPINSKI: You've got an unhomogenous mixture,  
6 you're keeping track of fuel and blanket materials, you're  
7 moving around in three-dimensional geometry. You're bringing  
8 us back and you're calculating ramp rates. How do you know  
9 that's being done right? You've got to know where the material  
10 distribution is in the three-dimensional geometry.

11 MR. THEOFANOUS: I'm saying that the controlling  
12 aspect that has given -- I'm saying that SIMMER does the best  
13 job we know of any code in that area.

14 MR. BELL: Actually, there has been some benchmarking  
15 done. NRC actually ran it through the critical experiment  
16 some years back where they took the core and slumped it altogether  
17 at the middle and slumped it out top and bottom, several  
18 different drastic configurations like that, and then various  
19 codes and various methods were used in an attempt to calculate  
20 those various configurations. And if I recall, I think SIMMER  
21 probably came out better than any other tool around in  
22 calculating those criticals.

23 You can obviously do some benchmarking against other  
24 codes that in some ways are neutronically different, constructed  
25 different, but ultimately, you go back to the same fundamental

1 cross-sections and fundamental data. Whatever residual  
2 uncertainties are in that data, of course, are in here, also.

3 MR. CURTIS: SIMMER did very well against those  
4 experiments, but I would say that any transport code that had  
5 a proper group structure in terms of the neutron energy and  
6 enough resolution of the spatial dependence also gives very  
7 good answers. And SIMMER relies very heavily on TwoTran  
8 and the established transport codes that have been in use for  
9 many years.

10 MR. BELL: Characteristically, we've had more confi-  
11 dence in neutronics over the years than we have in the ability  
12 to predict precisely where this fluid is and I think that's  
13 probably still true.

14 If there are no other questions about the energetics  
15 of the ramp rates that we see in the system, I will move on,  
16 then, to the area of fuel removal.

17 MR. CARBON: One question. Back there on your  
18 last slide, in the way you presented it you jumped rather  
19 rapidly from 100 dollars per second phase to 300 dollars per  
20 second phase. You just sort of jumped one to the other as if  
21 there were nothing in between. I don't think you're saying  
22 that here, but it's not clear.

23 MR. BELL: I think what I would really like to do  
24 is let this stand as a general perspective on where the real  
25 bounds are out here, but as I go through this next section

1 I think what you're going to see is that this is a state at  
 2 which we never expect to arrive, and, therefore, we don't  
 3 really need a high level resolution of what's going on.

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*End  
AR-3*

HEMLOCK  
ERASABLE  
COTTON CONTENT

1 (Slide.)

2 Let's look now at this area of fuel dispersement.  
3 If you remember Theo's early introductory chart of this  
4 continuous disruption process, there were two processes  
5 that were competing simultaneously for control of the  
6 accident, one of them being disassemblies off to the right  
7 side, which could bring about neutronic termination, and off  
8 to the left side was dispersal or fuel removal. And the  
9 dispersal or fuel removal process is a more or less continuous  
10 process that begins clear back in the initiating phase and  
11 continues all the way along, and ultimately the termination  
12 mode is not satisfied with one big discharge of material, but  
13 it's the accumulated effect of that discharge occurring  
14 all during this time and with some final point then where  
15 sufficient inventory has been removed that it renders  
16 a system subcritical.

17 So in dealing with this dispersal problem in that  
18 termination mode, there are really three fundamental aspects  
19 that have to be dealt with. The first one being that you  
20 have to have a fuel removal path available.

21 Now, what we will do is go through -- as we go  
22 through the disruption sequence, we will identify which  
23 paths are available, how many of them there are, and what  
24 their general characteristics are for fuel removal.

25 Now once you have a path, you have to be satisfied

ar4-2

1 that indeed you can move material through those paths with  
2 a sufficient throughput to make a difference.

3 In other words, if we put through one gram per  
4 second, it takes us 10 years to get the inventory out, that  
5 won't do us any good. We've got to get it out in the time-  
6 frame between the start of disruption and the achieving of  
7 this homogenous cylindrical pool. That's what we're taking  
8 as kind of the cut-off time window to see if indeed we can  
9 see the fuel removal taking charge before that occurs.

10 Now this then is fundamentally dependent on the  
11 mechanics of freezing and plugging of materials in this  
12 passage, and it is also dependent on pressure. We are in  
13 an environment that is energy-starved in the beginning.  
14 In other words, we have not integrated enough energy in the  
15 system to bring the bulk mass of the core up to a  
16 temperature that would sustain a discharge pressure.

17 MR. THEOFANOS: Any time you want me, I'm here.

18 MR. BELL: Come on. Fortunately, we are inter-  
19 changeable.

20 MR. MARK: This is the opposite of one of those  
21 movies where one guy plays six different characters in the  
22 same role.

23 (Laughter.)

24 MR. THEOFANOS: All right. Going on to the  
25 dispersal, then, next we want to look into each one of those

1 aspects before we integrate all of this together and see as a  
2 function of the core disruption how much fuel gets out of the  
3 core.

4 (Slide.)

5 The first thing is the pressure, and here is the  
6 result of the calculation for the whole-core integral  
7 calculation that Charlie mentioned before. This is a  
8 function of time, and you see the spikes and those  
9 correspond to the activity spikes that we showed before for  
10 the power spikes.

11 What is interesting is not so much the spikes as  
12 that the pressure never gets below 5 bar. As soon as the  
13 pressure tries to go over the material comes back together  
14 and gives it another boost and keeps up the pressure level  
15 at this 5 bar. So from the point of view of obtaining a  
16 perspective as far as how much fuel you can push out of  
17 this core, as the different paths become available, it is  
18 important to know that the pressure does not go all the way  
19 down to zero.

20 In addition, we want to remember that there are  
21 those pressure spikes that, of course, are very effective  
22 in themselves in pushing fuel out.

23 (Slide.)

24 Now something about the bundle availability. In  
25 the beginning, of course, the subassembly walls are intact,

1 and during the tail end of the initiating phase and early  
2 stages of the subassembly pool phase, the only paths  
3 available are the coolant passages.

4 Here I have portrayed the pins. You see the  
5 cladding and these are the coolant passages, and that's  
6 the only thing available for the fuel to get out.

7 Now we can make some very good arguments, we  
8 think, for claiming that for the case of irradiated cores  
9 where we have pressure coming in to interfere with the  
10 sodium, and together with the radial incoherency of melting  
11 of this cladding, we can make some very good arguments that  
12 no part of the core is going to be blocked in these areas  
13 here because of plugged locations. Even independent of core  
14 disruption, for example, for people who don't like to buy  
15 these kinds of arguments. So that as soon as the fuel within  
16 the core disrupts and the cladding in it reaches the  
17 temperature of the fuel and therefore provides the pressures,  
18 we would expect to have an axial motion of this core  
19 disrupted mixture into these areas here.

20 Now even under the worst conditions, however, at  
21 least half of the upper axial blankets is going to be  
22 unblocked, and three quarters of the lower axial blankets  
23 also is going to be unblocked. So even under the worst  
24 conditions we have a lot of space there.

25 (Slide.)



ar4-5

1                   Here is a calculation of penetration of this  
2 molten core disruptive material into the upper axial blanket.  
3 This is a result of the calculation which has been benchmarked  
4 against the experiments. This is where it is pushed, under  
5 pressure, into a structure similar to blankets. In this  
6 penetration in here, we have the picture of the pins. This  
7 is the blanket area, this is the spring area you heard about  
8 before. Penetration and distance, that's roughly 32  
9 centimeters on the blanket. This is a function of the  
10 injection pressure.

11                   Now the experiments we have are at high pressures,  
12 so we benchmark around here, and then we are able to go back  
13 by writing the calculations to see how much penetration we  
14 get at lower pressures.

15                   Well, what you find even under 5 bar pressure,  
16 which is this .5 mega Pascal, we have a complete penetration  
17 of the upper axial blanket.

18                   Another interesting thing that you see in th's  
19 calculation is that the degree of penetration is independent  
20 of the amount of superheat. Somebody might have thought  
21 the more superheat it is, the more it will penetrate, but  
22 you see the two results are very close together. For that,  
23 of course, the cladding is so thin that it very quickly  
24 gets hot, so the amount of superheat plays no great role.

25                   Now in this distance of the length of the upper

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ar4-6

1 axial blanket, by the time you melt the cladding in that  
2 area, and therefore you produce more space, you can  
3 accommodate essentially half of the core fuel in that area.  
4 So you think in terms of the 20 percent you need to become  
5 temporarily subcritical and the 40 percent you need to  
6 become permanent subcritical. You can render the system sub-  
7 critical.

8 MR. MARK: What is the delta TS?

9 MR. THEOFANOS: This is how much the temperature  
10 is higher for the melting temperature, so this is, for  
11 example, the fuel coming in just at its melting point, while  
12 this one is 200 degrees above that.

13 MR. MARK: And this is running into cold areas?

14 MR. THEOFANOS: Right.

15 (Slide)

16 Now let's assume that at this point there is still  
17 neutronic activity as you had before, so the next time that  
18 we are going to have some -- obviously the thing continues,  
19 that means somehow the exits were plugged. The next time  
20 around we have some more possibility when the subassembly  
21 walls begin to fall apart.

22 This is a cross section in the core, and we  
23 have taken a small part and blown it up. What we are  
24 showing here is the drivers with the disrupted fuel, and the  
25 green area here is the internal blankets. Now because the

ar4-7

1 subassembly walls are going to become heated, and as they  
2 become heated, they become more pliable, they will probably  
3 be pushing against each other so that even when they melt,  
4 that's going to cause merging rather than an escape path.

5           When you look at this junction over here, you find  
6 that indeed because the blankets are cold, there is a gap  
7 there, and as soon as the subassemblies crack at the edges,  
8 there will be a path that is open for fuel to go into this  
9 gap, travel axially downwards, and here we are looking at  
10 the lower axial blanket area. Also the driver subassemblies  
11 are not very hot. Therefore, the fuel can go not only  
12 axially but can start going around in a radial motion. So  
13 the fuel goes first down along the interblanket gaps, but  
14 then by the time it goes to the colder areas, it begins  
15 to spread off in all directions.

16           MR. KASTENBERG: Theo, why doesn't it freeze on  
17 the cold green surfaces before it drains down?

18           MR. THEOFANOS: It can freeze, and that is part  
19 of the analysis that we want to show you next. Here I am  
20 trying to show you the paths, and I want you to remember this  
21 number. There are 90 gaps, if we count only the gaps in  
22 the internal area, and we will consider that next --

23           MR. KASTENBERG: At what stage of life are you  
24 looking at of the core?

25           MR. THEOFANOS: If you are thinking in terms of

1 swelling, yes, sure, the swelling is going to be more  
2 at the center of the core, but eventually the subassemblies  
3 also are going to melt in areas. So that's really not a  
4 very big problem.

5 MR. CARBON: Would you go back to the preceding  
6 slide for a moment?

7 (Slide.)

8 If I understand this correctly, it's penetration  
9 vertically upward as a function of injection pressure  
10 and temperature and so on?

11 MR. THEOFANOS: Yes.

12 MR. CARBON: Really, with zero injection pressure,  
13 I presume that penetration upward would be zero?

14 MR. THEOFANOS: Yes.

15 MR. CARBON: And those curves don't indicate  
16 anything like that?

17 MR. THEOFANOS: That's why they didn't go all the  
18 way to zero, Max.

19 MR. CARBON: I know, but is there some explanation  
20 for why it would have almost a vertical slope there at the  
21 start? Or is it that these curves are numerically off and  
22 this isn't correct?

23 MR. THEOFANOS: No. All it shows here is that  
24 you need a very, very small amount of pressure to actually  
25 go through the process here of the cladding or the blankets

ar4=9

1 and then heat loss makes this freeze up; unless you are  
2 unable to do that, you are not able to plug up the channel.  
3 It takes some penetration to be able to do that. So the  
4 tendency here is not just to go in and freeze right away,  
5 but in the beginning you are going to see a very high  
6 sensitivity to the pressure because even if you have small  
7 pressures, you penetrate a lot. But then what you show here  
8 is a rather great sensitivity to plugging out of the  
9 disruption process.

10 There are two ways to look at that. From here  
11 to here (indicating) this thing shows very abrupt, because  
12 this one is very slow. If we plot it differently, this  
13 thing could show like this, and then showing almost like a  
14 horizontal line. See, it is kind of like portraying this.  
15 The reason this looks very abrupt is because that is very  
16 sensitive.

17 MR. DICKSON: What is the unit on the vertical  
18 access?

19 MR. THEOFANOS: This is centimeters. That is 30,  
20 40.

21 MR. MARK: It says meters.

22 MR. THEOFANOS: I know it says meters. That is a  
23 mistake. I am sorry.

24 MR. CARBON: So it should be centimeters?

25 MR. THEOFANOS: Right. That is what we have here.

1 This is the upper axial blanket, and that is about a foot,  
2 30 centimeters.

3 (Slide.)

4 Here we will go to the freezing dynamics.

5 What we have done here is to put in a pressure,  
6 three pressures and the temperature, and considered how much  
7 mass is going through those gaps as a function of time.

8 What you see here -- let us take the green one  
9 first. It increases all the way out to more than 22 and  
10 then very quickly goes down as a result of this still  
11 coming in from the wall of the subassembly and making a  
12 particulate that actually slows down the motion.

13 Now at this time you see that almost complete  
14 plugging has occurred except that it did not quite stop.  
15 It just oozed out and eventually the thing picked up again,  
16 and after this particulate comes out of that oozing stage,  
17 actually the passage has become now bigger and you have  
18 got a passageway much longer for fuel to go through that.

19 The blue line at 3.4 atmosphere and 3400 degrees  
20 Kelvin gives us up to about 10 or 15, and then slows down,  
21 again tails off a little bit, and then takes off like that.  
22 The 3.4 atmospheres and 3100 K goes up to 10 kilograms per  
23 second, and it just gradually slows down asymptotically.

24 The point here is that within one second or half a  
25 second, you might want to claim 20 kilograms or 15 kilograms

1 per second. You read those things for pressures of 3.4  
2 atmospheres.

3 MR. MARKS: These kilograms, 20 per second --

4 MR. THEOFANOS: Per gap. So we have to multiply  
5 those numbers times 90, times the total amount of kilograms.  
6 Just to do it very roughly, that's why I told you about this  
7 15 kilograms per second.

8 MR. BELL: I just want to point out that the red  
9 curve is sort of the classical fuel freezing and occlusion  
10 problem where it is pure fuel at its melting point going  
11 into the gap channel. That calculation was actually checked  
12 against a theoretical solution which it can attain for that  
13 situation and matches very, very well.

14 MR. THEOFANOS: Also, the other calculations,  
15 these other calculations are also benchmarked against the gap  
16 data at Argonne and they also -- in fact, that's how the  
17 numbers were fixed.

18 MR. LIPINSKI: How long are the gaps?

19 MR. THEOFANOS: Well, half of the core length;  
20 about half a meter. That's it. Most of the injection point  
21 is in the center. Like Bill said, if there is some  
22 swelling there, then it would be just maybe 10 or 20  
23 centimeters.

24 (Slide.)

25 Further into the core disruption stage, we have

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1 one more class of gaps, so to speak, coming in and they are  
2 coming in because of radial blankets. These are radial  
3 blankets all around the core and there is a lot of warmth  
4 associated with them.

5 This is a driver that opens up between these  
6 radial blankets and now the flow can go radially out this  
7 way, but of course also can axially down now into those areas  
8 through those gaps axially down. So you get a three-  
9 dimensional flow that quickly develops as soon as the  
10 corners open up here.

11 In fact, what we think is that by that point  
12 the gap areas become so many that the problem is completely  
13 overwhelmed by the number of passages that are available  
14 for the fuel.

15 This is indicated here by saying that the  
16 discharge area is 10 times greater than the area associated  
17 with the opening up of the internal blankets.

18 Now I also want to point out that the availability  
19 of those paths will be right in the beginning at the  
20 annular pool stage as soon as the subassembly walls are  
21 falling apart, and the annular pool phase begins to form.  
22 That is when those become available.

23 (Slide.)

24 Here is a summary. Because we not only wanted you  
25



1 to know how many gaps or how many parts are available, but  
2 also we need to know how much the coolant volumes can hold,  
3 and here this is summarized for you, together with the  
4 times at which we have access to those. The time zero  
5 right at the beginning of the core disruption phase following  
6 the initiating phase, upper and lower axial blankets within  
7 the pin structure, we are talking about capacity in terms of  
8 percent, how much is core pool, 12 percent in the upper and  
9 25 percent in the lower. The percent of removal now under  
10 5 bar pressure is on this list here, and you said that in  
11 order to remove these 12 percent, you need about two seconds  
12 to do it; to remove 25 percent, you need about two seconds.  
13 That's about the amount of time you need to begin to fall  
14 apart -- for the walls of the subassembly to fall apart, so  
15 the internal blanket gaps begin to open up.

16           Excuse me. I think I made a mistake. This  
17 capacity is consistent with this timing up here. The total  
18 capacity is more like 50 percent.

19           MR. BELL: I just want to clarify the difference  
20 between the upper and lower axial blankets. You notice that  
21 one is just twice the other and we purposely derated the  
22 upper axial blanket by a factor of two because in the long  
23 term that fuel could fall back into the core. When it falls  
24 back in with the blanket material coming with it, it has the  
25 neutronic effect of having half of that fuel removed. In

1 other words, that dilution effect of the blanket is the  
2 same as removing half of that fuel. So that is simply derated  
3 to take that into account.

4 MR. THEOFANOS: In the internal blanket areas, we  
5 have the time of access on the order of 2 seconds, and in  
6 the lower axial blanket area we can accommodate about 10  
7 percent of the fuel, and in a matter of this time we can put  
8 in 21 percent under 5 bar pressure, which is always going to  
9 be there. And the radial blanket gaps, the rate of removal  
10 and percent per second is 200 percent over here. In both  
11 cases, the access time is 2 to 4 seconds, and the capacity  
12 of the homogenous fuel that we can remove in the time  
13 that is available is about more than 10 to 40 percent.

14 Now remembering that we need only 25 percent at  
15 this stage here to keep it temporarily subcritical, and you  
16 need 40 percent at this stage here for permanent subcritical,  
17 you realize that any combination of those numbers is going  
18 to give us -- in those numbers over the periods that are  
19 indicated here -- to give us the room that we need in order  
20 to assure there will be permanent subcriticality at the time  
21 of entering the heterogenous pool.

22 (Slide.)

23 Here is one example of how we have tried to do  
24 that. Just to show you the bottom line in the upper axial  
25 blanket, an adjusted rate of 3 kilograms per second, taking

1 into account the effects you just heard, the discharge time  
2 of 5 seconds, we can remove in the lower axial blanket 6  
3 kilograms per second times the time of 5 seconds. The lower  
4 blanket is the only one that counts. About 10 percent in  
5 the radial blankets, 10 percent in the reflector, 40 percent  
6 in the control center. You're hoping it's more like 100  
7 percent.

8 That kind of gives you an idea of how much margin  
9 is there for removing fuel from the volumes that we are  
10 looking for.

11 MR. ZUDANS: In that case, up there where you  
12 list 40 percent --

13 MR. THEOFANOS: Here. This is the radial reflector.

14 MR. ZUDANS: You can leave that much to the gap,  
15 but if you bring that material to that location, that is  
16 associated with another tortorous path, so this is not the  
17 real number. I mean you have to get it sitting there to be  
18 discharged at the gap. It has to come to that location.  
19 That's half a core.

20 MR. THEOFANOS: That's only about that far apart.

21 MR. ZUDANS: Can you get there at the same time?

22 MR. THEOFANOS: Sure. Why not?

23 MR. ZUDANS: If you say so.

24 MR. BELL: Typically these discharge velocities --  
25 the material can move the distance of a meter in a second, and

ar4-16

1 and we are talking about a fraction of a meter, too.

2 MR. ZUDANS: But does it have the passage open to  
3 do that? In order for it to be open you would have to have  
4 the whole core, the whole thing.

5 MR. BELL: No, these gaps that he is talking about  
6 are in the radial blanket reflector which are not being  
7 heated. They are in fact open.

8 MR. ZUDANS: But to discharge 40 percent of your  
9 core through there, you have to bring that 40 percent to the  
10 gap?

11 MR. BELL: Oh, yes, it's sitting right there all  
12 ready to go, if you remember our presentation.

13 MR. THEOFANOS: It's about so far apart.

14 MR. ZUDANS: There's no further distinctions in  
15 the core?

16 MR. THEOFANOS: Maybe I can put the picture  
17 back on.

18 MR. BELL: The only thing is the blanket gaps,  
19 and the blanket gaps are also cold, and they are going to be  
20 filling up also first, and that is the first number that you  
21 saw there.

22 (Slide.)

23 MR. THEOFANOS: Now we have an overlay here.

24 MR. CARBON: If you feel this is the start of a  
25 new topic, perhaps it would be a good time for a break.

1 MR. THEOFANOS: No, this is the summing up of what  
2 we have got covered up to now, and then we are going to get  
3 to a new topic.

4 (Laughter.)

5 The completion of this detailed discussion. Now  
6 we want to go back to the origin, where we started from.  
7 This is the picture I showed you, and now we want to go and  
8 put those numbers that we talked about, the probability  
9 numbers, and since everything is fresh, we should do that  
10 before the break.

11 MR. KASTENBERG: Ours doesn't overlay, Theo.

12 (Laughter.)

13 MR. THEOFANOS: Next time I will remember.

14 MR. KASTENBERG: It looks like a "follow the dots."

15 (Laughter.)

16 MR. THEOFANOS: I have a lot of confidence in  
17 your imagination.

18 I will start from the beginning. The initiating  
19 phase, the disruption phase, as you have already -- we  
20 expect zero energetics. There is no way of producing  
21 energetics from that, and we therefore put on this  
22 probability  $10^{-3}$  because of the fact that we believe  
23 energetics in this stage are physically unreasonable. In  
24 attaining a permanent subcriticality, that is discharging  
25 40 percent of the fuel, we also consider that to be an

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1 unlikely event, not so much because we cannot see ourselves  
2 with the core not plugging up, but because we can see some  
3 situation under which some exits of the core might be plugged  
4 up. That is the reason that we give a number of  $10^{-1}$ . That  
5 means the main path will be going now from the initiating  
6 disruption of the pin structure to the subassembly pools.

7 Now, as you heard from Charlie a minute ago, the  
8 likelihood of obtaining significant energetics is very low.  
9 We assigned a  $10^{-1}$  number mainly because we think -- we  
10 might have factored in energetics in this assembly, but this  
11 path only indicates the potential for attaining a disassembly,  
12 not of failing the vessel. This over here is -- this is  
13 the age of spectrum condition, conservatively also under  
14 this step, the age of spectrum condition again, because we  
15 can see that maybe it will take more than this 1 second  
16 that is available from this point to that point before you  
17 fail all the walls to discharge 40 percent of the fuel. So  
18 that seems to be  $10^{-1}$ , so again the main path would be to go  
19 to the annular phase.

20 Now at this stage again the energetic potential  
21 is very similar to that, we will say maybe slightly higher,  
22 and again that is a spectrum. However, now, in going from  
23 this stage to this stage, we have the availability of the  
24 radial blankets. As you saw the problem, it was becoming  
25 completely overwhelmed by these openings, and that is the

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1 reason for the  $10^{-2}$ , assuming it is considered to be outside  
2 the spectrum.

3 So the main path, then, the way we see it, is going  
4 to go from the initial disruption of the pin into subassembly  
5 pools into the annular pool and exit into a milder  
6 specialty from there, and that is conservative, with less  
7 likelihood you could do, and with even less likelihood it  
8 would do that. Assuming that we got in a whole-core  
9 homogenous pool, still again, of course, we would have  
10 the ability to do the paths and the probability would be  
11 mainly this one, as a spectrum condition, to attain it for  
12 this stage.

13 Now I want to follow each of these energetic  
14 paths and see what the probability is and the corresponding  
15 outcome of the disenergetic assembly.

16 This part here is almost not here. Similarly  
17 this part here is considered to be physically unreasonable.  
18 The subassembly pool was all very well bounded by \$100 per  
19 second. We have all the additional probability of  
20 another \$100 per second, and we think it is unreasonable  
21 physically to fail the vessel. Similar things apply for  
22 the annular pool. Again, the outlays were bounded and  $10^{-3}$   
23 was for this path over here.

24 Finally, for the whole-core pool stage, as you saw,  
25 we had bounds. There is a level required to fail the reactor

1 head.

2 On the other hand, we recognize that really this is  
3 a rather involved kind of situation. Max, you mentioned that  
4 we only explored the one case that was two-phase. It is a  
5 multi-dimensional kind of a problem. We cannot really have  
6 the confidence that one has with those two stages for not  
7 failing the vessel, and it is for this reason that we gave  
8 the number of  $10^{-4}$  for the probability of failing the vessel  
9 for this kind of an event.

10 So this has completed this story. Obviously we  
11 have to multiply this  $10^{-6}$ , multiply this path  $10^{-4}$ , multiply  
12 this path  $10^{-4}$  and finally multiply this path here another  
13  $10^{-4}$  and sum them up, and the total is  $3 \times 10^{-4}$ . That is  
14 well below the  $10^{-3}$  that you considered physically reasonable,  
15 and therefore we come to the conclusion that for loss of  
16 flow accident, the conditional probability to fail the head  
17 is  $10^{-3}$ , but then converting it back to physical reasoning  
18 is physically unreasonable and not expected.

19 MR. DICKSON: Question, please Were those  
20 numbers made with the fuel pins in their present configuration?

21 MR. THEOFANOS: If you're asking me with respect  
22 to the gas compaction, you see, we put a  $10^{-3}$  up here.  
23 That means physically we wouldn't get any energetics in that  
24 stage. So that is assuming that we didn't have the plenum  
25



ar4-21

1 gas fission problem. If we did have this problem, we would  
2 have some problem knowing what to put here with a very high  
3 degree of confidence, and the way we want to bring this  
4 across is that we have a very high confidence level for all  
5 those numbers in the conservative side, and we can claim we do  
6 that in every step of the way here. If we had that problem,  
7 we might have some difficulty really assigning any numbers. We  
8 just wouldn't know what to put.

9 MR. KASTENBERG: Theo, just a couple of questions.  
10 Let me see if I can phrase this correctly:

11 If you end up in the disassembly box, you have to  
12 go somewhere, so do I interpret the green line as being one  
13 minus?

14 MR. THEOFANOS: Yes, that's what it is, but I  
15 didn't put the numbers on it to avoid the complication.

16 MR. KASTENBERG: So it is basically one?

17 MR. THEOFANOS: This plus that should be.  
18 That's what I did here. Those are all the leftovers from  
19 those parts. So it would be very close, even if it got through  
20 that path.

21 MR. KASTENBERG: Two other things. This assumes  
22 that if you, for example, went on the gamma gamma prime path,  
23 that those probabilities are not dependent on one another,  
24 yet they are in a sense? I mean they are end of spectrum and  
25 if you are going on that path, you'd have to complete that path,

1 wouldn't you? In other words, you wouldn't come across  
2 gamma and go down on the green line, you would go down on  
3 the red line which says gamma prime? No?

4 MR. THEOFANOS: No, no, no. Remember, this is only  
5 there for processes of qualifying for an exact determination.  
6 An exact determination means we have a high pressure  
7 developed because of an event, and as a result of that, the  
8 fuel somehow gets out of there and that is the end of it.  
9 Now having gone through this -- that's why this is in the  
10 process now.

11 Now at this point we can ask the question, is  
12 this a necessary event sufficient to fail the reactor  
13 vessel head? And if it is, we end up with this fact? Or  
14 is it not? And if it is not, we are not in this box which  
15 means the fuel is some place inside the reactor vessel and  
16 dispersed.

17 So in fact you go through this, then you split  
18 into two parts, depending how energetic the event was.

19 MR. KASTENBERG: Well, I have to think about it  
20 all, because it just seemed that once you embark on one of  
21 those Greek letters, your process becomes dependent upon  
22 what you have assumed to get to that stage and become  
23 independent.

24 In other words, if you're end of spectrum, you are  
25 end of spectrum all the way? No?

1 MR. THEOFANOS: No, no, no, because these  
2 numbers here, Bill, they are not -- well, those numbers  
3 really should be distribution probability. If they were,  
4 what we are saying is you could associate one part of the  
5 distribution with another part of the other. But those  
6 are boundings, so what we are saying is that out of all these  
7 events that can get us into this state, only we could have  
8 end of spectrum events to get us there. That means we count  
9 all the processes that actually qualify for hydrodynamic  
10 disassemblies.

11 Now that's all of them there.

12 Now we have to go there and pick out some of  
13 them, if there are any that have failed the vessel, and  
14 we have to count those against the previous ones to find  
15 out what is the likelihood of going through that. And all we  
16 are saying is that even though we see some edge of spectrum  
17 situations bringing us here, we think that we're not going  
18 to find any among all those disassemblies that can cause vessel  
19 failure.

20 So it is dependent, but we do that dependency on  
21 the basis of the technical material that was provided before  
22 we look at the whole picture and say, "Can we find any  
23 situation that can get us to that," and assigned a number.

24 In other words, we don't try to do that in a very  
25 detailed fashion because in almost all cases we couldn't find

1 any cases, so we couldn't limit it or identify one.

2 Any other questions on this?

3 MR. ZUDANS: You say you couldn't find any, but  
4 how do you dispose of that \$300 insertion?

5 MR. THEOFANOS: Because it's single phase and  
6 disassembles quickly. It's much less energetic.

7 MR. ZUDANS: In other words, single phase?

8 MR. THEOFANOS: Yes. In fact, that was much less  
9 energetic than some of those cases up here.

10 MR. ZUDANS: Then the highest number you got on  
11 the energetics was 18, something like that?

12 MR. THEOFANOS: We got about 80 or less, and really  
13 80 is ridiculously conservative. I think the numbers here  
14 should be more like 30 or 40, and I think the closest thing  
15 that we came to producing significant energetics was taking  
16 the whole-core pool homogenized, perfect symmetry, putting a  
17 source in the center, perfectly symmetric, then allowed to  
18 have a perfect symmetry of slosh, and on that gave us \$100  
19 per second. It's really putting a lot of things on top of  
20 the other, and even that one only gave us the equivalent of  
21 what would be required on the basis of our first discussion  
22 this morning to only fail and still be very far from doing  
23 anything to the vessel head.

24 MR. KASTENBERG: I don't want to harp on this, but  
25 I think it's important to understand something, because it

1 was your second major bullet, of certain things that were  
2 physically unreasonable by this thought process and I want to  
3 make sure I understand it.

4 If I go through and I look at what your best  
5 estimate, so to speak, I go one and one and I am at an  
6 annular pool, and your best estimate is that you will get  
7 dispersal with the annular pool, tell me again what does  
8 your gamma mean when you get into the minus one?

9 MR. THEOFANOS: That is the conservative best  
10 estimate. I'm bounding it along the way.

11 MR. KASTENBERG: But what does the gamma mean,  
12 going the other way?

13 MR. THEOFANOS: It means we expect it is not  
14 totally impossible that we might have an initiating event  
15 from this stage.

16 MR. KASTENBERG: What is the physical process  
17 that would move you along the  $10^{-1}$  gamma path? What is  
18 the end of spectrum process?

19 MR. THEOFANOS: That would be some criticality of  
20 the type you heard previously from Charlie. Any of those  
21 are criticalities. This \$50 to \$80 per second bounding  
22 estimates, they will be doing that, and I do believe there  
23 is some possibility for those things to happen.

24 MR. KASTENBERG: Suppose that did happen, then  
25 what is your physical process that would take me along the

1 delta prime?

2 MR. THEOFANOS: You mean the gamma prime?

3 MR. KASTENBERG: What is the physical process  
4 that would take me that way?

5 MR. THEOFANOS: That will take you here? The  
6 physically unreasonable process would be -- well, you know,  
7 physically unreasonable doesn't exist. That's why I can't  
8 tell you what the process is.

9 (Laughter.)

10 MR. KASTENBERG: This is what bothers me about  
11 the schematic is that I see the --

12 MR. THEOFANOS: If I had a way, Bill, of telling  
13 you that I had to assume this and that, and I will get the  
14 vessel to fail and we have told you what that was, but we  
15 don't have that. Now we say -- in fact, we have seen the  
16 whole-core pool coming in, giving us reasonable energetics,  
17 not very high or not very low, and we can see how one can  
18 maybe go back in the calculation and can push a few things  
19 back and forth and make energy levels come to this level  
20 that we need to fill the head. But in those cases we really  
21 don't see any physical process to cause that, and we have  
22 to represent that in some numerical fashion, so we assigned a  
23 number of  $10^{-3}$  for the process which we cannot see how this  
24 can happen. That is what we defined as probability level to it.

25 MR. DICKSON: Could I try to paraphrase it and see

1 if I've got it right? When you are going the horizontal  
2 gamma line at  $10^{-1}$ , that's the \$40-\$80 ramp rate that you say  
3 probably won't happen, but is within the spectrum? When you  
4 get on down the vertical gamma line, you'd have to have a  
5 couple of hundred dollars a second, and you don't see how that  
6 could happen.

7 MR. THEOFANOS: That's right. It's outside of the  
8 spectrum because I have bounded it before at this point, and  
9 in order to get \$200, I must do something outside of the  
10 real possibility.

11 MR. KASTENBERG: And what is the physical process,  
12 going back to your annular pool again? Because that's where  
13 you are most likely going to end up dropping that  $10^{-2}$   
14 to complete disruption. What is the physical process that  
15 might take you down that line?

16 MR. THEOFANOS: The physical process would be if  
17 somehow it was possible for these radial blankets to not  
18 open up. We cannot see how this can happen because the  
19 same process that is homogenizing the pool and melting all  
20 the internal blankets, not only the walls, but also the solid  
21 materials, and that is why we are putting outside the spectrum,  
22 because we can't see how this can happen. But you see  
23 what the process would be, that I postulate all the outside  
24 walls remain intact --

25 MR. KASTENBERG: Even with that, you feel you

1 would go to the dispersal?

2 MR. THEOFANOS: O'h, yes. That would mean more  
3 time available, because every time you are pushing your time  
4 limit, you must push harder and harder for getting material  
5 out.

6 By the way, as you saw from the whole core  
7 calculation, the transition from this stage to that stage  
8 is not a very instantaneous thing. It takes some time  
9 before you homogenize everything, and if you ask me, that's  
10 one of the most interesting things learned, was that part  
11 of it.

12 MR. BELL: Theo was pointing out one possible way  
13 to defeat the paths to the left. It would be those gaps  
14 where available to come down to the next stage, and to defeat  
15 that you'd have to also say, for example, the control rod  
16 removal wasn't available, or other modes of removal were not  
17 available. In addition, you're compounding the sort of  
18 incredible nature of the whole thing.

19 MR. THEOFANOUS: Any other questions?

20 MR. CARBON: Let's take a break.

21 (A short recess was taken.)

22

23

24

25



1 MR. CARBON: Move right on, Theo.

2 MR. THEOFANOUS: We have completed the discussion on  
3 this, and now we want to cover two more accidents, the transient  
4 overpower or the TOP, and the loss of heat sink.

5 The TOP, of all the accidents, it is looked at, and  
6 in fact the applicant has spent a little time on it. The loss  
7 of heat sink accident is all the more moderate, somewhat more  
8 neglected in the past, probably not for bad reasons. But we  
9 thought we would take a look at that accident also.

10 I think I will give you the bottom line from now that  
11 we see no energetic behavior in any of those accidents, but  
12 maybe you are still interested to know what the story is  
13 or what story we are putting together.

14 (Slide)

15 Now, then, this is the accident that is characterized  
16 by very, very low heating rates. Now, remember, about 1 degree  
17 per second, and because of that there is a lot of time for  
18 recovery, and this is something that is not really covered in  
19 the probabilistic risk assessment studies adequately, in our  
20 opinion. There is not enough credit taken for recovery from  
21 this accident. Therefore, maybe the probability of them getting  
22 into very extended core disruption is slightly overexaggerated.  
23 In this case, if one postulates, extends an unmitigated loss  
24 of heat sink, eventually the core will disrupt and the way that  
25 this is going to happen will be only following uncovering of

1 the core sodium. There is no way it can start melting while it  
2 is covered with sodium. And that is very similar to what is  
3 happening in the reactors.

4 Yes?

5 MR. ZUDANS: Could you quickly explain to me what is  
6 the difference between loss of flow and loss of heat sink?

7 MR. THEOFANOUS: Well, the loss of heat sink is in  
8 the -- going back to the first vuegraph I gave you -- is a  
9 protected accident. Remember, no power generation, very  
10 quickly within a few seconds is down to a few percent. And  
11 somehow you lost capability of removing heat from the system.  
12 That can happen because of a large earthquake and you sheared  
13 off the pipes or it could be a particular problem with the  
14 heat exchanger.

15 MR. ZUDANS: And loss of flow also as a reactor  
16 shutdown?

17 MR. THEOFANOUS: No. The loss of flow we have been  
18 discussing here is an unprotected accident. The pumps were  
19 closed down, shut down, the reactor protection system, the  
20 control system, has failed to act, so the power is at a high  
21 level. So that is a very, very unique, very, very different  
22 situation.

23 So as long as sodium is in the core, the core will  
24 remain intact, but that sodium will -- that will happen in a  
25 matter of a few hours or more like close to 10 hours. And

1 then if the primary system holds together at this very, very  
2 high temperature, I will take quite a long time to vaporize  
3 all the sodium in the upper plenum before the core actually  
4 begins to get uncovered. And that is 100 hours to do that.

5           However, at these high temperatures, other things  
6 happen structurally in the system that might cause the emptying  
7 of the vessel of the sodium before it has an opportunity, all  
8 of it, to vaporize. In the beginning -- and I think I brought  
9 up in the discussion in November -- one of the things we were  
10 concerned with is maybe the high temperature creep of the  
11 vessel wall might cause creeping and vessels fall off. We  
12 examined that, considering some recent data at the high  
13 temperatures for creep. And I found that the times were much,  
14 much longer than the time frames that other things would be  
15 happening.

16           On the other hand, there is something else that is  
17 of similar nature. As the vessel wall heats up, it will  
18 expand downwards because it is so far from the top. Similarly,  
19 the guard vessel is going to be heated up, and that will be  
20 expanded upward because it is so far from the bottom. There is  
21 not enough clearance here as this nozzle goes into the vessel,  
22 so the one vessel is going to be interfering with the other.  
23 And some preliminary assessments -- and I think the applicant  
24 has done work here, and we have come to the conclusion there will  
25 be some structural failure associated with this interference

1 between those two structures.

2 Well, if that is going to happen, the vessel is going  
3 to empty of sodium and the core will remain on the body. If  
4 that were not going to happen, there is not any difference  
5 because 100 hours later there is enough sodium. You talk about  
6 you get to 45 minutes and it does not affect the conclusions  
7 or the way that you go about analyzing the core disruption  
8 independently of what happens.

9 In any case, this is the range of the heating range  
10 that we can expect to be present at the time of core -- beginning  
11 of core disruption.

12 (Slide)

13 Now, I would like to follow and see how that core  
14 study begins to fall apart, and our interest is to see if there  
15 is any potential way. Obviously, in the absence of coolant,  
16 there is no way that one can associate any of that initiating  
17 energetics that were talked about before with this kind of  
18 scenario here, so here we have pictured inside the case, the  
19 structures, because of the slow heating up of the fuel, there  
20 is ample time for the cladding to keep pace with it and  
21 because of the 1,000-degree temperature difference between  
22 the melting of the fuel and the melting of the cladding, there  
23 is ample time for the cladding to melt and just drain out.

24 There is no forces to move it upwards, so all the  
25 cladding from the core is going to melt, is going to fall down,

1 and is going to form a very massive steel blockage in the  
2 entrance to the core. And in fact, if it's going to block up  
3 here, in fact, some of the cladding is going to make it all the  
4 way, and some of it will be inside the core region itself.

5 The other interesting aspect of this is that the  
6 upper axial plant cladding, because of the proximity to the  
7 active core region itself, is going to melt quickly and fall  
8 down. So the upper axial pellets, they are going to fall on  
9 the top of the rest of the core.

10 Now, the core at that point is very hot. The pellets  
11 are cindered together. They might stay together. But even if  
12 they don't, if they shake up a little bit, they might make a  
13 random rubble bed. But in any case, we have a solid packed  
14 configuration. It is like a packed bed and is far from being  
15 fully compacted.

16 We have typically a worth fraction of 40 to 50  
17 percent.

18 Then the next question that we had was will the upper  
19 core structure melt? Again, because of the long time duration  
20 of this thing and the very, very small heat, obviously, there  
21 will be radiation, heat transfer from the core into this melted  
22 -- into this steel area here.

23 Well, our conclusion was that this will not melt,  
24 because of two reasons: Number one, this layer of insulating  
25 pellets or the blanket that is not heat producing and has

1 very poor thermal properties. Then we did some radiation  
2 calculations and found out, in order to be able to produce the  
3 heat fluxes to melt the area here, one would have to have  
4 temperatures of the order of 5,000-6,000 degrees in the  
5 center.

6 Of course, that goes well beyond the fuel melting  
7 point. Therefore, the picture is one in which the steel melts  
8 out. That is why we call this steel melt-out. The steel heats  
9 up slowly. The upper axial blanket falls and forms a rubble  
10 bed on top of that. And the upper core structure remains, or  
11 even maybe might be resting on the top of this bed. We will  
12 show it here for clarity. But the whole thing might be resting  
13 on it.

14 And, of course, the UIS is still sitting up there.  
15 Now, then, the temperature continues to increase at this low  
16 rate, and the fuel will begin to approach the melting point,  
17 and it will begin to stack. So it will start becoming smaller  
18 and smaller, being absorbed by the siding of that oozing fluid,  
19 pellets. And all this time it will remain -- well, it's  
20 subcritical, but it will approach criticality because, as you  
21 know, we know that something of the order 10 centimeters of  
22 total reduction in height of the core would be sufficient to  
23 cause criticality with all of the control rods being inside the  
24 core, which will be the course in this accident. So that is  
25 one of the ways in which you approach criticality is by this

1 gradual siding.

2           The criticality approach will begin to produce heat,  
3 and you have already approached the melting point. Right around  
4 the flux area. The fuel begins to melt. And of course, this  
5 melting begins. The weight of all that bed is going to try to  
6 push its weight downward. The weight that we are picturing in  
7 this situation, the weight we picture is as follows.

8           (Slide)

9           Just like suddenly the center of this core becomes  
10 molten and it cannot carry the weight anymore, and all the rest  
11 will fall. It falls under gravity. Similiarly, the fact that  
12 in going from critical to prompt critical we require only 1  
13 or 2 centimeters' displacement. Therefore, there is only so  
14 much acceleration that you can achieve. And by figuring out the  
15 upper limit, the whole thing falling under gravity in this kind  
16 of a distance, and that by the worth curve which is about \$1  
17 per centimeter, we come up with a maximum to operate of \$60  
18 per second.

19           Now, as you know, this would not have any concern  
20 even were the sodium melted. There is no sodium to transmit  
21 any of the forces. Of course, that is of total neglect. Really  
22 is nothing about that at all.

23           If this was to happen, it obviously would be enough  
24 to dispense the core, and one can claim this would lead to a  
25 termination. On the other hand, remembering that the upper core

1 structure is still sitting on top and the UIS also on the top,  
2 and remembering that this melting is not likely to happen  
3 across the core altogether but more like in the center of the  
4 core because of the radial peaking, one would expect more like  
5 maybe a fraction of that materializing. Something like maybe  
6 \$10, \$20 per second. And if that was the case, one would not  
7 obtain termination at this point. And then the question is,  
8 as you go through this burst, the question is -- and of course,  
9 as you go through that, the whole thing becomes molten, so you  
10 have the whole core cool situation.

11 Can one obtain from that the higher activity rate  
12 than before? Well, the immediate reaction of one is that, as  
13 was ours, well, this is a case already for loss of flow case  
14 and so we can forget about that. Well, this is not the case.  
15 There is quite a bit better than the other one, because now we  
16 have quite hot radial blankets.

17 Now, the upper blankets are sitting up here, and  
18 because of this, the tendency to homogenize the radial blankets  
19 into this molten situation is much easier. And this will be  
20 enough to produce permanent criticality because from that --  
21 and I want to show that just from that.

22 So in the event we did not obtain permanent  
23 recriticality from this initial burst, there is a bubble.

24 (Slide)

25 It will tend to push things up for a time and then



1 bring them back down. In doing so, there will be a tendency  
2 for this blanket to come in and homogenize and the other radial  
3 blankets to come in and homogenize. To give you a little  
4 account that Dr. Carbon was asking about, here is the activity  
5 balance. If we homogenize the upper axial blanket, they are  
6 worth about minus \$20.

7 If you homogenize the internal blankets, they are  
8 worth about \$20. So if those two were to come in together, the  
9 effect will be not really very high activity state in the  
10 system. To homogenize half of the radial blanket, it will have  
11 the effect of minus \$40. To remove the control material has  
12 the effect of plus \$30. These two things will tend to happen  
13 at about the same time scale. The radial blankets will come in  
14 at the time that the other material is being vaporized. And to  
15 completely puddle the core, we require plus \$10.

16 As you see here, the sum is zero, and all of those  
17 processes happen together, and we cannot see that core  
18 achieving configuration stage in which the potential for higher  
19 activities is present following that initial process.

20 So the way that we see the process going on is  
21 following this series of mild recriticalities we can mix all  
22 the blankets in and put all the material together and being  
23 eventually all of the steel up there that, of course, will have  
24 to be melting and coming in and end up with a permanent  
25 subcritical puddle simply because radial blanket material can

1 mix well and doesn't separate, of course.

2 So the best-estimate discussion here, the best-estimate  
3 result is, of course, no significant energetics at all and  
4 permanent subcriticality by dilution with other materials coming  
5 into the core.

6 MR. ZUDANS: Theo.

7 MR. THEOFANOUS: Yes?

8 MR. ZUDANS: Do you assume that the sodium was lost  
9 completely?

10 MR. THEOFANOUS: Well, that is a long time ago. It  
11 has to be before anything melts.

12 MR. ZUDANS: Did the guard vessel go?

13 MR. THEOFANOUS: Regardless of what?

14 MR. ZUDANS: Did the guard vessel, did it go too?  
15 Because it is supposed to keep the sodium level about the core.

16 MR. THEOFANOUS: Do you remember, that was the  
17 interaction between the two? That is one of the possibilities  
18 is either the sodium gets out because of the failure, because  
19 of the interaction of the guard vessel and the vessel itself  
20 or the other one is if it doesn't happen, the core is going to  
21 sit there until all the sodium vaporizes and gets into the  
22 containment.

23 MR. ZUDANS: I am just wondering, is there a  
24 sufficient way to get it out of the guard vessel so that it  
25 ceases to flood the reactor?

1 MR. THEOFANOUS: If the guard vessel fails, you can  
2 get it out.

3 MR. LIPINSKI: You didn't fail it?

4 MR. ZUDANS: It still leaves sodium in the reactor  
5 tank. That's what I understood now.

6 MR. THEOFANOUS: Let me show you the picture here.

7 MR. LIPINSKI: All you did was dump sodium in the  
8 guard vessel.

9 MR. MARK: It's a place where sodium can run out.  
10 He will boil it off.

11 MR. LIPINSKI: Where is he boiling it to?

12 (Slide)

13 MR. THEOFANOUS: I think the first thing that we need  
14 to understand is unless sodium gets out of the core, we don't  
15 have to worry about core disruption. It doesn't melt. That  
16 is the point that you needed to remember.

17 MR. ZUDANS: Yes.

18 MR. THEOFANOUS: Point number two is, somehow that  
19 sodium isn't going to sit there forever. It will have to get  
20 out one way or the other. If it doesn't get out because of  
21 structural failure, it will boil off. When it boils off, it  
22 will boil over to the containment. If the loss of heat sink  
23 was caused by earthquake, already we have a broken part in the  
24 system that causes leakage.

25 The seals can fail because of the high temperatures

1 because you build very high pressures and you will get sodium  
2 vapor out of that, but it has to get out and only after it gets  
3 out then the core can melt.

4 (Slide)

5 Okay. We are done with that. Then, of course, we  
6 like to also ask the what-if questions, and here is one for  
7 you: What if we like to postulate an energetic? Of course,  
8 the thought here is that we don't have the sodium pool and  
9 therefore there is no direct means by which we can couple the  
10 high-pressure means to the head. And just as a matter of  
11 curiosity more than anything else, we carried out of this  
12 compilation -- and Professor Hawkins in November was curious  
13 about this situation, so he might like to see this result --  
14 from the same kind of compilation, \$200 per second compilation,  
15 that we ran to assess the vessel with sodium in. We ran it  
16 again except now we don't put the sodium in. So we get out of  
17 the expansion as before, and then we let the UIS displace  
18 upwards, and we let the venting process go on its way.

19 By doing that, we calculated the pressure history  
20 on the top of the UIS. By taking the pressure history of the  
21 UIS, we calculate, neglecting the effect of those because maybe  
22 they are warm and somebody might say that they are strengthless,  
23 everybody to accelerate under this kind of pressure. And we  
24 found that this UIS will be given total kinetic energy of 45  
25 megajoules, and considering the head capability that this kind

1 of energy is, from the point of view of the pressure to the  
2 head, direct pressure to the head, in the case of sodium  
3 in, you know that at the time of impact you have a hammer effect,  
4 and it goes to very, very high pressures, and they are  
5 shortlived.

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1           In this case, the pressures are of the order of 20  
2 bar, and I don't want anybody to get the idea to expect that  
3 20 bar pressures there. There is just the hypothetical  
4 compilation to show that margins available from the point of  
5 view of failing the head.

6           MR. ZUDANS: The 20 atmosphere, of course, the vessel  
7 couldn't take?

8           MR. THEOFANOUS: Of course, I don't want anybody  
9 to say that we expect this kind of pressure, but it is part of  
10 the hypothetical computation.

11          MR. MARK: Will 5 megajoules left of the UIS even up  
12 as far as the head?

13          MR. THEOFANOUS: The 5 megajoules are in the UIS.  
14 That is a free body.

15          MR. MARK: Yes.

16          MR. THEOFANOUS: Whether it will work against gravity  
17 up to this distance -- I have a feeling that it would.

18          MR. MARK: I see.

19          MR. THEOFANOUS: Yes.

20          MR. MARK: It might, because there is going to be  
21 gas ahead of it which will exert a downward pressure as well  
22 as gravity.

23          MR. THEOFANOUS: Yes. And to summarize the loss of  
24 heat sink, because of the long recovery times on it in hours,  
25 we really think the likelihood of getting into these situations

1 up to core disruption is considerably lower than one normally  
2 says.

3 In any case, the loss of heat sink event is  
4 nonenergetic. And even if it was, there is a very high  
5 tolerance for energetics because of the sodium pool.

6 And that summarizes our discussion of the loss of  
7 heat sink. Any questions on that?

8 MR. KASTENBERG: Could you put that back? We don't  
9 have that.

10 MR. BOEHNERT: Yes, you do.

11 MR. KASTENBERG: Yes.

12 (Slide)

13 MR. THEOFANOUS: By the way, maybe I should make a  
14 remark that there were a number of people in fact that were  
15 very much concerned about the loss of heat sink accident for  
16 some time now and the reason being is that they could see  
17 themselves how they could get into a whole core pool situation  
18 and then extracting from the literature about what was existing  
19 about the core pool situation and compiling that together with  
20 the higher probability expected for such accidents to happen,  
21 they felt maybe that could dominate even core disruption. And  
22 I think that we want to make a very strong point here that we  
23 really don't think that this is the case or should not be the  
24 case.

25 The last initiator that we considered is the power

1 accident. That has been cited by many, many people. And here  
2 I want to give you the essence of our argument for claiming  
3 that -- for the kind of TOP that we consider not completely  
4 incredible.

5 We don't expect to have energetic behavior. And, of  
6 course, the interesting question is how much of an operator  
7 do you expect in a situation. And we had a little bit of that  
8 this morning in our discussion with Walt Lipinski, who wanted  
9 to put some numbers from the TOP itself. We did not try to  
10 quantify the TOP. We do not consider it was part of the scope  
11 of our study. Rather, what we did was we said let's consider  
12 the most probable TOP that we can have there.

13 Now, let's consider higher or lower than that and see  
14 what is the relative probability. How much more additional  
15 systems have to fail in order to achieve this higher transient  
16 overpower condition. And what we found is the most probable  
17 between the -- if I remember correctly -- between 5 and 6 cents.  
18 That is the most probable.

19 Now, from that, we go down to 2 cents or go up to 10  
20 cents. You need another loss of probability for a factor of 2.  
21 And then from there to go on to something like 15 cents per  
22 second, you need another -- additional systems have to fail to  
23 produce that situation.

24 We decided therefore that this 3 order of margin  
25 between the most probable TOP, that the people consider all the



1 time, and this higher range would be adequate. And we put our  
2 thrust at 10 to 12 cents per second.

3 All right. That is the maximum to operate that we  
4 considered here. Now, what is the concern with the TOP is that  
5 if one postulates and if the beams were to fail in the center  
6 line, the same process that we were discussing before, in-fuel  
7 motion pushed by the gases can potentially cause prototypic  
8 behavior.

9 Now, the mechanics of the failure here are not very  
10 well understood. And that is why we are almost pushed to the  
11 limit of having to pursue it. We have no real reason for  
12 excluding it. On the other hand, it must be said that  
13 we have no real hope for expecting them to happen. The fact  
14 is interesting, because as one goes through the analyses, one  
15 is interested to know what is the rate of sweep-out of the  
16 fuel. And then one takes a look at the experiments and what is  
17 difficult to see is the experiments don't provide midplane  
18 failures, so you can follow the fuel for a long period of  
19 time, and that gives you a little bit of the discrepancies of  
20 conditions that we have to assume in order to get into this  
21 problem here.

22 On the other hand, it must be recognized that the  
23 operation increases, there is more of a tendency for the fuel  
24 to melt and pressurize more and more to the center of the core  
25 and therefore it becomes more and more likely for the failure

1 to happen at this point.

2 So that is the reason that we put a discrimination  
3 there between what we want to consider and what we consider  
4 outside of the spectrum with reasonable consideration.

5 So now then, having done that much, we like to know  
6 at this upper rate of 10 to 12 cents per second whether there  
7 is any potential for -- because of this mechanism here, and what  
8 are the processes as follows. The operator is introduced at  
9 full flow. Another part is produced and melting happens, and  
10 then the subsequent importance tells us which way the accident  
11 is going to go and depends on in-pin fuel motion at the failure  
12 point at the midplane. And the activity will be augmented.

13 If it is outside or farther from it, the activity  
14 will be reduced. This process of sweep-out is the process by  
15 which the fuel and the coolant mix and pressures generate and  
16 tend to push the fuel out and away from the point of failure.

17 This is not only the process from the coolant but  
18 also they can arise from fission gases that are coming out  
19 together with the fuel. This is a sweep-out, and in general,  
20 especially if this is the midplane, the effort will be negative.

21 Of course, if the fuel is going to move out by some  
22 pressures, also the sodium can move in the same way; and, and  
23 course, if moving sodium is the causative effect, it will be  
24 increasing, so in the analysis that you are looking for, has to  
25 properly weigh all three of the effects in order to fulfill the

outcome.

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RG  
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1           The other interesting part here is that not all  
2 processes here are coincident. This is a process that has  
3 to happen first. Then it has to follow. From that point of  
4 view, then, the coherence at which failures take place is  
5 important. If all the core was to fail coherently, you would  
6 have the whole core, the fuel moving inward, and if I wanted  
7 to still postulate, one would have a bad situation before  
8 anything else happened. But the other thing happens because of  
9 the power distribution.

10           Some part of the cores would fail first, and some  
11 motion is going to take place, and then as soon as the fuel  
12 gets into the tunnel, this due process is going to come in  
13 before even the rest of it begins to fail. So from that point  
14 of view, the timing between success and failure of the core  
15 is very crucial because that tells you how much time is  
16 available for the sweepout to come in and before the cause  
17 and effect.

18           So we are looking, then, at this process.

19           MR. CARBON: Question?

20           MR. THEOFANOUS: Yes.

21           MR. CARBON: Before you left it, does your plenum  
22 gas play a role there in pushing fuel in toward the center?

23           MR. THEOFANOUS: Well, no, because in this case  
24 the cladding is so strong and everything is in place.

25           MR. CARBON: But you assumed this morning that

j-4c-2 1 there was zero friction between the pellets and the --

2 MR. THEOFANOUS: Yes. You know, this disruption  
3 happens with molten material in the center and just rushing  
4 out. There is still, in fact, solid fuel around that cavity,  
5 so you cannot really compact that very well from that point  
6 of view.

7 Furthermore, the time scheme of this process is  
8 slowed up in the projection. By the time there is any chance  
9 for anything else to happen, it would be completely out.

10 MR. THEOFANOUS: Okay. So this here summarizes  
11 what are the three aspects that we are looking at. We are  
12 looking at coherence. We are looking at failure location and  
13 at sweepout. Or the coherence increases with the operation.  
14 We have already said, considering 10 to 12 cents per second,  
15 we consider this to be adequate. Also, the coherency in-  
16 creases with the power distribution.

17 Now, the end of Cycle 3 core is much more coherent  
18 than the other cycles because in this core the six subassem-  
19 blies that are driven fuel are refueled and replaced by  
20 blanket fuel, that is, with closer power and therefore more  
21 coherence for more support.

22 In the beginning of our work, in fact, the data for  
23 the analysis for this core were not available. We requested  
24 from the Applicant and we got them in good time, and in fact  
25 we were able to do the analysis as the Applicant was able to

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1 do the analysis. On the failure locations, again, it goes  
2 more to the midplane with a higher ramp and affected by burnup,  
3 and there are a lot of questions associated with that. If  
4 you remember, in November I brought up the discussion about  
5 the W2 test. That was showing, temporarily, at least, that  
6 the midplane failures are more likely to happen, but since  
7 that time the Applicant has taken a closer look at this test  
8 and found out that there were a very large number -- in fact,  
9 they make this test completely worthless from the point of  
10 telling us anything when the pins are going to fail.

11 MR. CARBON: Question. If that test is non-  
12 productive and you don't know it until afterwards, how many of  
13 the others are that way?

14 MR. THEOFANOUS: Well, why we even take a look at  
15 the test after it is run? First, the person wants to do a  
16 good job before he runs the test and decide whether the test  
17 is close enough reality, and then after running the test, I  
18 doubt that anybody can see any results. One doesn't take it  
19 at face value when cross-examining him.

20 MR. CARBON: Take it at face value if they are  
21 good results?

22 MR. THEOFANOUS: No. I think that may be -- I  
23 doubt that anybody takes any results at face value. At least  
24 in my experience I don't know of anybody that does that, but  
25 there are many different forces in this environment, and every

j-4c-4 1 force is looking together from its own perspective, and I  
2 don't think there is any chance at all there will be, but  
3 they are sitting there and all of these different people are  
4 going to look at it, and they can forget about some important  
5 and --

6 MR. CARBON: Usually the things that were wrong on  
7 the W2 weren't wrong on the other tests.

8 MR. THEOFANOUS: Yes, Paul?

9 MR. DICKSON: There wasn't anything wrong. That  
10 test wasn't meant to be prototypical of Clinch River fuel.  
11 You had a lot of experimental fuel elements. It was the first  
12 one in which we saw failure at the center line, and a lot of  
13 people got excited. Then when people looked at it, what does  
14 that mean relative to our fuel? There were a large number of  
15 non-prototypicalities. It wasn't because they didn't want to  
16 achieve their goal. They weren't trying to make it proto-  
17 typical.

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MR. THEOFANOUS: Well, yes. I think this is a little bit maybe why you go off on a tangent from what we are discussing, but I do want to say the obvious question to why does anybody do tests that are completely outside the range of interest here, why you were doing tests that do not support or are not relevant to the case at interest, so I can really make it both ways. But let's not pursue that any further.

9 The fact of the matter is we don't have any good tests to tell where the failure location is going to be under these kinds of conditions, and as a result of that, we are forced to assume mid-plane failures which might be a very conservative assumption.

14 On the sweepout we have the forces produced by fission gases as well as by hydraulic pressure. Remember that the flow is going full-blast through the core and that is -- so if you were to come in a tunnel, it would be carried away with the sodium, and we refer to that as hydraulic pressure, hydraulic force.

20 Now, on this one we have experiments, and we can make use of those experiments, as the Applicant has done, to quantify the sweepout. It is very important to quantify sweepout correctly, because as you see, the time margins vary process here, after we have assumed these processes are not that great. So one does do a good job in qualifying the

1 timing of the sweepout. The quality of the fuel involved  
2 in this is relatively small. We are about ten percent of  
3 the pin.

4 What happens is there is a molten fuel that builds  
5 up inside the cavity up to the failure point and then opens  
6 up, and all of it just comes out, and a little bit more melting  
7 you get after that, but it is not anything significant. So that  
8 kind of an operation is really limiting the amount of fuel  
9 that makes it out into the tunnel. So from the point of view  
10 of sweepout it is not as detailed and difficult and anything  
11 to examine as, for example, if you were going to melt half  
12 of a pin coming out, where you can visualize the first maybe  
13 5 percent coming in or 10 percent, interact, produce pressures,  
14 and get sweepout, and all the rest of the 40 percent coming in  
15 and get nothing except going out and not being able to move  
16 away there.

17 So because of this small quantity of molten fuel  
18 available, one can really be quite a bit more comfortable  
19 about this whole sweepout.

20 (Slide.)

21 That is the point I was trying to bring across,  
22 and here is what kind of a -- here is what you can do in  
23 calculating the experimental sweepout activity, and this is  
24 specific sweepout activity expressed in cents, subassembly  
25 per gram injected per peak, and the experiment is done in the



1 reactor, and it was a high -- it was on the order of \$7 to \$8  
2 per second, and the calculation is done with a PLUTO-2 code  
3 with a set of assumptions concerning very detailed physical  
4 processes. So I don't want you to get hung up with all --  
5 begin verification and all that the -- all those numbers  
6 forget the important point is that we are trying to match  
7 the experimental sweepout and the fact one could have done  
8 in this just as good a job on the back of the envelope or  
9 just by taking the experimental data and applying these to  
10 the reactor condition.

11           And here is the important point why one can use the  
12 PLUTO-2 tests to make those judgments. In the early part of  
13 the ejection of the fuel, the ejection is very, very similar  
14 between the L-8, which was at \$7 per second and this 10 cents  
15 per second, that if that was to be done either late or if  
16 it ever was going to happen in the reactor, and there is --  
17 and the reason is that as soon as the rupture happens, whatever  
18 molten material was there will just run out very quickly. And  
19 to the extent that the rupture is going to be a function of  
20 how much molten material is there, you see about the same  
21 quantity of material will be available to come out, and that  
22 is why it is about the same in all three cases.

23           What happened in this case is because it operates  
24 slower, the rate at which you are melting more material is  
25 much slower, of course, so you are not putting much more in

1 while here you are continuing on to melt very rapidly, and  
2 you are ejecting much more.

3 On the other hand, the process of sweepout is more  
4 related, especially for this water bridge, is more related to  
5 this small, up to 20 milliseconds on the time frame, and small  
6 quantities to 20 grams per pin.

7 MR. KASTENBERG: Theo, do you really mean written in  
8 the green PLUTO-2 L-8 or do you mean PLUTO-2 Clinch River?

9 MR. THEOFANOUS: No, no. L-8, we took the PLUTO-2  
10 code and applied to L-8, because another question might have  
11 been how does the L-8 --

12 MR. KASTENBERG: But that is what you read once.

13 MR. THEOFANOUS: No. That is \$7 per second.

14 MR. KASTENBERG: L-8 was \$7?

15 MR. THEOFANOUS: This is the L-8 experiment itself.  
16 It granted \$7 per second and analyzed with PLUTO-2 code. This  
17 green line is the PLUTO-2 compilation applied to a hypothetical  
18 L-8 experiment but ran \$7 per second rather than 10 cents per  
19 second, while this blue line is assessed for a PLUTO-2  
20 calculation. Together, the two together form the 10 cents  
21 per second.

22 There is some difference in the lengths of life and  
23 so on, and here is it. There is the most to say for the end  
24 of Cycle 3 with the new data that came in, which we used to  
25 generate independent inputs and to run the calculations with,

1 a new tape, and this was done very successfully, which has  
2 the possibility -- capability of very, very fine. As you  
3 see, the number is much more than the numbers in Article 15  
4 that are being used, for example. Here in parentheses is the  
5 number of subassemblies, so as you see, the group is very  
6 small.

7 Here is the fuel temperature versus time. Now, one  
8 might say we have a lot of series about the location for --  
9 were serious about the time of failure, and therefore, one might  
10 say how does this translate to your knowledge of the coherence  
11 of failures?

12 Well, the point there is that whatever it is has to  
13 be related, and in fact, in the report we have some good argu-  
14 ments that the failure cause or failure mechanism is the --  
15 can be related to the FOP in the fuel, and as long as that  
16 happens on a similar temperature, just by looking at the  
17 relative position of the curve we can learn about the  
18 coherence without necessarily being too much dependent if the  
19 failure happened at this point or this point or that point.

20 As you see, this is of the order of 100 milliseconds.  
21 As you see, we have something of the order between this group  
22 and this group here of more than 100 milliseconds of time  
23 available for sweepout to come in and to begin initial in-fuel  
24 motion that happened because of the failure of those two.

25 To relate you to that, here is again the experiment.

(Slide.)

A hundred milliseconds is way up here, and you see that the experiment gives us a specific sweepout of the order of .15 cents per the subassembly for 13 within 30 milliseconds, so there is ample time.

So what you do in the report is with this activity you know how much fuel is coming out and very quickly calculate what is the negative effect, and you can actually see and visualize the result that also one obtains from an actual analysis, multi-tunnel analysis of the all the effects going on together. And what you see here is one such analysis for a total activity feedback. Here is the input. It comes in and then this is the second ejection that you saw before. This is the sodium which is just under 1 cent per second per sub-assembly.

This should be per subassembly, and this is the sweepout activity, and the net comes in in this picture here. You see within 28 milliseconds sweepout was able to cancel the sodium, and the in-pin fuel motion, and from them on it keeps on being negative because it is very strong and brings another activity to -2 cents per second and a little later even more, clearly saying that the action terminates by this process before the next group fails to come in and gives its portion.

So based on the results then, we conclude that the -- operates less than 10 to 12 cents per second. There is

1 negligible autocatalyst potential, and that TOPs of this  
2 range or less are nonenergetic.

3 Operates greater than 15 cents per second. Maybe  
4 there is some potential autocatalyst. I think after one learns  
5 more about failure location and failure mechanisms, one can  
6 answer more definitely this question. But in any case, the  
7 probability of this type event is three times lower than this  
8 event, and in fact, it is in the range where you consider it  
9 is a negligible level.

10 MR. MARK: You said three times lower?

11 MR. THEOFANOUS: Three orders among the lower, thank  
12 you. And I think with that I have concluded what we had  
13 prepared here. Do you have any more questions?

14 MR. KASTENBERG: Yes. There are two questions I  
15 would like to raise. Last year, I guess back in May, I had  
16 asked the question about one configuration representative of  
17 a class, and I don't remember exactly which one it was, where  
18 following the initial TOP you looked back at the picture. Let's  
19 put it up.

20 (Slide.)

21 You ended up in a situation where you had, say,  
22 one range of assemblies failing and sweeping the fuel out.

23 MR. THEOFANOUS: Yes.

24 MR. KASTENBERG: But ending up in a situation where  
25 you were at some low power level below steady state power,

1 and you were a little subcritical or somewhat subcritical,  
2 but what it was that was causing the reactivity insertion  
3 still had a ways to go, and that you might start on a second  
4 TOP.

5 MR. THEOFANOUS: Yes.

6 MR. KASTENBERG: And at that time I guess we heard  
7 that someone was going to be looking at that, and I wonder if  
8 you have looked at that or if the Applicant has looked at it.

9 MR. THEOFANOUS: I think that the Applicants have  
10 not done any more in this area. This work was done for us  
11 by Harry Humble at Argonne, and I repeatedly questioned him  
12 about it, and he tells me it is clearly shut down and that  
13 there is no concern for really building up again to a situation  
14 where we have to consider even more coherence. But I think  
15 that you will find probably in more detail.

16 And, Bill, I think that is a good question, but I  
17 do want to give you a more specific answer to that. We don't  
18 have much of that in the report, but Harry Humble has an  
19 assembly of his own based on which we abstracted the important  
20 things in the report, and that summary is going into a  
21 computer which is all the additional information that we can't  
22 put in that big thing, and I will let you have that, and maybe  
23 that will answer the question for all those detailed analyses  
24 that we discussed. But I know that Harry has looked into it.

25  
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1 MR. KASTENBERG: A second point that has been  
2 raised also last year, I guess it was May, when we discussed  
3 this with respect to TOP. That was long-term cool bit, and  
4 there is some school of thought that when you do have sweep-  
5 up, the particles will cool as they get swept up and freeze  
6 and plug, blocking some channels, and this could conceivably  
7 lead to a secondary event, a secondary melt. Again, I wonder  
8 whether you have pursued that at all.

9 MR. THEOFANOUS: I know the Applicant talked to us  
10 about that and had some feeling of when they might be in  
11 trouble, but I wonder if you have looked at that independently.  
12 We have looked into that but not in a great detail, and the  
13 reason is that we feel we don't expect to have a very exten-  
14 sive luggage -- because of the small amount again. If it was  
15 \$8 per second, take the L8, I will give you a different answer.  
16 But if it is only 10 cents per second, I don't expect to have  
17 a lot of fuel moved out in the first place. Then even if that  
18 fuel was moved out, I don't expect it to form enough blockage  
19 to stop the flow.

20 I want to take it one step further. Even if the  
21 flow was going to stop, I don't think I can get very excited  
22 about local multi-ridges in the core, especially when I see I  
23 can move fuel all around the whole core and still somehow  
24 being able to get it out.

25 (Pause)

j-4E-2

1 MR. THEOFANOUS: I know it's not a direct answer,  
2 but I know it's a very, very detailed kind of a thing, and if  
3 you really try to approach it on the best estimate basis, it  
4 is a simple thing to follow wherever part of the fuel has gone.  
5 But in view of the small amount of fuel involved, I don't  
6 think we are very concerned about that.

7 VOICE: Talking about 4 or 5 percent fuel to shut  
8 it down?

9 MR. THEOFANOUS: Yes, sir. Very small.

10 MR. MARK: I don't have a question. Maybe I will  
11 later.

12 MR. THEOFANOUS: You don't have a question?

13 MR. MARK: I don't mean later this afternoon. I  
14 do have a slight comment. For example, I find in the summary  
15 at the end of Section 0.4 the expression "such events are of  
16 sufficiently low probability that they can be excluded from  
17 consideration." I don't believe that is the only place that  
18 a phrase of that quality is in the summary.

19 Now, this morning you explained -- and I think from  
20 my point of view rather satisfactorily -- what you intended by  
21 using that phrase. If I try to say what I carried away from  
22 that, I think of an event like an ATWS or an event like pump  
23 stopping, and you think of it as having some low level of  
24 probability, and that that actually happens -- perhaps  $10^{-3}$   
25 or  $10^{-5}$ , and when you use that phrase, you mean following that



j-4E-3 1 event you must have enough sufficient unlikely failures or  
2 sequels that you feel, judge, that you could put on that first  
3 probability, which you don't here concentrate on -- another  
4 factor like  $10^{-3}$ .

5 MR. THEOFANOUS: Right.

6 MR. MARK: I know this is probably the final time  
7 that you ever hope to see this report, certainly to carry it  
8 anywhere.

9 (Laughter)

10 MR. MARK: If you were ever to come back to it,  
11 it would seem to be useful to put in there very visibly what  
12 you mean by using that phrase. I can just hear Professor  
13 Okrent running on that phrase.

14 MR. THEOFANOUS: Yes. You can't imagine how much  
15 we sweated in choosing that phrase.

16 MR. MARK: I am not saying I think it is easy. You  
17 did it to my satisfaction this morning pretty well, and I  
18 don't find that in the report. It may be there somewhere else.

19 MR. THEOFANOUS: The way I said it early this  
20 mornning is in the bulk of the report. My original inclination  
21 would be in the summary to use something like "incredible"  
22 Instead of saying sufficiently low, saying incredible.

23 MR. MARK: You put it in a scale of incredibility.  
24 If you did, that would be much more satisfying than just using  
25 some words like "it's so small we don't care," or "it's

j-4E-4 1     incredible."

2             MR. THEOFANOUS: Yes.

3             MR. MARK: Those phrases don't sell well, and I  
4     don't mean only to Okrent, but to whoever reads it.

5             MR. THEOFANOUS: That is a good point. I think we  
6     will consider that.

7             Any other questions?

8             (Pause)

9             MR. KASTENBERG: Just a point of information.  
10    There is considered a final draft now --

11            MR. THEOFANOUS: This is the very final report,  
12    yes.

13            MR. MARK: See, I was assuming that maybe I should  
14    change it. I knew I was talking in a vacuum.

15            (Laughter)

16            MR. CARBON: Are you intending to say anything here  
17    today on peer review and agreement with this?

18            MR. THEOFANOUS: I can give you some off-the-cuff  
19    remarks if you like on this. As you know, we have been  
20    operating on a very tight time schedule. On the other hand,  
21    we had an extensive review. We had really two reviews, pretty  
22    much official reviews. One was done at Los Alamos about two  
23    weeks ago or one week ago -- I think it was two weeks ago --  
24    in which this review was conducted by the Los Alamos management,  
25    represented by McDowell, Mike Stevenson, Jim Scott, and Laron

j-4E-5 1 Smith. And as part of that, it was the first time that we  
2 actually presented the results of this study to a body of  
3 people. At that time we actually saw after the report and  
4 also this meeting was attended by Curtis Allen, Bill Morris,  
5 and Nelson Grace from the NRC.

6 Then following this meeting we went through and  
7 made some -- based on their comments, made some changes. Then  
8 we had a meeting last week which was like a review meeting  
9 conducted for the NRR at large and the Office of Nuclear  
10 Regulatory Service, so this was attended by Hal Denton, and  
11 the Office of Research was Denny Ross, Charles Killberg,  
12 Bob Curtis, Bob Wright, and a number of other people, and Phil  
13 Wood.

14 We gave basically a similar presentation to the  
15 meeting. So that was like a peer review, and in particular  
16 Denton encouraged everybody up to 6:00 or 7:00 or 8:00 to  
17 bring up any comments or reservations they had, and we did  
18 not hear anything negative. In fact, everything that we heard  
19 from that group as well as from the previous group was very  
20 positive.

21 Now, we also had, as you probably gathered from  
22 the cover page -- the main part of our team is from the  
23 Los Alamos Laboratory, and to the extent that these people are  
24 there and available, in fact I think they have read most of  
25 this in court over the last two weeks, and again, individual

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1 workers were familiar with what we were working, but it was  
2 the first time they saw the whole thing, and the comments I  
3 got were that everybody didn't have any problems with what we  
4 were saying here.

5 That leaves out only the group at Arconne, and that  
6 is Harry and Phil and leaves out the Sandia Laboratory. We  
7 have one member who acted as a consultant, Dick Bast, who has  
8 not seen it. I think that is about it.

9 Now, what we intend to do is get this report to  
10 Sandia and also to get it to the people that -- like was  
11 suggested in the previous presentation. We did send our  
12 document at that time around to the laboratories and asked  
13 for comments, and we did get some comments like that, and those  
14 letters are available. We can make a copy of them any time  
15 and get them to you.

16 We intend as soon as we finish this business here  
17 to respond to those comments. Some of the comments, of course,  
18 we took them and we incorporated them or did some work and put  
19 in the report. Some other ones we didn't feel they apply,  
20 and those we are going to respond to by letter at some later  
21 time. So the next thing we are going to do is follow the same  
22 procedure, take this report and send it around to the national  
23 laboratories in the community and ask them for a letter from  
24 them with their comments, hopefully before the 10th of the  
25 month so we can have their input by everybody before the next

j-4E-7  
1 committee meeting. Maybe Charley wants to add anything to  
2 that.

3 MR. BELL: I was trying to recall who else was at  
4 the NRR review last week. I knew Professor Reynolds was  
5 there.

6 MR. THEOFANOUS: Curtis Pierce was there.

7 MR. BELL: Alan Walter. I guess he is a consultant,  
8 but well-known in this area. Dick Ireland, I guess, was there.

9 MR. CARBON: Could you, Mr. Stark, go a step further  
10 and say are all the NRC people that Charley just mentioned  
11 in agreement with this?

12 MR. STARK: We don't know of anyone that has a  
13 problem. We certainly made several opportunities recently  
14 available to encourage as many people, and as Theo is saying,  
15 Harold in particular was trying to encourage people to dig  
16 down and ask questions, get them out now, don't be bashful.  
17 We made a very conscious and serious effort to do that, and I  
18 know of no one within the NRC right now.

19 MR. THEOFANOUS: But I do want to say, though,  
20 that we don't want to oversell the agreement now at this  
21 point because I think that it is not unreasonable to expect  
22 that somebody will go to a whole-day meeting and listen to  
23 this complicated story, and by just not voicing big disagree-  
24 ments, to take it for granted that that means a lot of  
25 agreement.

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1 I would like to see that everybody had a copy of  
2 that and actually had it with them for two weeks, and if  
3 then I don't hear anything, only then I will say that there  
4 are no problems.

5 MR. CARBON: So you will follow up?

6 MR. THEOFANOUS: We are following up by sending out  
7 and requesting comments from the next two to three weeks, and  
8 we will hope to be able to summarize the comments for you.

9 MR. CARBON: Bob Wright.

10 MR. WRIGHT: Bob Wright. I might just comment in  
11 that regard, speaking essentially for myself, I think I am  
12 the only one from Research here at the moment. I think the  
13 questions were did we have real objections with the conclu-  
14 sions, and I think all the Research people that were there  
15 went along and thought that the scoping analysis and all the  
16 conclusions followed well, and we so expressed that. There  
17 were some details, as one would expect, in such a substantive  
18 piece of work with which one could have some questions, and  
19 several of us, including myself, did have such, but I don't  
20 think they affected the conclusions particularly. That is  
21 the point I want to make in general.

22 The thrust of the report was very well-conceived.

23 MR. CARBON: Any more questions of Dr. Theofanous?  
24 If not, thank you very much.

25 MR. THEOFANOUS: Thank you.

1 MR. CARBON: The Applicant is on the agenda to  
2 comment. Dr. Dickson?

3 MR. DICKSON: We have no comments. Thank you.

4 MR. CARBON: May we ask, are you in agreement with  
5 what has been said here today except for the one point that  
6 you have only committed yourself to consider the plenum gas  
7 question that came up this morning?

8 MR. DICKSON: Well, of course we haven't had the  
9 opportunity that the Staff has to review this document. We  
10 just saw it. But I don't believe anybody has any basis for  
11 disagreement at this time.

12 MR. CARBON: Dr. Fauske represents you or is a  
13 consultant to you, is he not?

14 MR. DICKSON: He is.

15 MR. CARBON: Could he comment on his views of this?

16 MR. FAUSKE: I think generally speaking, as Paul  
17 Dickson pointed out, I think we agreed with what we have heard.  
18 I would just like to reemphasize that all day we addressed  
19 probability events, and in addressing such events, one should  
20 apply reasonable assumptions, and I think on the basis of  
21 applying reasonable assumptions, the project came to the  
22 conclusion that energetics are very benign, and I think  
23 listening to Theo and Charley, they apply reasonable  
24 assumptions, and it is my understanding that energetics is  
25 very benign indeed, and in fact, in order to develop energetic

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events, you have to stretch physical reality, and even if they stretch it pretty far, so to speak, the energetics still are within the capability of the machine.

So in summary, I think that we generally agree. I think in the particular area of compaction of the fuel by fission gas, I think we want to look at it a little more, and my only personal view on that is that I think that is of potential concern. It is a means by which you could get into energetics, and I think you want to look at the analysis. I think analysis of this regard may be more expensive than making the design change. That is something we will be addressing in the next year or so, so in summary, I would like to take the opportunity to congratulate Theo and Charley and all the rest of the consultants for, I think, very outstanding and a very independent piece of work in this most difficult area.

MR. THEOFANOUS: Thank you.

MR. CARBON: Does anyone have any more questions?

MR. THEOFANOUS: This morning you brought up the problem of yesterday having to do with the expansion process, that there was some confusion about how we are doing it and how the Applicant is doing it. Then he was hoping that my presentation will qualify this point. Is there any point of confusion left there as far as how the Applicant is counting energy and how he is doing this versus how we are doing it?

Since we are all here, maybe we can inquire.



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MR. ZUDANS: I don't think at this time I have any further problems.

MR. THEOFANOUS: Okay.

MR. ZUDANS: I think that you added to -- the question is -- I am a very primitive person. Unless I can do it, I am not satisfied. I think you have done a magnificent job.

MR. THEOFANOUS: Thank you.

MR. CARBON: I would also comment to you and to Charley, it appears to be a real fine job. We thank you. We thank all for the presentation.

(Whereupon, at 4:30 p.m. the meeting was concluded.)

END 4E