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ADVISORY COMMITTEE ON REACTOR SAFEGUARDS

(ACRS)

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STRUCTURAL ANALYSIS SUBCOMMITTEE

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FRIDAY

OCTOBER 23, 2015

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ROCKVILLE, MARYLAND

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The Subcommittee met at the Nuclear Regulatory Commission, Two White Flint North, Room T2B1, 11545 Rockville Pike, at 8:30 a.m., John W. Stetkar, Chairman, presiding.

COMMITTEE MEMBERS:

JOHN W. STETKAR, ACRS Chairman RONALD G. BALLINGER, Member DENNIS C. BLEY, Member-at-Large HAROLD B. RAY, Member PETER C. RICCARDELLA, Member STEPHEN P. SCHULTZ, Member

GORDON R. SKILLMAN, Member

DESIGNATED FEDERAL OFFICIALS:

CHRISTOPHER L. BROWN

KATHY WEAVER

ALSO PRESENT:

JON AKE

NILESH CHOKSHI

STEPHANIE DEVLIN-GILL

VLADIMIR GRAIZER

REBECCA KARAS

SEUNG MIN

CLIFFORD MUNSON

JOSE PIRES

LISA SCHLEICHER

DOGAN SEBER

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1	P-R-O-C-E-E-D-I-N-G-S
2	(8:32
3	a.m.)
4	CHAIRMAN STETKAR: The meeting will now
5	come to order. This is a meeting of the Structural
6	Analysis Subcommittee. I'm John Stetkar, Chairman
7	of the subcommittee meeting. ACRS members in
8	attendance are Harold Ray, Dick Skillman, Steve
9	Schultz, Dennis Bley, Pete Riccardella, and Ron
10	Ballinger. Christopher Brown and Kathy Weaver of
11	the ACRS staff are the Designated Federal Officials
12	for this meeting.
13	The purpose of this meeting is to
14	receive a briefing on the treatment of
15	uncertainties in probabilistic seismic hazard
16	analyses conducted for combined license
17	applications and early site permits. We'll hear
18	presentations from representatives of the Office of
19	New Reactors and the Office of Research. The
20	subcommittee will gather information, analyze
21	relevant issues and facts, and formulate proposed
22	positions and actions as appropriate for
23	deliberation by the full committee.
24	The rules for participation in today's
25	meeting were announced as part of the notice of
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6 this meeting previously published in the Federal 1 2 Register on October 13th, 2015. We've received no 3 written comments or requests for time to make oral 4 statements from members of the public regarding 5 today's meeting. A transcript of the meeting is being 6 7 kept and will be made available, as stated in the 8 Federal Register notice. Therefore, we request 9 that participants in this meeting use the 10 microphones located throughout the meeting room 11 when addressing the subcommittee. Participants 12 should first identify themselves and speak with 13 sufficient clarity and volume so that they can be 14 readily heard. I remind everyone to check your 15 little communications devices and please silence 16 them. 17 I believe we have a few folks on the 18 bridge line. We keep the bridge line open and 19 muted so that we don't get disturbed by crackles and pops. We'll open the bridge line at the end of 20 21 the meeting so any members of the public or anyone 22 who is out there can make comments. 23 To remind us all why we're here, these 24 exchanges, these briefings were prompted by our

observations during the Fermi COL application and

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1	the South Texas COL application where we raised
2	questions about why the uncertainties in the
3	seismic hazard curves seemed both fairly
4	constrained and quite uniform over the full
5	spectrum of the accelerations, from very low
6	accelerations to very high accelerations, and
7	followed that general practice over most of the
8	range of the ground acceleration frequencies,
9	hertz. It's always difficult in these discussions
10	to talk about frequency when we talk about things
11	that people call recurrence intervals and frequency
12	when we talk about hertz, but we'll try to keep
13	that straight.
14	So we had a meeting almost a year ago,
15	back in November of last year, to try to get us up
16	to speed on the methods and how they're used and
17	possible reasons for this behavior. And as I
18	characterize it, we got just to the point where it
19	started to get really interesting, and we ran out
20	of time. So, hopefully, we'll pick up from just
21	about that point today. I understand the staff has
22	a few things they'd like to clarify and explain
23	from questions we raised during the last meeting,
24	so we'll perhaps backtrack a little bit. But,
25	hopefully, by the end of today, we're going to be

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1	not experts but at least know enough to be less
2	dangerous.
3	We'll proceed with the meeting, and
4	I'll ask Becky Karas, Branch Chief of NRO, to make
5	introductory remarks. Becky?
6	MS. KARAS: Thank you. This is Becky
7	Karas from NRO, chief of geosciences and
8	geotechnical engineering. And we do hope this will
9	be a productive meeting. As you mentioned, there
10	are a few open items left from the last meeting
11	that we will be going through on the earlier
12	matter, as well as some items that we picked up on
13	the transcript. We've made available some of the
14	top experts in this area from the agency, from both
15	the Office of Research and NRO, and so we hope to
16	have a very good discussion and answer the
17	Committee's questions in this area.
18	At that, I'll turn it over to Dr. Dogan
19	Seber of my staff who will be the main presenter
20	and with Dr. Cliff Munson from NRO and Dr. Jon Ake
21	from the Office of Research also available to
22	answer questions and provide remarks. Thank you.
23	CHAIRMAN STETKAR: Thanks, Becky. For
24	the folks up front, if you haven't been up here in
25	recent times, when you're speaking, just push the

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base of your microphone so the little green light 1 2 When you're done, turn it off. is on. It helps 3 our transcripts. DR. SEBER: Thank you. Let's start with where we 4 were. And as the Chairman said, we had several hours of discussions 5 Unfortunately, time wasn't 6 last time when we met. 7 enough to finish everything. And of one the 8 was at the time that we had several reasons 9 questions, so we'd like to address those, as we 10 feel we didn't do a proper job perhaps. Perhaps we 11 were not ready with several slides that we needed 12 to explain some of those. So we have them, at this 13 point, available to you, and these are the three 14 questions that we had that was the main purpose of 15 the meeting. These were summarized by the Chairman 16 already, so I'm not going to go into it. 17 The outline today, we have, in a sense, 18 The first one-third we're hoping three parts. 19 we're going to address the questions raised in the 20 earlier meeting and discuss -- these are based on 21 from our notes, as well as the transcript, that we 22 looked very carefully and tried to identify all the unresolved issues we identified. 23 And we developed additional slides for those. 24 25

And then the middle part we already

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10 likely, we're going to 1 discussed, so, most qo 2 pretty fast, unless there are additional questions But that's the plan. 3 we'll be happy to answer. And the third bullet here is where we left off, in 4 a sense, that I'm hoping we're going to spend most 5 6 of the time today. 7 So now let's go into the summary of past meeting discussions. 8 This was one of our 9 earlier slides that got a lot of attention at the 10 time, and the questions raised from what are these 11 red dots, do you get two, what is this how 12 seismogram what does represent, on top, it and 13 things. 14 At the time, we were imagining this to 15 be a cartoon showing very basic concepts, but we 16 realize that we needed to talk more technically 17 about those. So the next series of slides will 18 just summarize how we got to this slide, what it 19 represents. 20 So what we are showing here, what we 21 had identified as seismograms in generic terms that 22 we use to estimate the ground motion prediction 23 equations that predict future SO we can 24 earthquakes' ground motions. We primarily rely on 25 observed records, which we call seismograms. An

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11 example here is shown. It's a time series record, 1 2 usually several second duration, depending on the 3 magnitude of the earthquake. And these are sitting on the ground at some location. It could be a 4 building, it could be what we call in the field 5 somewhere that has its 6 own energy source and 7 The black box here represents the sensor. things. 8 are very sensitive instruments measuring These 9 acceleration. There should be ground some 10 vibration. It doesn't even need to be an 11 earthquake. Any type of vibration. 12 And they're connected to what is shown 13 here as the orange box. It's a recording device. 14 different sample They record at intervals, 15 traditionally 200 sample score second, depending on 16 how we want to go. You can go higher. These are 17 all engineering aspects of it. 18 is recorded is But what the 19 acceleration of the ground at any given point. The 20 point is fixed, time is changing obviously. An earthquake happens somewhere. Seismic waves

21 22 shakes travel, the point, and that shaking is 23 recorded by the sensor and then the orange box, in 24 a sense, puts them in digital so we have them. And 25 that record represents that.

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1	MEMBER BLEY: And you do something like
2	Fourier analysis on that to identify the frequency
3	
4	DR. SEBER: Exactly. Eventually, we'll
5	get to that. This is the whole record. It has all
6	the frequencies.
7	MEMBER RICCARDELLA: And for X, Y, and
8	Z?
9	DR. SEBER: Yes, that is, actually, one
10	point I wanted to make. Most of these instruments,
11	they come already built in three components: X, Y,
12	and Z. Traditionally, people prefer, obviously,
13	vertical everywhere, north/south and east/west
14	orientation. But you can orient them, of course,
15	anyway you want. Some structural engineers would
16	like to see differently and things. But typical
17	traditional seismology folks, they like to install
18	it as north/south, east/west, and vertical. In
19	fact, you have three components, and you can always
20	define the real motion if you have that.
21	And the last bullet here in bold is
22	trying to address one of the questions that was
23	raised last time, and that was does this instrument
24	have an answer that we pay attention to? The
25	answer is, yes, of course, but it is a very low-

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noise instrument that when you record 0.5g or even 0.1g, that is minute. So typical PSHA does not look into that, take it into account. It could be something wrong with your instrument, you're recording wrong. Those are exceptions. I'm not these are well-calibrated getting into it, but and people maintain these and pay instruments attention to what's happening there.

9 MEMBER SKILLMAN: Dr. Seber, several 10 times you said the device is recording ground 11 motion. Would it be more accurate to communicate 12 device is recording the motion that the at its 13 point of location? For example, many reactor 14 buildings have their sensors up on the rim, SO 15 those sensors are up two to three-hundred feet on 16 the structure and they're actually measuring the 17 motion at that point location. Is that what you 18 really meant?

19 DR. SEBER: This one, no. But, yes, it 20 It is done to get a structure's response is done. 21 for example. That's, you know, spectra, verv 22 important for SSI analysis and people do that. But 23 what we are trying to do, since we are trying to 24 qet to а prediction equation, a model, that 25 represents the ground shaking at a certain point,

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1	we are now specifically talking about the ground
2	movement.
3	MEMBER SKILLMAN: Yes, thank you.
4	Thank you for that clarification.
5	MEMBER RICCARDELLA: And these aren't
6	at a plant. These are just generally located
7	around the country.
8	DR. SEBER: Correct.
9	MEMBER SKILLMAN: These are like
10	geodetic survey points.
11	DR. SEBER: Exactly. If you're
12	familiar with GPS, it's similar to that kind of
13	observations, available in many parts, like
14	California, Japan, like Central Eastern U.S.
15	MEMBER SKILLMAN: Thank you. Yes, sir,
16	thanks.
17	DR. SEBER: So when a ground shakes,
18	this instrument records it. What you're recording,
19	of course, is a function of the distance here
20	represented by the recording station, the magnitude
21	of the earthquake and source of the earthquake, and
22	earthquake parameters that we talk about, like how
23	the fault moved, is it like a normal fault that's
24	showing here.
25	So this is just to help us understand

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	15
1	that these recordings vary even from short
2	distances, depending on where you are along the
3	fault of distances and the magnitudes. And we
4	briefly discussed this one. Each recording is
5	three components. Here, an example, in a sense,
6	cartoon but real data, showing north/south, which
7	is the horizontal. Two of them, obviously, are
8	horizontal components, one vertical component.
9	East/west at the bottom and a vertical component.
10	Most of the ground motion prediction
11	equations that engineers or seismologists develop,
12	they are full of horizontal components. Then,
13	usually, we have conversion relationships that we
14	estimate vertical. Hence, you remember GMRS
15	horizontal, GMRS vertical, GMRS and scaling factors
16	and things. That scaling factor makes it vertical.
17	MEMBER RICCARDELLA: Excuse me.
18	Centimeters per second squared. 1g is about a
19	thousand of those?
20	DR. SEBER: Yes, 980.
21	MEMBER RICCARDELLA: Nine hundred and
22	eighty. Okay, thank you. MEMBER BLEY:
23	This is just a curiosity question. I'm sure
24	multiple government agencies and research
25	institutes and universities have sensors of their

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	16
1	own put out. Is there one central place where they
2	all feed information to?
3	DR. SEBER: There are several places
4	that data flow to. USGS has it. The academic
5	community, like IRIS, now they maintain motion
6	records. International communities maintain it.
7	Usually, if someone wants to develop something,
8	they need to communicate with those sensors
9	MEMBER BLEY: So they're kept
10	separately. Okay, thanks.
11	DR. SEBER: So if I may, I'm just going
12	to go back to the original figure. This was slide
13	eight showing you have multiple points that are
14	represented here by red dots. And, ultimately,
15	you're trying to curve to it, which is the red, and
16	then you try to estimate the sigmas. And one of
17	the questions was what do these red dots represent?
18	How did we get them?
19	And since this is PGA, which is a peak
20	ground acceleration, so we have some slides here to
21	
22	MEMBER RICCARDELLA: But each red dot
23	is one of those sensors?
24	DR. SEBER: One seismogram, yes. One
25	measurement, one earthquake, one distance. So,

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say, magnitude 6.0 at 25 kilometers at point B, 1 SO 2 this is the curtain for that. And the recorded 3 seismogram shown here since, in this specific 4 example, we're looking for peak ground 5 acceleration, it's easier. You just find wherever the peak measurement is, carry it to a chart. 6 That 7 represents one data point at this point. But if 8 there's an earthquake, it's likely recorded by 9 multiple stations. 10 MEMBER SKILLMAN: Sir, let me ask you 11 this. In our country, we have a term called NAVD 12 or NAGD, which is the height altitude data for the 13 country, and there are data stations all over the 14 United States that give the NAVD or NAGD. Is there 15 a requirement in our country for a certain matrix 16 or number of these devices so that in any ground 17 motion event there are sufficient points to develop 18 that curve, or all those points simply the result 19 of random devices that just --20 DR. SEBER: Not that I know --21 MEMBER SKILLMAN: -- happen to be there 22 because people are really smart? 23 DR. SEBER: The best thing that I know 24 of is what USGS is doing in their national seismic 25 But there are a lot of academia, a lot networks.

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	18
1	of state geology surveys, and a lot of people who
2	do that.
3	MEMBER RAY: But those points aren't
4	all just one event on the preceding slide.
5	DR. SEBER: This is an earthquake
6	happening at a point, one station.
7	MEMBER RAY: That. That's not just one
8	event.
9	DR. AKE: Actually, that is actually
10	one event.
11	MEMBER RAY: All right. Does it need
12	to be always for this process to work?
13	DR. AKE: No, this particular example
14	slide that we show here is for one particular
15	event, which is, you know, an exceptionally rich
16	data set for
17	MEMBER RAY: I was going to say.
18	DR. AKE: Usually, if you go through
19	the next couple of slides that Dogan was showing
20	you, that's more of what it's like where you have a
21	data point or two, not a very rich data set.
22	MEMBER RAY: Okay, thank you.
23	DR. SEBER: This was I guess I was
24	mistaken initially even in the first presentation.
25	This was supposed to be a cartoon showing the

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concepts, but now we understand the interest from That's why we tried to build this one the members. from one earthquake, one recording, to one earthquake, multiple recordings. This is similar to slide eight. But then if I go one more, to represent the concept, multiple earthquakes, multiple recordings, producing a big ground motion database. There's one comment in the back.

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9 MR. GRAIZER: Vladimir Graizer, 10 seismologist. Sorry. I spent probably 25 years in 11 strong motion. То answer your question about 12 distribution, yes, there is no special requirements for the whole country. But in California, there is 13 14 an agreement between California Geological Survey 15 which is main organization instrument in strong 16 motion stations, and United States Geological 17 Basically, the requirement is at least one Survey. 18 instrument in ZIP code. This is the minimal 19 requirement in California. But, of course, this is 20 a minimal. Let's say you are in the Mojave Desert. 21 This will be one instrument. But, realistically, 22 in Los Angeles, it can be a hundred let's sav 23 instruments in one ZIP code if it's а reallv 24 highly-populated area. But the minimum requirement 25 California is in one instrument. But,

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realistically, in all urban areas, there are many 1 2 more stations and they are close to the bridges, to 3 the big buildings, to make so-called input motion. 4 This is, again, to answer your question because we count structural instrumentation 5 don't in this 6 discussion that is presented today. But almost all 7 important structures have what is called free 8 field, which is counted here. 9 MEMBER SKILLMAN: Thank you very much. 10 MEMBER BALLINGER: So just to close in 11 on that question, are there cases where there's a 12 monitor on one of these devices on the ground in 13 front of a building and one on the top floor of the 14 same building? 15 DR. SEBER: That's pretty common for 16 big structures. 17 MR. CHOKSHI: This is Nilesh Chokshi. 18 There is specific requirement for seismic 19 instrumentation, particularly for the fault lines. 20 MEMBER BALLINGER: Okay, thank you. 21 DR. SEBER: This little break reminded 22 me, it probably became obvious, but we have little 23 icon on top that says open item to remind you 24 that's a new slide trying to address one of the 25 questions or concerns raised at the last time. Ι

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	21
1	should have said that at the very beginning.
2	So to go back to number eight,
3	basically, when you have all of these measurements
4	available to you, then that forms your database.
5	From your database, you can try to fit
6	representative curve to represent them all. Of
7	course, it is never going to be perfect.
8	Observations are never perfect. So much
9	variability exists in the system. And people try
10	to fit different kind of forms. If want to fit,
11	you can fit linear curve. If you want to do
12	anything you want to do, you can do it.
13	There's standards that people use these
14	days that we can go into it. We have in the
15	backup slides some of the examples.
16	MEMBER RICCARDELLA: Just to be clear,
17	GMPE is ground motion prediction equation?
18	DR. SEBER: Correct.
19	MEMBER RICCARDELLA: Yes. That wasn't
20	as I was going through the slides, I never saw
21	that.
22	DR. SEBER: Again, this is day and
23	night for us, so we assume certain things that we
24	should
25	MEMBER BALLINGER: One more. I'm

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	22
1	looking at the slide. Are there cases where there
2	are two instruments exactly at the same location?
3	In other words, is there a way to compare
4	instrument to instrument one foot away from one
5	another? You have two instruments
6	DR. SEBER: Vladimir may know the
7	examples. I'm not familiar with it. I would
8	probably do it as a test case.
9	MR. GRAIZER: Yes, we do such kind of
10	tests. And, of course, you're right. The first
11	test is basically for reliability of the
12	instrumentation. And, yes, we did for example,
13	when we transitioned from so-called analog type
14	instruments, which were recorded from the field, to
15	digital, the first question was are you getting
16	same results?
17	But, generally speaking, for purposes
18	of saving money, nobody put two instruments nearby,
19	especially after all this testing is performed.
20	Sometimes, we put what is called arrays when you
21	have close-by instruments, maybe few meters away
22	from each other. But this is for scientific
23	purposes. For testing purposes, basically this is
24	already done many times.
25	MEMBER BALLINGER: Thank you.

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23 DR. SEBER: So the final points here is 1 2 to make that, once you obtain the database and 3 create a reasonable model in a functional form, 4 then you're estimating the median ground motions 5 for the region, representing your database. Then, of course, you need to represent the uncertainties. 6 7 In this case, we'll talk about more later. This is 8 model randomness in the data set. 9 This uncertainty, at the end of the 10 PSHA actually, is going to be incorporated into the 11 PSHA. So it's not going to contribute as much to 12 fractiles in that sense because it's going to be 13 represented in the mean hazard codes. When we qo 14 through the equations and things, this will become 15 But I just wanted to -clear. 16 MEMBER BLEY: And there's kind of a 17 plot. It kind of says regardless of the magnitude 18 of the earthquake or the detailed time history, 19 there is a relationship between PTA and distance. 20 DR. SEBER: No. This is what slide is 21 missing. This is a specific magnitude. 22 So all of these --MEMBER BLEY: Okay. even if it's multiple earthquakes, it's all the 23 24 same magnitude. 25 DR. SEBER: Functional forms are

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1	developed for a
2	MEMBER BLEY: Okay. That makes sense
3	to me.
4	DR. SEBER: specific magnitude and
5	distance. And you may have other parameters,
6	depending on the developers. But given the
7	magnitude and distance, this curve goes up and
8	down, depending on what distance you're at and what
9	your magnitude is.
10	MEMBER BLEY: Thanks. That helps.
11	DR. SEBER: This is just representing,
12	in this case, say M1, whatever that M1 is versus
13	distance.
14	MEMBER BALLINGER: So one more naive
15	question. Because it's by ZIP code and because the
16	population density is a function of ZIP code, there
17	may be cases where you've got many, many more
18	instruments in particular ZIP code than in another
19	and power plants are not necessarily in the center
20	of Los Angeles, so does that affect the uncertainty
21	by ZIP code, I guess?
22	DR. AKE: I think want to be clear
23	here. That ZIP code discussion that Vladimir was
24	bringing up is really just for California. That
25	doesn't apply to anywhere else. The instruments in

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25 the rest of the country are located based on the 1 2 interest of particular projects. You see some that 3 are, as Dogan noted a moment ago, that are put out by the USGS in various different places to sort of 4 5 in a general sense, populate spatially the help, 6 country. Others are put out by nuclear power 7 large dams. plants in some cases, Some large 8 bridges and things like that, you'll have them. So 9 it's an extremely heterogeneous distribution of 10 instruments across the country. 11 DR. MUNSON: One other point to 12 is we remember have some very large earthquakes 13 that are recorded by multiple stations, and we have 14 into account, we have to to take worry about 15 whether that earthquake is unduly biasing our 16 So we have to do some weighting. So model. 17 there's more sophistication that we're not showing 18 to develop that curve, that median curve. So 19 there's kind of often a two-step regression where 20 you do an initial weighting to make sure that a 21 single event isn't biasing the results. 22 So we didn't want to get into that much

23 detail with this, but this is the general concept.
24 MEMBER RICCARDELLA: And generally
25 speaking, would you characterize this uncertainty

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1	as aleatory?
2	DR. SEBER: Yes.
3	MEMBER RICCARDELLA: No matter how good
4	your equations got
5	DR. SEBER: Usually, researchers tried
6	hard to reduce it. No matter which angle we look
7	at it, it seems to be staying similar, but there's
8	still hope that we'll reduce them at some point.
9	MEMBER SKILLMAN: Is your use of the
10	term median important in the context of what we're
11	doing here versus mean or average? When we look at
12	these curves, should we be thinking median is the
13	gold standard for our understanding of this
14	information?
15	DR. SEBER: I'll say what I think and
16	others can jump in. The median is essential here
17	because for how you represent the uncertainties,
18	then you make an assumption of Gaussian
19	distribution on the aleatory variability. And when
20	you have the median, when you have the sigma
21	defined for that, you're defining all levels of
22	aleatory that you may want to use. If you do that
23	for a mean, then you have other issues that you
24	need to deal with.
25	MEMBER SKILLMAN: Thank you.

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1	CHAIRMAN STETKAR: Before we leave
2	this, we're getting a sense of how much we know
3	and how much we don't know. Just remember we don't
4	know anything. This plot here at least gives me
5	some confidence that, if I do the simple how many
6	lines on a piece of paper does the uncertainty
7	span, the closer in to the epicenter, the
8	uncertainty is lower than further away from the
9	epicenter by, like, two lines on a piece of paper
10	versus four lines on a piece of paper. So this
11	behavior, the way you've plotted it at least,
12	reinforces the way I'd expect things to go, so
13	that's good.
14	DR. SEBER: But also remember this is
15	just a cartoon and observations.
16	CHAIRMAN STETKAR: I understand that.
17	DR. SEBER: You may see differences.
18	CHAIRMAN STETKAR: Don't become a
19	researcher on me. This is the type of behavior
20	that I would expect, so, if we go on, I'm still
21	curious about why the behavior that we're observing
22	this apparently isn't the source.
23	DR. SEBER: By the end, hopefully we'll
24	be able to explain.
25	MEMBER BLEY: And this is too simple to

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1	mention, but I'm going to mention it for other
2	people. We're talking, the way John's talking is
3	he's dropping a vertical line at a particular
4	distance and looking at the uncertainty at that
5	distance, not the other way where they look
6	parallel.
7	CHAIRMAN STETKAR: No, I was going to
8	say but the other way doesn't kind of make much
9	sense on this
10	MEMBER BLEY: It doesn't.
11	CHAIRMAN STETKAR: No, no, no, I have
12	this engineering device, and it's narrower at the
13	left end than it is on the right end by, like, a
14	factor of two on this engineering device, and
15	that's all I care about. It doesn't make any
16	difference. It's got the right shape and it's got
17	the right behavior to it.
18	DR. AKE: Just to amplify a little bit
19	on the Chairman's comment, if you think of it as a
20	vertical line, it sort of gets back to what Dogan
21	was saying a moment ago. Think about running a
22	vertical line at a given distance. Those
23	observations are log-normally distributed and,
24	hence, they can be most effectively represented by
25	a median and a standard deviation to represent

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1	CHAIRMAN STETKAR: You shouldn't have
2	said that because well, maybe it's okay, but you
3	know that it's log-normally distributed or you just
4	assume that it is?
5	DR. AKE: Actually, all of the analyses
6	that have been done to date suggest out to at least
7	three standard deviations. It appears to behave as
8	a log-normal distribution.
9	CHAIRMAN STETKAR: Okay. That's good.
10	DR. SEBER: So this was for the peak
11	ground acceleration option, as we had already
12	discussed. And we're going to remind you again
13	that the current ground motion equations, the EPRI
14	2004/2006, as well as the revised ones now, we have
15	the updated versions, 2015, they're defined at
16	seven different frequencies. So it was easier to
17	do the PGA example. It's a little bit harder, but
18	we're going to attempt to show you how other
19	frequencies of interest.
20	And there was an earlier question,
21	okay, this is your observed seismogram, how would
22	you do the other in-between frequencies?
23	Obviously, you can do filtering transformation and
24	things. But there is a catch in between because,
25	ultimately, engineers are interested in spectral

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acceleration, not the ground acceleration. We had shown this slide last time, too. Here is just a reminder what you're measuring is X-double-dot. What everybody else wants to know, you double dot. 5 The acceleration of a single degree of freedom 6 system responding to the ground acceleration. Usually, the damping here is five percent. That seems to be the case. At least all the ground 9 motion prediction equations that I'm familiar with use damping spectral acceleration. So this was one of the figures that we

put out last time and got a lot of questions, so we have some clarifications on this one, too. And we tried to clarify the figure and see what it meant, and then we have additional slides. On top, what you're seeing N/S, representing a north/south, that is the input motion from a specific earthquake that represents the ground acceleration.

But we are interested in seeing how the spectral acceleration look like from this given earthquake. So then, basically, you need to solve the differential equation here of U-double-dot and filter it for the frequency ranges of interest to you.

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In this specific case, similar to EPRI

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1	ground motions, we are showing 0.5 hertz, 1 hertz,
2	5, 10, and 25. PGA we just did, which we assign
3	usually to 100 hertz.
4	DR. MUNSON: We did leave out two and a
5	half hertz but
6	DR. SEBER: Okay, okay, sorry, yes.
7	MEMBER RICCARDELLA: You use the term
8	filter it, but really what you're doing is
9	adjusting the spring
10	DR. SEBER: Right.
11	MEMBER RICCARDELLA: mass bigger
12	than that frequency.
13	DR. SEBER: Yes, I was trying to tie it
14	to your comment earlier.
15	MEMBER RICCARDELLA: And that's
16	different, as I understand it, than a Fourier
17	transform, right? If you did a Fourier transform
18	of this, would you get essentially the same thing?
19	DR. SEBER: I would say if we get U-
20	double-dot under the Fourier transform, it should
21	be representative of this one. But you can, I
22	mean, of course, I would filter it out and put an
23	inverse Fourier transform, and that should be
24	pretty much the same.
25	MEMBER RICCARDELLA: But U-double-dot

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1	or X-double-dot?
2	DR. SEBER: Well, now we're in the U-
3	double-dot domain. If you do the X-double-dot
4	Fourier, you're going to get yes, but you're going
5	to get Fourier amplitude of the acceleration, and
6	that's not what engineers seem to like, a more
7	force-oriented, I guess. That's why they like to
8	get the spectral acceleration and a typical
9	representative of a typical building.
10	DR. MUNSON: Right. And the Fourier
11	analysis will give you a velocity, so if you take
12	the Fourier spectra of the X-double-dot, you'll get
13	a velocity in units of centimeters per second in
14	your Fourier spectra.
15	MEMBER RICCARDELLA: And then you'd
16	have to, then you'd have to differentiate.
17	DR. SEBER: Yes. Well, if you do it in
18	the Fourier domain, it's easier.
19	MEMBER SKILLMAN: In this example, if
20	you go back a slide, what is "M?" M is the mass
21	that you're trying to excite from the ground
22	motion, and so U-double-dot, which is an
23	acceleration, is really a function of M.
24	DR. SEBER: Right. M in the spring and
25	the

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1	MEMBER SKILLMAN: And so what is M on
2	the next slide? Does that image on the right apply
3	to everything?
4	DR. SEBER: That's the frequency.
5	MEMBER SKILLMAN: No, that's
6	acceleration for an M, for a given M.
7	DR. SEBER: And stiffness.
8	MEMBER SKILLMAN: Okay. And so if I
9	have the George Washington Bridge, it has an M
10	that's very different than if I have a small
11	concrete building that's anchored at the same
12	location.
13	MR. PIRES: May I make a comment?
14	Excuse me. I'm Jose Pires. I'm from the Office of
15	Research. This is a way of writing a single degree
16	of freedom, like, for instance, representing a very
17	simple representation of, for example, a building.
18	If it is a single-story building, then the M would
19	be the mass of the story at the top. But what
20	happens there, there is something missing there.
21	That is the stiffness of the spring. So if you
22	divide the stiffness of the spring by the mass, you
23	get the frequencies, the square of the frequencies.
24	So the properties of these structures,
25	so to speak, are defined in terms of the stiffness,

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1	which is K, and the mass, M, but the duration of
2	the two is the square of the frequency. So it's
3	embedded there. It's normalized.
4	MR. CHOKSHI: I think to this is
5	Nilesh Chokshi. To further clarify what Jose said,
6	I have, as you said, two examples, one way massive
7	and one small. What this figure shows that for
8	ground motion, the structure with a particular
9	frequency will see this much amplification. So for
10	my bigger mass, I will multiply it for the force,
11	mass, time, acceleration, and I will get a much
12	bigger force. With a smaller mass, I'd still
13	multiply by that amplification, but I get a much
14	smaller force. But as long as the frequencies are
15	same and the dampings are same, you can apply from
16	this analysis.
17	MEMBER SKILLMAN: Thank you, thank you.
18	DR. SEBER: As I said last time when we
19	showed this figure and the number of seven
20	frequencies and how we picked those frequencies
21	became one of the topics of discussion and the
22	figure that we had shown here at significant peaks
23	at, I think this was like 2 hertz or so, maybe 4,
24	they were not matching with the sample points. We
25	tried to explain, probably didn't do a good job at

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1	the time, that those are specifically site effects
2	that given sites amplifies or de-amplifies certain
3	frequencies.
4	In this case, there are obviously two
5	major amplifications at around 2 hertz and maybe
6	around 4 - 4 2 hertz. And, of course, when you do
7	that to a more generic ground motion prediction
8	equation frequencies, you don't capture them. It
9	doesn't mean that we don't capture them ultimately
10	in a nuclear power plant application because that
11	comes at the end with the site response correction.
12	We wanted to emphasize here that the
13	ground motion prediction equations, which is
14	misspelled here as GPME it should be GMPE
15	obviously and, therefore, generic rock
16	conditions, and that's why one needs to be careful
17	about looking at observed spectra, response
18	spectra, and where it is coming from and if there
19	is any soil impacts on that, too.
20	So just to clarify our point a little
21	bit more, we have one more example, and this is an
22	example more representative of generic rock
23	recordings. As you can see, it doesn't have those
24	big peaks and downs that we were seeing in the
25	other figure. But in this case, seven frequencies

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represent the overall range, and this is also consistent with everything else that we know out there in how engineers, researchers quantify ground motion prediction equations.

Having said that, and Vladimir and Jon 5 heavily involved, well Cliff 6 are as as here, 7 there's a new effort in ground motion prediction generation models 8 equations for the next for 9 Central Eastern United States, and the number there And I don't know the exact number 10 is a lot higher. 11 now, but it was 20 - 30, something like that, 12 different frequencies. So researchers are thinking 13 about improving these and obtaining more samples. 14 I cannot speak much to it because I do not know 15 is going on in the research area at this what 16 point, but Ι just wanted to put it out as а 17 reference that, in certain cases time \_\_\_ next 18 here discussing with you perhaps we are NGA's 19 models -- you may see more frequencies represented in the ground motion prediction equations. 20

21 MEMBER RICCARDELLA: Theoretically, you 22 could tune that spring mass amp system enough by a 23 tenth of a hertz and move it through the whole 24 spectra, right? It's just how much data can you 25 keep track of.

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1	DR. SEBER: Yes, the question is do you
2	have observations to justify the range?
3	DR. AKE: And are we gaining anything
4	by that, you know, doing that? Just because we can
5	doesn't mean we should.
6	MEMBER RICCARDELLA: Yes, yes, I
7	understand.
8	DR. AKE: The other thing, I think, to
9	keep in mind was if you look back at the slide, the
10	previous slide that had the very scalloped nature
11	to the spectrum there, the ground motion prediction
12	equations essentially are ensemble averages of all
13	of these. So when you perform that ensemble
14	averaging over a large number of records, it
15	automatically smooths these things out in a sense.
16	So a more limited number of frequencies is
17	representative of those smoother spectra that
18	represent the
19	MEMBER BLEY: But maybe you lose a
20	touch of the uncertainty you ought to have in that
21	ensemble by having, you know, missing some of the
22	peaks along the way. If it's only a few, it
23	probably makes no difference. If it's more than a
24	few, it might.
25	DR. AKE: The way that is captured in -

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1	- if you go back and think about the figure we had
2	a few moments ago, which was the cluster of dots
3	with the line through them, that is being captured
4	in that part of it the various
5	MEMBER BLEY: Where you did the actual
6	peaks, yes.
7	DR. AKE: Yes. Because you have a
8	curve like this for each one of those different
9	frequencies, so we're picking that up as that's
10	what we're really characterizing as the aleatory
11	part of this.
12	MEMBER RICCARDELLA: Yes. And what
13	that plot is is really just the same thing, only
14	with a very, very stiff spring, right?
15	Essentially, you put a rigid in that spring mass
16	dampening, and you get the ZPA.
17	DR. AKE: Exactly, yes.
18	MEMBER BALLINGER: Another naive
19	question. You've got a ground motion prediction
20	equation, and so you make predictions for a site
21	using that equation. But I'm assuming there have
22	been cases where not only have you made the
23	prediction, but there are sensors on the site for
24	the same earthquake for at least, and that's been
25	compared. How do they compare?

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1	CHAIRMAN STETKAR: Let's let him go on
2	because we're not talking about uncertainties in
3	things that we can measure. We're talking about
4	uncertainties in earthquakes that ain't never
5	happened before. That's my big concern. I
6	understand but
7	MEMBER RICCARDELLA: Well, I think
8	we're a long way and a lot of steps from that
9	comparison that you were asking for.
10	DR. SEBER: This is where we jump into
11	some of the, I'll call minor level of concerns and
12	questions that we got at the last meeting. And
13	this is just a reminder slide that we're still
14	looking at EPRI 2004/2006 ground motion prediction
15	equations. The key thing that I would like to
16	bring that's going to make, hopefully, a difference
17	is at the end, that these models come as composite
18	models. It is not just one equation that defines
19	EPRI. It is a cluster of equations. Each one is
20	composed of multiple components, so we'll go
21	through that a little bit and that will make a
22	difference at the end in the fractiles and how we
23	distributed, how we calculated and things.
24	So what we're trying to show here, we
25	always say EPRI 2004 ground motion prediction

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1	equations. Sometimes we make it sound like it is
2	something simple, one formula or both formulas. It
3	is, obviously, it is done through SSHAC level three
4	study. People sat down and looked at what's
5	available out there, and they listed all
6	reasonable, reliable models published by
7	researchers. That's what is shown here in these
8	models. And then they grouped them based on the
9	model or the assumptions or the methodology that
10	they used into four different groups, which they
11	called clusters. In each cluster, there are a
12	varying number of original papers published that
13	they need to evaluate and come up with a central
14	model that represents all these models.
15	In cluster one, for example, in this
16	slide, you have two, four, six alternatives, in a
17	sense papers published, opinions published, on what
18	should a ground motion prediction equation look
19	like. And based on the database available or that
20	was available to them at the time, 2000 - 2003,
21	they did some quick search how many of these models
22	match with the reality, i.e. the observations are
23	almost similar to what you said. And based on
24	that, they assigned certain weights to these
25	models. Those are the ones that are represented on

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1	the last column.
2	And depending on the methodology, like
3	cluster four, we have only one paper, one research,
4	one opinion, and they maintain that as just one.
5	It makes a difference at the end because,
6	ultimately, for each cluster, you're calculating a
7	median ground motion prediction equation and its
8	standard deviations. This is going to lead to the
9	question that I think the Chairman raised last
10	time, but they're also smoothing the uncertainties
11	and would this be a problem at the end? So we'll
12	come to that. That's what we're building up to
13	right now.
14	MEMBER RICCARDELLA: Just to help me
15	understand, the differences in these ground motion
16	models is in the treatment of the soil conditions
17	between the earthquake and the sensor?
18	DR. SEBER: I will say differences of
19	opinions, how we treat you have a database.
20	Everybody could start from a same database of
21	ground motions recorded, but you do it
22	independently, I do it independently. Because of
23	the approaches we use, because of the methodologies
24	we use, we may not come up with the same answer.
25	The ballpark is the same, but yours is stiffer,

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42 mine is shallower, because I heavily weighted the 1 2 distances, you heavily weighted the far short 3 So it varies. distances. 4 DR. AKE: Yes, let me clarify just a These are published models that 5 little bit here. you see here in the columns as models that were 6 7 available in the early 2000s. Some were by 8 academic researchers, some by people that were 9 working on specific projects. The Frankel et al. 10 model was one that was developed by one of the 11 researchers at the USGS for use in the national 12 seismic hazard maps. A variety of different people 13 doing this. 14 At that point in time, there was not a 15 deal, as Dogan said, а deal great great of 16 standardization in the available database. It's 17 like, you know, I knew about some data that maybe 18 you didn't know about, Jon knew about some other 19 data, so we all had sort of disparate data sets 20 that we started from. And then in addition to 21 these were typically done for firm rock that, 22 conditions in the Central and Eastern United 23 These are the relationships we see here. States. individual researchers 24 All the had 25 different functional forms that they used to fit

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Thinking back a moment ago to the plot that data. we had with all the data dots, and then we had a curve going through it, you can represent that curve by a number of different functional forms. slightly different everybody had functional So they used to fit relationships that this data, which is something we'll come back to probably in a moment here when we discuss how we came up with the median models for each of these clusters. DR. MUNSON: One other point.

10 11 Obviously, we're in the Central and Eastern U.S. 12 That's a stable continental region with not many 13 earthquakes, SO each of these different models 14 represents a different approach and strategy to 15 dealing with a lack of data that we have for the 16 stable continental for the U.S., so our data set is 17 primarily magnitude 6.0 and below, just our 18 recorded data. So we're interested in predicting 19 higher-magnitude earthquakes at shorter distances, 20 each one of these models has a different SO 21 strategy to try to develop what we call synthetic 22 data to come up with a map.

23 MEMBER RICCARDELLA: But what I'm not 24 understanding, if you and I did this on a 25 particular seismograph, we'd get the same answer,

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1	right? So it's not that the prediction is doing
2	something else besides taking that seismograph and
3	coming up with
4	DR. AKE: Right, these are all working
5	with the spectral accelerations, so it's the output
6	of that calculation, the response spectra, that's
7	the data that everyone is working with.
8	DR. SEBER: In a sense, those are the
9	red dots in the charts that we showed earlier.
10	Once you have those red dots, what I call the
11	database, not anymore wave form records but the
12	measurement of calculated points based on the
13	single-degree freedom systems. And then how you
14	fit a curve to that is more variable based on the
15	assumptions you make, based on the assumptions I
16	make.
17	MEMBER RICCARDELLA: All right. So
18	it's the curve that goes through the red dots that
19	is the results from the model.
20	DR. AKE: That's what we're referring
21	to here as the GMPE.
22	MEMBER RICCARDELLA: And you do these
23	red dots for all those different spectral
24	frequencies, as well?
25	DR. AKE: That's correct. And there's

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something we probably need to make very clear here, 1 2 and it really sort of informs the different model 3 types that are listed here. As Cliff alluded to a 4 ago, we're relatively data poor in moment the It's an area we don't see 5 Central and Eastern U.S. It hasn't been heavily 6 lot of earthquakes. 7 instrumented, so we don't have a tremendously rich 8 empirical data set. So what constitutes our data 9 for many of these models, especially the first, 10 second, and fourth clusters here, are the use of 11 developing synthetic seismograms. There is 12 processes that one can use to develop that. The 13 model at the bottom, Finite Source/Greens 14 is probably the most seismologically Functions, 15 robust and defensible way to do it, but you do that 16 at the penalty of it requires, it's greatly more 17 difficult to come up and do those calculations. The first two clusters are represented by development processes for synthetic seismograms that are much more simplified, what are called single corner, double corner point source models.

18 19 20 21 22 They represent the seismic source as a point. And 23 that really is what the data set is for everything 24 larger than about magnitude 5.5 that is used to 25 develop these models, and clusters one, two, and

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synthetic data. it's four are So based on synthetic size parameters, and that's what really is the main focus by which we broke those up into four different model types, four different clusters.

The third approach referred to as the 6 hybrid, this really relies on recognizing that, well, really what's the difference between places 9 like Japan and California that have lots and lots of data and the east is really just the crustal properties. And so they're taking those already empirically-derived established relationships 13 developed for active tectonic regions, like Japan and California, and modifying them just based on 15 the difference in cluster properties. it's So 16 essentially applying a scaling to those active tectonic region relationships.

18 MEMBER SKILLMAN: this, Let me ask 19 In the past, when I've looked at one of please. 20 those dots, I did not appreciate how much work had 21 gone in to, if you will, certifying the accuracy of 22 that dot. Let me ask this: on a regular dot, when 23 we look at a dot on one of these plots, how many 24 calculations are involved and is there a particular 25 that might have had 10,000 calculations and dot

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some that only had three? When we look at the dot, how should we think about the homework or the intensive effort that went in to identifying that one piece of data?

DR. SEBER: I don't think I can give you a number of calculations for each. But like you stated, it is a long process, and you have your wave forms. That is the only common thing you have. But in certain cases, you process the data. Again, how you process it and how I process it may be different, and there's always going to be some differences in the databases.

13 We have seen this in practice and 14 examples that, you know, in a certain year а 15 database is published using existing data, and then 16 database is revised, database red dots I'm а 17 talking about, because there is a new methodology. 18 Somebody is doing something slightly different and 19 perhaps more assumptions in how you process the 20 Processing is a big important component, how data. 21 you prepare the data to become the red dot.

DR. MUNSON: It's not just the processing, though. It's also, you know, some of that data is synthetic. It's not actual --

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DR. SEBER: In this example, yes, it is

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1	also synthetic.
2	DR. MUNSON: Yes. So the ones that are
3	real, actually, are quite simple. You just have
4	to, you know but the synthetic ones, each of
5	those involve a different source theory that is
6	applied to develop the synthetic data. So you're
7	right, some of them involve quite a number of
8	calculations to get.
9	MEMBER SKILLMAN: Thank you. That was
10	very helpful. Thanks.
11	MEMBER SCHULTZ: The weights are
12	developed on the basis of the richness of the data,
13	the faith in the models? What comes into play?
14	DR. SEBER: In this case, based on the
15	match to existing observations, how well a given
16	model matches the observation.
17	DR. MUNSON: So the EPRI in 2004, they
18	said, okay, this is our data set, we're going to
19	use this data set. These other researchers had
20	varying different data sets, so they wanted to
21	compare to fit with these different models to the
22	data set they were using and EPRI was using in
23	2004. So that's what those weights are.
24	MEMBER SKILLMAN: Did the weights come
25	out of the SSHAC process?
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1	DR. SEBER: Yes, this is a SSHAC TI
2	team's assessment of
3	MEMBER SKILLMAN: So there's an expert
4	solicitation type thing to set those weights.
5	MEMBER SCHULTZ: To that level of
6	accuracy. I mean, I don't understand why the SSHAC
7	process wouldn't have discarded some of the models
8	if it wasn't based on the richness of the data or
9	something additional to
10	MEMBER BLEY: Well, part of what
11	happened let me try this and you guys tell me
12	that I'm screwed up because I was trying to follow
13	it back then. Before they did it this way. They
14	sat down and tried to assign weights and could
15	never agree. You know, one guy wanted to give all
16	the weight to his, and the other guy and,
17	eventually, and what the SSHAC process tries to do
18	is say is there any case where the other model
19	might fit the real world, and they finally
20	acknowledged that, yes, maybe some of the time or
21	for some earthquakes that model fits, or the
22	general consensus in the community is that maybe
23	that's applicable. So this set of weights was
24	based on trying to reach a consensus kind of thing.
25	MEMBER RICCARDELLA: But, still, to

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1	have the weights to three Steve's point is you
2	have the weights
3	MEMBER SCHULTZ: It's more than
4	retention, but I can understand that viewpoint.
5	DR. AKE: Yes, you know, those weights
6	are, you know, essentially, they're obviously up to
7	sum to one within each cluster and they're derived
8	on, you know, they're inversely proportional to the
9	degree, the misfit if you will, the variance
10	between the observational data that they were using
11	that EPRI had in their database versus the
12	individual predictions of each of those equations.
13	So if you had a high degree of misfit, you got a
14	low weight. And, you know, you could have rounded
15	them to two figures instead of three.
16	MEMBER BLEY: They don't add to one,
17	and people give you grief about that.
18	DR. SEBER: So now we're one step
19	closer to addressing the earlier question. Each
20	cluster, we talked about multiple models, in one
21	case just one model. Now, how do we develop a
22	representative median ground motion that represents
23	that cluster? And then when you develop that,
24	well, this is not fully representing all the
25	epistemic knowledge, and the last bullet, or the

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second bullet in this case, says at the time the TI team decided that we're going to develop epistemic models representing the 5th and 95th percentile of that median based on the sigma that we're going to see next time and calculate and use them as weighted average. I have one logic tree, but it looks like it's coming later.

just to summarize here, 8 the So EPRI 9 models are composite models, four main clusters. 10 Each cluster is representative of several, in most 11 cases, alternative models' median. And each one, 12 and once you qet a median, you get 5th 95th 13 percentile. That's, again, the decision has been 14 made at the time they're going to represent the 15 epistemic uncertainties in that median.

16 So then cluster one gets one median, 17 two epistemics. In a sense, three different models 18 represent cluster one. And same thing for cluster 19 two, cluster three, and cluster four. That's why 20 when we talk about EPRI ground motions at the end, 21 dealing with 12, 9 for we are 9 or typical 22 background sources because, again, ΤI the team 23 decided if you have very large sources able to 24 produce very large earthquakes, RLMEs in this case 25 that you have heard about, repeated large magnitude

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1	earthquake, sources. Then you use all four
2	clusters. If you are talking about Central Eastern
3	U.S., generic background, you use three clusters,
4	nine equations, and each one becomes an input to
5	hazard calculations. Each one will become a
6	contributor to the fractile estimates at the total
7	site hazard that will come to a ten.
8	DR. AKE: So I just want to reiterate
9	this just a little bit, and this is something where
10	we recognized that there's not an easy way to
11	explain this because of terminology issues a little
12	bit here, and that is that, you know, when we refer
13	to the 5th and 95th percentile within cluster
14	models, that's representative, these are still
15	median models. These are median models. These are
16	representing within cluster uncertainty in the
17	median models, okay? And we've used the word sigma
18	here in a couple of different ways, but this is not
19	the aleatory variability signal we're talking about
20	that represents that distribution in the dots we
21	were trying to fit, right? This is the
22	uncertainty, and where that red line or that blue
23	line goes amongst those dots we were showing for
24	where's the best place for that? So it's trying to
25	represent that. This is the epistemic uncertainty

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1	in the median models, and that's, I think,
2	something that is easy to miss with the way we were
3	forced to use the terminology that is sort of
4	DR. SEBER: We have a quick example
5	next few slides. I think that will clarify some of
6	the confusion that, you know, we may be introducing
7	into the system by using the same definition for
8	multiple things. And this is a scenario, these are
9	hypothetical things. We did show them last time,
10	but I think it's worth to go into it because we
11	have additional slides to address the question or
12	issue that John was talking about.
13	We're going to be looking at two
14	hypothetical earthquakes, one far away source, New
15	Madrid in this case, recorded at site Fermi, and
16	one very local recorded to, again, nearby Fermi
17	magnitude 6.0. These happen to be the numbers that
18	the Fermi COL identified as controlling earthquakes
19	for their sites based on the PSHA results and
20	things. And one representing 6, which is the high-
21	frequency probable contributor, and one is the low
22	frequency, which is a far-away event.
23	So if you use the EPRI ground motion
24	prediction equations, these are the spectral
25	accelerations that you would get for this

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deterministic with case red representing the distant earthquake because it is so far away, even though magnitude is almost 8.0, it's very low impact at the site. The other one is magnitude 6.0, but it is very close. Obviously, the spectral acceleration is pretty high across all frequencies, low, high, medium.

8 But we just discussed these EPRI models 9 are composite model clusters. So now let's break 10 it into its components. So now what's shown on the 11 top figure is the medium, the one that we just 12 previous figure, slides 26. showed in the And 13 three of the clusters, because this is a local 14 event, because this is what I'll call the medium 15 level and it's the background source, we're not 16 using cluster four for that calculation. So we 17 have cluster one, cluster two, cluster three with 18 the weights shown on the slide on the right --

19 CHAIRMAN STETKAR: You don't use it 20 because, by rule, cluster four only applies for 21 magnitudes greater than 6.0, and this one is 22 exactly 6.0, right?

23 DR. SEBER: Six is not the magical 24 number. It is more like repeated large magnitude 25 earthquakes or characteristic earthquakes.

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1	CHAIRMAN STETKAR: How different would
2	it be if you used cluster four for that?
3	DR. SEBER: We have not done that
4	calculation. I could do it. I mean, it's an easy
5	calculation. Obviously, you just have to add it.
6	But
7	DR. AKE: It lowers the hazard.
8	CHAIRMAN STETKAR: It lowers the
9	hazard? Okay.
10	MEMBER SKILLMAN: But John's question,
11	I think, is important because, at least in my mind,
12	it raises the question of completeness. So is
13	cluster four excluded based on administrative
14	guidance or regulatory guidance or
15	DR. AKE: It's excluded because, in the
16	discussion, it was the, in the report itself and
17	the TI team, the cluster four was developed using
18	simulations really only for larger-magnitude
19	events. So it's really unconstrained at lower
20	magnitude events.
21	MEMBER SKILLMAN: So the cutoff at 6.0
22	was just an admin?
23	DR. AKE: The developer, when he was
24	doing the simulations, cluster four was the one
25	that was done for Finite Source. Greens Functions
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calculations were only done for larger-magnitude 1 2 events, so they didn't attempt to try and develop 3 anything between 5.0 and 6.5. Basically, they 4 didn't do simulations down there, SO you're 5 essentially extrapolating to those lower values and the team felt that was really not a robust way to 6 7 it. So in the calculations, typically, when do 8 we're doing the PSHA calculations, we do not 9 implement cluster four in the calculations for just 10 the distributed seismicity because the majority of the hazard in the distributed seismicity is coming 11 12 from the moderate magnitude events. 13 MEMBER SKILLMAN: Thank you. 14 MEMBER RICCARDELLA: Why did you pick 15 these particular sites particular and these 16 magnitudes for this example? 17 CHAIRMAN STETKAR: This is an example 18 because it's for Fermi where the we started 19 question. I understand that. 20 MEMBER RICCARDELLA: 21 But, I mean, the 7.9 must be an RLME location. 22 DR. SEBER: It is in New Madrid, yes. It is an RLME location. 23 24 MEMBER RICCARDELLA: But the 7.9, have 25 they had a 7.9?

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1	DR. SEBER: Yes, in New Madrid they
2	did. But this is I tried to explain, but let me
3	say a little bit more. These example scenarios
4	that we defined, those are the output that
5	obtained, I guess, Fermi obtained based on de-
6	aggregating the total hazard and then what we call
7	the controlling earthquakes, the most likely
8	earthquakes that could affect the site. And at the
9	high-frequency and low-frequency range is
10	represented by usually 1, 2 2 hertz being the low,
11	510 being the high. And those are the numbers that
12	they identified and staff had agreed to. In this
13	example, we use them, as the Chairman said,
14	representative since the whole issue started with
15	the Fermi review.
16	MEMBER RICCARDELLA: Thank you.
17	MEMBER BLEY: Let me try something
18	because I'm listening to John and to you guys.
19	What I hear you saying is that cluster C4 was not
20	used not strictly because M wasn't greater than 6.0
21	but because the underlying model that's built in to
22	cluster C4, in your judgment, didn't fit for this
23	particular
24	DR. SEBER: It is not our judgment. It
25	is the TI team that initiated the effort.

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1	MEMBER BLEY: But you find it a
2	reasonable
3	DR. SEBER: We found it, at the time
4	when this was reviewed, a reasonable
5	representation.
6	MEMBER SCHULTZ: I thought I heard it
7	was the model developer that put bounds upon the
8	magnitude of application.
9	DR. SEBER: That is correct. And based
10	on that information, the TI team techs acknowledge
11	and says, well, I cannot apply to these lower-
12	magnitude sources because, like John was saying
13	here, we don't even know if it applies to that, is
14	it scaled correctly if it's only for larger? Based
15	on all these inputs, the TI team apparently decided
16	we're going to split this into two levels.
17	Anything that you have background sources and
18	things, you're going to use three clusters. When
19	you have very large sources, like New Madrid or
20	Charleston, we're going to incorporate the fourth
21	one because that applies to that. In other cases,
22	you don't.
23	Magnitude 6.0 is a little bit somewhat
24	misleading because, when we do the PSHA for local
25	earthquakes, if the source maximum magnitude is
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1	7.0, PSHA goes all the way to 7.0 and uses one,
2	two, three clusters, not the fourth one still
3	because of definitions of how the original model is
4	set up.
5	CHAIRMAN STETKAR: Let's dwell on this
6	a little bit, though, because I got to the point,
7	and I've got my head screwed on in a certain
8	direction, that if I looked at all of the results,
9	what I was seeing is the behavior in the
10	uncertainties was, as I expect, for sites that are
11	highly influenced by the RLME sources. It was not,
12	as I expect, for sites that are highly influenced
13	by the distributed sources. In other words, sites
14	like Fermi, that are primarily influenced by the
15	distributed sources are showing relatively uniform
16	uncertainties. If I go down to Chattanooga or some
17	place, because that's an example in the report,
18	that are influenced more by RLME sources closer to
19	the site, I see an uncertainty behavior, as I would
20	expect, increasing uncertainty as a function of
21	increasing acceleration across all frequencies,
22	hertz.
23	Now, if I look at these pictures here,
24	they start to explain why that difference is
25	because if I look at the top picture, which his

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1	kind of the distributed way of thinking about
2	things, the uncertainties are pretty doggone narrow
3	up there. If I now add that cluster four for that
4	RLME source, despite the fact that it's a long
5	distance away, I start to see larger uncertainties.
6	And is it only because we're adding NC4 for the
7	large acceleration RLME sources that's driving all
8	of that?
9	DR. SEBER: It is not only C4, but it
10	definitely contributes to. And we'll show some
11	examples
12	CHAIRMAN STETKAR: Okay. I really want
13	to understand why because, if it's this notion of
14	including one cluster for some subset of sources
15	versus others, I still want to understand why the
16	world works in the top for the vast majority for
17	the Central and Eastern U.S.
18	DR. AKE: Actually, one thing you need
19	to look at here just a little bit, too, is look at
20	the y-axis. See where that is down there? It's
21	0.01g for the distant earthquake and it's 1g for
22	the nearby earthquake. That will figure into this,
23	as well, and we'll get to that in a few minutes.
24	CHAIRMAN STETKAR: Okay.
25	DR. AKE: The other thing we would note

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1	here, and I think it's something that is a little
2	more difficult to illustrate, but the difference
3	between clusters is less than the within cluster
4	variability, okay? In other words, the different -
5	_
6	MEMBER BLEY: That we can't quite see
7	in the picture.
8	CHAIRMAN STETKAR: That we can't see
9	here because you don't have the
10	DR. AKE: Yes, it's the difference in
11	the median models within a cluster is bigger than
12	the difference between, in the central tendency,
13	the highest weighted model between the different
14	clusters.
15	CHAIRMAN STETKAR: But, in principle,
16	if we could see that, it would certainly expand the
17	uncertainty if we could see the variability in the
18	medians. But it still doesn't answer my question
19	about why the uncertainty when we're looking at the
20	ground motions from the RLME sources behaves the
21	way I'd expect it and the uncertainty for the
22	distributed sources doesn't. I don't care about
23	the absolute magnitudes in a sense because part of
24	the observation was the absolute magnitudes seem to
25	be fairly narrow but, more importantly, the

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62 uncertainties didn't seem to be behaving, 1 in а 2 relative sense, the way you'd expect them to. 3 And I think just a preview DR. MUNSON: 4 what you'll see in the end is, for example, 25 5 hertz, which was tightly grouped together, as you get out to higher accelerations, then that starts 6 7 to also broaden. The lower frequencies, say half a 8 hertz, they start to broaden much earlier --9 CHAIRMAN STETKAR: Yes. 10 DR. MUNSON: \_\_\_ and the reason, the 11 main reason is there's not enough sources that are 12 contributing larger spectral accelerations. For a 13 site like Fermi where the RLMEs are very distant, 14 just lot of can there's not а sources that 15 contribute 1g spectral accelerations, so they start 16 to spread out earlier, whereas for 25 hertz there's 17 a lot of sources that can contribute 1g, 2g, SO 18 they stay tighter together. And we'll get there. 19 CHAIRMAN STETKAR: We'll get that. Ι 20 don't want to get too far ahead but . . . 21 Hopefully, we'll be able to DR. AKE: 22 represent this in a clear way on that. And I've heard that 23 CHAIRMAN STETKAR: 24 argument before. Thinking a head a little bit is 25 that I kind of get that, except that all of the

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1	sources that were closer in for Fermi were much
2	less capable sources. So when you say they can
3	produce 1g earthquakes
4	DR. AKE: Remember we're doing this for
5	rock, so we're going to have that factor also. But
6	25 hertz, 1g, that's not that high, actually, for a
7	rock site. But we'll get there in the end.
8	CHAIRMAN STETKAR: Okay, okay.
9	MEMBER RAY: Well, let me make a I
10	wasn't at the earlier meeting. Are we at all times
11	here, although there may be some things that apply
12	elsewhere, we're just talking about CEUS?
13	DR. MUNSON: Yes.
14	MEMBER RAY: That's it?
15	DR. MUNSON: For the time-being, yes.
16	But I don't want to go there because I don't know
17	which ones apply and which ones don't.
18	DR. SEBER: Keep it Central and Eastern
19	U.S. focus at this point. You're not going to go
20	wrong. A lot of the concepts will be applicable to
21	
22	MEMBER RAY: And, ultimately, when this
23	is used by licensees or applicants, it will be
24	explicitly clear that this is CEUS so people don't
25	misconstrue what we're talking about?

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1	DR. SEBER: Yes. The GMP 2004/2006,
2	so-called EPRI GMPs, those are Central and Eastern
3	U.S. specific so
4	MEMBER RAY: I'm trying to correlate
5	what you're talking about with what I'm used to,
6	and I'm not going to keep doing this, but I just
7	want to make sure we're just talking about CEUS.
8	DR. SEBER: Yes, yes, that is correct.
9	If I may, I just want to add one more topic to this
10	discussion. What also you're seeing at the low-
11	frequency ground motion prediction equations, an
12	ability of the research community come together and
13	understand and uniquely define what the ground
14	motions should be at low frequencies. There are
15	multiple reasons, not many observations. How far
16	you want to go, what, you know, seismic module
17	parameters you use. So that is actually shown in
18	this slide right away. When you add the, you know,
19	just take the four out and just look at the first
20	three, at the low frequencies, things are higher
21	uncertainties.
22	CHAIRMAN STETKAR: And that's because -
23	- because I did see the lower frequency results
24	behaving a little bit more like I'd expect them,
25	compared to the high frequencies.

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1	DR. SEBER: And that represents the
2	communities', I don't want to say disagreement but
3	lack of data, and, you know, you do something, I do
4	something, we end up quite significantly different.
5	CHAIRMAN STETKAR: It's just the state
6	of knowledge. I'd put it that way. And that's
7	fine
8	DR. SEBER: It's not state of knowledge
9	is not same in the low frequency versus high
10	frequency where you have much more observations
11	that you can focus on, so that impacts also all
12	theses issues.
13	So now we're breaking down one more
14	level. We show the generic or final model. We
15	show the clusters and their variations. One more
16	example, we talked about cluster, in this case
17	cluster one, the central point, which, in this
18	case, the median, that's the best matching model.
19	But then through the SSHAC process, the TI team
20	decides to add additional epistemic uncertainty
21	because certain seismological parameters I'll
22	put some names out: stress drop and propagation
23	effects and the things are not fully taken into
24	account. And they say, okay, I'm going to look
25	into all these five - six models in cluster one,

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1	get a sigma on those, and, using that sigma, I'm
2	going to estimate 5th and 95th percentile and I'm
3	going to assign them as epistemic uncertainties to
4	this median. In a sense, cluster one becomes three
5	curves, three different equations, related. And
6	they decided to use this kind of weighting. So
7	they're not these equal weights. The median gets
8	almost two-thirds and the remaining split into
9	between 5th and 95th because median supposedly
10	represents more of the published literature but
11	still not capturing the uncertainty available in
12	that. And then they put this one. And this
13	applies to cluster one to cluster four.
14	DR. AKE: I think, just a second here,
15	let me walk through this again just with slightly
16	different wording just to make sure because this
17	is sort of a fundamental point to some of the
18	questions you've had. Within cluster uncertainty
19	in these median models is developed first by
20	looking, you know, stepping back, for cluster one
21	we saw we had six different models. One of those
22	models that had the highest weight, the functional
23	form of that particular model was selected in each
24	of the different clusters as being sort of the
25	backbone model you were going to fit. So

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1	CHAIRMAN STETKAR: Jon, let me stop you
2	there because somehow I missed what those two
3	sentences said. So rewind and tell me what that
4	means again.
5	DR. AKE: Okay. Go back to slide 22
6	for a second here.
7	CHAIRMAN STETKAR: So I have stay
8	within cluster one. So I have
9	DR. AKE: Let's do the cluster one
10	CHAIRMAN STETKAR: I have six models.
11	DR. AKE: Right. And we described how
12	the weights were derived in the far-right column,
13	okay? They were based on a fit to the available
14	empirical data, okay? And so the Silva et al. 2002
15	single-corner variable stress drop, which is the
16	fourth column there, the highest-rated one, the
17	formulation of that model was selected as the
18	formulation for this cluster. So for each one of
19	these individual attenuations or GMPEs here, I
20	think they did 15 different magnitudes and 63
21	different distances. They exercised each one of
22	these for hard rock conditions for 15 different
23	magnitudes at 63 different distances and created a
24	synthetic data set.
25	CHAIRMAN STETKAR: For each one of the

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1	six models?
2	DR. AKE: Yes.
3	CHAIRMAN STETKAR: Got it.
4	DR. AKE: So you have this huge amount
5	of it's not synthetic in the way it's a
6	synthetic seismogram. It's the spectral
7	acceleration values, the output of the GMPE for
8	each of those, so this big data set. Then you take
9	that selected functional form and you fit it to
10	that data.
11	CHAIRMAN STETKAR: Which selected
12	DR. AKE: In this case, it was the one
13	
14	CHAIRMAN STETKAR: Only that one? Why
15	is it done that way? Why isn't it the functional
16	form of each of the individuals
17	DR. AKE: Because if you do it that
18	way, you can't get to the next step, which is
19	developing the epistemic uncertainty in the within
20	cluster uncertainty, okay? I'll explain how we get
21	there in a second, okay?
22	CHAIRMAN STETKAR: I mean, you can
23	because you can have six sets of three curves, if
24	you will. The three curves
25	DR. AKE: Which three curves? You're

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1	just going to get six curves out of this if you
2	just exercise each one of these, right? So the
3	only way what you're assuming then, if you do it
4	that way, is that the epistemic uncertainty, the
5	within cluster epistemic uncertainty is uniquely
6	represented by what's given in those six equations.
7	CHAIRMAN STETKAR: Yes. That's yes,
8	that's what I would assume.
9	DR. AKE: And that under-represents, we
10	feel, under-represents the epistemic uncertainty in
11	the within cluster variability uncertainty.
12	CHAIRMAN STETKAR: Well, I guess it
13	depends on what they look like because I could
14	take, as I started to say, six sets of models, each
15	of which have their own internal uncertainty,
16	weighted
17	DR. AKE: No, their only uncertainty is
18	the aleatory variability. Each one of these
19	represents a unique median model, okay?
20	CHAIRMAN STETKAR: Okay.
21	DR. AKE: So you have six unique median
22	models here.
23	CHAIRMAN STETKAR: All right.
24	DR. AKE: And so
25	CHAIRMAN STETKAR: You have six models.
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1	DR. AKE: Right.
2	CHAIRMAN STETKAR: You use terms like
3	median and uncertainty. I'll call them six models.
4	DR. AKE: They're six models of the
5	median.
6	CHAIRMAN STETKAR: No, you have six
7	models of the way things
8	DR. SEBER: And you're seeking one
9	model out of that six for cluster one to help you
10	in the calculations.
11	CHAIRMAN STETKAR: Well, that's what
12	I'm trying to understand what that selection
13	process does in terms of either reasonably
14	representing the uncertainty within each model and
15	overall of the six models versus artificially
16	constraining that uncertainty by some rules that
17	you may apply.
18	DR. AKE: Actually, what you see is
19	that the results of this, and I'm now sitting here
20	thinking of different ways we always do this
21	different ways to represent this, but the within
22	cluster variability represented by those six models
23	is much less than the within cluster variability
24	that comes out of the process that was used in
25	developing this. In other words, the median, best-

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1	fit median and 5th and 95th that come out of this
2	are broader than those six models.
3	MEMBER BLEY: I understand what both of
4	you are saying, so I'm having a little trouble I
5	mean, there are multiple ways to skin this cat, and
6	one is to just take the whole mass of the stuff in
7	cluster one, spread it all over, and try to sort it
8	out, which is a nasty job. And what they've done
9	is taken the most likely one, heavily the most
10	likely one, found its median, and now they're
11	trying to take the results from all the others to
12	lay out the epistemic uncertainty around that one -
13	-
14	DR. AKE: You're taking that functional
15	form and you're fitting it to that broader range of
16	data that was the synthetic
17	MEMBER BLEY: In all six.
18	DR. AKE: That was developed from all
19	six. That's then going to be what is the central
20	or highest weighted median model in cluster one,
21	okay? Now, what you have to recognize then is now
22	you have a range of values that represents the
23	within cluster variability that comes from those
24	six different models.
25	MEMBER BLEY: From all six.

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1	DR. AKE: From all six. So you can
2	calculate, if you will, a standard deviation in
3	those results around that best-fitting model, and
4	that's one piece of the uncertainty that goes into
5	developing the within cluster variability, okay?
6	The other parts, another part, a very small part,
7	is obviously there's a misfit in taking the
8	formulation of Silva et al. 2002 and fitting it to
9	this data. It's not a perfect fit. There's some
10	degree of misfit. That has to be captured in this.
11	And then there's also the pieces that give to the
12	question you had about smoothing, and it has to do
13	with representing the epistemic uncertainty in the
14	source parameter, the stress drop, as well as the
15	Q, or attenuation at distance. And those are
16	pieces that are dealt with in different parts of,
17	the latter parts of the EPRI report here, and those
18	are the things they smooth. And we can go into
19	that in a little bit more detail. Dogan's got a
20	nice
21	DR. SEBER: We have slides on that.
22	DR. AKE: But it's basically looking
23	then at those different pieces of that
24	uncertainty are what you use to develop the sigma,
25	but then you're going to simply calculate the 5th

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and 95th. And those become then, for cluster one, three alternative median models. There's actually four different pieces of that that's used to develop the within cluster variability, and it's greater than, simply by the arithmetic, it has to be greater than the variability that's represented by the number of models in the cluster because you're adding more variability to it.

9 DR. MUNSON: The only thing I would 10 add, though, they look at those six models and say 11 what are the parameters that those six models use, 12 those developers use, to come up with those models, 13 and they say, you know, that doesn't really span 14 the whole distance. You know, we would think that 15 the Q in the Central and Eastern U.S. could be from 16 this value to this value, and these six models are, 17 you know, fairly narrow in their Q selections, so 18 they actually broaden that uncertainty for the 5th 19 and 95th model. So as John said, you end up with 20 three models, 5th, median, and 95th, that are 21 actually wider than if you use the six as they are. 22 MEMBER RICCARDELLA: The O selection? 23 What is the Q? 24 DR. MUNSON: 0 is attenuation, the 25 quality factor, in the propagation. So each of

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74 those six have a different Q value, and so 1 they 2 look at that and they say, you know, that's a 3 pretty narrow range for those six, we think it 4 should be this wide, so they actually broaden it 5 even more. One other thing that we want 6 DR. AKE: 7 point here that we neglected to point to out 8 previously, virtually all of these models were, the 9 developers went out, you know, from near field 10 close in out to approximately 500 kilometers or 600 kilometers. 11 That's approximately the distance 12 range over which the models were developed, okay? 13 And that becomes an important point we're going to 14 get to in this discussion that Dogan has coming up 15 now. 16 MEMBER BLEY: I lost track of one thing in this, Cliff, at least one. 17 18 CHAIRMAN STETKAR: You can fill me in 19 at lunch because I'm depressed. 20 MEMBER BLEY: That's why I said at 21 The "they" has got me in Cliff's last least one. 22 discussion. 23 DR. MUNSON: The TI team. The SSHAC 24 team. 25 We have several SSHAC MEMBER BLEY:

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teams along the way here, right? They have a SSHAC 1 2 team that worked to put together the EPRI clusters. That was a SSHAC team, as I understood it. 3 Now we 4 have another SSHAC team that's doing a particular 5 analysis, or did that team that built the EPRI come 6 up with the uncertainty on the within cluster? 7 Okay, so it's that same, it's the same SSHAC team 8 who had people representing all of these models 9 involved. So when you came up with saying that it 10 was too narrow, that was the same people who were 11 advocating the various models and evaluating them 12 here in this SSHAC team? 13 DR. MUNSON: Right. 14 MEMBER BLEY: Thanks. That helped. 15 John, is it reasonable MEMBER SCHULTZ: 16 or can you tell me what is the, to assume you 17 called it the backbone model that was used for 18 cluster one. What was used for clusters two and 19 three? 20 DR. AKE: In each of those, it would be 21 the most highly weighted model in that cluster. In 22 other words, that was --23 MEMBER SCHULTZ: So look at three and tell me what's chosen. 24 25 DR. AKE: It was the Atkinson --

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1	MEMBER SCHULTZ: It was the 0.363
2	because that was the largest. I don't mean to be
3	humorous about it, but there was some process that
4	the team used to select. And Atkinson was chosen
5	for cluster two.
6	DR. AKE: Yes, and I believe it was the
7	Atkinson model for the hybrid model but
8	MEMBER BLEY: But even if it were the
9	other one, they look back at all three to
10	DR. AKE: And it would be right, yes.
11	MEMBER RICCARDELLA: So if I go to
12	cluster one and you got those six or seven
13	different models, and you said you calculate for
14	each one the median and the standard deviation for
15	each one of those models, right? For each one of
16	those
17	DR. SEBER: Median models. Because
18	standard deviation, at this point, you don't look
19	at that
20	DR. AKE: Yes, at this point, we're not
21	even looking at the standard deviations on these.
22	Right now, we're simply calculating the median
23	predicted ground motion from each one of these
24	models over a broad range of magnitudes and
25	distances to compute a synthetic data set that

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1	we're going to then fit our backbone model to.
2	MEMBER RICCARDELLA: But then you
3	talked about this I'm trying to understand the
4	within cluster uncertainty. Okay. So then for
5	each model, you have a 5th and a 95th percentile,
6	right?
7	DR. AKE: No. That's where I say that
8	our nomenclature is misleading a little bit. We're
9	calculating a median, you know. We're calculating
10	simply a single value, a spectral acceleration
11	value, say, for example, 10 hertz for the Silva et
12	al. 2002, the second line there let's say. We're
13	going to calculate a 10 hertz spectral acceleration
14	value for 15 different magnitudes and 60 different
15	distances, okay? So I got a big set of data there.
16	Those are simply a median estimate of the median
17	ground motion predicted by that equation, ignoring
18	the aleatory variability term in that equation for
19	the moment. We're going to do that for each and
20	every one of these. That's a huge amount of data.
21	If you think about to our earlier plot, this is a
22	big cloud of data then.
23	MEMBER RICCARDELLA: And that's still
24	just at one frequency?
25	DR. AKE: Yes. And we do this
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frequency by frequency. And we'll take that backbone model, the fourth one there, and we're going to take that functional form, Y equals MX plus B plus something. It's basically a term in natural log space. And we're going to fit that, we're going to solve for the coefficients that best fit that data set.

8 So now I end up, if you think back to 9 what we had in our earlier plot, we had a bunch of 10 dots and we had a curve going through there. We've now stuck that curve using that functional form to 11 So this 12 big data set, okay? this is just а the median value, 13 representation of the median 14 predictions. Just think of these as, this is data, 15 okay? These are data.

So now I have some range of variability that Jon pointed out about that median curve. That represents a standard deviation, okay? So that's the within cluster standard deviation.

20 MEMBER SCHULTZ: It sounds like what is 21 being done is the SSHAC team is, they've come up 22 with a weighting function, and now they're trying 23 to determine, not determine, they're trying to 24 build one model for each cluster.

DR. AKE: Three for each cluster.

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1	MEMBER SCHULTZ: Well, the three would
2	be the new median, the 5 and the 95. And you've
3	got it. You've got the median and you've got the
4	uncertainty.
5	DR. AKE: This is why I think we're
6	still not quite there. Each of those three models
7	within each of those clusters, and that's what
8	we're showing here. This is the implementation of
9	the epistemic uncertainty in the median ground
10	motion model, okay? So each one of these, there's
11	three median models within each one of these
12	clusters, and each with those three models in each
13	cluster are designed to capture the uncertainty in
14	what the true median should be for that.
15	MEMBER BLEY: And this is a new "they"
16	now, right? Because we're about
17	DR. AKE: SSHAC level three technical
18	innovations in 2004.
19	MEMBER BLEY: Okay. We're still down
20	there. Okay.
21	DR. AKE: This is the process they went
22	through.
23	CHAIRMAN STETKAR: Part of the way, I
24	mean, the problem is, for me, people tend to
25	confuse things by putting that 50th, 5th, and 95th

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1	on it. This thing to me says that they, whoever
2	they be, is they said there's a 63-percent
3	probability that one of those median models is the
4	way the world works. There's an 18.5 percent
5	probability that another one works, and that, for
6	some reason, is, let's call it worser, and there's
7	an 18.5 percent probability that another median
8	model works, and we're going to call that betterer.
9	And saying that that's a 50th percentile or a 95th
10	percentile or a 5th percentile of some sort of
11	distribution doesn't make any sense to me. So it's
12	three models that are assigned three ways. Is that
13	and they're characterized as median models about
14	which there's aleatory uncertainty.
15	DR. AKE: Each one of those would then
16	have an aleatory variability associated with them.
17	CHAIRMAN STETKAR: Right, right, right.
18	MEMBER RICCARDELLA: But then this plot
19	gives something more, right?
20	CHAIRMAN STETKAR: This is the aleatory
21	uncertainty about that median model for cluster
22	one, right?
23	DR. AKE: No, this is a representation
24	of the epistemic uncertainty
25	CHAIRMAN STETKAR: I'm sorry, I'm

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1	sorry, you're right. You're right. So this is the
2	three if you can get the colors a lot closer
3	together, this is the three 18.5, 63, and 18.5 for
4	cluster one?
5	DR. AKE: Right, right.
6	CHAIRMAN STETKAR: And there's aleatory
7	uncertainty around each of those.
8	DR. AKE: Each of these models have an
9	aleatory variability term, which will then
10	associated with that to then use in the hazard
11	calculations. And I think that's what we're
12	CHAIRMAN STETKAR: And that aleatory
13	variability is characterized, for whatever reason,
14	by log-normal uncertainty distribution.
15	DR. AKE: Right. Actually, we're
16	impressed because this is pretty opaque and you
17	guys are asking
18	CHAIRMAN STETKAR: No, no, no, no.
19	Pretty opaque?
20	MEMBER SKILLMAN: Just let me get a
21	calibration check. Slide 22. Okay. So for
22	cluster one, Silva et al. 2002 SE-VS 1.56. They
23	took that, they developed what is slide 27, which
24	shows the median plus the 5 and 95. When I take
25	that, I back up to 26 and what is on 27 is really

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the blue curve at the distant earthquake magnitude 1 2 7.9 and that's where those pieces fit. And you did 3 that for cluster two and three. But in this 4 particular case, because the magnitude was greater 5 than 6.0, you also did cluster four. But these are 6 the median curves, but you're also prepared to 7 present these on an array of all four curves. That would then give the distributions that Jon 8 is 9 talking about. 10 DR. MUNSON: So slide 26 only shows the 11 median, the middle one. 12 MEMBER SKILLMAN: But it's the blue that is on 26. 13 14 DR. MUNSON: Right. 15 MEMBER SKILLMAN: And you're going to 16 that four times because the earthquake do is 17 greater than 6.0. 18 DR. MUNSON: Right. 19 MEMBER SKILLMAN: And then when you do 20 the full data presentation, you're going to have 21 three curves for each median. 22 DR. MUNSON: We're going to have three curves for each cluster. 23 MEMBER SKILLMAN: 24 Yes, excuse me, per 25 Got it. cluster. Thank you.

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1	DR. SEBER: Those will become the
2	epistemic models that we use in hazard curves.
3	Each one will end up giving us a different hazard
4	curve from which we're going to
5	MEMBER SKILLMAN: I'm with you.
6	Thanks. Thank you.
7	MEMBER RICCARDELLA: What I'm trying to
8	understand between 26 and 27, in 26 you're showing
9	some variability because of the difference in the
10	three models.
11	DR. SEBER: And we showed a median. In
12	this case
13	MEMBER RICCARDELLA: Different models
14	within the cluster.
15	CHAIRMAN STETKAR: You show the thing
16	that's given away to 63 percent in slide 20, the
17	middle one. We'll call it the middle one.
18	MEMBER RICCARDELLA: But somehow you
19	combine I mean, are you somehow combining those
20	two uncertainties, the uncertainty on page 26 and
21	the uncertainty on page 27, or not?
22	DR. SEBER: No. What it's trying to
23	do, 27 is a breakdown of one of the curves called
24	Cl in 26
25	MEMBER SKILLMAN: It's on the bottom
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1	half of 26 and it's the blue curve on 26. The blue
2	curve on the bottom of 26 is the red median on 27.
3	DR. SEBER: So let me say it this way.
4	Maybe it will help. To get blue curve on 26, the
5	bottom one, you get, you need three of these
6	weighted average. Orange, three times three.
7	Thick red line times 0.18 times the upper and 0.18
8	times the other one. When you add them up, it's
9	going to give you the blue. This one.
10	DR. MUNSON: No, it's just the middle
11	one
12	DR. SEBER: No, it's the breakdown of
13	C1.
14	MEMBER SKILLMAN: No, it says median.
15	CHAIRMAN STETKAR: Jon, Jon, Jon, turn
16	your mike on so we get you on the transcript.
17	DR. AKE: I apologize, John. Yes, what
18	we've plotted here, I believe, Dogan, for each of
19	those clusters is the result of weighting the three
20	alternative median models in each cluster with
21	their weights. It's a little bit misleading in a
22	sense. This is simply trying to illustrate, you
23	know, Dogan was trying to illustrate a point here,
24	and I think he said it a moment ago. This isn't
25	how we actually use those. We don't combine them

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1	like this.
2	CHAIRMAN STETKAR: You keep, you keep,
3	you keep
4	DR. AKE: We keep them separate.
5	CHAIRMAN STETKAR: You keep,
6	essentially, three blues and three oranges and
7	three reds and three purples and three greens.
8	DR. AKE: So in other words, each one
9	of these nine for the clusters one, two, and three,
10	will each individually produce, when combined with
11	an aleatory variability term, individual hazard
12	curve. We're going to get to that a little bit
13	later. You'll see when we start to combine these
14	things
15	MEMBER BLEY: I think that just helped
16	me a lot. Let me say what I think. It will solve
17	one of my problems. On slide 28, the weights we
18	see here are the weights that are used in the
19	process you just described, Jon, where you keep all
20	this together.
21	DR. AKE: Yes, sir.
22	MEMBER BLEY: When you call these, for
23	the life of me I'm not quite sure why, the 5th and
24	95th and 50th and show me picture on page 27,
25	that's an illustration, but that's not what you

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1	use. You just told me that's not what you use.
2	You go back to the thing that's based on the judged
3	weighting of each of these. This seems confusing.
4	The red one seems confusing and not helpful, but
5	that's just me. Because you don't use those. It's
6	just the 95th.
7	DR. SEBER: Let me correct it. This is
8	the three that you use in cluster one. When they
9	say cluster one model
10	MEMBER BLEY: Here's the three curves.
11	DR. SEBER: These individually.
12	MEMBER BLEY: That sounds
13	DR. SEBER: You say another PSHA is in
14	the
15	MEMBER BLEY: Weighted by these weights
16	that are on
17	DR. SEBER: Weighted
18	MEMBER BLEY: Which is
19	DR. SEBER: fractiles.
20	MEMBER BLEY: But they are not weighted
21	by 0.5, 0.95, or 0.05 or whatever. You know what I
22	mean. Based by their weights, yes. That makes
23	sense, that makes sense.
24	DR. SEBER: That's why they consider
25	the epistemic all three, even though top and bottom
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1	resulted from the middle one that they estimated.
2	MEMBER SCHULTZ: I don't want to
3	confuse things, but can we go back for a moment to
4	26 and can you explain one more time how that
5	relates to 25, 27, and 28?
6	DR. SEBER: Let me even go here
7	because, remember, this one, we called it
8	earthquake deterministic approach, so we're now
9	trying to understand what these EPRI ground motion
10	models are.
11	MEMBER SCHULTZ: All right.
12	DR. SEBER: So now we have two
13	examples. Let's focus on the far away one. 7.9
14	earthquake, several hundred kilometers away from
15	one site. And then that gives you, when you do the
16	full calculation, it gives you the red curve here
17	as the best representative spectra from an
18	earthquake at that distance at that magnitude.
19	DR. MUNSON: So that red curve is the
20	weighted average of 12, right? Because four
21	clusters, three models in each cluster, so that's
22	the weighted average of 12.
23	DR. SEBER: Now we're starting to break
24	it down to show you how it was put together.
25	MEMBER SCHULTZ: How it came oh,
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1	that helps a lot.
2	DR. SEBER: Now, again, look at the
3	bottom figure.
4	MEMBER SCHULTZ: So rather than show us
5	the other ones first, you showed us that as an
6	example. And now you're showing in these slides
7	how it was built.
8	MEMBER BLEY: So each of these is the
9	weighted average of three.
10	DR. SEBER: Now we're going to show
11	only C1 how it was composed. And when you do
12	deterministic, of course, you need the first figure
13	I showed you. When you do PSHA calculations that
14	we're interested, these three become the dominant
15	branches and plus the other because this is just
16	C1.
17	DR. AKE: Just cluster one.
18	MEMBER SCHULTZ: I was trying to follow
19	it from a different direction. Now I understand
20	that you came together and broke it apart to show
21	us how it's developed. I get that now.
22	DR. SEBER: That's why later on you're
23	going to hear us talk about nine ground motion
24	prediction equations, so that nine refers to three
25	times three. And if you do it 12, you know, all

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1	four plus three. So that is a distinction.
2	MEMBER BLEY: Before the break, I want
3	to take you back for a question.
4	DR. SEBER: Sure.
5	MEMBER BLEY: Twenty-eight I think I
6	get, those weights and what they mean and where
7	they came from and how they're used. I think I get
8	that. Back on 22, which is the EPRI clusters
9	themselves, there's another set of weights that are
10	different weights produced by the same team. What
11	do these weights mean, and where are they used or
12	are they even used?
13	DR. SEBER: They're used to get the
14	median for the cluster one because, remember, Jon
15	was explaining we have six different alternatives.
16	MEMBER BLEY: We had cluster two here.
17	We were looking out on this one. You have cluster
18	two up here, and it's got weights.
19	DR. AKE: All the weights assigned to
20	each of the individual component models that they
21	use to develop cluster two for
22	MEMBER BLEY: I get that, that sounds
23	reasonable to me.
24	DR. AKE: Right. Now, the weights in
25	the right-hand column, remember we discussed this.
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1	Those were based on the fit of these particular,
2	that Atkinson-Boore relationship, let's say, the
3	fit to the available empirical Central and Eastern
4	United States data. They measured like a global
5	misfit, and they did it for each frequency and
6	essentially loaded up over
7	MEMBER BLEY: But these weights are a
8	reflection of the fit.
9	DR. AKE: The degree of fit. So
10	CHAIRMAN STETKAR: So Atkinson-Boore
11	DR. AKE: They fit the data the best.
12	CHAIRMAN STETKAR: Atkinson-Boore won
13	this game. They got 71 percent weight assigned to
14	them.
15	MEMBER BLEY: But when you come over
16	and assign weights which to use as the real state
17	of the world, if they're in the same order, they
18	only get 18 percent.
19	DR. MUNSON: No, that's not true.
20	That's different. That's not the 95th right there.
21	That's just the three individual models that we're
22	going to use to form the median. Three median
23	models, basically.
24	MEMBER BLEY: Oh, okay, okay.
25	DR. MUNSON: Those aren't the three
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1	MEMBER BLEY: These are the models that
2	came in that led to so now we have three median
3	curves in cluster two, which aren't individually
4	associated with these.
5	DR. AKE: They've lost their
6	association.
7	MEMBER BLEY: They've lost their
8	association.
9	DR. AKE: There's no longer direct
10	association.
11	MEMBER BLEY: Thank you. That helps a
12	lot.
13	MEMBER SCHULTZ: The team is developing
14	a conceptual model to predict the future.
15	DR. SEBER: Should I make it a little
16	bit more complicated? They do come to get the what
17	we call 5th and 95th percentile epistemics because
18	whatever the final cluster one median is, how it
19	varies from the other models establishes its own
20	sigma. From that sigma, then you get what they
21	call the 95th and 5th percentile and TI team
22	assigns them as alternative epistemics. So it's
23	disconnected, but there's still a little bit use
24	because that goes to the Chairman's question at the
25	end because that is the sigma that they're

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1	smoothing out at the end.
2	MEMBER BLEY: Let me try my own, to
3	repeat maybe my understanding of it. So on this
4	figure here, these weights are used to establish
5	the four curves on page 26, right? Okay. So curve
6	one, cluster one is the blue curve. That's based
7	on that whole group of seven weights on each one,
8	right? And then
9	DR. SEBER: Well, there's
10	CHAIRMAN STETKAR: They're saying no so
11	
12	DR. SEBER: The data distributions, you
13	take the median of that. That's where that stops,
14	and that median becomes C1.
15	DR. MUNSON: These weights here are
16	another level of weighting. So this is saying I
17	believe cluster one, the guys in cluster one,
18	actually cluster two, I believe their models in the
19	aggregate are a little bit better than cluster one
20	and a little bit better than cluster three. So
21	this is a TI team going
22	DR. AKE: There are three different
23	sets of weight in this, and we can go over that
24	again. But the first set of weights that are here
25	are simply taking a look at those particular

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1	relationships, how well do they fit the
2	observational data, the empirical data set that the
3	team had to work with, okay? That simply sets it
4	out. Those weights then get carried along and used
5	in developing the within cluster variability, as
6	well.
7	MEMBER RICCARDELLA: But aren't these
8	calculations with these weights, weren't they used
9	to produce the green curve on page 26?
10	DR. SEBER: No. The green one?
11	MEMBER RICCARDELLA: Cluster one is the
12	green curve.
13	DR. SEBER: Oh, cluster one is
14	DR. AKE: Those weights are used to
15	develop the weighting on the different models
16	within cluster that led to this. So it's not, you
17	know, it's not quite as direct as one would like,
18	but, yes, they inform the weighting that was used
19	to develop the cluster one
20	MEMBER RICCARDELLA: Green curve.
21	DR. AKE: Yes.
22	MEMBER SCHULTZ: You get that by
23	circling back through 28 and 27 to 26. So that's
24	the order that I go through in my head to get to
25	26. They presented 26 and then they showed it in

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1	27 and 28 how they got the median curves in 26.
2	But those are only one curve, and there are three.
3	Only one is shown of three for each cluster on 26.
4	DR. AKE: I apologize for kind of going
5	silent here because I'm thinking about the best
6	ways to illustrate this.
7	CHAIRMAN STETKAR: I'll tell you what.
8	It's 10:20. Let's take a break. Maybe you guys
9	can think about a better way, you know, during the
10	break of trying to walk us backwards or forwards
11	through this. Conceptually, I think it's
12	important. I'd really like to get through this by,
13	you know, midnight tonight. But on the other hand,
14	if we're hanging up on some of the basic stuff, we
15	need to understand it. So let's take a break and
16	reconvene at 10:35.
17	(Whereupon, the above-referred to
18	matter went off the record at 10:21 a.m. and went
19	back on the record at 10:36 a.m.)
20	CHAIRMAN STETKAR: We're back in
21	session. I don't know where we left off. We left
22	off with people scratching their head up front. I
23	know there was a lot of discussion during the
24	break, so where are we?
25	MEMBER RICCARDELLA: We were at slide

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DR. SEBER: I think we can go back and summarize the sequence and the rationale why we show what we show. It will help. Okay. So let me go back to this example.

Here the goal is to show how the EPRI 6 7 ground motion models work. Ultimately, you can do 8 a deterministic calculation if you know where your 9 earthquake is going to be, what magnitude it's 10 going to be, which means you know the magnitude, 11 you know the distance. And two examples we show near-distance 12 representing here: blue а local earthquake and the other one representing the far 13 14 much larger-magnitude earthquake.

talked about many times already We these clusters. We have a total of four are clusters. Each cluster has its own components, and now we're trying to explain how these clusters look You take any of the either blue or the red, like. you break it down to three clusters for the top figure, 26 slide, and four clusters and their distribution in a sense for the bottom one.

What we then eventually said, yes, these are clusters, but they are also made up of three different models. And here the median model

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and the two epistemics representing it. 1 I don't 2 want to say 5th and 95th because it confuses, but 3 these are epistemic models representing the range slide 22 4 of possibilities that we see on for cluster one, six alternative models. 5 This is the 6 ΤI team's best effort to capture all the 7 uncertainty in that and show up as three models 8 shown on slide 27 representing cluster one. And 9 you do this same example for cluster two, cluster 10 three, and cluster four and, ultimately, end up 11 either nine if you're using background sources or 12 large-magnitude earthquake 12 if you use very 13 sources in the calculations when you do the PSHA 14 down the road. 15 The weights that we talked about here 16 on slide 22 are used to help you get a single 17 cluster value because not every published data 18 point, I'm going to call it in this case, by these 19 others are treated equally. They are heavilv 20 treated as seeing some of them 0.56 or less heavily 21 treated, 0.034, in other cases. That guides you to 22 get what I would call the average mean for that one 23 later on representative of cluster one. 24 But one thing they do, we keep that

sigma, in this case, the variations from individual

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models in cluster one. Hwang-Huo and Silva, 1 the 2 alternative models of Silva. We keep track of it 3 somewhere. We keep it in our hands because that 4 sigma guides us. That's why the original team 5 represented these 5 and 95th percentile as That sigma guides us to establish 6 representation. 7 Eventually, we drop the 5th and 95th. these. We 8 just call them as part of the epistemic, part of 9 the range of possibilities covering the 10 uncertainties, covering the unknowns that we don't 11 know at this point. 12 CHAIRMAN STETKAR: And, again, without giving those percentile connotations, as it's used 13 14 in the actual calculation, there are three sets of 15 results that are weighted 0.185, 0.630, and 0.185. 16 DR. SEBER: Correct. And that 17 represents each cluster's epistemic range. 18 CHAIRMAN STETKAR: Yes. 19 DR. SEBER: On top of it, of course, is 20 the aleatory part that we're going to perhaps talk 21 a little bit later. We're not even discussing 22 that. We're iust sticking with epistemic 23 uncertainties at this point. 24 DR. MUNSON: So I boil down the six 25 models for to three cluster one, those six

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1	different alternative models. And now I have three
2	models for cluster one, three models for cluster
3	two, and three models for each cluster.
4	DR. SEBER: Any other questions?
5	MEMBER BLEY: For the first three
6	clusters? Also for cluster four?
7	DR. SEBER: Yes. These are the
8	questions that we were able to identify from the
9	transcript and tried to be as accurate possible.
10	These were the questions raised at the end of the
11	last meeting: how much does this impact at the end?
12	And just to go a little bit background on this one,
13	remember when we talked about, when we had a lot of
14	interest, the sigma within each clusters,
15	alternative models representing a range of
16	opinions? And we said we're going to get,
17	ultimately, one and, at the end, you get the three
18	epistemics for each cluster. This one represents
19	most of that sigma. Within cluster one, you had
20	six individual researcher's results. We tried to
21	estimate eventually, and we ended up three
22	representing the epistemic. And this one is the
23	sigma showing how much variations you have among
24	these six models based on your final model that you
25	came up with, the thick red line that we saw in the

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1	other picture.
2	Plus, because these are total sigmas,
3	then, as Cliff was talking about earlier, the TI
4	team, when they developed these models, they said,
5	well, but some of the seismological parameters that
6	even the six teams use, they don't capture the
7	uncertainty or range of values here for the Q
8	quality factor. So then they estimated those
9	additional sigmas, and they added them up, and they
10	got this total sigma.
11	So if I were to look at it, cluster
12	four, I don't think I have a slide for it yet, but
13	cluster four will have sigma, as well. But,
14	remember, cluster four has one model input because
15	that sigma for cluster four comes from
16	seismological uncertainties the TI team
17	established. They said it should be within that
18	range. And from that range, they calculated the
19	sigma. So all four clusters ultimately have sigmas
20	like these.
21	MEMBER BLEY: And we're still talking
22	epistemic?
23	DR. SEBER: They're all epistemic. So
24	this is, I just put the two slides together to
25	answer the Chairman's question to show the

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1	variations. I have cluster one and cluster two.
2	The one on the left you see is the original
3	calculated sigma based on mismatch to published
4	data set or published models, plus the added
5	uncertainty due to seismological parameter ranges.
6	And the one on the left is the original. The one
7	on the right is smoother. And at some point in the
8	game, TI team decides it is hard to match a model
9	to the uncertainties because it is more apparent
10	here in cluster two. All these bumps and things in
11	the curve, they said we're going to smooth it and
12	we're going to make an assumption that that
13	smoothing does not impact the results that much,
14	and then they stayed with that.
15	CHAIRMAN STETKAR: And let's focus on
16	this one because it's more dramatic and a
17	representation of the question that I had. If you
18	look at the original uncertainty, and I'm not
19	arguing about the bumps and wiggles. What I'm
20	concerned about is if you look at the, you're
21	calling it filtered but smoothed, the stuff on the
22	right, and if I look at high-frequency, the
23	uncertainty in the let me back up and start the
24	incoherent sentence again.
25	If I look at the uncertainty as a

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function of distance, which is the x-axis, and then 1 2 if I also look at the uncertainty as a function of 3 frequency, so if you think of it that way, when I 4 look at the smoothed stuff at higher frequencies 5 longer distances, the smoothing process and tremendously reduces those sigmas compared to the 6 7 original. If you look at those things that are 8 shooting -- essentially, from the graphics, if you 9 look at the stuff that's shooting up rapidly at 10 lonq distances on the left-hand side -yes, 11 there's a good one. Pick the one where the cursor 12 And if you look at the comparable smooth is going. stuff beyond 10 hertz, 25 hertz is not a good one, 13 14 but if you look at 10 hertz it's a dramatic one. 15 That process seems to have reduced, the sigma that 16 you're using from that smooth process seems to have 17 reduced the uncertainty a lot compared to what I 18 the original people thought involved in this 19 process was trying to convey. And that's a bit of 20 affect my concern because it may those high-21 frequency effects as we see them in the overall 22 results. 23 DR. MUNSON: So this is the smoothing, 24 in effect, is making the 5th and 95th a little bit 25 closer to the median.

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1	CHAIRMAN STETKAR: Well, when you say a
2	little bit closer, I'm sorry, this is sigma and a
3	little bit closer, you know, small changes in sigma
4	can make a big difference, depending on how you're
5	interpreting, you know, your
6	DR. AKE: There's a couple of things.
7	Let's go back and think a moment ago when we
8	described the original models that were developed.
9	Typically, nobody paid much attention, especially
10	at 5 hertz, 10 hertz, 25 hertz, in PGA beyond 500
11	kilometers. And so nobody really spent much time
12	doing simulations or anything else out there, so
13	these models are relatively unconstrained at those
14	distances, okay? The reason is those are not
15	ground motions of engineering interest, quite
16	frankly. Five hundred kilometers away from a
17	magnitude 6.5, you get about 0.02g or 0.001g or
18	something like that.
19	So these are not ground motions of
20	particular engineering interest, so there wasn't a
21	great deal of effort put into constraining the
22	models at those large distances.
23	MEMBER RICCARDELLA: Is that why, if I
24	go to, like, the third one down on the left, it has
25	that unusual hump there coming back down to almost

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1	zero. Is that because it's zero acceleration and
2	the variability of zero
3	DR. AKE: Just the variability.
4	Actually, the variability, because this is
5	representing the variability is actually decreasing
6	at far distances. The other
7	MEMBER RICCARDELLA: Well, because if
8	it's zero, we can predict that
9	DR. SEBER: Let's look into that
10	because that's a good example. Let's break these
11	curves in. The black line here is for a magnitude
12	5.0 earthquake, which happens to have more
13	disagreement among the publishers of ground motion
14	prediction equations. And what does that mean? As
15	Jon was saying, this is about 300 or 400 kilometers
16	away, magnitude 5.0. People will use different
17	kind of data sets because now we are getting the
18	margins of recordings and whether or not actually
19	your instrument has been triggered. And that
20	creates the uncertainty or I'll call it differences
21	of opinion in the models in this case because not
22	everybody sees that this is a marginal magnitude.
23	When you increase the magnitude, go to,
24	I think the last one is eight, which is the bottom
25	one. The models seem to be merging together, and

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1	then they're consistent because at that range, at
2	that magnitude, you actually do have good data or
3	good observations that you can constrain models and
4	says all the models coming together. In the
5	magnitude 5.0, all the models are going all over
6	the place because sigma is becoming very high.
7	MEMBER RICCARDELLA: I was just trying
8	to understand. There's a distinct difference in
9	the shapes where some of them have humps and some
10	of them take off asymptotically. I was just trying
11	to understand that.
12	CHAIRMAN STETKAR: Okay. I'm not
13	trying to get into the humpy stuff. Let's take the
14	10 hertz example. We've got the 10 hertz example.
15	It's the top one on the right side. If I look at
16	this in the 10 hertz example for I don't care
17	what the moment magnitude is. In the original
18	stuff on the left, there seems to be, once I get
19	out past about 500 kilometers or so, a dramatic
20	increase in the uncertainty. All of them, all of
21	them are shooting up, okay? Way above 0.5.
22	The smooth stuff on the right-hand side
23	is well below 0.5 for most of them. So the
24	question is why are we constraining the uncertainty
25	that way if the people said there's very large

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uncertainty for a 10 hertz response at large Large uncertainty. I'm not talking distances? I'm not talking about anything. about damage. I'm talking about the uncertainty in the people who put the epistemic uncertainty. together the models, smoothing process constraining Whv is the that uncertainty in the way that it does?

DR. AKE: This is dominated by, if you go through and look at the different pieces of chapter four here, this is totally dominated by the path effects part of this. So it's Figure 4-6, for example, in that document.

The overwhelming amount of the total 13 14 epistemic uncertainty here that's being represented 15 as this sigma for each cluster is dominated by this 16 And there is a significant amount of path term. 17 variability that the ΤI team added into these 18 models. They perturbed the models by using 19 different Q models in them to try and develop this 20 estimate of what the sigma is. These aren't a 21 representation necessarily of the within cluster 22 variability just between the models. This is bv 23 perturbing terms in path effects. those the 24 There's no simple way to explain --

MEMBER BLEY: Well, I think the first

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1	thing is I don't think the technical integration
2	team drew these curves as-is. These are the
3	results of calculations on the things they did
4	apply their uncertainty to.
5	DR. AKE: The first set of curves.
6	MEMBER BLEY: Yes, even the first set
7	of curves.
8	DR. AKE: Yes, the first set of curves
9	on the left. That's correct.
10	MEMBER SCHULTZ: That's the result.
11	MEMBER BLEY: What I think I heard Jon
12	say earlier, and let's stay with that 10 hertz
13	curve because that's the most extreme of the bunch.
14	First I have a question you probably answered
15	earlier, and I'm still not clear on it. Was it
16	then, after the fact, looking and seeing these or
17	running tests and seeing the effects of these that
18	the TI team, the same TI team, the same "they,"
19	decided they hadn't thought much about that and
20	what they had done in setting up the models led to
21	results they think are unphysical. It was that
22	same team that did whatever we're calling filtering
23	or smoothing, probably to pull down those few cases
24	that are extreme out there. And then whatever else
25	happens is probably the consequence of applying

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1	that same operation across the board.
2	DR. AKE: I think that's a very good
3	way to put it. It's discussed a little more
4	clearly actually in the 2013 EPRI ground motion
5	model update when they describe this process. They
6	went through a similar process. It's not exactly
7	the same but a similar process. And, you know, one
8	of the things that they looked at was, for any
9	given cluster, if I take the weighted average, like
10	Dogan was showing a moment ago, for a given, you
11	know, the weighted average of the plus and minus
12	epistemic uncertainty within a cluster, that should
13	predict monotonically decreasing amplitude with
14	distance beyond 500 kilometers. We kind of think
15	we know enough about the physics of propagation to
16	buy that.
17	In some cases, when you get these
18	really large sigmas, then when you're calculating
19	the 95th percentile, it doesn't behave that way.
20	MEMBER BLEY: And this was an effort to
21	correct that consequence
22	DR. AKE: This is an effort
23	MEMBER BLEY: of the complexity of
24	the model maybe.
25	DR. AKE: Right. To try and make sure

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1	that we had physically realistic behavior at
2	distance for some of these models.
3	MEMBER BLEY: Was their 2013 team the
4	same TI team as in the original, or is it different
5	people?
6	DR. AKE: No, no, I think there's only
7	
8	MEMBER BLEY: But they did the
9	smoothing, they did the smoothing in the original
10	document, as well?
11	DR. AKE: This is from the original
12	document.
13	MEMBER BLEY: This is from the
14	original, so the same people did both sides of
15	this. It kind of makes sense for me. I'd like to
16	understand better what they did.
17	DR. MUNSON: So I think, conceptually,
18	what's happening is that the TI team looked at the
19	models and they looked at the parameter, for
20	example this Q parameter, and said these six
21	individuals didn't have a wide enough span, so
22	we're going to make it wider. And then you're
23	developing some kind of wild results here as you
24	get out to very far distances. And so they're
25	saying, as Jon just explained, hey, we need to rein

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1	that in a little bit, so we're going to smooth and
2	come up with the curves you see on the right.
3	MEMBER BLEY: But those, again, are a
4	consequence. The smoothing was back at the level
5	of the 12 models, the 12 median models. Where did
6	they do the smoothing?
7	DR. AKE: This is a smoothing that was
8	used to develop the 5th and 95th or, sorry, the
9	upper and lower epistemic for each of the
10	MEMBER BLEY: For each cluster. So you
11	go back there and then this is the result that you
12	get after you make that change.
13	DR. AKE: That's the parking place for
14	this. That gets you this sigma here
15	MEMBER BLEY: I guess, on the one hand,
16	any time you do an expert elicitation, you want to
17	go back and show the experts the consequences of
18	their estimates, and this is a case where they say,
19	gosh, we didn't mean it to do that. But the model
20	is very complex, and we didn't know it would do
21	that.
22	MR. CHOKSHI: This is Nilesh Chokshi.
23	I think the classic example was the Lawrence
24	Livermore. If you remember the expert file and we
25	went back, and, basically, his conclusion was that

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1	you can't apply my model. I think similar things
2	that you then do and use this this far out, and I
3	think the expert will say no. I think that's just
4	what this is reflecting.
5	MEMBER RICCARDELLA: These are plotted,
6	I think, as normalized standardized deviations,
7	right?
8	DR. AKE: They are standard deviations
9	in natural log units.
10	MEMBER BLEY: Well, since they're
11	constrained at one, they must be normalized.
12	MEMBER RICCARDELLA: But I just wonder
13	if maybe you plotted them in g's you might not see
14	so dramatic an effect.
15	DR. AKE: Well, you can't, it's not,
16	you can't plot them in g's. They represent the,
17	they're in natural log units and sort of, and this
18	sort of gets to the Chairman's comment a moment
19	ago. When you get over here in some of these where
20	you see a natural log sigma of 0.5 or something
21	like that, that means that 95th percentile is a
22	factor of about four or three or something above
23	the median. So that means what's now going to be
24	an alternative median model is really far from the
25	central rate of your observations here. Is that a

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1	fair way to characterize it?
2	MEMBER BLEY: I think so, but I want to
3	come back to what Jon said earlier. I think it's
4	kind of not fair to say the curves on the left, the
5	original epistemic uncertainty curves, were what
6	the original experts intended because they put
7	their estimates at a different level of this
8	process, and this is a consequence of
9	DR. AKE: This is
10	MEMBER BLEY: and it's one that they
11	said, God, I wouldn't have, if I were asked to draw
12	this curve, it wouldn't look like this.
13	CHAIRMAN STETKAR: I get that, in a
14	sense. What I'm curious about, though, is there's
15	discussion in the EPRI documents as, therefore, the
16	values were smooth with a Gaussian smoothing
17	operator defined by a following equation. It
18	sounds like somebody just applied an equation. It
19	doesn't sound like the process that you're trying
20	to say that you went back and actually talked to
21	the people who developed the epistemic uncertainty
22	and then were presented with the results on the
23	left and they said, oh, my God, no, I don't want to
24	do that. It sounds like somebody went in and post-
25	processed things and had an equation that they

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1	applied.
2	MEMBER BLEY: And never went back to
3	the original experts to say is this reasonable.
4	MEMBER SCHULTZ: I think there might
5	have been discussions, as Nilesh suggests
6	DR. SEBER: One clarification there,
7	though.
8	MEMBER SCHULTZ: let them go where
9	they went.
10	DR. SEBER: One clarification. When it
11	says original epistemic uncertainty, it is not
12	referring to individual publications. It is a
13	product of the TI team. They get these medians
14	that we talked about, three clusters, three
15	equations within each cluster and things. And they
16	calculate themselves this epistemic uncertainty
17	based on the data that they use. And they look at
18	it and they say, oh, I want to fit a curve to it,
19	but it's too much wiggling and I cannot do it. And
20	then they say, yes, I lose some sigma which they
21	created, but I can come to that later maybe. How
22	much it will impact and what frequencies and what
23	distances are affecting it? This is my presumption
24	of what they did at the time. They say this will
25	be inconsequential at the end because it affects

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1	the high frequencies at large distances. High
2	frequencies do not travel large distances.
3	So if you say my sigma is 20. Out of
4	nothing, it's still nothing. So that is the logic
5	I think they're using, that they're okay with
6	reduction in sigma because, ultimately, there's no
7	contribution to hazard at distances. You cannot
8	have 10 hertz based on 6 at 800 kilometers which is
9	significant and meaningful because earth absorbs
10	frequencies, and ground motion prediction equations
11	reflect that.
12	MEMBER BLEY: I guess I'm going to have
13	to go back and read some of those, especially the
14	update in the EPRI. I'm much more comfortable
15	thinking they saw something they didn't like, and
16	somebody said what if I make this change and you
17	get this, and they're comfortable with it? Then,
18	gee, do you care what I do with things out beyond a
19	certain distance? Because it doesn't make much
20	difference. Then they'd have to make sure they're
21	happy with what happened everywhere else.
22	CHAIRMAN STETKAR: Because, indeed,
23	part of the Fermi seismic hazard may be influenced
24	somehow by that.
25	MEMBER BLEY: In ways they haven't
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1	thought about.
2	CHAIRMAN STETKAR: In ways they haven't
3	thought about it because of the ways that the
4	models are
5	DR. SEBER: I would not isolate Fermi.
6	If it is the case
7	CHAIRMAN STETKAR: No, I'm just taking
8	Fermi as an example because it was the first
9	outlaw. I'm not focusing on one firm, it's just
10	the first one we looked at. The second one we
11	looked at was South Texas, so I could also say
12	South Texas.
13	DR. SEBER: Sure.
14	MEMBER SCHULTZ: Yes, I wouldn't be
15	surprised if they tried to apply it and found that
16	they were getting results that weren't reasonable
17	or they just determined by looking at the right-
18	hand set or the left-hand set of uncertainty and
19	said this is not going to work out the way we
20	thought and it's unreasonable to, these results are
21	unreasonable. We ought to represent them
22	differently. And I think, Dennis, there must have
23	been that feedback to some degree, once they
24	determined to do it, either based on their own
25	CHAIRMAN STETKAR: I'm a cynic. I've

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1	talked to a lot of people, and they don't
2	necessarily think much about uncertainty, which is
3	why we're having this meeting in the first place.
4	So, you know, all of these experts who know
5	everything tend not to think very much about
6	uncertainty, other than some sort of abstract
7	statistical parameter that might be characterized
8	as 95 percent, but it's really just the weight of
9	0.185 on a curve. So that's what I'm trying to
10	probe here is how much thought really went into
11	that, or was it just somebody sitting back and
12	saying we need to smooth this stuff out and let's
13	apply some sort of equation and it seems to be
14	working okay.
15	MEMBER BLEY: And the text.
16	CHAIRMAN STETKAR: The text reads that
17	way very much.
18	DR. SEBER: I agree it does.
19	DR. AKE: There is a little better
20	discussion on this in the 2013 model discussion
21	about the development of the epistemic uncertainty
22	in that, and I can point to that part of the
23	document, as well. MEMBER BLEY: That would
24	be really helpful, for me anyway.
25	CHAIRMAN STETKAR: Yes, I didn't read

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1	the, I don't think we had the 2013 when we started,
2	did we? I read the original and I read the 2006
3	one. Okay.
4	DR. AKE: Actually, I would note that
5	the 2004 model was really the, in a sense, the
6	prototype for a level-three study. And I think we
7	learned quite a lot that we've captured in some of
8	our guidance subsequent to the recently-published
9	guidance on doing these kinds of studies. And I
10	think the 2004 - 2006 update model here, the 2013
11	model, does reflect an increased emphasis on good
12	quality documentation, and I think there is a
13	better discussion of some of these things in there.
14	DR. SEBER: We had one example of how
15	that would impact one of the three epistemic curves
16	or two epistemic curves for one cluster. And just
17	a quick example of increasing sigma by 30 percent
18	instead of what they did, and it does impact it, of
19	course. It will increase the range, how much at
20	the end this is going to contribute. You need to
21	remember the weighting for the upper and lower
22	bends is about 18 percent, and the middle one is
23	still the same, still two-thirds of the weight.
24	And for the same issue that I talked about, this is
25	seeing mostly high frequency part for distances

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1	which we don't observe. Considering all of those,
2	personally at least, I can speak and the other
3	gentlemen can speak. I'm not concerned about that
4	sigma reduction. I'm not sure that is not the
5	problem that we are having everything steadily
6	coming down and not separation as we go higher
7	frequencies, but I do not have full calculations to
8	shore it up.
9	MEMBER RICCARDELLA: But this was for a
10	particular magnitude earthquake, your plot.
11	DR. SEBER: This is the same
12	earthquake, that distant earthquake that we've been
13	talking earlier in the day.
14	MEMBER RICCARDELLA: Magnitude 6.0.
15	DR. SEBER: Pardon? 7.0, 7.9.
16	MEMBER RICCARDELLA: 7.5, okay.
17	DR. SEBER: And it is the same that I
18	think was 28 or something. And now instead of it
19	was a 3 red curves, now if you just increase the
20	sigma, it's by 30 percent randomly.
21	MEMBER RICCARDELLA: This is still just
22	cluster one.
23	DR. SEBER: One example. And one of
24	the rationale, we picked this one because cluster
25	one and cluster two, they dominate the hazard

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1	overall, about 75 percent. And this had the
2	variation.
3	DR. MUNSON: So just where we are right
4	now, after this slide we jump into the material
5	that we previously covered that kind of gives the
6	background on how we do the PSHA and develop, you
7	know, use a logic tree and, eventually, how we
8	developed the fractile hazard curves. So this is
9	the jumping point now to where the material that we
10	already covered previously.
11	CHAIRMAN STETKAR: So we're going to
12	skip that?
13	DR. MUNSON: You want to skip that?
14	DR. SEBER: There is one question on
15	the earthquake rates we received, I don't know
16	which member it was, whether or not the calculated
17	earthquake rates, are we doing something to make
18	sure that they're going to be valid for the future
19	or how do we estimate that? And in the yellow box,
20	we basically summarize it. Rates are calculated
21	based on the historical records, and we make a
22	little assumption that that is going to be
23	representative of the future earthquakes, at least
24	within the time of any engineering structure. And
25	then we use those rates to represent the future

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earthquakes. And here I put some rates and how 1 2 they vary because, ultimately, these will become 3 We'll be discussing these when we important. qo into fractiles. We have three different rates, and 4 each has eight different alternatives and things. 5 6 So we'll come to that, but Ι just wanted to 7 highlight this one on the record that we have an 8 answer to it. 9 So I'm not going to qo into this 10 because we talked about these. These are 11 representing the uncertainties. 12 DR. MUNSON: I think we do need to -- I 13 don't know. Now we're going to jump from ground 14 motion models into hazard curves, and do we need to 15 Go from where you 16 CHAIRMAN STETKAR: plan to go from. And if we need to go backwards, 17 18 backwards, we'll qo okay? Let's try that. 19 Because, otherwise, we're going to get to slide about 75 around 4:00 this afternoon. You sort of 20 21 had a plan of where you were going to take off, so 22 let's start that and see how much we have to retrench from there. 23 24 DR. SEBER: Let's go then guickly over 25 fast the slides just to very some of remind

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1	everybody. And this is how we calculate a PSHA.
2	It's a probability of a certain earthquake
3	happening, a certain distance, a certain magnitude,
4	and exceeding certain ground motions. You add them
5	up for all equations. So each one will contribute
6	to the uncertainties that we are talking about.
7	Each three components will contribute. Those are
8	seismic sources, earthquake rates, and ground
9	motion prediction equations that we talked about.
10	Ultimately, there are 9 or 12 alternative options.
11	I'm not going to go into that.
12	Here, just to go into a little bit more
13	detail, all the three that we talked about, source,
14	rate, and ground motions, the primary uncertainty
15	comes in because we have either lack of data or we
16	have quite a bit. As a consequence, we have a lack
17	of full understanding of what is going on. And
18	then it becomes a problem of defining what we know
19	and what we don't know. That's what we call the
20	uncertainty, and that one contributes the most in
21	our logic trees and the uncertainties that we
22	model. And we'll go to examples for that. And
23	here I'm talking like seismic source geometries.
24	Is it square or is it rectangle? That's an
25	uncertainty. And some people say it's square, and

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1	some people say it's rectangle. Okay.
2	And seismic sources largest maximum
3	magnitudes, that is purely personal in some level
4	based on what you know about your region and based
5	on typical persons' assumptions what is valid based
6	on the evidence that we have, which is very
7	limited. Can this region have a magnitude 6.0 or
8	can this region have a magnitude 7.0? A group says
9	6.0, a group says 7.0, and the TI team building the
10	model says, well, I'm going to put this in an
11	uncertainty tree and I'm going to give 50 percent
12	6.0, 50 percent that becomes an uncertainty.
13	And the other one is thicknesses, this
14	buzzword seismogenic thickness basically says how
15	deep the earth waves could go within the location
16	within the source you're interested in. If I say
17	10, Jon says 15, okay, here we go, another
18	uncertainty. That goes into a logic tree, as well.
19	And the examples of the seismicity
20	rates, we have uncertainties in earthquake
21	locations that indirectly goes into rate
22	calculations. Our earthquake catalogs are not
23	complete, meaning that, you know, I wish I had
24	hundreds of thousands of years of earthquake data.
25	We don't. A certain level. The earth works a lot

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slower time frame, so then you need to create some 1 2 mechanism how long in time history you have a good 3 complete record of earthquake, given the magnitude 4 earthquake. Usually, the larger the magnitude more complete record you have. 5 You can go deep in time. The smaller the magnitude level, like 3.0s, 4.0s, 6 7 you may be going on only like 30 years of complete 8 records. 9 MEMBER RICCARDELLA: Isn't there some 10 geological studies that go into this where they 11 look at --12 It will be only valid for DR. SEBER: 13 large-magnitude earthquakes. It's not going to 14 give you an indication of 3.0s, 4.0s, and 5.0s. 15 MEMBER RICCARDELLA: Yes, I understand 16 that. 17 DR. SEBER: But, yes, I mean, we do 18 look into instrumental, basically, recordings and 19 go back to, you know, by the century. And geology 20 records can go thousands of years back, so we try 21 to make best use of all the available data. But 22 the point is available data is very limited for the 23 time frames that we are looking at. 24 So this one just lists those 25 uncertainties, and this is, the point in this slide

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1	is the models, just like the ground motion
2	prediction model, models define the uncertainties
3	and put in a structure. This, at the end, results
4	in an objective look within the model. So if you
5	were to do a calculation and I would do a
6	calculation, the expectation is that, because
7	everything is defined in the model for you, you're
8	going to get similar answers. And that is an
9	important point to make because these models are
10	used uniformly for all new reactor applications,
11	either ESP or COLs.
12	That eliminates personal judgments in a
13	lot of the places because the model provides where
14	you should put uncertainty into the system and what
15	that number should be. In the example of seismic
16	sources, the model says for this given source, site
17	source A, the uncertainty for maximum magnitude,
18	meaning the largest earthquake that we can have on
19	that, the uncertainty ranges from 6.0 to 7.5, and
20	the model even provides the weighting how 6.0 is
21	going to be contributing, say 15 percent or
22	whatever, and where the dominant weight is and
23	where the extreme end, say maximum, say 7.5 or so.
24	So those are important concepts for us
25	to understand because neither of us subjectively

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1	reevaluate tools. The only exception is if there
2	is new information, a new publication that says,
3	oh, your model says maximum magnitude here, 7.0, we
4	just had a 7.5 earthquake. Then, of course, we
5	modify the uncertainty parameters in the system
6	prior to running the PSHA calculations.
7	So in a sense, having the models
8	structured and the models giving us what
9	uncertainty we use, which parameters we should use,
10	it makes it more objective and it makes it more
11	robust and it makes it easier for us to evaluate.
12	In the absence of new information, the output
13	becomes, in a sense, a mathematical calculation.
14	DR. MUNSON: So what Dogan is referring
15	to is we published this in 2012, but we got
16	together with the Department of Energy and EPRI and
17	developed seismic source models for the Central and
18	Eastern U.S. so that defines the possible sources,
19	seismic sources, that could impact different sites
20	in the Central and Eastern U.S. And what Dogan is
21	referring to it was a SSHAC level three process
22	where they got together and, you know,
23	parameterized the model, developed logic trees,
24	captured the uncertainty, so that there's a
25	structured framework for new reactor applicants,

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1	they will all be using this model. And I think we
2	handed out CDs of this at the last meeting, so you
3	probably have it somewhere.
4	MEMBER BLEY: We've gotten it many
5	different ways.
6	MEMBER RICCARDELLA: You're not
7	referring to NUREG-2115, all seven volumes.
8	DR. SEBER: Yes.
9	DR. AKE: I think it's also useful to
10	point out here, too, that, in addition, any future
11	applicants, obviously, in addition to utilizing
12	this model, they would still be responsible for
13	looking at the, you know, immediate vicinity of
14	their site to ensure that there is nothing that,
15	any new recent relevant information that would be
16	brought to bear in the immediate vicinity of their
17	site.
18	MEMBER RICCARDELLA: But on your
19	previous slide, the second bullet says they used
20	either 2115 or the older EPRI-SOG model.
21	DR. SEBER: Some applicants and we're
22	going to come before you when those come in. They
23	still maintain EPRI-SOG models as their design
24	basis models.
25	MEMBER RICCARDELLA: But then aren't
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1	they going to have to update? I mean, the
2	operating plants have to update.
3	DR. SEBER: Well, we look at it during
4	the review, and we'll provide you the full we
5	do, they do look at it. They need to confirm that.
6	The old model is at least as conservative as the
7	new models and justify its
8	MEMBER RICCARDELLA: So if they're SSE
9	by the old model, it bounds the ground motion
10	response
11	DR. SEBER: Even though the models they
12	used are not the most recent models. Having looked
13	at the new models, calculated the results, still
14	the differences are nothing or old models are still
15	pretty conservative. Hence, we'll come to those
16	kind of decisions at some point. So don't be
17	surprised when you see new COL reviews with using
18	still EPRI models as the design basis, and there's
19	going to be rationale for that.
20	MEMBER RICCARDELLA: You know, I guess
21	I could see how you'd make that simple comparison,
22	and it would be adequate for design. But if you're
23	getting into trying to do a seismic PRA or even a
24	margin, I think you'd need the more recent models,
25	wouldn't you, with all the uncertainties?

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1	DR. SEBER: Just for the sake of the
2	old models, the old models had broader
3	uncertainties because they were average of six
4	different models, each one similar to this NUREG-
5	2115, and they all produce their own uncertainties.
6	You can imagine when you put six people together or
7	six groups, in that sense multiple people, and they
8	produce varying results when they get the fractiles
9	up, all those.
10	CHAIRMAN STETKAR: Is there any reason
11	to believe just because different people looked at
12	it in a different year that the fundamental
13	uncertainties ought to get smaller?
14	DR. SEBER: No, absolutely not.
15	CHAIRMAN STETKAR: But they have.
16	DR. SEBER: Yes. I mean, I'll tell you
17	my personal perspective on that. In the EPRI
18	models, which I had major concerns and problems
19	with validity and scientific accuracy of those,
20	many of the models developed in '89. In the
21	reviews, we looked at it many times. We updated
22	most, if not all, of those models. But it was
23	still a built model that six different groups
24	looked at it. Now we're reduced down to one very
25	large effort. And looking at all available data,

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1	it is natural that you're going to reduce the
2	uncertainty in that.
3	MEMBER BLEY: If you treated it well
4	the first time.
5	DR. SEBER: Well, still we need to look
6	at what we know in science and which ones are
7	justifiable. No question 2115 model is much more
8	justifiable from a scientific perspective and what
9	we know about the earth and U.S. and tectonics.
10	The other ones are not as justifiable anymore.
11	MEMBER BLEY: So to Pete's point,
12	before one of those new plants operates, it has to
13	do a complete site-specific PRA. If they base
14	their seismic PRA on the older models, it would be
15	incumbent upon them, I suppose, to show that their
16	uncertainty at least includes the newer one, that
17	they are, in that sense, conservative. We haven't
18	none of those have been done, so we're talking .
19	
20	MR. CHOKSHI: I think not only the
21	uncertainty but also the ground motion
22	characteristic, that's where it matters. So, yes,
23	you're to think about that.
24	CHAIRMAN STETKAR: That's for the new
25	plants. The existing plants, many of them will
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1	need to do some sort of seismic PRA to get out, I'm
2	sorry, to address the current post-Fukushima
3	concerns. So I don't know and we haven't seen
4	any of those, obviously, either. So it's not just
5	new licensees going forward. It's a fair number, I
6	believe, of operating plants. We'll have to do
7	some sort of, about a third
8	MEMBER RICCARDELLA: About one-third of
9	the CS plants I think the new GMRS exceeds the SSE.
10	CHAIRMAN STETKAR: Right. And they're
11	going to need to do some sort of probabilistic
12	analysis.
13	MEMBER RICCARDELLA: They're using the
14	term seismic margins assessment or seismic PRA.
15	That's what
16	DR. MUNSON: They're all doing seismic
17	PRA.
18	MEMBER RICCARDELLA: They're all doing
19	seismic PRA? Okay, thank you.
20	DR. SEBER: I don't think I need to
21	focus on this one. This is the logic tree. We
22	showed these last time, but I want to just
23	reiterate some of the key things because this is
24	getting closer to the fundamental question we are
25	trying to address. Here there are three examples

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1	that we're going to be showing you and what is
2	shown is what it says it is titled, the sensitivity
3	study to M-max in this case. And we're going to
4	show 1 hertz and 10 hertz curves for three main
5	parameters.
6	This is the source parameter, and all
7	it does, everything being the same for a given
8	source, how the seismic hazard curves, in this case
9	now we're in the hazard curve domain, changes given
10	spectral acceleration. If you look at these two,
11	10 hertz being tighter and 1 hertz being broader
12	range, so this is one we know that M-max
13	contributes at least at the lower frequency to the
14	spread of these hazard curves more than it does to
15	high frequencies. These are test results from the
16	same report that Cliff was showing, NUREG-2115.
17	The other one, this is everything being
18	the same for a specific site, 1 hertz hazard curves
19	calculated at these nine different using nine

18 19 calculated these nıne different, using nıne at 20 different ground motion prediction equations that 21 we just talked about earlier in the morning. And 22 that provides the spread of that. And this is the 23 same for the 10 hertz frequency. And the same 24 story at the higher frequencies. The range is 25 tighter than the lower frequencies just due to the

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1	change in the ground motion prediction equations.
2	This is the third example we have. The
3	rate changes. And we didn't go into too much, but
4	in the new models, NUREG-2115, earthquake
5	recurrence rates are described as three alternative
6	sets, case A, case B, and case E, and their
7	collective contribution is shown on this slide for
8	1 hertz and 10 hertz. In this case, the spread is
9	larger high frequency at the higher parts or lower
10	accelerations. And in this case, 1 hertz
11	sensitivity, it is almost uniform across the range.
12	So these three cases tell us what
13	individual parameter contributes the spread of
14	these hazard curves at the end and some, obviously,
15	like M-max, more than the others. That was the
16	point to make.
17	And now we are starting to build the
18	seismic hazard curves for a more complex, more
19	realistic example. And this slide shows very high
20	levels still how one component of the seismic
21	source model is set up. This is the distributed
22	seismicity model. We have two high-level models,
23	distributed seismicity and we split it into M-max
24	and seismic tectonics as two separate models.
25	MEMBER RICCARDELLA: The top is

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1	DR. SEBER: This is parallel to it.
2	This is one part of it. I have slides coming to
3	that. Ultimately, you do the distributed
4	seismicity models. You do the RLME and add them
5	up. But this is one logic branch of these total
6	sources.
7	CHAIRMAN STETKAR: Dogan, let me, I
8	have to apologize because I was trying to follow
9	two things at once. You went through slides, I
10	don't know, I don't have the number, 49 through 55
11	pretty quick, a bunch of curves.
12	DR. SEBER: Yes, that was
13	CHAIRMAN STETKAR: What was the point
14	of doing that for the moment?
15	DR. SEBER: To show which parameter
16	impacts the spread that we are looking at in the
17	hazard calculations the most for the level of
18	spread for specific examples. This is, as the
19	title says, some sensitivity study to check
20	parameter space and see the hazard curves and how
21	much spread we observe by keeping everything the
22	same except that parameter that's being analyzed.
23	MEMBER RICCARDELLA: So these are all
24	one fixed distance from the source, right?
25	DR. SEBER: Differences. On this one,

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1	slide 50, what I'm showing here, one site, one
2	source, and what we are changing M-max within that
3	source.
4	MEMBER RICCARDELLA: Yes, I understand.
5	But that one source is a fixed distance from that
6	site, right?
7	DR. SEBER: Well, I can show this is
8	the source and your site is here and this is the
9	source. The source is fixed, but then you start
10	using the alternatives. And one alternative view
11	says this PEZ-N source could have an earthquake of
12	maximum magnitude 6.0. The alternative view is,
13	no, 6.5, 6.7. So then that's what this one
14	represents. In this case, it goes from 5.9 to 7.9.
15	And you did the calculations for that specific
16	source at that specific location, and you calculate
17	the seismic hazard curves at 10 hertz and 1 hertz
18	on the previous slide, and you'll look at the
19	outcome. The outcome is the hazard curves and
20	their spread across spectral acceleration.
21	MEMBER SCHULTZ: What's represented in the
22	parenthesis here?
23	DR. SEBER: Those are the weights that
24	you're going to be using eventually in the larger
25	tree to get the median, to get the mean, as well as

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1	the fractile, curves. Those are something that we
2	keep track of for every seismic hazard curve that
3	you calculate. Eventually, they're going to be
4	very helpful in estimating the fractiles.
5	MEMBER SCHULTZ: For this set that you
6	selected to plot, is that what
7	DR. SEBER: In this set, they're just
8	referenced only. They are not contributing to any
9	of the hazards.
10	MEMBER SCHULTZ: Right.
11	DR. SEBER: Because they are shown as
12	if weights are 1. And I you were to use magnitude
13	5.9, that's the blueish-cyan type curve that you
14	get. And if you were to use 7.9, that's the hazard
15	curve that you get from that source.
16	MEMBER SCHULTZ: But other examples add
17	up to one. That's why I'm trying to understand
18	whether this is a partial set
19	DR. SEBER: Yes, this is probably part
20	of the logic tree. This is just one snapshot of
21	one thing.
22	MEMBER SCHULTZ: I got it. Thank you.
23	MEMBER RICCARDELLA: But there's also a
24	big difference in the horizontal scale between
25	DR. SEBER: That is actually going to

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1	be
2	MEMBER RICCARDELLA: 49 and 50.
3	DR. SEBER: That's one of the
4	conclusions that we're going to end up with.
5	MEMBER RICCARDELLA: Like, if I'm, you
6	know, 1g on page 50, it's relatively tight. If I'm
7	lg on page 49, it's way off the end of the scale.
8	DR. SEBER: Because of the annual
9	frequency of accidents. Yes, it's very important.
10	MEMBER RICCARDELLA: Because the spread
11	is going to be huge. Okay.
12	DR. SEBER: And the other 51 and 52
13	where the GMRS and GMPs and the last one was the
14	earthquake recurrence rates alternatives.
15	DR. AKE: Something to sort of keep in
16	mind is, looking at 51 and 2, is that, you know,
17	for a given ground motion value, the variability
18	introduced by, in this case the nine different
19	median models we described and spent the morning
20	talking about, is really very large. You look at
21	this example, which is 1 hertz, and you look at,
22	say, 0.2g, you have a full two orders of magnitude
23	difference in the annual exceedance frequency at
24	0.2g for 1 hertz at this site. Two orders of
25	magnitude just from that one parameter.

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1	MEMBER RICCARDELLA: Okay. But at 10
2	hertz, it's much tighter.
3	DR. SEBER: Yes. It goes back to
4	these, you know, low ground motions and which
5	cluster you use and the uncertainty in the low
6	frequencies versus uncertainty that we have, or
7	confidence I should say, prediction of low
8	frequency versus confidence in the high frequency,
9	that's reflecting itself in these hazard curves.
10	No further questions?
11	I think we were discussing this one,
12	and we just highlighted M-max sources and their
13	logic trees. The point we are trying to make in
14	this series of slides maybe I should say that so
15	that everybody is with us, the logic trees
16	representing the seismic source models are
17	extremely complex. There are thousands and
18	thousands of branches when you open them up and put
19	them together, which means you're going to be
20	calculating thousands and thousands of individual
21	hazard curves for a given frequency that you're
22	going to be calculating all fractiles from and
23	ultimately establishing the final answer in the
24	hazard calculations.
25	So now we're going deeper and deeper

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into the logic tree just trying to understand the components. And in this one, it says single source This is an alternative that versus double source. the SSHAC team came up with, and we not only keep track of these models but also we keep track of each individual weights assigned to these models so that at the end we can get an accurate estimate of the fractile calculations. And this is the lower branch of slide

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61, distributed seismicity logic tree. This is developed as an alternative to M-max in the absence of full knowledge of which one is correct.

13 MEMBER BLEY: When you review а 14 submittal based on these kinds of calculations and 15 development of these logic trees, how deep do you 16 dig in and do they give pretty good justifications 17 for, like on the one you have up there for the 18 wide 80/20 split between and narrow 19 interpretations?

These are approved models, 20 DR. SEBER: 21 so we take the numbers as they are, unless there is 22 23 MEMBER BLEY: Oh, this is only 2115. 24 That's right. 25

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DR. SEBER: 2115. So NRC endorsed the

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1	starting models. The keyword is starting models.
2	MEMBER BLEY: Starting models, yes.
3	DR. SEBER: If there's an update, we
4	expect any applicant to revise it and change those
5	and now we're actually asking them to do some at
6	least SSHAC level two processing the update so we
7	have some confidence in the updates, as well.
8	Let's assume a site that you pick and
9	they do analyses and they see that there's really
10	no need to update the models, everything looks
11	good, then we never look into, oh, did they use 0.4
12	or 0.6? We don't question it. We do look at that.
13	We expect them to use 0.4 here and M-max as 0.6.
14	Of course, they have the option of changing them,
15	but then they need to give justification and we
16	need to be okay with that, why that justification
17	is valid.
18	MEMBER RICCARDELLA: Could you just
19	give me a one- or two-sentence description of the
20	difference between an M-max source and a
21	seismotectonic source?
22	DR. SEBER: Sure. I mean, M-max source
23	makes the assumption that, tectonically, you don't
24	know much about the region, so you just pick a big
25	range, big area, maybe split into in this case,

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1	everything is uniform in this slide. I call this a
2	single source. Entire Central Eastern U.S. can be
3	characterized as a single source. That's what this
4	one says.
5	MEMBER RICCARDELLA: Okay.
6	DR. SEBER: And this is part of the M-
7	max. Within that M-max, alternative representation
8	is, well, you know what, there is some tectonic
9	boundary in the middle of the Central Eastern U.S.,
10	let's represent those as two alternative
11	representations. So you don't go into, oh,
12	Appalachian Mountains in this one or Virginia is
13	over here and Tennessee there and things. You
14	don't get into those details. That represents one
15	line of thinking in the process.
16	MEMBER SCHULTZ: Dogan, it seems like
17	the decision and the proposal to change the
18	approach by an applicant should be a very high
19	hurdle. I mean, you mentioned you would require
20	them to go and get some, another team or a team to
21	look at this and provide that kind of
22	justification.
23	DR. SEBER: They need to do the
24	research and provide that there is a need to
25	update.

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1	MEMBER SCHULTZ: In the first place.
2	DR. SEBER: In the first place.
3	MEMBER SCHULTZ: And then in the
4	process
5	DR. SEBER: Once that is established
6	and we ask them to do SSHAC level two kind of
7	updates because now we consider this as a SSHAC
8	three level, we have new guidance, 2117 I think,
9	and there's going to be even revision to that soon.
10	Jon knows
11	MEMBER SCHULTZ: So the height of the
12	hurdles that they need to jump are relatively well
13	defined.
14	DR. SEBER: Yes, it is. It is well
15	defined.
16	MEMBER SCHULTZ: Okay, all right. So
17	we can
18	DR. SEBER: All expectations. I mean,
19	as you know, everything is guidance. They don't
20	need to follow directly and etcetera, etcetera.
21	But at least the procedure is set up for how to
22	proceed forward if you fall into this category and,
23	if you go that path, easier review.
24	MEMBER SCHULTZ: Yes, I understand. So
25	the guidance provides a set path of hurdles to jump
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1	and there can be variation from that, but that also
2	needs to be further justified.
3	MEMBER RICCARDELLA: So on the previous
4	figure with all the dots, you randomly just assume
5	earthquakes at all those dots? Is that what you
6	do?
7	DR. SEBER: These are actually real
8	observations. The little dots that you see on this
9	figure, those are earthquakes that happen in this
10	part from the catalog that was published as part of
11	NUREG-2115. And when it becomes important, using
12	those earthquake locations and their distributions,
13	you calculate recurrence rates. So even though we
14	say single source, every point within the source is
15	not equal in the sense how often you could see an
16	earthquake in that part. If you look at this slide
17	in the upper part, there are very few earthquakes
18	identified.
19	So, naturally, when you calculate the
20	earthquake rates per time for a unit area, you're
21	going to get very low values. Contrary, if you
22	look at New Madrid or Eastern Tennessee or parts of
23	Central Virginia and things, more earthquakes and,
24	hence, you're going to get higher earthquake
25	recurrence rates. It will eventually contribute

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1	more to your seismic hazard in those regions.
2	MEMBER RICCARDELLA: Okay.
3	DR. SEBER: This was the alternative to
4	within M-max sources. Now, alternative to M-max is
5	the seismotectonic sources. Again, these are all
6	defined by the model that we now endorsed as the
7	weights and their geometries, and these are now
8	we dig a little bit deep into seismotectonics.
9	In this case, you're starting to use
10	more geology, geotechnical information, tectonics
11	information, sorry, tectonics information to draw
12	the boundaries of your potential sources that
13	you're interested. You know, each one has certain
14	names, you know. Each one assigned certain
15	geometries. And with that, each source that you
16	see come with its own distribution of potential
17	maximum magnitude earthquake definitions. Those,
18	again, will become part of the logic tree.
19	MEMBER RICCARDELLA: And frequencies?
20	DR. SEBER: And the recurrence rates.
21	Each one will have its unique recurrence,
22	earthquake recurrence calculations.
23	MEMBER RICCARDELLA: And this was all
24	done with these weight factors in 2115
25	DR. SEBER: Yes.

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1	MEMBER RICCARDELLA: for each
2	location, right, or each site?
3	DR. SEBER: That is the stable part of
4	the model, you know. Once you are confident that
5	there are no updates, you can literally take the
6	model and put them in your computer and calculate
7	the whole thing. So these are just showing
8	alternative models, and I'm not going to go much
9	into it.
10	And this I think we talked about at
11	high level M-max versus seismotectonics.
12	MEMBER RICCARDELLA: Wait. Go back to
13	RLMEs.
14	DR. SEBER: This is actually a good
15	point because I was wondering why we have this.
16	This is to introduce us to RLME, as we discussed
17	very briefly. M-max zones. Alternative to M-max
18	is seismotectonics. But as we talked a little bit
19	earlier, M-max zone is not enough. Then you need
20	to add to it RLME sources and RLME, of course,
21	either this option or that option, you still need
22	to incorporate RLME because they are additions to
23	your seismic models. Either distributed M-max or
24	seismotectonics are alternatives. Each one needs
25	to have the RLMEs built into that.

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1	And RLMEs, just to refresh you, those
2	are the repeated large magnitude earthquake
3	sources. In these regions, there are at least
4	historically-defined minimum two earthquakes that
5	happened with magnitudes 6.5 or more. The most
6	famous ones are New Madrid and Charleston.
7	Through this work that we discussed,
8	2115, they identified, they delivered more work,
9	paleoseismology work, and they identified
10	additional RLMEs. And now this model incorporates
11	all these seismic sources available in the Central
12	Eastern United States.
13	MEMBER RICCARDELLA: What about the one
14	that recently hit near the North Anna site? That's
15	not an RLME?
16	DR. SEBER: That is not an RLME. That
17	is a known seismic source.
18	MEMBER RICCARDELLA: It wasn't greater
19	than 6.5?
20	DR. SEBER: It was not. And whether or
21	not repeated. To be an RLME, it's a pretty high
22	bar. You need to have at least well-documented two
23	earthquake 6.5 or larger in the past, and that
24	source doesn't qualify for that. But it still is
25	used as a background source, of course.

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Then this goes into what Fermi did. This is now again setting up ourselves, we're setting up ourselves to get into the final hazard calculations using the Fermi example. We have Mmaxes, recurrence rates, GMPs. We identify how many sources impact at a given site.

7 I think I have, yes, this slide. For 8 the case of Fermi again, they identified, they took 9 the 2115 model with all the sources defined in it, 10 they selected the ones that impact their hazard calculations. 11 You can imagine very far, small 12 seismic sources will not contribute anything to the 13 seismic hazards, so that's excluded by default very 14 far, very low contributing sources. You know, for 15 example, if you're doing Fermi, you don't need to 16 use Oklahoma RLME source because, by definition, 17 it's a very low hazard contributing source for that 18 If you're in Texas, of course you would do site. 19 that.

So this is the list of seismic sources 20 21 that Fermi used. And to us, what it means, and to 22 well, you're going to take all them as these 23 sources, calculate following seismic the logic 24 trees, seismic hazard curve for all these options 25 up be thousands and thousands of which end to

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1	hazard curves at the end. And then from that,
2	we're going to calculate the fractile hazard curves
3	that we are questioning versus broader range.
4	MEMBER RICCARDELLA: Now, you said this
5	is what Fermi used, but my understanding is that
6	they didn't do this. Didn't they just refer to the
7	ones that are in NUREG-2115?
8	DR. SEBER: They selected out of that.
9	See, if this is a complete model, they selected a
10	subset that impacts their sites.
11	MEMBER BLEY: So it's kind of like
12	having an overlay tree that puts zero and ones for
13	all of the sources kind of for your site, right?
14	DR. SEBER: Well, zero if
15	MEMBER BLEY: If it doesn't apply, yes.
16	CHAIRMAN STETKAR: Each site develops
17	this list for their site. So for example, Fermi
18	developed this list. When we looked at Calvert
19	Cliffs, they had, you know, somewhat different. We
20	looked at South Texas, and they had somewhat
21	different.
22	DR. AKE: And the characteristics for
23	all of those sites that Dogan was going over, you
24	know, the recurrence rates for each of these
25	different sources, the M-max distributions, and all

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1	of those sorts of things are all, those inputs are
2	part of this report.
3	CHAIRMAN STETKAR: That's the catalog.
4	DR. AKE: Right. So all of those
5	various input parameters, in terms of the
6	characteristics of these sources, are part of that
7	report.
8	CHAIRMAN STETKAR: South Texas set its
9	sights on, you know, South Texas' list would
10	obviously be somewhat different than Calvert
11	Cliff's list. But once you've selected your list,
12	then the catalog tells you
13	MEMBER RICCARDELLA: But there were
14	examples that were run all the way through within
15	2115. As I recall, South Texas just came in and
16	said, well, we used them because we used the one
17	CHAIRMAN STETKAR: South Texas, South
18	Texas is an aberration because they said we're just
19	going to use the Houston example but Fermi not so
20	much. We looked at Calvert Cliff's also. They did
21	their own.
22	MEMBER BALLINGER: And the licensee has
23	to justify the choice, of course.
24	DR. SEBER: Absolutely.
25	MEMBER BALLINGER: There's no rubric
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1	that's universal.
2	DR. SEBER: No.
3	MEMBER BALLINGER: Okay.
4	DR. SEBER: Either we do internal
5	calculations or they provide adequate justification
6	and we say, oh, this sounds good, we don't need to
7	do it.
8	DR. MUNSON: And, of course, we have
9	the software to run all this ourselves, so we do
10	confirmatory
11	MEMBER RICCARDELLA: Yes, and you did.
12	I mean, I have the, I think it's the October 2012
13	letter where you did the preliminary GMREs for all
14	the operating plants. And then on some of those
15	plots, you had the applicants' or the licensees'
16	values compared to your values, right?
17	DR. SEBER: Correct.
18	MEMBER RICCARDELLA: And so, I mean
19	
20	DR. SEBER: So we're going to pick one
21	source and go through a little bit of an example to
22	explain some of the complexity that we are facing
23	and everybody else. So we picked among the seismic
24	time sources, you know, mid C, mid-continent source
25	here, and that is actually where Fermi resides.

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149 That's what we call the host source. 1 This is the 2 seismic source that includes that site of interest. 3 And that source that we're talking 4 about has assigned maximum magnitude uncertainties 5 I'll call it at this level, ranges of maximum magnitudes, starting from 5.6 to 8.0. 6 These are 7 all alternative views. One view says the maximum 8 earthquake you're going to have is 5.6, nothing 9 more than that. The alternative view is, no, 10 maximum is 8.0 and then anywhere in between. And, 11 again, the numbers below are the weights that we're 12 going to keep track of every time we do these. 13 MEMBER RICCARDELLA: And these are 14 different experts within the SSHAC team? 15 DR. SEBER: And in the larger 16 community. 17 CHAIRMAN STETKAR: This is part of the 18 catalog, though. Fermi did not assign these 19 weights. That's part of the catalog. 20 Right. DR. AKE: And there was а 21 quasi-objective process used to develop these 22 distributions in that, you know, the largest, you 23 know, prior to the Prague earthquake last year, 24 Praque, Oklahoma earthquake, the largest event in 25 this region was a 5.5 something moment magnitude.

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That earthquake was about a 5.7, like 0.68 or something like that.

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3 But prior to that, that was the largest event that had occurred historically that we had 4 instrumental measurement for or even historical 5 evidence for in this region. 6 But these events, 7 these big events occur very infrequently, so we 8 used a process, or a process that was used, Ι 9 shouldn't say we, that looked globally at stable 10 continental regions like this and said what are the 11 biggest earthquakes we see globally in these types 12 and used of settings, that was to develop а 13 distribution on maximum magnitudes for these types 14 of situations. And that was used in sort of a 15 Bayesian updating way in conjunction with the 16 observations in this area to develop a distribution 17 model.

18 MEMBER RICCARDELLA: And these are 19 assumed, these maximum magnitudes are assumed where 20 in that think zone?

21 They occur anywhere. DR. AKE: Each 22 individual point is capable of producing а 23 magnitude at least as large as each one of these 24 branches, depending on which branch you're on.

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MEMBER BLEY: And when you're not

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1	instrumented, when you're relying on historical
2	record, you don't even know for sure where that
3	was, but
4	DR. AKE: No, you have much less
5	certainty in where it was, and that's sort of what
6	Dogan was getting to a few moments ago, as well as
7	how big it was. I mean, obviously, when you're
8	relying on interpretation of intensities from
9	newspaper accounts, it's a much more uncertain
10	thing.
11	MEMBER BALLINGER: So you're saying
12	that, in this case, the maximum ever seen is now
13	5.68?
14	DR. AKE: Something like, yes.
15	MEMBER BALLINGER: And, yet, with an
16	equal probability, there's an 8.0 here and you got
17	to that or they got to that from the fact that they
18	looked at some global catalog of similar tectonic
19	regions, regions with similar tectonic
20	characteristics, and that was the biggest event
21	MEMBER BLEY: WE haven't been
22	collecting data for a long time.
23	MEMBER BALLINGER: Yes, but that's
24	three orders of magnitude.
25	MEMBER BLEY: We haven't been

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1	collecting data for many years. We have the
2	reports from centuries but
3	DR. MUNSON: The rates on a 5.0, 5.0s
4	are occurring much more frequently than 8.0s, so
5	the hazard isn't dominated by having a magnitude
6	8.0.
7	MEMBER RICCARDELLA: But we still
8	haven't talked about frequencies, right? We're
9	only talking about magnitudes and weight functions
10	on the magnitudes.
11	DR. SEBER: This is just the magnitude.
12	The way you refer to frequency, I'm going to refer
13	to earthquake recurrence rates. That makes it a
14	big deal because you can't say a magnitude 8.0
15	happening, oh, could potentially happen but rate is
16	near zero. The contribution will be zero because
17	it's
18	MEMBER BALLINGER: I've factored that
19	in. Thank you.
20	DR. SEBER: So now let's go back to
21	this rate concept that we briefly talked earlier.
22	And, again, I'm going to refer to the 2115 model.
23	It provides 24 alternative rates. This is to cover
24	all types of uncertainties in the range. And the
25	reason they do that let's go one more slide

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they have three subclasses of rates. They call it case A, case B, and case E. We can't go into it, but there's a good rationale. Initially, if you read the report, they did look into it. Then they decided that it does not meaningful any A, B, and E will be sufficient to contribution. get the uncertainty, and they use it.

8 So what do they mean, A, B, and E? 9 This, again, goes back to our earlier discussion on 10 how you calculate earthquake rates. To calculate 11 earthquake rates, I have a slide. If you don't 12 mind, I can just -- well, let's do this one. This is a simpler math and easier text for them. 13 And 14 we're making assumption what an that this 15 Gutenberg-Richter recurrence level is valid. What 16 it says, it says there is a relationship between 17 number of earthquakes you observe in a region and 18 the magnitude of that earthquake. So to read this 19 chart, the bottom one, what is plotted in this? Ιt 20 says any region -- it doesn't matter, this is a 21 magnitudes, earthquakes cartoon here with \_\_\_ 22 magnitudes 5.0 and more in this region is expected 23 to occur at this annual rate. The smaller the 24 magnitude, higher the rates. Larger the magnitude, 25 It's just a linear relationship lower the rates.

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1	that, you know, scientists observed a long time ago
2	and now heavily used as a basis to get the rates.
3	It's been modified, and we can talk about it.
4	MEMBER RICCARDELLA: But, I mean,
5	wouldn't this curve be different in California than
6	
7	DR. SEBER: It is. That is why we have
8	different rates and rate maps even within the
9	source because what controls the rate is the
10	distribution of earthquakes. But this is the, you
11	know, very basic fundamental theory.
12	MEMBER RICCARDELLA: But you use this
13	to extend this plot to different magnitudes. I
14	mean, this is rates for 5.0 and above.
15	DR. SEBER: In a sense, you know, what
16	I was trying to say here, 5.0 and above in this
17	zone, assuming a linear relationship, annual rate
18	in this range, basically it says every hundred
19	years we should expect one 5.0. That is the rate.
20	Five or higher.
21	DR. AKE: Within the spatial area
22	covered by those data points that were calculated.
23	CHAIRMAN STETKAR: For comparison, if
24	you looked at a different range, geographic range,
25	if you assumed a linear fit, the slope would be

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1	different. True or not?
2	DR. AKE: The slope doesn't change
3	much. It's the scaling up and down that mostly
4	changes.
5	DR. SEBER: Global slope is one, B
6	equals one. But you do see some certain regions
7	where it varies, but it is a lot more stable than
8	A. A goes all over the place.
9	MEMBER RICCARDELLA: But do I
10	understand that you would use this curve and you'd
11	pick a spot here, and this gives you a certain
12	frequency. And then for that point, you go to this
13	curve to say what the probability of higher and
14	lower
15	DR. SEBER: Yes, this gives you, the
16	curve here gives you the rate. Go back to, like,
17	our 5.0 example, 5.0 and above let's say average
18	every hundred years. Then you have the area that
19	that, you know, that is a gridded map and each grid
20	will have its own area and its own time frame,
21	which is your record catalog time frame. So then
22	if you look at the caption, I think it says rate
23	per year per degree squared or something. And it's
24	in the legend.
25	MEMBER SKILLMAN: What is the

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difference between a case A realization and a case 1 2 E realization? DR. SEBER: And there's even one case B, so three different alternatives. 3 4 This is obviously a very cartoonish general-5 looking map, and let's assume you have the red dots it, 6 from your catalog, you obtain and you're 7 looking for a linear relationship. And it's never 8 like this, straight like that. And then you say 9 what represents the true slope of this? Everything 10 I have between magnitude 2.0 to magnitude 8.0 or 11 maybe 2.0s and 3.0s or biased, one reason, the 12 other data bias. Maybe I should not use 2.0 as a Maybe I should start using 3.0. Case A 13 beginning. 14 says take everything in the catalog. That's one 15 Case B says do not take the lowest alternative. 16 start somewhere, I think 3.5 or so, ones, and 17 that's another alternative. And case E says, no, 18 no, even drop that, don't take the smallest ones, 19 2.0 pluses, 3.0s. I have those numbers. That 20 represents basically how much of this low magnitude 21 do you use to control the rest of the slope? 22 MEMBER SKILLMAN: Okay. Let me ask one 23 presentation more. In the same of the 24 realizations, the color bar is the size or the 25 magnitude per square degree.

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1	DR. AKE: It's the rate per square
2	degree.
3	MEMBER SKILLMAN: Rate per square
4	degree. So I happen to come from a place where I
5	had to memorize a dot on earth at 404843.7 north
6	and 734578.47 west. That's the flag pole at my
7	alma mater. Is that square, is the degree of that
8	location the degree that I see on Long Island on
9	this map? Is that the square degree? Is that what
10	that means? That's the north latitude, west
11	longitude.
12	DR. AKE: So there's a square degree
13	that's maybe not centered on that, but it includes
14	that. It includes that. So the predicted rate of
15	magnitude 5.0 or greater events per year in that
16	square kilometer or in that corner degree cell is
17	given by the rate with that color there.
18	MEMBER SKILLMAN: Got it. So once we
19	look at one of these applications and we know what
20	their coordinates are, we can come to these curves,
21	excuse me, to these representations and say for
22	this magnitude earthquake this is the approximate
23	rate at that plot of land. Okay, I got it. Thank
24	you.
25	DR. SEBER: Now we talk briefly about

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1	case A, case B, and case E. Those are three
2	alternatives. Each case has what the report calls
3	it eight different realizations. This is to
4	represent the statistical variations and how one
5	would calculate the rates within that area, given
6	the case A or B or E.
7	So then case A has eight different
8	alternative representations. And so that's case B
9	and case E. Then this becomes a total of 24
10	alternative rates that one must use to calculate
11	seismic hazard at any given
12	MEMBER BLEY: And after, just looking
13	at the pictures, after that process, we've gotten
14	rid of, there are no more areas left that have zero
15	frequency in them. On one of these curves, every
16	spot has got some color in it.
17	DR. SEBER: Actually, even the worst-
18	case scenario, you have something, what we call the
19	base or the flow value, and that comes from the
20	methodology that SSHAC team used to calculate
21	these. There are no regions in the Central Eastern
22	United States with zero rates. There's always
23	very, very low rates, yes, and perhaps even not
24	contributing in most places, yes. But no zero.
25	CHAIRMAN STETKAR: I hate to cut this

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1	off but I'm going to because we're rapidly losing
2	bodies who had commitments at noon.
3	MEMBER RICCARDELLA: I have a noon
4	meeting.
5	CHAIRMAN STETKAR: Yes, I think we all
6	do. So what I'm going to do is recess now. We'll
7	come back at 1:00 and pick up from wherever the
8	heck we are.
9	(Whereupon, the above-referred to
10	matter went off the record at 12:01 p.m. and went
11	back on the record at 1:03 p.m.)
12	
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1	A-F-T-E-R-N-O-O-N P-R-O-C-E-E-D-I-N-G-S
2	(1:03 p.m.)
3	CHAIRMAN STETKAR: We're back in
4	session. You can either go forward or you can go
5	backward or you can
6	(Laughter.)
7	CHAIRMAN STETKAR: do whatever needs
8	to be done to
9	MEMBER BLEY: Check all.
10	CHAIRMAN STETKAR: Now seriously, are
11	we around the table here okay with understanding
12	the Case A, Case B, Case, whatever they are, E
13	realizations and what they mean?
14	MEMBER BLEY: I think I've reached the
15	point I can probably understand the documents I've
16	been trying to understand for some time now. So in
17	that sense, yes.
18	CHAIRMAN STETKAR: Okay.
19	DR. SEBER: Okay. Let's start slide 81
20	then. And this is going to be an example and we're
21	going to go into details slowly, slowly. Of a
22	single source we're going to make an assumption
23	that we're going to be using nine ground motion
24	prediction equations; that is, the C1, C2 and C3
25	that we discussed, plus all the three alternatives.

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1	And we just talked about 24 earthquake recurrence
2	models for a given source and we have 5 M-max
3	values, again for a single source. So we're going
4	to run this series of calculations, 9 times 25
5	times 5, which is a little over 1,000 separate
6	hazard calculations that we're going to calculate
7	for a single source contributing to a site and for
8	a given frequency.
9	CHAIRMAN STETKAR: Now, don't each of
10	those 1,080 curves each of those have a weight
11	associated with them?
12	DR. SEBER: Absolutely. And all the
13	weights that we saw in the logic trees, that we've
14	been asked some of them what are those numbers
15	below these curves and things
16	CHAIRMAN STETKAR: Right.
17	DR. SEBER: secretly we're keeping
18	track of those. And then now when you get these
19	180, each one will have its own assigned total
20	weight.
21	CHAIRMAN STETKAR: Right.
22	DR. SEBER: And then we're going to use
23	them. And hopefully, for a single source they're
24	going to end up as 1.0.
25	(Laughter.)

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1	CHAIRMAN STETKAR: And, but the
2	important thing is when the analyses are done, each
3	one literally just has a weight associated with it
4	such that when I add all of those weights together,
5	they sum to 1.00000.
6	DR. SEBER: Right. And the
7	CHAIRMAN STETKAR: And they're
8	interpreted anything else as other than weights in
9	the math?
10	DR. SEBER: Right.
11	CHAIRMAN STETKAR: Okay.
12	MEMBER RAY: Okay. I got drawn out to
13	a couple of meetings, so I missed this, and I don't
14	want you to go back, but I just looking at this
15	I'm going to ask you a simple I hope a simple
16	question. You chose the words "all plausible." We
17	struggle with words sometimes when they get
18	imbedded in the process. Is that a thoughtfully
19	considered choice of words or was it just what got
20	written on the slide?
21	DR. SEBER: It is referring to given
22	the inputs that go into the model. You have 1,080
23	options.
24	MEMBER RAY: Right.
25	DR. SEBER: That's what it's referring
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1	to.
2	DR. MUNSON: But it doesn't mean that
3	that's every plausible possibility. In other
4	words, we're not considering magnitude nine
5	earthquakes.
6	MEMBER RAY: Well again, I'm just
7	asking is that the terminology we should think
8	about now as compared with what we've used
9	traditionally in the qualitative world of seismic
10	hazards? I guess the word it would be
11	equivalent of using the word "reasonable," or
12	something like that. I'm just
13	asking
14	DR. SEBER: Yes.
15	MEMBER RAY: is it a deliberately
16	chosen adjective?
17	DR. SEBER: I'll say yes and no. In a
18	sense if you look at from the higher level, sure,
19	there may be certain alternatives available, but a
20	bunch of people getting together under SSHAC
21	process, they're creating a model that they believe
22	represents the community knowledge and the center
23	and the ranges, and they lock it in in that sense.
24	And within that model this is the all plausible
25	options.

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1	MEMBER RAY: Well, I've been through a
2	lot of SSHACs before, again in the West, not CEUS,
3	but that's neither here nor there. I just wanted
4	to ask that question because sometimes words begin
5	to
6	DR. SEBER: Sure. No, I understand.
7	MEMBER RAY: assume meanings that
8	they were not intended to have, or people don't
9	understand what they mean and so on.
10	DR. SEBER: It is not meant to say 10
11	years, 20 years from now that will not change.
12	This is based on
13	(Simultaneous speaking.)
14	MEMBER RAY: Yes, or more importantly,
15	it's not meant to say that something else is
16	implausible. It's just what, as you described, a
17	SSHAC process deems to be everything that needs to
18	be considered.
19	DR. AKE: Yes, it's a very good point
20	you raise. I think the way we might characterize
21	that might be to say that these curves represent
22	the range of plausible alternative hazards. Again,
23	sort of what Dogan is saying, it's difficult to
24	assign the word "all" to that, but it would try
25	or the attempt is to capture the range of plausible

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1	alternative hazard curves.
2	MEMBER RAY: I just give you that
3	feedback as a caution to think about.
4	DR. AKE: Yes, it's a good point.
5	Appreciate it.
6	MEMBER RAY: What word choice you make
7	here.
8	DR. AKE: Yes, it's a very good point.
9	MEMBER RICCARDELLA: So when you say a
10	single source, do you mean a single one of these
11	degree points?
12	DR. SEBER: Like in that I think if
13	I read correctly from here, that is the source.
14	DR. AKE: Yes, if you go back to
15	slide
16	MEMBER RICCARDELLA: The whole thing?
17	The whole MIDC?
18	DR. SEBER: Correct.
19	DR. AKE: Yes.
20	DR. SEBER: The whole thing.
21	MEMBER RICCARDELLA: One source.
22	DR. SEBER: The polygon in a sense that
23	represents that area, that is one source.
24	DR. AKE: Yes, the setup to this was on
25	slide 75 where this is the example that I think

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1	Dogan
2	DR. SEBER: I can put that up.
3	DR. AKE: is describing that we were
4	(Simultaneous speaking.)
5	DR. MUNSON: The shaded.
6	DR. AKE: It's the pink shaded there is
7	MIDC-A. That's the example source we're going to
8	run for this site.
9	CHAIRMAN STETKAR: But that source is
10	characterized if I understand it, it's
11	characterized as something, but the
12	characterization is derived from the process that
13	you
14	DR. AKE: Right.
15	CHAIRMAN STETKAR: described
16	earlier, right?
17	DR. AKE: Right. The characteristics
18	of this in terms of recurrence are as shown in the
19	figure you were
20	(Simultaneous speaking.)
21	CHAIRMAN STETKAR: Right.
22	MEMBER RICCARDELLA: But if you're
23	applying a ground motion equation, I mean, you have
24	one earthquake up here, one little zone up here
25	with one frequency and magnitude that could affect

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1	the site. You could have another one here, another
2	one here. How do you not consider every single
3	little
4	DR. AKE: You do consider every single
5	it takes hours.
6	MEMBER RICCARDELLA: But then that's
7	more than 1,080.
8	DR. AKE: No, you're going to sum up
9	those contributions from each of the cells.
10	MEMBER RICCARDELLA: So it's 1,080 for
11	each cell?
12	DR. SEBER: Ten-eighty
13	MEMBER RICCARDELLA: One thousand
14	eighty cases?
15	DR. SEBER: No, 1,080 refers to how
16	many possible alternative hazard curves you have
17	from that source. Each one has maybe millions of
18	calculations considering all these sources, cells
19	in this case within the source, which has the
20	varying rates and distance. Using that distance
21	rate you get contribution part A plus contribution
22	part B. So you scan all the cells
23	MEMBER RICCARDELLA: You integrate over
24	every cell in that
25	DR. SEBER: Right, integrate over

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168 DR. AKE: You integrate over 1 every 2 magnitude plot you ranged in that cell and then 3 integrate over all of the cells. You do both that. 4 You integrate --DR. SEBER: Meaning all distances. 5 DR. AKE: Yes. 6 7 MEMBER RICCARDELLA: Okay. Ι 8 understand. 9 DR. AKE: So that you capture the full 10 range of magnitudes in each cell and the 11 contribution from every cell. 12 MEMBER RICCARDELLA: Okay. MEMBER BLEY: That old equation with 13 all the sums in it. 14 15 DR. SEBER: Yes, that is the generic 16 equation that we show. 17 MEMBER RICCARDELLA: But it's 1,080 18 calculations -- well --Not based on cells. 19 DR. SEBER: Based 20 on the properties of the source. 21 MEMBER RICCARDELLA: Okay. 22 Right. So for each of those DR. AKE: 23 1,080 curves it represents a couple loops. Ιt 24 represents --25 MEMBER RICCARDELLA: All the cells. **NEAL R. GROSS** 

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1	DR. AKE: Yes, it represents the
2	integration over magnitude
3	MEMBER RICCARDELLA: I got it now.
4	DR. AKE: and the integration over
5	distance within each cell.
6	MEMBER RICCARDELLA: Yes. Yes, yes,
7	yes. Okay. Understand.
8	DR. AKE: Here's what it looks like.
9	DR. SEBER: So now, when you do it for
10	a single source
11	CHAIRMAN STETKAR: Just before you get
12	to this spaghetti thing, those are characterized as
13	median hazard curves, right?
14	DR. AKE: They're just hazard curves.
15	CHAIRMAN STETKAR: Okay.
16	DR. AKE: They're neither medians or
17	they're just the hazard curves.
18	CHAIRMAN STETKAR: Okay. We'll use
19	that. But there's still aleatory uncertainty
20	around
21	DR. AKE: Well, they are calculated by
22	the aleatory term is what gives you that curve.
23	DR. MUNSON: The aleatory curve and the
24	GMPE is what
25	CHAIRMAN STETKAR: Oh, I'm sorry. You
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1	have nine GMPEs. That had nothing to do with
2	aleatory uncertainty.
3	DR. AKE: Not so far.
4	CHAIRMAN STETKAR: It's three curves
5	for each of the three
6	DR. MUNSON: And each of those has an
7	aleatory
8	(Simultaneous speaking.)
9	CHAIRMAN STETKAR: And those are simply
10	a representative of epistemic?
11	DR. MUNSON: Right.
12	CHAIRMAN STETKAR: That's how you got
13	nine?
14	DR. SEBER: Correct.
15	CHAIRMAN STETKAR: Each of those has
16	aleatory uncertainty around
17	(Simultaneous speaking.)
18	DR. MUNSON: Right, the same aleatory
19	uncertainty.
20	DR. AKE: Well, yes, that's associated
21	with it, yes.
22	CHAIRMAN STETKAR: Okay.
23	DR. AKE: Because you have
24	CHAIRMAN STETKAR: But these 1,080
25	things that; we can call them hazard curves, are

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1	derived from those 9, which are interpreted as
2	median curves with weights.
3	DR. SEBER: Let me put the PSHA
4	(Simultaneous speaking.)
5	DR. AKE: You have to go back to the
6	PSHA thing, I think.
7	DR. SEBER: This is probably a good
8	one. So, this is source. This is distances, M-
9	maxes or magnitudes larger than certain maxes.
10	This is the one now we're talking about. What it
11	says, given a distance and magnitude, what is the
12	likelihood of exceeding certain ground motion at
13	that point? In this case this is where we
14	introduce the aleatory into the system. And we
15	didn't talk about the aleatory calculations because
16	that is the one we're using, or that have been
17	used, 2006 report that
18	(Simultaneous speaking.)
19	CHAIRMAN STETKAR: reports.
20	(Simultaneous speaking.)
21	DR. SEBER: 2004.
22	CHAIRMAN STETKAR: I care about
23	understanding how the math is done.
24	DR. SEBER: Yes, let's go through it.
25	You have the median ground motion prediction

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1	equations that we went through. And then there is
2	a generic term that defines what aleatory sigmas
3	are for those. And that's what you use and you
4	assume log-normal distribution on that median for
5	the curve.
6	CHAIRMAN STETKAR: I'm sorry. You said
7	normal distribution? How do we know that
8	DR. SEBER: Log-normal.
9	CHAIRMAN STETKAR: Oh.
10	DR. SEBER: Log-normal.
11	CHAIRMAN STETKAR: Okay. That's
12	normal.
13	DR. SEBER: I said log-normal, yes.
14	CHAIRMAN STETKAR: So the mean is not
15	the median anymore?
16	DR. SEBER: No, absolutely not.
17	CHAIRMAN STETKAR: Okay.
18	DR. SEBER: Yes.
19	CHAIRMAN STETKAR: So in the ground
20	motion prediction equations do you use the entire
21	distribution amount? Do you use the mean? Do you
22	use all I know right now is this picture that's
23	on slide 27.
24	DR. SEBER: Yes.
25	CHAIRMAN STETKAR: And I'm not going to
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1	go further than cluster 1. I know for cluster 1 I
2	have three curves each of which has a weight, and
3	those are characterized as median values. And I
4	know there's aleatory uncertainty around each of
5	these. And I know the mean value is not the value
6	on this curve because you've already told
7	DR. SEBER: Right, it's the median.
8	Yes. Right.
9	CHAIRMAN STETKAR: I know the mean
10	value is not the value on this curve because you've
11	told me that it's a log-normally distributed
12	uncertainty. Now explain to me again how you
13	account for the aleatory uncertainties and these
14	epistemic uncertainties in your 1,080 whatever
15	you want to call
16	(Simultaneous speaking.)
17	DR. AKE: Okay. Just think about a
18	single curve. You want to walk through
19	(Simultaneous speaking.)
20	CHAIRMAN STETKAR: I'll pick a curve.
21	Sure.
22	DR. AKE: Yes, just we're going to do
23	one
24	CHAIRMAN STETKAR: Pick a curve.
25	DR. AKE: trip through this tree.

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1	CHAIRMAN STETKAR: I got one curve.
2	DR. AKE: Actually you have three.
3	CHAIRMAN STETKAR: Well, no, you told
4	me to pick one, so I'm going to pick one.
5	DR. AKE: Okay.
6	CHAIRMAN STETKAR: I'll pick the
7	DR. AKE: Pick the middle one.
8	CHAIRMAN STETKAR: I'll pick the no,
9	I'll pick the bottom one
10	DR. AKE: Okay.
11	CHAIRMAN STETKAR: that's got a
12	0.185 weight assigned to it. How is that treated?
13	That's one of the nine on your slide 81, right?
14	DR. AKE: Right. That represents a
15	median for a distribution.
16	CHAIRMAN STETKAR: That does.
17	DR. AKE: Okay. For a ground motion
18	prediction equation
19	CHAIRMAN STETKAR: It does.
20	DR. AKE: when I'm doing my
21	calculations, this is going to rub this is a
22	single ground motion prediction equation.
23	CHAIRMAN STETKAR: Right.
24	DR. AKE: And it's characterized by a
25	median
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1	CHAIRMAN STETKAR: Yes.
2	DR. AKE: which is given that curve
3	represents one realization for a particular
4	magnitude, right?
5	CHAIRMAN STETKAR: Yes.
6	DR. AKE: So for each frequency that
7	represents a median. Okay?
8	Now there's an aleatory variability
9	term we haven't really even talked about yet.
10	CHAIRMAN STETKAR: That's what I'm
11	trying to ask the question about, I guess not very
12	effectively.
13	DR. SEBER: That's the slide. It's
14	that one.
15	DR. AKE: Right. So this is the
16	aleatory variability term, which is that slide
17	right there.
18	DR. SEBER: That is defined on the top
19	of the curves that we've been discussing today.
20	And that's what I was trying to say earlier, that
21	this is a separate report that overwrote the
22	discussions in 2004 report which had the aleatory
23	descriptions that said, no, this doesn't make
24	sense. Two years later EPRI came with a corrected
25	aleatory uncertainties applicable to 2004, which

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1	are shown here. And those are the ones that are
2	used in DPSHA integral. And these will include
3	that.
4	DR. MUNSON: These aleatory
5	uncertainties come from the West. We're actually
6	using Western data for the aleatory uncertainty.
7	Okay. So we have a sufficient range of magnitudes
8	and distances for the West that and we use the
9	aleatory models that were developed from Western
10	U.S. ground motion prediction equations. We use
11	those for the East. Okay. So the natural log
12	sigma values, which are dependent on magnitude,
13	those values 0.7 straight across and then it
14	dives down a little bit those are values from
15	the West.
16	DR. AKE: Okay. But let's talk through
17	your discussion a second.
18	CHAIRMAN STETKAR: I'm back to I'm
19	still pointing to this curve here. I'm going to
20	keep doing this. I got this curve. And it's got a
21	weight of 0.182 to it and it gives me the response
22	in terms of spectral acceleration as a function of
23	frequency. And I can see this curve where I have
24	now the aleatory uncertainty about that curve as a
25	function of frequency, so that if I pluck this

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1	curve off at 10 hertz, I get a value of a median
2	and I see that I ought to use
3	DR. AKE: 0.7.
4	CHAIRMAN STETKAR: a log sigma of
5	0.7
6	DR. AKE: Right.
7	CHAIRMAN STETKAR: in my aleatory
8	uncertainty. Okay. That's some sort of rule. How
9	is that
10	DR. MUNSON: Then I go back
11	(Simultaneous speaking.)
12	CHAIRMAN STETKAR: realized when I
13	do the calculations?
14	DR. MUNSON: So then I calculate that
15	probability. I calculate a Z score. For each
16	spectral acceleration value I start at 0.01. I'm
17	going to go all the way up to 5 g or whatever value
18	I want. I calculate the Z score using that median
19	and that sigma, and that gives me that probability
20	right there. Then I multiply it by the probability
21	mass function for the magnitude, probability mass
22	function for the distance, that rate, and I have a
23	hazard curve.
24	CHAIRMAN STETKAR: You said Z score.
25	I'm not familiar with that term anymore. What is
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1	that?
2	DR. MUNSON: That's the cumulative
3	distribution function. The area.
4	MEMBER RICCARDELLA: The number of
5	standard deviations above or below the
6	(Simultaneous speaking.)
7	DR. MUNSON: Yes, number of standard
8	deviations.
9	DR. AKE: If you want to go back for a
10	second, it might be a little more illustrative. Go
11	back to the figure you were just at where yes,
12	right no, that one right there.
13	Think of it as you see in the upper
14	right we have a distribution, right? So what we're
15	really the Z scores really just tell what you're
16	really looking at. So in this case the PGA for a
17	magnitude six, the median value is what is that,
18	3.3 g or something like that?
19	DR. SEBER: 8.3.
20	DR. AKE: So, it's 0.3 G. So and let's
21	say that looks like it's the median. So 50 percent
22	of that distribution with a probability of
23	exceeding that is 0.5, right? As I go farther out,
24	or as I go farther up that Y axis to higher and
25	higher ground motion values, I have a lower and

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1	lower probability of exceeding.
2	CHAIRMAN STETKAR: You don't have to
3	lecture me on how the probabilities work. I want
4	to see how the math is done
5	DR. AKE: Well, I'm trying to explain
6	that.
7	CHAIRMAN STETKAR: when you
8	convolute all of the distributions.
9	DR. AKE: That's the piece that Cliff
10	was just explaining
11	CHAIRMAN STETKAR: Okay.
12	DR. AKE: is that piece right there.
13	That's exactly what that is.
14	CHAIRMAN STETKAR: So I understand. I
15	have a distribution about this. Now, how is it
16	actually done when the final uncertainties are
17	represented? How is the calculation done in the
18	computer? Do I sample from all of those and save
19	the weight? Because I know I have a 0.182 on this
20	now
21	(Simultaneous speaking.)
22	DR. AKE: That's simply the weight,
23	right?
24	CHAIRMAN STETKAR: I understand that.
25	I have a 0.182 on this, what I will call an

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1	infinite family of curves, because each point here
2	I have an uncertainty distribution characterized by
3	your log-normal sigma. And the width of that
4	uncertainty distribution changes a bit depending on
5	the frequency according to that little rule that
6	you have up there. Right?
7	DR. AKE: Right. The aleatory changes
8	as right.
9	CHAIRMAN STETKAR: Yes. How is that
10	family of curves treated when you do this
11	combination process to preserve the combination? I
12	understand if I do this discretely how I account
13	for the epistemic uncertainly because I have this
14	0.182 weight on this curve. How do I account for
15	the effects of the fuzziness, if you want to call
16	it that, around this curve due to the aleatory
17	uncertainty?
18	DR. MUNSON: That function, that simple
19	function is the aleatory value, right, for a given
20	frequency.
21	DR. AKE: That's epistemic that you're
22	showing us right there, right?
23	CHAIRMAN STETKAR: Yes.
24	DR. AKE: That individual realizations
25	that are epistemic then is combined with aleatory
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1	variability to produce that distribution that we're
2	going to then calculate the probability of
3	exceeding in the first part of that equation.
4	CHAIRMAN STETKAR: What I'm trying to
5	get at is you're talking about a value it is a 12
6	significant figure probability of exceeding value
7	that is not convoluting the uncertainty
8	distributions. And if the calculation treats it as
9	a single point value, then you're not capturing the
10	uncertainty.
11	DR. AKE: I'm sorry. I'm not following
12	you. I'm sorry.
13	CHAIRMAN STETKAR: You see what I'm
14	talking
15	DR. MUNSON: So we use the aleatory
16	uncertainty to calculate an individual hazard
17	curve, right? That's the only uncertainty that
18	comes in when we calculate an individual hazard
19	curve. Now I'm going to have 1,080 hazard curves,
20	but each of those carries a weight from the
21	epistemic modeling I did on the source. So that's
22	how I'm capturing the epistemic on the source. I
23	have epistemic on which ground motion model I'm
24	using. But, and individual hazard curve by itself
25	only factors in that aleatory term. That's how I

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1	calculate what I call the Z score in that
2	cumulative probability. So I'm going to calculate
3	that for a range of spectral accelerations for all
4	magnitudes and all distances. So it's a triple
5	summation. That will give me a hazard curve.
6	MEMBER BLEY: Which only has aleatory
7	uncertainty in it.
8	DR. MUNSON: Right. And each hazard
9	curve it has a weight assigned, an epistemic
10	that came from my epistemic modeling of either the
11	ground motion model or the sources. And I'm going
12	to combine all those together and come up with a
13	mean hazard curve at the end, which then I use to
14	calculate my GMRF.
15	MEMBER RICCARDELLA: Let's talk about
16	this plot. What are all those different lines in
17	there? Maybe that will
18	(Simultaneous speaking.)
19	DR. AKE: Each one of those is a hazard
20	curve.
21	MEMBER RICCARDELLA: And each one of
22	those reflects both aleatory and epistemic?
23	DR. AKE: Yes, each one of them just
24	represents the aleatory. The range of those,
25	that's what's captured in the epistemic.

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1	MEMBER RICCARDELLA: Okay.
2	MEMBER BLEY: So you have the set of
3	curves that you just pointed to are all epistemic
4	uncertainty?
5	DR. SEBER: Correct.
6	DR. AKE: Right.
7	MEMBER BLEY: Now that's one member of
8	a family of similar curves that model the epistemic
9	uncertainty. Is that what you said? Or maybe you
10	used a different word for these things that cover
11	the epistemic uncertainty. I've always thought of
12	it as
13	I'm sorry?
14	CHAIRMAN STETKAR: Aleatory or
15	epistemic?
16	MEMBER BLEY: Those are all aleatory.
17	DR. SEBER: No, the spread is
18	epistemic. Each
19	(Simultaneous speaking.)
20	MEMBER BLEY: Your family of curves are
21	epistemic?
22	DR. SEBER: Each one
23	MEMBER BLEY: Is a different
24	DR. SEBER: incorporates aleatory
25	already
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1	(Simultaneous speaking.)
2	MEMBER BLEY: Yes, in the curve. Okay.
3	DR. SEBER: Yes, that, you're not going
4	to see it after that point.
5	MEMBER BLEY: Yes. Okay.
6	DR. SEBER: That comes in the first
7	term that Cliff was trying to explain. Using that
8	sigma, using the log-normal distribution you track
9	for what is the probability of exceedance given a
10	single point in the
11	(Simultaneous speaking.)
12	MEMBER BLEY: And then the family is
13	covering the epistemic uncertainty?
14	DR. SEBER: Yes. So at that point
15	you're splitting. Aleatory is not contributing
16	anymore to fractiles in that sense. It just
17	contributes to individual hazard curves as shown on
18	that figure.
19	DR. MUNSON: And then each hazard curve
20	in that figure has a weight.
21	MEMBER BLEY: Yes.
22	CHAIRMAN STETKAR: Yes. And those
23	weights account essentially for the epistemic
24	uncertainty.
25	DR. MUNSON: Right.
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1	DR. SEBER: Correct.
2	DR. AKE: And this simple example is
3	only for a single source within this range of
4	different sources we would include in the
5	calculations for a site like Fermi or the many
6	thousands of more curves above and beyond the
7	model. Yes.
8	MEMBER RICCARDELLA: And there's a
9	different set of curves like this for the various
10	frequencies, right?
11	DR. AKE: That's correct.
12	DR. SEBER: Correct.
13	MEMBER RICCARDELLA: But this just
14	happens to be 0.5 hertz?
15	DR. SEBER: Is 0.5, correct. In all
16	seven frequencies you have that. And again, this
17	is a very simplified case. One source. And we'll
18	go through an example of multiple sources and
19	multiple rates just to make things a little bit
20	more complicated.
21	MEMBER RICCARDELLA: I'm struggling
22	with how a single curve incorporates the aleatory.
23	Help me with that.
24	DR. MUNSON: I wish we could draw.
25	DR. AKE: Yes, if we could draw
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1	pictures, this would be easier. But that's
2	essentially
3	MEMBER RICCARDELLA: Pardon me?
4	DR. AKE: what
5	DR. MUNSON: If we had a whiteboard or
6	something.
7	DR. AKE: If you think about it, on the
8	Y axis of that plot we were showing a moment ago
9	that had the distribution about the median for the
10	ground motion
11	(Simultaneous speaking.)
12	DR. MUNSON: Let me go back to that.
13	MEMBER RICCARDELLA: Yes, I know the
14	one you mean. Thirty-five.
15	DR. AKE: The spread there, that's the
16	aleatory component, okay?
17	MEMBER RICCARDELLA: But how is that
18	factored into the individual curve?
19	DR. AKE: That's what your
20	DR. SEBER: I'm going to go to the
21	beginning.
22	DR. AKE: Yes.
23	DR. SEBER: Because beginning has a
24	bigger one.
25	DR. AKE: Okay.
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1	MEMBER RICCARDELLA: Page 35, right?
2	DR. SEBER: Well, one of the first ones
3	that we said created this is probably better.
4	DR. AKE: Right. So that's essentially
5	the first term in that integral is calculating the
6	area under that curve.
7	DR. MUNSON: Go to the hazard equation.
8	DR. AKE: And so that blue area there
9	in that slide, that represents the aleatory
10	variability.
11	CHAIRMAN STETKAR: It is the result of
12	an integral. It does not represent the actual
13	uncertainty.
14	DR. MUNSON: It's a representation of
15	the aleatory variability in observed ground motion.
16	CHAIRMAN STETKAR: It is the mean value
17	from integrating under the curve. It's not the
18	actual uncertainty, right?
19	MEMBER BLEY: The line is.
20	CHAIRMAN STETKAR: The line is.
21	MEMBER BLEY: But that curve is the
22	CHAIRMAN STETKAR: But they don't use
23	the curve. I don't believe they use the curve. I
24	believe the use the curve to calculate the line and
25	then treat the line as the line from there on out.

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1	So line has zero width, I believe. That's why I'm
2	trying to understand.
3	MEMBER RICCARDELLA: But is the line
4	the mean or the median and you assign this sigma to
5	that line?
6	DR. MUNSON: No, the sigma comes from -
7	- I mean, so you go across to the Y axis.
8	MEMBER RICCARDELLA: Yes, I understand
9	that. No, I'm just saying this line. One line
10	here is
11	DR. MUNSON: No, go back to the
12	calculation, hazard curve.
13	MEMBER RICCARDELLA: All right. The
14	equation?
15	DR. MUNSON: The equation. So I'm
16	going to take on that hazard curve with all the
17	I'm going to let's say I want to calculate the
18	value for 0.1 G, okay? I'm going to take the
19	probability so I'm going to take the natural log
20	of 0.1 G, subtract it from the median. I'm going
21	to divide by the aleatory sigma. So I'm
22	calculating a Z score. And then that will give me
23	the and then I'm going to look and that will
24	give me the probability in that tail. Okay?
25	There's a distribution. So that gives me the

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189 cumulative distribution. So that's how I use the 1 2 aleatory, to calculate that probability term right 3 there. The first probability term. 4 DR. AKE: So that probability term 5 DR. MUNSON: 6 was calculated for each acceleration as you qo 7 the X axis. I'm going to take across that 8 particular acceleration, take the natural log, 9 subtract it from the median, divide by the 10 aleatory. That will give me a Z score. And then 11 I'm going to calculate the area, the cumulative 12 And then that area under the curve is that area. 13 probability. I'm going to multiply that by the 14 mass function for the magnitude and the distance 15 and the rate. And that will give me the hazard 16 curve. 17 MEMBER RICCARDELLA: It's built into 18 that Z score, right? And that's how --19 DR. AKE: That's just telling you, yes, 20 what fraction of the tail is above where your 21 interest is. 22 MEMBER BLEY: Kind of the place we're 23 coming loose here is back to figure 8. And if just 24 type number and return, you'll jump to 8. That 25 blue is the aleatory uncertainty in that particular

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Now what John's been raising is a concern 1 curve. 2 that you're forgetting about that in the future and 3 you're just using that median curve. But you're 4 capturing that and you aren't showing the picture 5 way we had done it in the past. You're the 6 capturing that in a sigma and you're carrying the 7 sigma on through the calculation. 8 DR. MUNSON: The sigma is used to 9 calculate a hazard curve, an individual hazard 10 curve. That aleatory sigma is used to calculate a 11 hazard curve. 12 It's not the median anymore DR. SEBER: 13 at that point because you're looking for 14 probability of exceedance certain given g value. 15 That's what we're trying to say. And at that point 16 just a reference point to calculate median is And sigma sets the blue Gaussian 17 things from. 18 distribution. 19 MEMBER BLEY: And a hazard curve runs 20 from the frequency of an earthquake of this us 21 kind. There we go. 22 CHAIRMAN STETKAR: I'm not getting this 23 part of it. 24 MEMBER BLEY: Yes, but --25 CHAIRMAN STETKAR: And it shows on

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1	this.
2	MEMBER BLEY: if we drew this as a
3	hazard curve against
4	DR. SEBER: This is not you can't.
5	DR. MUNSON: This is a ground motion.
6	MEMBER BLEY: I know it's a ground
7	motion curve, but a hazard curve thinking back
8	to the original
9	(Simultaneous speaking.)
10	DR. SEBER: You're going to go to
11	probabilistic domain from this deterministic curve
12	that we are showing. You need to get the
13	probability of an earthquake happening, distance X
14	with magnitude X. Given those two what is the
15	probability of exceeding 0.001 G, 0.01 G, 0.1 g and
16	all those things.
17	CHAIRMAN STETKAR: Okay. Let me
18	MEMBER BLEY: You have frequency versus
19	
20	CHAIRMAN STETKAR: Let me try this.
21	Let me try this, because people throw the term
22	"probability" around too loosely. What is your
23	confidence in that probability, the thing that
24	you're calculating? What is your confidence in
25	that probability? Because you are not 100 percent

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1	confident in that probability. There's uncertainty
2	in that probability, if you will.
3	DR. AKE: I don't know if you want to
4	try this, but it's I mean, that calculation is
5	exact. I mean, it's assuming a log-normal
6	distribution with this mean or median; excuse
7	me, and this log-normal standard deviation. That
8	probability in and of itself is an exact
9	calculation. I mean, that's
10	(Simultaneous speaking.)
11	CHAIRMAN STETKAR: You can calculate a
12	mean value from that, absolutely.
13	DR. AKE: Well, as we change the ground
14	motion values we're marching along and calculating
15	the change in the area in that tail given this
16	median and the standard deviation. Right? That's
17	all we're doing in this particular part of the
18	calculation.
19	DR. MUNSON: So if you look at the
20	bottom curve, it's the complementary cumulative
21	distribution function. As I get out in higher
22	spectral accelerations, my probability in that term
23	is falling off to zero if you look on the Y axis.
24	So the probability that I'm going to exceed 0.0001
25	for a given magnitude and distance is 1. And then

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1	as you exceed in acceleration as you come down the
2	probability starts decreasing. But that
3	probability is calculated using the median and the
4	aleatory sigma.
5	CHAIRMAN STETKAR: What are the three -
6	- the blue, the red, the green down there? What
7	are those?
8	DR. MUNSON: Different distances.
9	DR. AKE: For the same magnitude.
10	DR. MUNSON: So my GMPE is going to
11	give me a different median value for an R of 25 and
12	a magnitude of 7.6. An R of 200 a distance of
13	200 and a magnitude of 7.6 I'm going to get a
14	different median value and a different median value
15	for a distance of 320 and a magnitude of 760.
16	CHAIRMAN STETKAR: Okay. So let's pick
17	a distance. Let's pick 200. I don't know which of
18	the colors is it.
19	DR. SEBER: Two hundred is the red.
20	CHAIRMAN STETKAR: Good. Tell me
21	exactly how that's calculated and why there is no
22	spread on that curve, why there is no uncertainty,
23	why is that the mean of the convolution? Is it
24	a number?
25	DR. MUNSON: No, that's a cumulative
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1	distribution function, backwards cumulative
2	distribution function.
3	DR. AKE: Yes, I think maybe you're
4	making this a little more difficult than it needs
5	to be.
6	CHAIRMAN STETKAR: Maybe you're not
7	making it as a difficult as it ought to be.
8	(Laughter.)
9	DR. AKE: Well
10	CHAIRMAN STETKAR: That's my concern.
11	DR. AKE: Well, the
12	CHAIRMAN STETKAR: Maybe you're trying
13	to do so many little point calculations and losing
14	the larger structure of keeping track of the
15	uncertainty.
16	DR. AKE: Well
17	CHAIRMAN STETKAR: And that's what I'm
18	trying to understand.
19	DR. AKE: Well, I think this given
20	the assumption that this is a log-normal
21	distribution and we feel that it could be
22	represented by that behavior, then this is the
23	result of that calculation. The uncertainty we're
24	trying to capture by the fact that we have
25	obviously different representations for what the

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1	median value would be at a magnitude of 7.6 and
2	it's a distance of 200 kilometers here. So, we'll
3	get a whole family of these based on those
4	different median values and the different aleatory
5	variabilities that we combine with those median
6	values. So I think that's where the uncertainty
7	you're talking about is coming in. I mean, there
8	is the fundamental assumption. I think maybe
9	that's the point you're trying to make is that we
10	are assuming that this thing does act like a log-
11	normal distribution.
12	MEMBER BLEY: That's not the source of
13	argument here.
14	CHAIRMAN STETKAR: That's not the
15	source
16	MEMBER BLEY: It has a distribution.
17	CHAIRMAN STETKAR: It has a it could
18	be anything, but I don't care.
19	MEMBER RICCARDELLA: But there's a
20	certain confidence level associated with that, that
21	that distribution applies.
22	CHAIRMAN STETKAR: I'm trying to
23	understand
24	MEMBER BLEY: Yes, but that's tertiary,
25	I think to the basic point they're arguing.

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1	MEMBER RICCARDELLA: But I think what
2	they're saying is once you pick a g level, you go
3	into this blue curve and that gives you your
4	probability. And that's what goes into the first
5	term of that equation.
6	DR. SEBER: Exactly.
7	MEMBER RICCARDELLA: It's different at
8	different distances, but that blue curve defines a
9	probability versus g level. Okay? And so once he
10	says my g level is 0.6 g, I read across here and
11	that gives me a probability on the blue curve that
12	goes into the first part of that equation, right?
13	DR. MUNSON: So it's from 0.6 up. The
14	area into that curves me a probability.
15	MEMBER RICCARDELLA: Yes. Yes,
16	cumulative. Right.
17	DR. MUNSON: And that's what I plot
18	here. MEMBER BLEY: It's probably
19	okay, but it's formulated very differently.
20	CHAIRMAN STETKAR: And you do lose over
21	all of those.
22	DR. AKE: Yes, I think that's I just
23	want to make sure we're clear on that hazard
24	curve is that full triple summation, right?
25	MEMBER RICCARDELLA: And the second two

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1	terms are the epistemic?
2	DR. MUNSON: No epistemic here, yes.
3	MEMBER RICCARDELLA: Oh.
4	DR. MUNSON: Okay. So we're just doing
5	one hazard curve.
6	MEMBER RICCARDELLA: Okay.
7	DR. MUNSON: We're integrating over all
8	distances, all magnitudes and the rates.
9	MEMBER RICCARDELLA: Okay. What slide
10	number is that?
11	DR. MUNSON: This one, 41.
12	MEMBER RICCARDELLA: Thanks.
13	DR. MUNSON: Yes.
14	CHAIRMAN STETKAR: By the way, is this
15	rather than taking up a lot of is this
16	actually ever explained anywhere, like to walk
17	people through it or
18	DR. MUNSON: Oh, yes. I mean, there's
19	textbooks that have this.
20	CHAIRMAN STETKAR: Huh?
21	DR. MUNSON: Textbooks that have this,
22	geotechnical
23	(Simultaneous speaking.)
24	MEMBER RICCARDELLA: It's probably
25	somewhere in the seven volumes of 2115, too, if you
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1	want to wade through them, right?
2	MR. CHOKSHI: Yes, I think this classic
3	equation appeared in 1968, Alan Gardner's paper.
4	MEMBER RICCARDELLA: Yes.
5	MR. CHOKSHI: And I think at that time
6	epistemic uncertainty was not really so what you
7	see is this and basically the probability of given
8	ground motion incorporates the uncertainty through
9	the distribution.
10	MEMBER BLEY: But just a few years
11	later we had the one with Allen and Kaplan and a
12	couple others. That one
13	MR. CHOKSHI: Right, and then we
14	started letting epistemic.
15	MEMBER BLEY: that it broke it into
16	the two pieces and
17	MR. CHOKSHI: Two pieces, yes.
18	MEMBER BLEY: laid it out. I'm very
19	familiar with that formulation and it's not quite
20	the way this calculation is being organized.
21	That's why I'm having trouble.
22	MR. CHOKSHI: Yes, you're right, but I
23	think the central idea is to the
24	(Simultaneous speaking.)
25	MEMBER BLEY: Well, but see

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1	MR. CHOKSHI: equation.
2	MEMBER BLEY: Yes, that's what I'm
3	trying to find, to see that it is.
4	DR. MUNSON: But we can get you some
5	material that has
6	DR. AKE: Actually we have a couple of
7	little spreadsheet things that might be helpful for
8	this that actually just do this very the very
9	much simple one, which might be helpful to
10	understand.
11	CHAIRMAN STETKAR: That might help.
12	DR. AKE: I think most of us eventually
13	went through that process trying to understand this
14	ourselves, so that's probably fair to assume that
15	others might
16	MEMBER BLEY: I don't think we're going
17	to get there here. I think we're going to have to
18	go study.
19	CHAIRMAN STETKAR: Anyway, let's go
20	forth.
21	MEMBER BLEY: Maybe since Nilesh
22	remembers he can help me find one, because I'm
23	still stuck in that formulation and I'm not seeing
24	that they're the same.
25	CHAIRMAN STETKAR: I am, too. And

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1	that's I don't know whether in the overall grand
2	scheme of things it's important, but that's a
3	different issue.
4	MEMBER BLEY: But it can be.
5	CHAIRMAN STETKAR: It could be if it
6	could be.
7	MEMBER BLEY: But I think the way
8	you just talked through it I think the things we're
9	concerned about are getting covered and I'm just
10	not
11	I can't get my fingers on exactly where in this
12	process.
13	MEMBER RICCARDELLA: Something that
14	would help me, on page 35, which is the log sigma
15	curve, what is the significance of having two
16	different curves. One's at M6 at 10 kilometers and
17	one's M7.9 at 742 kilometers.
18	DR. SEBER: This is just to show these
19	are the same earthquakes that we did look at
20	earlier in the morning
21	MEMBER RICCARDELLA: Oh, I see.
22	DR. SEBER: that show how the
23	aleatory would work given the definitions of
24	aleatory sigma in 2006 definitions. It basically
25	says it's not as much magnitude-dependent, but it

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1	is more frequency-dependent.
2	DR. MUNSON: It's not distance-
3	dependent. He just kept the two distances from the
4	scenario earlier.
5	DR. SEBER: This is the same two
6	scenario controlling aspects at six
7	(Simultaneous speaking.)
8	CHAIRMAN STETKAR: This was just an
9	example to show that indeed the aleatory
10	uncertainty does to some extent depend on the
11	frequency and it's not sensitive to magnitude,
12	right?
13	MEMBER RICCARDELLA: But since it's
14	normalized it would be for any distance and
15	magnitude, right? Would you get the same thing for
16	any distance and magnitude?
17	DR. SEBER: What did you say
18	normalized?
19	DR. AKE: No, it's magnitude-dependent,
20	slightly. The newer ones are even more so.
21	DR. SEBER: Yes, the update
22	CHAIRMAN STETKAR: They are more you
23	see
24	DR. SEBER: Yes.
25	DR. AKE: The newer, the later versions

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1	
2	CHAIRMAN STETKAR: Okay.
3	DR. AKE: are not only frequency-
4	dependent, but there is a magnitude dependence to
5	the aleatory, the sigma. This sigma is the
6	aleatory variability, obviously, and that is much
7	more magnitude-dependent when you
8	(Simultaneous speaking.)
9	CHAIRMAN STETKAR: When you say "much
10	more"
11	DR. AKE: Well, there's
12	(Simultaneous speaking.)
13	CHAIRMAN STETKAR: a lot of white
14	space or some visible white space?
15	DR. AKE: Some visible white space
16	CHAIRMAN STETKAR: Okay.
17	DR. AKE: would be a way to
18	characterize it.
19	CHAIRMAN STETKAR: Okay.
20	DR. AKE: There's obviously almost none
21	here.
22	CHAIRMAN STETKAR: Yes, there's almost
23	none, but I mean is it instead of having 0.700
24	across, is it like 0.66 and 0.7?
25	DR. AKE: Yes, I'm not going to I

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1	can't remember. All I know is there's a little bit
2	of spread in there.
3	CHAIRMAN STETKAR: Oh, okay.
4	DR. AKE: We can do that in no time.
5	CHAIRMAN STETKAR: It's just a minor
6	DR. AKE: But the other thing that sort
7	of is sort of one of the takeaway points here; I
8	think Dogan mentioned it and we touched on some of
9	the questions we brought up before, is noticing in
10	that figure that the aleatory variability at the
11	low frequency not only do you have a broad
12	spread in the epistemic amongst those different
13	median models
14	CHAIRMAN STETKAR: Yes.
15	DR. AKE: but the aleatory
16	variability term
17	CHAIRMAN STETKAR: Is larger
18	DR. AKE: is also larger at the
19	CHAIRMAN STETKAR: at the lower
20	frequencies?
21	DR. AKE: Yes.
22	CHAIRMAN STETKAR: Yes.
23	DR. AKE: Which we're going to kind of
24	do a little foreshadowing here. That aleatory
25	variability affects the slopes of the hazard

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1	curves. Okay?
2	CHAIRMAN STETKAR: I'll let you keep
3	going. Looking at all the pictures I looked at I
4	could understand why some of the effects that I was
5	seeing differed from low frequency to high
6	frequency. Let me just say that. Because I did
7	see that difference and I kind of got that part of
8	it. Hertz frequency, not recurrence rate
9	frequency.
10	MEMBER SKILLMAN: On this curve the
11	yellow lines are the median, the 95 and the 5
12	percent?
13	DR. AKE: I think on this one when I
14	made this this is simply the median 15th and 85th
15	on this particular one.
16	DR. SEBER: Fifteen? Okay. Sorry.
17	DR. AKE: I think that's correct.
18	MEMBER RICCARDELLA: Of the epistemic?
19	DR. MUNSON: This is your 9 ground
20	motion models, your 24 different recurrences and
21	your
22	DR. SEBER: This is why we have 1,080
23	curves.
24	DR. AKE: And so one thing; and this is
25	something that the Chairman brought up, in one of

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205 these curves -- now we've plotted them on the 1 \_ \_ 2 all the same here, okay? But with representing the 3 results of that triple summation we just saw in the 4 previous equation each one of these has a weight associated with it. And that's going to be real 5 6 important in а couple minutes when we do а 7 discussion about how you actually calculate 8 fractiles. 9 MEMBER RICCARDELLA: So this doesn't 10 have the weights in it yet? 11 DR. MUNSON: It's not shown. 12 DR. AKE: Not explicitly, yes. 13 DR. MUNSON: We're carrying them along, 14 but yes. 15 DR. AKE: Yes, we'll get to how we actually 16 do this calculation to produce those 17 yellow lines in a moment. 18 MEMBER RICCARDELLA: Okay. 19 DR. AKE: The weights are integral to 20 that --21 MEMBER RICCARDELLA: Okav. 22 DR. AKE: -- calculation. 23 DR. SEBER: So let's see, the 25 hertz 24 one, same style, same frames. And basically the 25 common observation that we see --

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1	MEMBER RICCARDELLA: Yes.
2	DR. SEBER: are a range. And
3	CHAIRMAN STETKAR: But at least let me
4	just make the observation. To me, big picture
5	stuff, both of these plots behave as I would expect
6	them to. If I step way I don't care about the
7	math now. If I step way back, even at low
8	frequencies the uncertainty increases very the
9	spread increases very dramatically in your previous
10	slide.
11	DR. AKE: Yes, in the half a hertz one,
12	yes.
13	CHAIRMAN STETKAR: At half a hertz.
14	DR. AKE: Yes.
15	CHAIRMAN STETKAR: Even at 25 hertz
16	it's increasing a lot because that's a logarithmic
17	plot over there. So both of these are indeed
18	behaving the way I would expect things to behave as
19	we increase spectral acceleration. They both
20	however they're calculated, they have that trait.
21	DR. AKE: Right. And I think as Dogan
22	was pointing out, in the yellow box there they
23	what you see for 25 hertz, almost everything gets
24	you a pretty decent ground motion from these nearby
25	local sources. You're getting lots and lots of

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these different combinations that are producing 1 2 substantial ground motions. Less so with the half 3 a hertz. DR. SEBER: This is 4 where we had 5 stopped last year. So now we wanted to show you an these fractile calculations 6 example of how are 7 And we're going to see how those little done. 8 weights that kept track of throughout we the 9 calculations impact the final results. 10 We are still looking at in this example 11 single source, five alternative M-maxes, 24 а 12 9 equations. rates, Given the spectral acceleration value, now we're taking one slice. 13 14 Now we're going to make use of those little weights 15 create a chart. now frequency and And of 16 exceedance, basically the value that we measure and 17 the cumulative weight. So now we start adding the 18 weights. And there are again 1,080 points, all the 19 way up here. Total weights add up to one. Fiftv This is five. This is 95 fractiles. 20 percent. 21 Those are the ones that are shown here. 22 Well, first of all, CHAIRMAN STETKAR: 23 are they the 5th and 95th, or are they 15th and 82nd and a half? 24 25 DR. SEBER: I remember it's --

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1	CHAIRMAN STETKAR: Yes, well, sometimes
2	
3	DR. SEBER: 95th.
4	CHAIRMAN STETKAR: that makes a
5	difference.
6	DR. SEBER: Sure.
7	DR. AKE: Yes, and I apologize. These
8	are figures I made and I apologize. If I had my
9	computer out I could tell you the answer to that
10	question.
11	CHAIRMAN STETKAR: The yellow things to
12	me don't look like they're the 90 percent
13	confidence interval from the
14	DR. AKE: Yes, but I believe that set
15	of curves is
16	CHAIRMAN STETKAR: passive
17	DR. AKE: 15th and 85th, not 5th and
18	95th.
19	CHAIRMAN STETKAR: You can't tell.
20	MEMBER BLEY: I mean, the inside is
21	black. You don't know how dense it is.
22	CHAIRMAN STETKAR: Yes, that's true,
23	but I get a bunch of black, at least the
24	(Simultaneous speaking.)
25	DR. SEBER: I think the next one is the

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1	yes, this may be easier to see. And these are
2	the same
3	CHAIRMAN STETKAR: But you can't tell.
4	DR. SEBER: You cannot tell because
5	it's not equally weighted, so this makes a
6	difference. I mean, if it is 1,000, you're not
7	going to have like 15 one side.
8	MEMBER BLEY: Plus the resolution isn't
9	such we can see how many lines are in there, yes.
10	DR. SEBER: Yes, you're going to
11	yes.
12	PARTICIPANT: Yes, this is one of those
13	if you were on the old calc comp plotter you'd have
14	bled through all the paper while you were making
15	these. Yes.
16	CHAIRMAN STETKAR: But the only reason
17	I say that is, doggone it, be careful. Don't be so
18	glib when you're throwing around probability
19	values. So, okay. We're going to draw a line and
20	we can see how you did the
21	(Simultaneous speaking.)
22	DR. SEBER: Yes, now I'm still thinking
23	about your I mean, we tried to explain. I don't
24	know. We can think about it and hopefully get back
25	to you guys if we need to do it.

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1	Given the parameters, there is one
2	probability value that comes out of it. That is
3	what is used. That is what is plotted.
4	CHAIRMAN STETKAR: This on the other
5	hand, the way I think about life is that that
6	slice, that red slice, the distribution, the cum
7	curve that you're showing on the next is indeed my
8	understanding of the uncertainty in the seismic
9	hazard for that red slice.
10	DR. AKE: That's the way we're trying
11	to
12	(Simultaneous speaking.)
13	CHAIRMAN STETKAR: And that's good,
14	because at least so far in everything you've shown
15	me that uncertainty seems to be behaving as I would
16	expect it to. So I haven't seen yet something that
17	would explain why the uncertainty in those examples
18	that we brought up in Fermi isn't behaving the way
19	I'd expect it to. So far everything is I have
20	these questions about how did you account for the
21	aleatory and how did you do the integral and stuff?
22	But in fact these plots are behaving as I would
23	expect them to.
24	DR. MUNSON: I think with the Fermi
25	example I'm not sure we went all the way from 10 to
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1	the minus 3 to 10 to the 10 gs basically. So
2	I'm not sure if there was a chance for the Fermi
3	curves to start to show their spread at 24 hertz.
4	CHAIRMAN STETKAR: It might. We asked
5	South Texas to extend them out, and they looked
6	like they might I will tell you they did for
7	South Texas as low frequency, low hertz. They
8	didn't seem to be changing much at high hertz.
9	MEMBER RICCARDELLA: If you'd go back
10	to the previous one, does that have the no, the
11	one with the blue curve on it. Does that have the
12	weights in it?
13	DR. AKE: No, that is the weights.
14	What you're doing
15	DR. SEBER: Horizontal axis.
16	DR. AKE: Yes. Specifically, yes.
17	Yes, you've sorted these from lowest to highest
18	with their associated weights to develop that
19	cumulative curve.
20	DR. SEBER: And as they climb from the
21	bottom at the first curve, get the value.
22	MEMBER RICCARDELLA: First curve is
23	over here.
24	DR. SEBER: From the bottom. No, no.
25	The one on the side.
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1	MEMBER RICCARDELLA: No, I understand,
2	but
3	DR. SEBER: Yes.
4	MEMBER RICCARDELLA: I mean, as I
5	read across, that's about a little above 10 to the
6	minus 7th, which is where the blue curve starts,
7	around 10 to the minus 7th.
8	DR. SEBER: That's correct. And then
9	you keep adding as you encounter a hazard curve,
10	keep track of weights. That calculates
11	MEMBER RICCARDELLA: As you go higher -
12	_
13	DR. AKE: Yes, you're summing the
14	weights, because that's what has to add to one. So
15	that's what's on the X axis there. So you sort
16	these from smallest to largest annual frequency of
17	exceedance, making sure you keep the association
18	there, and you just start summing these sums up
19	because they're a cumulative distribution CDF.
20	MEMBER RICCARDELLA: But you build in
21	these weights as you sum it?
22	DR. AKE: Yes. So the weights on these
23	are really small.
24	(Laughter.)
25	DR. SEBER: Very, very small.
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1	DR. AKE: Because you have 1,080
2	(Simultaneous speaking.)
3	CHAIRMAN STETKAR: Oh, yes. No. But I
4	mean, that's what computers do.
5	DR. SEBER: Yes.
6	CHAIRMAN STETKAR: Okay.
7	DR. SEBER: So those are two specific
8	examples given a single source.
9	CHAIRMAN STETKAR: Right.
10	DR. SEBER: Now we go from single
11	source to multiple sources, more realistic case,
12	what is done in the new reactor applications. And
13	just a reminder again, this is Fermi's example of
14	many, many sources, each one having alternative
15	representations. Each logic tree represents a
16	certain level of uncertainty. And at the end
17	I've never done the calculations, but this ends up
18	thousands of thousands, if not tens of thousands of
19	alternative curves. When you add to it, when there
20	are RLMEs contributing to hazard, they add their
21	own complexity into the system.
22	So there's really no way of us showing
23	in this room how this specifically works, but we
24	developed a very quick example. Make the problem a
25	little bit more complicated to show how to combine

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1	at least these sources that one would get the final
2	hazard curves in this simplified example. And in
3	this case we have GMPEs instead of 9 or 12.
4	CHAIRMAN STETKAR: Okay.
5	DR. SEBER: Two M-maxes instead of five
6	per source. Two seismotectonic sources with two
7	alternatives instead of multiple maybe a dozen
8	sources and multiple alternatives. Instead of 24
9	earthquake rates in this example we have two
10	earthquake rates. And we're not contributing
11	anything from the RLMEs to make life a little bit
12	simpler. And this is the one that chart that
13	shows how one would put together. The top part
14	representing the Ground Motion Prediction Equation
15	1, similar to C1, middle case. Bottom is C3,
16	middle case.
17	MEMBER RICCARDELLA: But you assigned
18	different weights
19	(Simultaneous speaking.)
20	DR. SEBER: They're assigned already
21	for us
22	MEMBER RICCARDELLA: Yes.
23	DR. SEBER: in the existing
24	documentation. So we just use them. But in this
25	example I think Cliff just assigned a number

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1	MEMBER RICCARDELLA: Yes.
2	DR. SEBER: because this is a
3	simplified example. And in this case it is 30
4	percent of C1, 70 percent of C3, whatever the total
5	contribution is.
6	MEMBER RICCARDELLA: Yes.
7	DR. SEBER: And then alternatives.
8	Remember, hazard contributions are all sources
9	summed up. In this case, first we are saying two
10	alternatives exist, messy sources and non-messy
11	sources. One in the north, say one in the south.
12	Both contribute a single point in the hazard. So
13	you add them up, the contributions. But as an
14	alternative to that you have two options. It gets
15	a little bit more complicated. MIDC with Rift
16	options. Doesn't matter. Say source A, or source
17	B. An alternative to source A and source B, source
18	A prime and B prime. In this case MIDC-D and RGC.
19	Instead of saying all those letters, we can
20	contribute something.
21	In each case each source having two
22	rates as this is how we define it, right? This
23	is rate 1, rate 2. This in a sense case A, case B.
24	And case A is defined by two M-maxes, because
25	that's going to make in the rate calculations if

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216 you remember the formula, that drops off and then 1 2 M-max is making a contribution there. Those are 3 two M-maxes. And again, assigned weights are 0.4, 4 0.5. So when open this up, you end up 16 5 alternative hazard curves representing this, the We split it into three top branches. 6 top branch. 7 So each one has 16 alterative curves. And we show 8 the combinations in the next slide. 9 And in a sense for one ground motion 10 prediction equation in this simple model one would 11 create 48 alternative hazard curves. And since 12 the same except different ground motion this is 13 prediction equation, here you would create another 14 So in this specific example you would have 96 48. hazard curves, but from multiple sources, multiple 15 16 rates. 17 I'm not sure if we should go into it, 18 but this the possible combinations that one would 19 Rate 1, maximum magnitude 1, rate 1 as an create. 20 alternative to maximum magnitude 2, and all the 21 combinations of all those fours. And then that's 22 what you get, 4 times 4 is 16. We even put the 23 weights to keep track of each hazard curve's total 24 weights, basically a summation of all the weights

25

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along one branch, and listed them.

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Clearlv we're

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1	not going to through all the number and things.
2	But at the end when you do this top portion to
3	bottom portion here, you get total of 96 possible
4	curves, summation of weights end up at 1. And you
5	would get instead of 180 that we showed you, you
6	would have 96 hazard curves and you would do the
7	same fractile analyses and estimate 5/95 or 15/85
8	percentiles, whatever.
9	MEMBER RICCARDELLA: So the 96 replaces
10	the 1,080?
11	DR. SEBER: Yes.
12	MEMBER RICCARDELLA: I got you.
13	DR. SEBER: Because we have
14	(Simultaneous speaking.)
15	MEMBER RICCARDELLA: So this would
16	reduce the number of cases?
17	DR. SEBER: Because we don't have any
18	more nine equations here. We said two. Not 24.
19	We said two rates, not 24 rates. Just to show what
20	would happen if it is multiple sources contributing
21	to the hazard. And you can imagine when you have
22	so many different sources, in some cases a dozen-
23	plus with all the 24 rates in each and 9 equations
24	in each, and then the issue becomes very complex,
25	of course. That's why we say thanks to computers.

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1	None of us do it manual like this.
2	MEMBER RICCARDELLA: And none of this
3	is Monte Carlo.
4	DR. SEBER: No, this is just brute
5	force. all the possibilities.
6	CHAIRMAN STETKAR: Now what you didn't
7	show though is from this 96 you didn't show us
8	that neat little bunch of black curves.
9	DR. SEBER: No, we did not calculate
10	(Simultaneous speaking.)
11	CHAIRMAN STETKAR: You haven't?
12	DR. MUNSON: No, we were just doing
13	this to illustrate the epistemic.
14	CHAIRMAN STETKAR: Okay.
15	DR. MUNSON: We didn't actually make
16	hazard curves.
17	CHAIRMAN STETKAR: Okay.
18	MEMBER RICCARDELLA: Does that
19	spreadsheet allow us to do something like that, the
20	spreadsheet you were referring to?
21	DR. AKE: It would allow you to do like
22	a single source, a single fault source or
23	something. It just shows for a single seismic
24	source how you calculate a hazard curve. I mean,
25	it's really kind of the basic building block we

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1	were talking about.
2	MEMBER RICCARDELLA: That's fine.
3	DR. SEBER: And we have I mean, we
4	talked about it many times, but basically what
5	we've been telling you in three hours in the past
6	that I don't know if I should read it, but it
7	just says, repeats what we just said.
8	DR. MUNSON: So, basically if you go
9	back to question 1
10	MEMBER RICCARDELLA: Yes, let's go to
11	question 1.
12	DR. MUNSON: what we're saying is
13	that they do eventually all start spreading out,
14	but the higher hertz hazard curve, the 25 hertz
15	hazard curves in this case, they start spreading
16	out as you you have to start getting to very
17	high accelerations on the X axis before you start
18	seeing them spreading out, whereas the half a hertz
19	hazard curves start spreading out earlier. Oh,
20	maybe that's question 2.
21	Question 1 what is question 1? In
22	does in other words, we are you will see that
23	behavior for 25 hertz. You just have to keep going
24	out on the X axis on the acceleration.
25	And then 2 was a similar question.

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1	MEMBER RICCARDELLA: Yes. Yes, that's
2	pretty much the same question. It's just that if
3	on this lower curve, if you went out further, you'd
4	probably get a 741
5	(Simultaneous speaking.)
6	DR. MUNSON: And question 2 is more in
7	depth, but asking the same things.
8	MEMBER RICCARDELLA: Is there something
9	about the physics that creates this difference
10	between 0.5 hertz and 25 hertz, or is it something
11	about our ability to calculate that?
12	DR. SEBER: No, it's not our ability.
13	Given the information we have, this is the natural
14	fallout of the calculations. Given the uncertainty
15	that we have for low frequencies, as we discussed,
16	estimating the low frequency ground motions is more
17	problematic than estimating higher frequencies.
18	And that is one and one of the parameters that
19	we three of this that we just came through,
20	maximum magnitudes contribute how they are defined
21	more to this spread than other parameters. So when
22	you incorporate the knowledge that you have, the
23	uncertainty that's defined in the models, apply the
24	math and this is the outcome. That's not going to
25	change.

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1	If you change input parameters
2	instead of using three, for example, ground motion
3	prediction equations like the SSHAC team
4	recommended, if you use five, add one more of
5	course it will be at the end. It will be
6	different. But the current thinking in the
7	scientific community/research community is that
8	those are representative of the current knowledge.
9	And if we'd learn more, if something changes like
10	NGIE is trying to do, you don't know how much
11	change we're going to see, but then we will have to
12	repeat these calculations just to see the impact of
13	the new models. And we'll look at the fractiles.
14	We'll see whether or not it makes a difference.
15	DR. AKE: And I think as Dogan said the
16	two real key things here for these lower spectral
17	frequency values, like half a hertz and one hertz -
18	- the range in the epistemic uncertainty and the M-
19	max values has a big influence on that spread and
20	the fractiles.
21	MEMBER RICCARDELLA: Yes.
22	DR. AKE: The other is, as we sort of
23	went over this morning, there was a greater
24	epistemic uncertainty in those different median
25	ground motion models at the lower frequencies than

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1	there were at the higher frequencies. Okay.
2	That's going back to the curves that showed the
3	smoothing. As you get and then you have to
4	compare from plate to plate, I guess. But there is
5	a higher sigma, if you will, in the lower
6	frequencies there, so the range in the epistemic
7	uncertainty in the median models was larger.
8	CHAIRMAN STETKAR: And that comes a bit
9	back though to my concern about did the smoothing
10	artificially reduce the sigmas in the high-
11	frequency response region?
12	DR. AKE: I'm not sure there's an
13	absolute answer to that. I mean, it's a good
14	question. I don't think
15	(Simultaneous speaking.)
16	CHAIRMAN STETKAR: Because a lot of
17	this stuff I see what you're saying, I see where
18	you're going. If indeed that smoothing process
19	artificially reduced the uncertainty in the high-
20	frequency ground motion equation, then you would
21	not necessarily see the same type of behavior,
22	because you'd see larger uncertainties at higher
23	frequencies. In other words, you'd start to see
24	those curves spread
25	DR. AKE: Spread earlier.

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1	CHAIRMAN STETKAR: earlier.
2	DR. AKE: And I think that's the point
3	that Dogan was trying to make earlier, in that the
4	largest effect of the smoothing for the high
5	frequencies is at very long distances
6	CHAIRMAN STETKAR: Yes.
7	DR. AKE: but very little effect in
8	relatively close. So
9	CHAIRMAN STETKAR: But what you'd see
10	is
11	when I compare the examples in 2115 and I
12	pick Chattanooga because it's good it's driven
13	by fairly close-in large sources.
14	DR. AKE: Moderate sources, yes. Yes.
15	Yes. Yes, okay.
16	CHAIRMAN STETKAR: Your moderate is my
17	
18	DR. AKE: Yes, okay.
19	CHAIRMAN STETKAR: Its uncertainty
20	behavior across the whole range of frequencies is
21	as I would expect. When you get to a site like
22	Fermi that does not have any influence to speak of
23	from something else we'll call a large source, all
24	of the high frequency stuff essentially from the
25	large sources gets attenuated.

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1	DR. AKE: Yes.
2	CHAIRMAN STETKAR: And that's I think
3	why you're seeing the uncertainties collapse.
4	DR. AKE: That's our interpretation
5	also.
6	CHAIRMAN STETKAR: You still get the
7	closer in ones, but they tend to be at any kind
8	of recurrence interval of interest
9	DR. AKE: Right.
10	CHAIRMAN STETKAR: they are moderate
11	accelerations and they feed
12	DR. AKE: So, everybody
13	(Simultaneous speaking.)
14	CHAIRMAN STETKAR: the lower hertz
15	stuff
16	DR. AKE: Yes.
17	CHAIRMAN STETKAR: but if the
18	smoothing process artificially reduced the
19	uncertainty in the high frequency response, you
20	might not see that same behavior for sites like
21	Fermi. I think, but I'm not sure.
22	DR. AKE: Well
23	CHAIRMAN STETKAR: Follow me. Because
24	
25	DR. AKE: Yes. Yes, I do follow. Yes,
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1	I
2	CHAIRMAN STETKAR: you'd see
3	DR. MUNSON: Yes, I'm thinking our
4	focus tends to be from the 10 to the minus 4 to 10
5	to the kind of minus 6 range, so I don't know if
6	we're starting to see
7	CHAIRMAN STETKAR: I know that's your
8	focus, but I'm thinking broader in terms of I
9	know what the staff is concerned about, I know what
10	the licensees are concerned about in terms of
11	regulatory compliance. I'm thinking about are we
12	appropriately characterizing the hazard today for
13	things like full-scope probabilistic seismic hazard
14	analyses because
15	DR. MUNSON: That's
16	(Simultaneous speaking.)
17	CHAIRMAN STETKAR: everybody's going
18	to following this temporal. The recurrence
19	interval; I'll use that term, of concern for
20	regulatory processes and for sites demonstrating
21	whether they meet their design basis seismic
22	that's okay. I get that. But that's the
23	regulatory compliance area. I just want to make
24	sure that there's nothing in this methodology that
25	is now accepted that artificially constrains let's

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1	say a more full scope seismic hazard analysis that
2	would go out to 10 the minus 8th 2 g earthquakes,
3	for example.
4	DR. MUNSON: Yes, right.
5	CHAIRMAN STETKAR: Especially for
6	people who are saying that the risk from their
7	plants for large early releases is like 10 to the
8	minus 12th
9	DR. MUNSON: Yes.
10	CHAIRMAN STETKAR: frequency per
11	year. Okay? Because new plants are going to be
12	using these things. New passive designs are going
13	to be using them. Small modular reactors are going
14	to be using them. Everybody's going to be using
15	these things, not just the current operating fleet.
16	DR. MUNSON: Yes, we've thought about
17	that quite a bit and we appreciate the comment.
18	And I think it's an extremely important one,
19	because that is the ultimate consumer of the
20	fractiles themselves is going to be the PRA in a
21	sense.
22	I mean, but the other thing I think
23	that we didn't maybe touch on quite enough that's a
24	reasonable question to ask you know, recognizing
25	that those smooth sigma values were used to develop

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the alternative median ground motion models in the different clusters. Again, when we go back then and look at plotting just simple attenuation curves of the six -- say for cluster 1 the six different components CGMPs that went into the development of that, they're subsumed well within the three curves that represent the alternative median models in cluster 1.

9 So in terms of the absolute, is it 10 broad enough? I think that's going to be a tough 11 question, because really what you're trying to 12 answer is in the development of these alternative 13 curves are I'm trying to represent the uncertainty 14 for models that haven't yet been developed. So 15 that's the effect that we're trying to capture 16 And my own personal opinion; this is just there. 17 my opinion, that's a tough one to answer in an 18 absolute sense given the constraints we have on this. 19

CHAIRMAN STETKAR: That's why I've been trying to understand what actually is done and how the weights, if you want to call them, are treated, not just mathematically; I think I know the math, but conceptually what they mean. And I don't know either. I don't know whether the difference makes

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1	a difference in the sense that if you I don't
2	know. Yes, I just don't know.
3	DR. MUNSON: Well, the latest ground
4	motion models, instead of having nine median
5	models, now we have 30 median models. Of course
6	that's a Western model, but people are definitely
7	looking at that, capturing it, the epistemic,
8	making sure we capture the full breadth of the
9	epistemic.
10	CHAIRMAN STETKAR: Let's push through
11	the end.
12	DR. SEBER: The whole presentation
13	basically was how these are treated, and since then
14	we've discussing that. And this is it. We have
15	some back-up slides and things, but if there's a
16	need we can go through them.
17	CHAIRMAN STETKAR: This is one of those
18	things where we were discussing earlier, whether or
19	not we're going to have to meet again in November
20	and then suddenly a miracle happens and we're all
21	done. So give me a moment to look through some of
22	the back-up slides.
23	(Laughter.)
24	DR. AKE: I mean, it occurred to me as
25	this discussion went on the one thing I should have
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1	grouped this by was by all of the five percent
2	models being
3	(Simultaneous speaking.)
4	CHAIRMAN STETKAR: Jon, if you're
5	talking, you're going to be on the record here, so
6	just be aware of that.
7	MR. CHOKSHI: May I John, this is
8	Nilesh.
9	CHAIRMAN STETKAR: Nilesh, yes you may.
10	MR. CHOKSHI: Okay. I think the
11	question about aleatory uncertainty and it would
12	take like think about fragility. If I only had
13	my do the math calculation as a log on the
14	distribution and beta-R, I could have one curve
15	which defines the shape of the fragility curve.
16	CHAIRMAN STETKAR: That's right.
17	That's interpreted as the mean fragility curve, but
18	it
19	does
20	MR. CHOKSHI: No, not necessarily.
21	CHAIRMAN STETKAR: I'm sorry. That's
22	the way
23	MR. CHOKSHI: It's one curve, one
24	it gives probability, community of distribution.
25	CHAIRMAN STETKAR: It's interpreted as

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1	it is used as the mean fragility curve, Nilesh.
2	MR. CHOKSHI: Not when we add the
3	epistemic. And then the mean is after you combine
4	both beta-R and beta-U. Then you call at least
5	in calculation that prove the mean fragility curve.
6	CHAIRMAN STETKAR: Oh, okay. You just
7	said beta-R.
8	MR. CHOKSHI: Yes.
9	CHAIRMAN STETKAR: Okay. Sorry. All
10	right.
11	MR. CHOKSHI: Okay. So there's no
12	CHAIRMAN STETKAR: Yes.
13	MR. CHOKSHI: it's not a mean. It's
14	a point estimate of probability. I think same
15	thing in the hazard space. Analog is for the
16	attribute. My alu-iridium uncertainty will be G1
17	realization given a probability of exiting that
18	given G. And then you're adding beta-U, because
19	fragility is easy to understand because it's simple
20	calculation. But it's same concern. It's a log-
21	normal distribution, one with a beta-R, and the
22	other of these, they're both beta-R and beta-U.
23	But then you start generating iridium several
24	iridium fragility curves, each with a beta-R. And
25	you've got a family of fragility curves. And

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1	normally we assign points, same grid to fire work.
2	That's how we used to do, is if we could go and get
3	it. I remember seeing some of those that you
4	generate 5 and assign 0.2 in the calculation. So I
5	see the same
6	(Simultaneous speaking.)
7	CHAIRMAN STETKAR: Before we figured
8	out we were doing it wrong. But, yes, that's
9	MR. CHOKSHI: But that's the way I
10	think the hazard I think that was the this is
11	a simple explanation, but the situation is exactly
12	same concept.
13	MEMBER RICCARDELLA: I'm struggling
14	with question 3. And as I go through revisiting
15	the questions, you said here's question 1 and you
16	gave us an answer to question 1. And then you said
17	here's question 2 and you gave us an answer to
18	question 2, but you never told us where question 3
19	is.
20	DR. MUNSON: It's at the very
21	beginning.
22	DR. SEBER: It is. Yes, we can go to
23	there. It is how did you answer these in DPSHA?
24	MEMBER RICCARDELLA: Okay.
25	DR. SEBER: How was it answered and
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1	developed in DPSHA, that question? Basically whole
2	day's discussion.
3	MEMBER RICCARDELLA: Oh, okay.
4	CHAIRMAN STETKAR: Let me come back to
5	what prompted all of this meeting. I'm holding in
6	front of me this is all publicly available, so I
7	can cite it. I'm holding in front of me figures
8	from the Fermi-3 COL application. That is in fact
9	Revision 6, which is the revision of record. And
10	I'm looking at 10 hertz spectral acceleration
11	exceedance I'm looking at hazard curves. And
12	I'm looking at the fact that they go out to like,
13	oh, about, I don't know, 5 g, which is pretty big.
14	And I'm not seeing a big difference in the
15	uncertainties. If I take the vertical slices and
16	I'm looking at 25 hertz and I'm seeing the same
17	and I'm looking at and they go out to like, oh,
18	5 g. So they're not truncated at 0.5 g. They're
19	out at 5 g.
20	So I get the fact that you're telling
21	me if I extended them out to way big gs, they might
22	start spreading out, but I'll tell you a 5 g
23	earthquake is a pretty darn big earthquake in terms
24	of our experience in the world, and it's not clear
25	to me why our uncertainty about that type of hazard

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233 is basically the same uncertainty as I have a 0.001 1 2 g earthquake. That's a little bit broader. I have 3 to be careful here, but it's not a lot --4 MEMBER RAY: You're talking PGA at all times? 5 CHAIRMAN STETKAR: No, this is spectral 6 7 acceleration. 8 MEMBER RAY: Okay. Well, that's --9 MEMBER RICCARDELLA: Which frequency 10 are you talking about? 11 MEMBER RAY: That's why I'm trying to 12 make sure what we were comparing. It's figure --13 CHAIRMAN STETKAR: if 14 you've got the FERMI COL or -- you don't, or do 15 you? What are you looking at? 16 MEMBER RICCARDELLA: I'm just looking 17 at this --18 CHAIRMAN STETKAR: Yes, well --19 MEMBER RICCARDELLA: I'm just looking 20 at what we have here, which is 25 hertz. And I 21 mean, I can --22 (Simultaneous speaking.) 23 CHAIRMAN STETKAR: Yes, but I don't 24 know. 25 MEMBER SCHULTZ: That's not what John

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234 1 has. 2 MEMBER RICCARDELLA: I know that. Ι 3 just --CHAIRMAN STETKAR: Twenty-five hertz --4 It should be Fermi. 5 DR. SEBER: Yes, but this 25. It's not 10. 6 7 CHAIRMAN STETKAR: Okay. I'll pick up 8 my 25 hertz. Here it is. That looks like -- and 9 the plot that I have -- yes, the end is -- the 10 right-hand edge of this is 10 g. So, yes, you're 11 right, it goes out. 12 Now, if I look at the -- if I take the 13 vertical slices through this thing, it gets 14 somewhat broader at 1 g compared to 0.01 g, but not 15 It's a range of about -- it's a heck of a lot. 16 about 2 to about -- what is it, 2, 3, let's say 40, 17 2 to 40. Or above it's about 3 to about 20. 18 That's difference biq in of not а terms 19 uncertainty. 20 If I look -- take the vertical slice at 21 1 q, take the vertical slice at 0.01 q, do the 22 ratios between the two blue things. 23 MEMBER RICCARDELLA: Yes, 0.01, I did 24 that earlier. Point oh-one q is about 10 X from 25 the top one to the bottom one.

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1	CHAIRMAN STETKAR: Yes, roughly.
2	MEMBER RICCARDELLA: And then if you go
3	to the far right, which is
4	CHAIRMAN STETKAR: Well, just take
5	MEMBER RICCARDELLA: with 10 g it's
6	about 60 X. Well, I'm extrapolating down
7	CHAIRMAN STETKAR: Yes, but take it at
8	one g where you can
9	MEMBER RICCARDELLA: One g is two to
10	four, so it's about two to forty. All right. So
11	it's about 20?
12	CHAIRMAN STETKAR: Yes.
13	MEMBER RICCARDELLA: Twenty X.
14	CHAIRMAN STETKAR: That's
15	MEMBER RICCARDELLA: It's about 20 X.
16	CHAIRMAN STETKAR: Or about a factor of
17	two in uncertainty.
18	MEMBER RICCARDELLA: Yes.
19	MEMBER BLEY: But then it starts
20	getting bigger.
21	CHAIRMAN STETKAR: It's starts getting
22	much bigger.
23	PARTICIPANT: Above one, yes.
24	MEMBER RICCARDELLA: There's something
25	in the physics here that I think I can't quite get
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1	my arms around it, but the higher we're plotting
2	them all as spectral accelerations, g, and the
3	lower frequencies don't give you high gs. They
4	give you high displacements because you got the
5	omega-squared term in there. I wonder if you
6	plotted some of these as displacement spectra if
7	you might see some differences.
8	DR. AKE: Well, you're hitting on the
9	key point here. You can't make these comparisons
10	at the same places on the X axis for these two
11	different spectral accelerations for that precise
12	reason.
13	MEMBER RICCARDELLA: Yes.
14	DR. AKE: And, yes, I think when I did
15	these slices here I chose these to be basically g
16	values on the X axis that were something close to
17	the median 10 to the minus 5 value or something
18	(Simultaneous speaking.)
19	MEMBER RICCARDELLA: Yes.
20	DR. AKE: So in an sense that would be
21	more like an apple to an apple comparison in terms
22	of that spread at two different X values for 25
23	hertz and half a hertz.
24	CHAIRMAN STETKAR: I'm sorry. What did
25	you say? You lost me when you said something about

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1	yellow values and median 10 to the minus 5
2	exceedance frequencies or something.
3	DR. AKE: Well
4	CHAIRMAN STETKAR: Because the Y axis
5	is
6	(Simultaneous speaking.)
7	DR. AKE: if you just think about
8	and it's precisely what you were just saying here,
9	is 1 g at half a hertz isn't equal to 1 g at 25
10	hertz
11	CHAIRMAN STETKAR: Twenty-five hertz.
12	DR. AKE: obviously, just because of
13	the spectral shapes we see. So if we want to make
14	a comparison of how this I mean, your point is
15	very true; that is, the 25 hertz is not expanding
16	as rapidly as it is at the half a hertz. But
17	that's partly because we need to compare things at
18	similar annual frequencies of exceedance in a sense
19	to make that comparison, although
20	(Simultaneous speaking.)
21	MEMBER RICCARDELLA: Or do horizontal
22	lines instead of vertical lines? Is that what you
23	said?
24	DR. AKE: Well, you're looking at the
25	distribution in the verticals.
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1	CHAIRMAN STETKAR: I'm sorry. It's
2	going to behave the same way.
3	MEMBER RICCARDELLA: Yes, but do you
4	understand what I was saying about the physics,
5	John? You calculate the acceleration. It's the
6	displacement times omega-squared in a sine curve.
7	And so, as you go to higher and higher frequencies,
8	you're seeing much more acceleration. It's the
9	low-frequency stuff that gives you high
10	displacements that causes the large forces on the
11	structure.
12	CHAIRMAN STETKAR: But isn't that
13	accounted for when you do your seismic response
14	analysis? This is not seismic response though.
15	MEMBER RICCARDELLA: Yes, I understand
16	that.
17	CHAIRMAN STETKAR: It's just a forcing
18	function.
19	MEMBER RICCARDELLA: Ultimately it is,
20	but it's
21	DR. AKE: And that's part of I think
22	a way to look at this and I think a good way to
23	make this comparison is if you look at 0.1 g on the
24	curve on the right, the 25 hertz curve, the ratio
25	is around 15 I think is what Sarah came up with

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1	here. And I believe this is the 95th here. That
2	range. A comparable comparison for that is
3	probably for 25 hertz. For a half a hertz we
4	should probably be looking at somewhere between
5	0.001 and 0.01.
6	MEMBER RICCARDELLA: Yes.
7	DR. AKE: On the X axis would be the
8	place to make that comparison.
9	MEMBER RICCARDELLA: Yes.
10	DR. AKE: Simply the physics of the
11	source, you don't get the same amplitudes at half a
12	hertz that you get at 25 hertz for these kind of
13	in terms of acceleration you're not going to get
14	the same amplitudes at these two different spectra
15	frequencies. And if you make the comparison
16	somewhere between 0.001 and 0.01 on the curve on
17	the left, it's going to agree somewhat better with
18	that ratio of 15 that you see on the right curve
19	there. But for the reasons you were just
20	elucidating a moment ago, the greater variability
21	and impact of M-max and the greater variability in
22	the component GMPEs we're still going to have a
23	bigger spread probably at half a hertz than we have
24	at 25 hertz. That's still probably going to be
25	true.

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1	CHAIRMAN STETKAR: And I've got that.
2	DR. AKE: Okay.
3	CHAIRMAN STETKAR: I still just find it
4	curious and this is a good example. I mean,
5	this is the one that you brought up, that if on the
6	right-hand side if you slide the red stuff out to
7	now 1 g, 10 times higher acceleration, the ratio is
8	still whatever peak
9	(Simultaneous speaking.)
10	DR. AKE: Twenty-something. Forty
11	something. Yes.
12	CHAIRMAN STETKAR: It's not much
13	different.
14	DR. AKE: Yes.
15	CHAIRMAN STETKAR: And I'm not
16	comparing 20 to 741, or 15 to 741 in an absolute
17	sense. I'm comparing 7 to 741 versus 6 to 15.
18	DR. AKE: Yes.
19	CHAIRMAN STETKAR: And still not
20	completely understanding why I get a factor of 100
21	increase over my 90 percent confidence interval on
22	the left-hand side; and I understand the
23	uncertainties are larger, but like a factor of two-
24	ish
25	DR. AKE: Yes.

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1	CHAIRMAN STETKAR: increase on the
2	right-hand side.
3	DR. AKE: And I think there's the
4	reasons we've discussed. And it's also you can see
5	a little bit of what's happening. We're beginning
6	to get some very unusual behavior in the half a
7	hertz model. You can see the mean now is deviating
8	very, very strongly from the median and approaching
9	the 90th percentile.
10	CHAIRMAN STETKAR: But that's okay.
11	That's the characteristics of the log-normal models
12	that you've used and that's fine. We all kind of
13	know that.
14	The only concern I have is somehow is
15	the uncertainty in the high-frequency stuff being -
16	_
17	DR. AKE: Under-represented.
18	CHAIRMAN STETKAR: under-represented
19	by what might be part of that smoothing process,
20	because I saw a larger, what I thought, effect from
21	that smoothing process in particular in the high
22	frequency models where there were wherever that
23	slide was you showed this morning, where they were
24	taking off at large distances and high frequencies.
25	DR. SEBER: Well, it's not going to

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1	happen. I'll tell you why. Because you're looking
2	at 1 g at 25 hertz, and we're talking about
3	800,000-kilometer distances. You'll never get 1 g
4	from an earthquake that distance at 25 hertz. It
5	will not contribute anything.
6	CHAIRMAN STETKAR: Okay.
7	DR. SEBER: So that's why
8	CHAIRMAN STETKAR: But bring it into a
9	100 kilometers and give me
10	DR. SEBER: Oh, sure. But then 100
11	kilometers, that Gaussian smoothing did not change
12	that much. That's what the little calculation we
13	did like 30 percent increase just to see how it
14	goes.
15	CHAIRMAN STETKAR: Well, but you
16	increased sigma by 30 a 30 percent increase in
17	sigma is not a 0.3 change in sigma, right?
18	DR. SEBER: Yes.
19	CHAIRMAN STETKAR: You made 0.3 sigma
20	to 0.3 whatever it is. One.
21	MEMBER SCHULTZ: I would think the
22	rationale for smoothing would have included that
23	argument or its inverse in terms of why the
24	smoothing was required, because the models didn't
25	capture

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1	DR. SEBER: I mean, I don't think any
2	one of us was there, but
3	MEMBER SCHULTZ: what you've just
4	described. No, I know, but
5	DR. SEBER: Based on the reading we
6	have it is done purely for simplification purposes
7	
8	CHAIRMAN STETKAR: Well, but see that -
9	-
10	DR. SEBER: in the calculation.
11	CHAIRMAN STETKAR: See, that statement
12	bothers me because I've read a lot of that in
13	(Simultaneous speaking.)
14	DR. SEBER: I agree with you. It would
15	bother me, too, but then when I think about the
16	consequences of that simplification, I do not see
17	any change at the end. We don't have unfortunately
18	full calculations to show you, but based on what we
19	have done, what we have seen, it will not be a
20	significant contributor. Hence, it eliminates that
21	problem of convenience.
22	MEMBER RICCARDELLA: Up to 100
23	kilometers the smoothing didn't do much.
24	CHAIRMAN STETKAR: Up to 100 kilometers
25	it didn't do much, but if I get out past 100

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1	kilometers and in particular the higher frequency
2	stuff, it does a lot.
3	MEMBER RICCARDELLA: But you don't get
4	much acceleration for the higher frequencies
5	CHAIRMAN STETKAR: That's why I said I
6	don't
7	MEMBER RICCARDELLA: at 100
8	kilometers.
9	CHAIRMAN STETKAR: know whether the
10	difference makes a difference.
11	MEMBER RICCARDELLA: Okay.
12	CHAIRMAN STETKAR: I don't know.
13	DR. SEBER: I think we all agree that
14	smoothing impacts the results. I think the
15	question in hand is like how much impact does it
16	have at the end? And at least I personally think
17	it's not going to change that much, but I don't
18	have calculations, as I said.
19	MEMBER RICCARDELLA: Let me try
20	something with this group here. Just eye-balling
21	it, if I draw a horizontal line on the left-hand
22	side at 10 to the minus 4th and 10 to the minus 6th
23	from the blue curve to blue curve, I get about a
24	factor of 10 at both of those. Okay? From the top
25	to the bottom, about a factor of 10. If I go over

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1	on the right curve at 10 to the minus 4th and 10 to
2	the minus 6th, I get a factor of 9 and 8. So
3	they're not if you go with the frequency that
4	the rates or the frequency of exceedance that we
5	think are important, there isn't that huge a
6	difference.
7	CHAIRMAN STETKAR: The interesting
8	thing, since you said it on the record, is you get
9	nine and eight, which means the uncertainty is
10	decreasing on the right.
11	MEMBER RICCARDELLA: A little bit.
12	CHAIRMAN STETKAR: As we well
13	MEMBER RICCARDELLA: Yes, but I'm eye-
14	balling the curve.
15	CHAIRMAN STETKAR: Yes.
16	MEMBER RICCARDELLA: It doesn't have a
17	grid on it.
18	CHAIRMAN STETKAR: My bigger concern is
19	not over the frequencies of interest from
20	regulatory compliance is are we characterizing the
21	hazard correctly such that if I do
22	MEMBER RICCARDELLA: Got it. PRA.
23	CHAIRMAN STETKAR: a risk assessment
24	
25	MEMBER RICCARDELLA: Yes.
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1	CHAIRMAN STETKAR: and I'm
2	interested in offsite releases that might be driven
3	completely by seismic events and people are
4	calculating exceedingly low frequencies of offsite
5	releases from everything else, are we appropriately
6	accounting for the seismic contribution once people
7	start to do that stuff. Because out in the tails,
8	out in those high accelerations it's all driven by
9	the uncertainty, if you do the uncertainty right.
10	DR. AKE: I just want to make sure I
11	understand. That's a very good point. I want to
12	make sure I understand that. The way I see these
13	being used down the road, the uncertainties here,
14	the results of the fractile calculations is we will
15	have some fragility calculations, we will have
16	discretized this X axis into bins
17	CHAIRMAN STETKAR: Yes.
18	DR. AKE: that represent relatively
19	constant behavior bins in terms of fragility. And
20	then we're going to assign maybe a three-point
21	distribution off of the Y axis to represent the
22	rate really
23	CHAIRMAN STETKAR: You assign an
24	uncertainty distribution.
25	DR. AKE: Yes. Well, right. Yes.

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1	CHAIRMAN STETKAR: Not a three-point.
2	DR. AKE: Yes, however many points you
3	are an infinite number, if you want. But you're
4	representing that Y axis uncertainty
5	CHAIRMAN STETKAR: Right.
6	DR. AKE: in that. Then that's the
7	rate turn that goes in
8	CHAIRMAN STETKAR: Yes. Right.
9	DR. AKE: with your calculations,
10	right?
11	CHAIRMAN STETKAR: That's exactly
12	right, yes.
13	DR. AKE: So we want to make sure we
14	haven't in some sense come up with a biased
15	estimate.
16	CHAIRMAN STETKAR: Because of the fact
17	that, as you've observed on the left-hand side
18	where it's really clear, the mean value is driven
19	by the uncertainty, that mean value on the left-
20	hand side here, the solid green curve, is very
21	rapidly approaching perhaps the 95th percentile.
22	And it can even exceed the 95th percentile. And if
23	you artificially constrained the uncertainty, the
24	mean value, your best estimate value will be
25	correspondingly constrained. And in a broader

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1	sense the overall characterization of the
2	uncertainly will not be represented correctly, but
3	the mean value might be artificially constrained
4	because of the way
5	you've
6	(Simultaneous speaking.)
7	MEMBER RICCARDELLA: Why are you saying
8	"mean" and is there significance to the fact
9	that you're saying "mean" and not "median?"
10	CHAIRMAN STETKAR: There's absolute
11	significance in the fact I'm saying "mean."
12	MEMBER RICCARDELLA: In a PRA the
13	CHAIRMAN STETKAR: Absolutely.
14	MEMBER RICCARDELLA: In a PRA it's the
15	mean
16	CHAIRMAN STETKAR: Absolutely.
17	MEMBER RICCARDELLA: and it's not
18	the median?
19	CHAIRMAN STETKAR: Yes. No. Median is
20	simply some parameter of a distribution.
21	MEMBER RICCARDELLA: So the mean
22	DR. AKE: It's a place to hang the
23	standard deviation, if you will. And I guess what
24	I want to come back to
25	CHAIRMAN STETKAR: In truth it's the

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1	whole when you convolute the results, it's the
2	result of the convolution, but the best estimate
3	value is the mean of
4	(Simultaneous speaking.)
5	MEMBER BLEY: The average value is the
6	one that affects the average risk, and that's why
7	you care about it. The weighted average is the
8	thing generally that's most important for most
9	applications. Median tells you the center, but it
10	doesn't tell you anything about how bad it can be.
11	MEMBER RICCARDELLA: But you're
12	convoluting the whole distribution.
13	MEMBER BLEY: Yes, that's what John
14	said.
15	CHAIRMAN STETKAR: That's what I said.
16	(Simultaneous speaking.)
17	MEMBER RICCARDELLA: distribution,
18	it's going to be
19	MEMBER BLEY: It's going to be a
20	distribution.
21	CHAIRMAN STETKAR: It's going to be a
22	distribution. But the value that you select when
23	you start to do evaluations of what's important to
24	risk is the mean value of the results of that
25	process. It's not the median value. And the only

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1	point here is the mean value because these are
2	not normal distributions, they're log-normal
3	distributions, the mean value is determined by the
4	uncertainty. And if you're constraining the
5	uncertainty, that means you're constraining the
6	mean. And even if you do so-called point estimate
7	value here you're multiplying mean hazard by mean
8	fragility at a given acceleration, you're going to
9	get the wrong number.
10	DR. MUNSON: I think the point I want
11	to make though is I feel like some for example,
12	with this movie we showed, I don't feel like that
13	was a reasonable constraint to get behavior that
14	you would expect at those distances from those
15	CHAIRMAN STETKAR: And that I don't
16	know. I still come back to the results that I see
17	for Fermi that are I think dominated by distributed
18	sources versus the examples that I see in NUREG-
19	2115. We can't see a COL submittal for a site
20	that's dominated by RLME sources. We just haven't.
21	I mean, at South Texas is Houston. We've seen
22	Calvert Cliffs and we've seen Fermi. Those are the
23	three that we've seen. They all have this same
24	general behavior.
25	But if I go to NUREG-2115 and I look at

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1	the examples for Savannah and I compare them with
2	the examples for Manchester, I now have one site
3	that's fairly heavily influenced by RLME; Savannah,
4	and I have another site that isn't. And I see the
5	differences in behavior.
6	In other words, Savannah across all of
7	the frequency seems to behave as I'd expect, like
8	the left-hand side of this. Whereas Manchester,
9	the sites that seem to be driven by the distributed
10	stuff behave more like what we see here, a broader
11	uncertainty at lower hertz, at lower frequencies,
12	but a fairly narrow uncertainty and not as
13	dramatically changing at the higher hertz and
14	MEMBER RAY: Can I intervene while
15	you're thinking? You used a term very early in the
16	discussion of "firm site," which I paid attention
17	to that because earlier this week we were dealing
18	with a hard rock site. Now this has not got to do
19	with what he's talking about, but what keeps coming
20	back to me is the huge difference between a hard
21	rock high frequency spectrum and in this case an
22	AP1000-certified design response spectrum. This is
23	not probabilistic and the uncertainties aren't part
24	of it, but they're so different that I guess I'm
25	pondering whether that's got something to do with

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1	the spread that we're talking about here. Because
2	it's just huge, the difference for the basically
3	the same source depending on what we're talking
4	about in that case.
5	Like I say, you used the term "firm"
6	for as it's hard rock and it just seems to me like
7	the characteristic of the response curve is so
8	different between the two for the same source term
9	that it might somehow be implicit in what's being
10	shown here. Now, that shows you how ignorant I am,
11	so I'll stop.
12	DR. MUNSON: But, no, you do get that
13	shape. For example, at 25 hertz you will see very
14	large accelerations for a rock site. And Fermi is
15	a definite rock site. So a lot of sources of the
16	distributed seismicity sources around Fermi for
17	different magnitude and distance combinations are
18	able to contribute this high spectral acceleration
19	you know, accelerations approaching 0.5 to 1 g
20	because of the shape of the
21	(Simultaneous speaking.)
22	MEMBER RAY: Yes. No, I understand the
23	physics of how it works perfectly well, but I just
24	don't know if it has any connection to what you
25	guys have been talking about for the last 20

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1	minutes here, which is basically the spread that
2	you get between 25 hertz and half a hertz.
3	MEMBER RICCARDELLA: The difference in
4	the spread?
5	MEMBER RAY: Yes.
6	MR. GRAIZER: Please, may I just
7	MEMBER RAY: Sure.
8	MR. GRAIZER: I think you are
9	CHAIRMAN STETKAR: Just
10	MR. GRAIZER: Oh, Vladimir Graizer,
11	seismologist. I think you are referring to what we
12	were hearing yesterday or couple of days ago.
13	MEMBER RAY: Yes.
14	MR. GRAIZER: And it's a hard rock
15	site. But I would be cautious about comparing
16	Western and Eastern United States because
17	MEMBER RAY: I'm not meaning to do that
18	at all. I'm sorry to interrupt you, but
19	MR. GRAIZER: In a way, yes, because
20	when you are talking about Western regular
21	spectrum, regular spectrum in a way is
22	representative of our Western behavior
23	MEMBER RAY: No, I
24	MR. GRAIZER: which is very flat at
25	high frequencies. Eastern spectrum is different.
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1	It is not flat a high frequencies.
2	MEMBER RAY: No, I'm sorry. I'm not
3	talking about that at all. Let's not
4	MR. GRAIZER: Okay. Maybe I'm
5	confusing. Sorry.
6	MEMBER RAY: All I'm doing is saying
7	for the same peak ground acceleration you have a
8	very, very different response spectrum. We all
9	know that to be true. There's nothing profound
10	about that. But the upshot of it is that because
11	of the difference, when I'm looking at these two
12	different frequencies here, I'm just asking the
13	question has that got anything to do with what's
14	troubling John in terms of this uncertainty spread?
15	That's all. I'm not talking about Western sites.
16	I know the difference there.
17	CHAIRMAN STETKAR: There's silence. Do
18	any of the members have any more questions for the
19	staff? I'm still personally a bit confused, but I
20	can do more reading. I'm not as confused as I was
21	
22	(Laughter.)
23	CHAIRMAN STETKAR: but I'm still
24	curious about the high frequency stuff. But I can
25	do some reading. And by the way, everybody sit

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1	down, it's helped me an awful lot. But first of
2	all, before we do anything else, anybody else have
3	any questions on sort of the technical presentation
4	for the staff? And if you do, we've got time.
5	MEMBER RICCARDELLA: Yes, I've got a
6	whole bunch, but I need to go back and think
7	CHAIRMAN STETKAR: Okay.
8	MEMBER RICCARDELLA: and formulate
9	them.
10	CHAIRMAN STETKAR: That's where I am.
11	MEMBER RICCARDELLA: This was very
12	educational for me. It really was. I appreciate
13	it.
14	CHAIRMAN STETKAR: Because I want to do
15	something later, but I want to get some
16	administrative things out of the way here first.
17	So where are you, Kathy?
18	MS. WEAVER: I'm right here.
19	CHAIRMAN STETKAR: I don't know if
20	there's anybody on the bridge line, but let's go
21	get that open.
22	While we're waiting for that, let's see
23	if there's anybody else in the room who'd like to
24	any comments. And we'll see if we can get the
25	bridge line open and see if there's

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1	MEMBER BLEY: I think we were too
2	esoteric for a large audience.
3	CHAIRMAN STETKAR: That's fine, but
4	it's still a Subcommittee meeting. See if we can
5	open up the bridge line and see if anyone out there
6	has any comments. And then I want to go around the
7	table, but before we do I'll ask us to consider
8	some things.
9	I heard a pop. I hate always to say
10	this, but I believe the bridge line is open. If
11	someone's listening in, please just say hello or
12	something like that to confirm it.
13	MR. LEWIS: Hello.
14	CHAIRMAN STETKAR: Thank you.
15	MR. LEWIS: I'm a member of the public.
16	CHAIRMAN STETKAR: So we know we have
17	the bridge line open. And now I'll ask if there's
18	anyone on the bridge line who would like to make a
19	comment, please identify yourself and do so.
20	MR. LEWIS: Marvin Lewis, member of the
21	public.
22	CHAIRMAN STETKAR: Okay. Marvin,
23	you're up first. I heard another voice. So, we'll
24	take this in order. Marvin?
25	MR. LEWIS: Well, first of all, enjoyed

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1	today's meeting very much, although it appears that
2	there's a big question coming up with the latest in
3	earthquake technology. And I hope there are a few
4	data points behind the curves being offered today
5	and I'll be contacting Mr oh, I forget his
6	name, the contact person, I'm sorry
7	CHAIRMAN STETKAR: Chris Brown.
8	MR. LEWIS: about getting the
9	handouts for today.
10	CHAIRMAN STETKAR: Good.
11	MR. LEWIS: Thank you.
12	CHAIRMAN STETKAR: And, Marvin, it is
13	Christopher Brown on our staff, and we'll make sure
14	we get you those.
15	MR. LEWIS: Thank you very much. My
16	email is M-A-R-V-I-N, L-E-W-I-S at J-U-N-O dot C-O-
17	M. marvinlewis@juno.com.
18	CHAIRMAN STETKAR: Great. We'll make
19	sure you get those, Marvin. Thank you.
20	Now, I heard another voice in the
21	background. If someone else would like to make a
22	comment, please identify yourself and do so.
23	MR. MARRONE: Yes, this is Jim Marrone
24	from Bechtel in San Francisco.
25	CHAIRMAN STETKAR: Good.

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1	MR. MARRONE: And I had a couple
2	questions. My when we got to the end of the
3	slides
4	CHAIRMAN STETKAR: Jim? Jim,
5	you're
6	MR. MARRONE: Yes?
7	CHAIRMAN STETKAR: breaking up a
8	little bit, so I don't know if you're on a speaker
9	phone or
10	we're catching about every fourth word you say.
11	MR. MARRONE: Okay. How is this?
12	CHAIRMAN STETKAR: That is better so
13	far.
14	MR. MARRONE: Okay. Sorry about that.
15	Maybe it's related. My phone failed as we got to
16	the end of the slides and I don't know if you've
17	got to the back-up slides, but I had a couple
18	observations that I wondered if they had been
19	considered in terms of looking at the character of
20	the distribution of hazard curves, the high
21	frequency versus low frequency. And the first
22	couple items that I see in the back-up slides is
23	first off for the low frequency the maximum
24	magnitude model certainly is showing some effect on
25	spreading out those hazard curves for the low

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1	frequencies.
2	CHAIRMAN STETKAR: Yes.
3	MR. MARRONE: And the other character
4	that I don't know if we've evaluated is looking at
5	the de-aggregation characteristics. The low
6	frequencies can be bimodal where you'll have
7	contributions from both nearby small magnitude
8	events and of course with large magnitude and
9	distant events, and you don't get that in the high
10	frequencies. So I'm wondering if there's a
11	characteristic of some of that lack of spread-out
12	in the high frequencies because the distance range
13	is much more compressed of the contributing events.
14	CHAIRMAN STETKAR: Thanks. For your
15	benefit we didn't talk about the de-aggregation in
16	our discussion. We didn't get to those back-up
17	slides. Your question is on the record. It's a
18	good question. We don't typically try to get into a
19	dialogue with members of the public, so we'll note
20	your question for the record. Anything else?
21	MR. MARRONE: No, that will do it.
22	Thank you very much.
23	CHAIRMAN STETKAR: Great. Thank you.
24	Anyone else on the bridge line who'd
25	like to make a comment?

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1	(No audible response.)
2	CHAIRMAN STETKAR: Interpreting five
3	seconds of silence as an opportunity, we'll close
4	the bridge line from that direction so we don't get
5	the pops and crackles that we get.
6	And now, as we usually do in a
7	Subcommittee meeting, I like to go around the table
8	and see if the members have any final comments.
9	Now, part of the final comments I'll ask you do we
10	at least at this time think that we need another
11	briefing? And I will at the risk of having the
12	staff start hurling things at me
13	(Laughter.)
14	CHAIRMAN STETKAR: say that we have
15	a window of opportunity that's opened up in the
16	middle of the month of November, our Subcommittee
17	week. Things have fallen apart. So there's that.
18	Other than that, we're talking about long-term in
19	the future. So keep those in mind. If you have
20	any final comments, what are they and do you feel
21	that there's a need for another briefing at least
22	based on what we know right now? And we don't have
23	to make a decision about that.
24	Dr. Ballinger?
25	MEMBER BALLINGER: I thought this was
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1	extremely informative. I had previously tried to
2	get up to speed, and I'm getting there a little bit
3	more and more. I think I understand enough about
4	what you're doing, not nearly as much as these guys
5	need to understand, but I do. So I think from my
6	perspective I don't need another meeting.
7	CHAIRMAN STETKAR: Okay. Dr.
8	Riccardella?
9	MEMBER RICCARDELLA: First, I'd like to
10	thank you guys. I mean, this was just a great,
11	great session. I really appreciated it. And as
12	chairman of the Structural Analysis Subcommittee
13	I've been trying to work my way through NUREG-2115,
14	and I might as well have been reading it in Greek.
15	(Laughter.)
16	MEMBER RICCARDELLA: But after this
17	session I'll go back again and I think I might be
18	able to get
19	CHAIRMAN STETKAR: Now it's only in
20	Latin, right?
21	(Laughter.)
22	MEMBER RICCARDELLA: Italian. I speak
23	a little Italian.
24	(Laughter.)
25	MEMBER RICCARDELLA: Anyway, I
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appreciate it very much. As for another meeting, I 1 2 would kind of like to keep it -- if we do have some space in November, maybe just for some questions 3 4 and answers. As we go back and try to absorb this 5 stuff, we might come up with some questions and 6 maybe just a short session to ask those questions. 7 Thank you. 8 CHAIRMAN STETKAR: Dr. Bley? 9 MEMBER BLEY: Thank you. I, too, want 10 to thank you guys. You put in a lot of effort to 11 try to make this clear and helpful to us, and we 12 really appreciate it. 13 And I think I'm at the point Pete 14 described. I don't think by mid-November I'd be at 15 the point that there would be any real call for a 16 meeting at that point. After I dig some more, 17 maybe at a later date. But I appreciate what you 18 brought to us today a lot. Thank you. 19 CHAIRMAN STETKAR: Dr. Schultz? 20 MEMBER SCHULTZ: Yes, I appreciate the 21 lead-through that you've done today. In terms of 22 providing the information and its description it 23 was quite an improvement over where we tried to go 24 a vear ago. So there's been a lot of information 25 that's been filled in for me and I think there's

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1	enough information to go forward to learn more from
2	what we've been provided from EPRI and enough to go
3	back to the history and try to piece it together.
4	I would not be ready for a meeting in November.
5	And I won't be here next year for this Committee,
6	so
7	CHAIRMAN STETKAR: Oh, sure, take the
8	easy way out.
9	MEMBER SCHULTZ: I'll take the easy way
10	out.
11	(Laughter.)
12	MEMBER SCHULTZ: But thank you for the
13	education.
14	CHAIRMAN STETKAR: Dick?
15	MEMBER SKILLMAN: I thought I
16	understood this quite well before this meeting, and
17	through the course of the day I've identified in my
18	own mind things that I'm not as strong at as I
19	thought I was. But I'm with Dr. Riccardella. I
20	would like a briefing, a short one in November,
21	because with what we've learned from you today, for
22	which I thank you very much, I'll be thinking about
23	this through the next couple of weeks and I'll have
24	more questions, but probably not significant ones.
25	But a short briefing would give me a chance to

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1	reconnect with you and to calibrate and fill in
2	some blanks, and for that I would thank you in
3	advance. But this has been very good and I
4	appreciate it. Thank you.
5	CHAIRMAN STETKAR: Harold?
6	MEMBER RAY: Nothing. This is tutorial
7	for me and I appreciate it very much.
8	CHAIRMAN STETKAR: Okay. Thanks. All
9	right. And as pleasant and tactful as I typically
10	am, I'd like to thank you all. You did a very good
11	job at answering our initial concerns. As I said,
12	I need to do a little more homework on, even in
13	real time here, looking at where I was pointed to
14	Section 7.7 of the 2013 EPRI document. And I'm not
15	seeing anything different in terms of the
16	smoothing, but I haven't read all the words, for
17	example.
18	With regard to something in the middle
19	of November, my sense is that we probably don't
20	know right now what questions, and given the time
21	between now and the middle of November, I doubt
22	that we could formulate enough questions in time to
23	get them to the staff so that they would have
24	enough time to come back and actually provide
25	meaningful answers.

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So I think the most productive 1 path 2 forward would be for us to spend a little more time 3 -- we now know a little bit better what we don't know and we know some of the stuff that we do know. 4 5 And if we do have questions, then think about the efficient 6 most productive and path forward. 7 Whether that's a full Subcommittee meeting or 8 whether it's one or two members interacting with 9 the staff in the sense of; we do this occasionally, 10 fact finding to develop more focused answers and 11 bring it back to the Subcommittee. That may be a 12 heck of a lot more efficient, both for us and 13 certainly for the staff. 14 So let's keep that in mind, that we'll 15 little more homework and if we do do have а 16 questions, we'll get them to you some way and try 17 to get them resolved. Ιf the people with the 18 questions feel after initial interaction that it 19 merits discussion in front of the full 20 Subcommittee, we'll do that. If not, can get an 21 answer and the members can sort of fill in the rest 22 of us with those answers. We'll take it that way. 23 So thanks again. Appreciate it. And I 24 know it's grueling. I know it's like trying to 25 train monkeys how to write novels, but with enough

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1	pounding eventually we might get there.	
2	And with that, we are adjourned.	
3	(Whereupon, the above-entitled	matter
4	went off the record at 2:57 p.m.)	
5		
6		
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12		
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United States Nuclear Regulatory Commission

Protecting People and the Environment

# Treatment of Uncertainties in Probabilistic Seismic Hazard Analyses (PSHA) Conducted for Combined License Applications and Early Site Permits

Dr. Dogan Seber (NRO) Dr. Clifford Munson (NRO) Dr. Jon Ake (RES)

October 23, 2015

# PART 2

This presentation is continuation of staff's presentation took place on November 17, 2014 on the same topic (ML14345A885)

# Where were we?

Addressing the ACRS Subcommittee questions on the following technical areas:

- Why doesn't the uncertainty increase appreciably as you go from small spectral accelerations to very high spectral accelerations?
- 2) Why is the uncertainty small for very high spectral accelerations at 25 Hz when compared to the uncertainty for very high spectral accelerations at 0.5 Hz?
- 3) How is uncertainty developed in PSHA calculations?

1. Why doesn't the uncertainty increase appreciably as you go from small spectral accelerations to very high spectral accelerations?



2. Why is the uncertainty small for very high spectral accelerations at 25 Hz when compared to the uncertainty for very high spectral accelerations at 0.5 Hz?



Annual Frequency of Exceednace

# Outline

#### Past Meeting Discussions

- Summarize essential fundamentals/concepts
- Re-visit many of the questions raised during the first meeting
- Clarify some of the technical issues raised

#### PSHA and Uncertainty Management

- Continue with uncertainty definitions in seismic hazard estimates
- Describe how they are used in PSHA
- Use of the CEUS-SSC Model and Uncertainties in COL and ESP Applications

#### Mean and Fractile Calculations in PSHA

- Fractile calculations with examples
- Impact on the GMRS

## **Summary of Past Meeting Discussions**

# **Development of Ground Motion Models**





# **Ground Motion Recordings**



- Ground motions at a point on the ground are recorded by accelographs
- Recordings are usually in three components
- Recording instruments are very sensitive with very low noise levels.
  - Uncertainty in the recorded ground motions is much lower than the amplitude of the ground motion recorded. Hence, this uncertainty is not incorporated into the PSHA calculations.



## **Earthquake Ground Motions**



**Recording station** 

Recorded earthquake ground motions vary due to magnitude, source type, earthquake depth (source effects), distance, the material in which earthquake waves travel (path effects), and recording site conditions (site effects).

Ground motion recording at the recording station

#### **Open Item**

# **Ground Motion Recordings – Three Components**

















#### **GMPE Development**

#### Observations Models

- Median and the variability (uncertainty) are modeled using a functional form.



**Open Item** 

#### How about other frequencies of interest?





# **Ground Acceleration vs Spectral Acceleration**

Recorded ground motions (seismograms) represent acceleration time history experienced by a particular point on the ground as seismic energy propagates through.

Ground motion models predict spectral accelerations, the response of a single degree of freedom (SDOF) system to a given earthquake ground motion.



## **Example of Spectral Accelerations**



# Are seven frequencies sufficient to represent the observed spectral accelerations?



This recording has strong site effects. Peaks observed near 2 and 5 Hz likely represent amplifications due to near surface soil/rock properties at the recording site.

GPMEs are developed for "generic rock conditions", which is observed usually at some depth. Hence, at the reference rock level, spectra are smoother than what is shown on left. Impacts of near surface rocks are taken into account in the site response calculations and corrections. Spectral acceleration of a different record obtained from a site with characteristics closer to generic rock conditions

**Open Item** 



EPRI (2004) ground motion models sample the spectra at seven frequencies. This is consistent with recent empirical GMMs (5-10 spectral frequencies). Future GMMs such as NGA East may use larger number of frequencies (>20 spectral frequencies).

# EPRI (2004, 2006) Ground Motion Models

- Developed using a SSHAC Level 3 process and approved to • be used in new NPP applications
- Defined at seven distinct ground motion frequencies •

– 0.5Hz, 1Hz, 2.5Hz, 5Hz, 10Hz, 25Hz, 100Hz (PGA)

- Composite model including four sub-models or clusters. Each cluster represents a different modeling approach.
  - Within each cluster three different median models capture the epistemic uncertainty
  - Cluster 4 used only for sources with significant large magnitude contributions
  - Alternative aleatory variability (sigma) models included
- Recently updated (2013) using a SSHAC Level 2 study for Fukushima NTTF 2.1 recommendations
- To be completely replaced with new NGA East Models currently being developed (2016?) 21
## Make up of EPRI (2004) GMPE Clusters

Cluster No.	Model Type	Models	Weights <sup>1</sup>
1	Spectral, Single Corner	Hwang & Huo (1997)	0.037
		Silva et al. (2002) - SC-CS	0.192
		Silva et al. (2002) - SC-CS-S	0.148
		Silva et al. (2002) - SC-VS1	0.560
		Toro et al. (1997)	0.029
		Frankel et al. (1996)	0.034
2	Spectral, Double Corner	Atkinson & Boore (1995)	0.714
		Silva et al. (2002) DC	0.154
		Silva et al. (2002) DC-S	0.132
3	Hybrid	Abrahamson & Silva (2002)	0.336
		Atkinson (2001) & Sadigh et al. (1997)	0.363
		Campbel/ (2003)	0.301
4	Finite Source/Greens Function	Somerville et al. (2001) 1	1.0

## **Developing Within-Cluster Models**

• Develop median model for each cluster

**Open Item** 

- Weight each GMPE based on fit to CEUS dataset
- For each weighted GMPE calculate predicted acceleration for a range of magnitude and distance pairs
- Calculate median acceleration for each magnitude & distance pair
- Select functional form for median model
- Develop 5<sup>th</sup> and 95<sup>th</sup> percentile within cluster models
  - Calculate standard deviation for each magnitude & distance pair
  - Calculate total standard deviation by adding uncertainty to capture wider range of source and path model parameters
  - Smooth total standard deviation and calculate 5<sup>th</sup> and 95<sup>th</sup> percentile models

# e.g., Two Scenario Earthquakes (How different are their spectral accelerations?)



### Deterministic Scenario Results (Examples from Fermi)



Spectral Acceleration (g)

#### **Epistemic Uncertainties in the EPRI Ground Motion Models**



## Within Cluster Epistemic Uncertainty in Cluster 1

Open Item



## Implementation of Epistemic Uncertainty in GMM





# Does the epistemic uncertainty smoothing described in EPRI (2004) result in the narrow range in fractiles, especially at high frequencies?

What is the potential impact of reducing the sigma at large distances due to smoothing?



## **Cluster 1**



- Within each cluster a sigma is calculated for magnitudedistance pairs, using original models defining the cluster
- Based on seismological uncertainties, additional sigma is applied
- Values of sigma are highly variable with distance causing problems in model fitting
- Prior to developing 5<sup>th</sup> and 95<sup>th</sup> percentile models sigma is smoothed

Open Item

#### EPRI (2004) GMPE Cluster 1





#### EPRI (2004) GMPE Cluster 2

#### **Epistemic Uncertainty Filtered Original Epistemic Uncertainty** PGA Smoothed Sigma PG/ .8 5.0.8 5.0 Sigma 6.0 .6 .6 7.0 8.0 .4 .4 .2 .2 C 10 Hz Smoothed Sigma 25 Hz 25 Hz 10 Hz .8 .8 Sigma .6 .6 .2 .2 0 2.5 Hz Smoothed Sigma 5 Hz 2.5 Hz 5 Hz .8 .8 Sigma .6 ,6 .4 .4 .2 .2 Ð 0 0.5 Hz Smoothed Sigma 1 Hz 0.5 Hz Hz .8 ,8 Sigma.6 .6 .4 .4 .2 .2 0 0 105 $10^{3}$ 100 10 1 10 100 1 103 10 100 103 10 100 1 Distance (km) Distance (km) Distance (km) Distance (km)

## Test the impact of sigma reduction through Gaussian filtering (Increase sigma values by 30%)





 Increasing sigma does not impact the outcome significantly

**Open Item** 

- 5<sup>th</sup> and 95<sup>th</sup> epistemic branches are weighted 18.5% each, while the median is weighted by 63% in the model
- Clusters 1 and 2 collectively account for 75% of the final ground motion model
- High frequency ground motions at large distances are very low.
  - Any potential problem related to reduction of sigma at such distances will not impact the high frequency hazard curve fractiles, as their contribution is near zero or zero.

## **Ground Motion Model Aleatory Uncertainties**





## Earthquake Recurrence Rates And Uncertainties

#### **Gutenberg-Richter Recurrence Law**

 $\text{Log } \lambda_m = a - b m$ 

where

 $\boldsymbol{\lambda}_m$  : Annual rate of earthquakes with magnitudes greater than m

a, b are constants



## Introducing upper bound on magnitude



Gutenberg Richter Law implies unbounded upper limit.

Physical bounds are introduced using the largest possible earthquake in a given seismic source M<sub>max</sub>.

$$\lambda_{m} = v \frac{\mathbf{e}^{(-\beta \ (m-m_{min})} - \mathbf{e}^{(-\beta \ (m_{max} - m_{min}))}}{\mathbf{1} - \mathbf{e}^{(-\beta \ (m_{max} - m_{min}))}}$$

where

$$v = e^{(\alpha - \beta m_{min})}$$
  
 $\alpha = 2.303 * a$   
 $\beta = 2.303 * b$ 

**Open Item** 

## **Earthquake Rates – Alternative Models**

NUREG-2115 (CEUS-SSC model)





Calculated earthquake rates are assumed to be representative of future earthquakes in the region

Case	M 2.9–3.6	M 3.6-4.3	M 4.3–5.0	M 5.0-5.7	M 5.7-6.4	M > 6.4
A (wt = 0.3)	1	1	1	1	1	1
B (wt = 0.3)	0.1	1	1	1	1	1
E (wt = 0.4)	0	0.3	1	1	1	1

## **Seismic Hazard Curves**

- Obtained through a PSHA
- Show annual rates of exceedance as a function of spectral acceleration
- Calculated for a given ground motion frequency (e.g., 1Hz)
- Includes contributions from all possible earthquakes in all seismic sources affecting a site
- Are used to obtain uniform hazard response spectra and GMRS



## **PSHA and Uncertainty Management**



## **Components of PSHA**

- Seismic Sources and Uncertainties
  - Source Geometries, Maximum Magnitudes
- Earthquake Occurrence Rates and Uncertainties
- Ground Motion Prediction Models and Uncertainties

$$\lambda_{g*} = \sum \sum v_i P[G > g*|m_j, r_k] P[M = m_j] P[R = r_k]$$



Uncertainties encountered in PSHA calculations are primarily related to alternative views and/or lack of certain knowledge and differ from uncertainties encountered in measurements



Two main types of uncertainties:

**Open Item** 

- Aleatory Natural variations (randomness) in observations (e.g., ground motion amplitudes)
- Epistemic Alternative conceptual models (e.g., seismic source geometries and magnitudes) Epistemic

### **Seismic Source Uncertainties**

**Open Item** 

- Arise from "differences in scientific opinion" and/or "multiple possible interpretations due to limited data"
  - Seismic source geometries
  - Source parameters, such as Mmax
  - Seismogenic thickness

## **Uncertainties in Earthquakes/Seismicity Rates**

- Earthquake location uncertainties
- Uncertainties in completeness records
- Uncertainties related to de-clustering
- Uncertainties in earthquake recurrence rate estimates
- Uncertainties in earthquake source mechanisms

## **Ground Motion Model**

- Variability (uncertainty) in ground motion recordings
- Model development uncertainties
  - Arise from "differences in scientific opinion" and/or "multiple possible interpretations due to limited data" <sup>45</sup>

## **Treatment of PSHA Uncertainties**

**Open Item** 

- Models used in new reactor PSHA calculations incorporate uncertainties to be used in seismic hazard calculations, making the uncertainty management a robust, and objective process.
  - CEUS COL and ESP applicants used either the older EPRI-SOG models or the newer NUREG-2115 source characterization models and the EPRI (2004, 2006) GMPEs in their PSHA.
- Each of these models incorporate extensive uncertainties in their model parameters (in the form of logic trees) minimizing any subjectivity in the treatment of uncertainty

## **Logic Trees in PSHA Uncertainties**

Because of incomplete knowledge, seismic source models and GMPEs always incorporate alternative assessments in PSHA inputs. Logic trees are used to represent alternative views with their assigned confidence indicators (weights)



## **Examples of Impacts of Parameter Uncertainty in PSHA Calculations**

#### **Uncertainties Related to Maximum Magnitude:**

The maximum magnitude value ( $M_{max}$ ) of a seismic source model represents the largest possible earthquake that should be expected to occur within that source. Because knowledge is limited to determine the exact value of  $M_{max}$ , it is essential to represent it not as a single value, but a range of possible values.



A range of potential **M**<sub>max</sub> values for the PEZ\_N seismic source: (5.9, 6.4, 6.8, 7.2, 7.9)

### **Example 1a: Impacts of M<sub>max</sub> Variations on PSHA Results**



NUREG-2115

### **Example 1b: Impacts of M<sub>max</sub> Variations on PSHA Results**



#### **Example 2a: Impacts of GMPEs on PSHA Results**



Spectral Acceleration (g)

#### **Example 2b: Impacts of GMPEs on PSHA Results**



### Example 3a: Impacts of Alternative Earthquake Rates on PSHA Results



## Example 3b: Impacts of Alternative Earthquake Rates on PSHA Results



## Use of the CEUS-SSC Model in COL and ESP Applications

CEUS-SSC, an accepted starting model in NPP PSHA calculations, incorporates two different types source models:

- 1) Distributed-seismicity models (Includes two submodels
  - A. Mmax Sources (large areas, minimal tectonic info)
  - B. Seismotectonic Sources (smaller sources, identified based on tectonic characterization)
- 2) Repeated Large Magnitude Earthquake (RLME) models (sources mostly identified by paleoseismology studies)

Each source model incorporates extensive uncertainty information. Uncertainties are represented by logic trees.



## **Distributed Seismicity Models**


# Single Source (Study Region)



# **M**<sub>max</sub> Source Zones (Narrow) Geometries



# **M**<sub>max</sub> Source Zones (Wide) Geometries



# **Distributed Seismicity Models**











Applicants select which seismotectonic sources to use based on source distances to their site and the level of hazard contribution to their site.

For example, in the case of Fermi Unit 3 PSHA, nine out of 12 seismotectonic sources were used.



### **High-level Logic Tree of the CEUS-SSC Model**



# CEUS-SSC: Repeated Large Magnitude Earthquake Sources



Applicants select which RLME sources to use based on the distances and sources' contribution levels to the total seismic hazard at their site.

For example, Fermi Unit 3 PSHA used a criterion of 1% or more contribution to the total hazard; resulting in the selection of NMSZ, Charleston, Charlevoix, Wabash Valley RLME sources in their PSHA analysis.



# Putting It All Together (1/2)

Once all seismic sources are selected, and all alternative geometries are identified, the next step is to calculate the seismic hazard from each seismic source (including alternative representations of the same source) using the established model parameters.

Model parameters to be used include:

- M<sub>max</sub> values
- Earthquake Rates
- GMPEs

# Putting It All Together (2/2)

In the CEUS-SSC model each seismic source is assigned 5 <u>alternative M<sub>max</sub> values</u>, each value representing a viable alternative.

CEUS-SSC model characterizes earthquake occurrence rates for distributed seismicity sources using three viable alternatives (Case A, Case B, and Case E). Each alternative case is represented by eight different "realizations" representing potential uncertainties. Hence, <u>a total of 24</u> <u>alternative rates exist per seismic source.</u>

EPRI (2004, 2006) GMPEs include four clusters, each cluster having three separate median models (capturing epistemic uncertainty), <u>a total of 12 possible models.</u>

# Example Hazard Calculations: (Identified Seismic Sources: Fermi Unit 3)

### Mmax Sources

- Study Region
- Mesozoic Extension (W/N)
- Non- Mesozoic Extension NMESE (N/W)

### **RLME Sources**

- New Madrid Fault System
- Charleston
- Charlevoix
- Wabash Valley

### Seismictectonic Sources

- Atlantic Highly Extended Crust (AHEX)
- Extended Continental crust Atlantic Margin (ECC-AM)
- Great Meteor Hotspot (GMH)
- Illinois Basin Extended Basement (IBEB)
- Midcontinent Craton (MIDC-A/B/C/D)
- Northern Appalachian (NAP)
- Paleozoic Extended Crust (PEZ-N/W)
- Reelfoot Rift (RR) and Reelfoot Rift-Rough Creek Graben (RR-RCG)
- St. Lawrence Rift (SLR)

# Example Hazard Calculations: (Fermi Unit 3)

### Mmax Sources

- Study Region
- Mesozoic Extension (W/N)
- Non- Mesozoic
  Extension NMESE (N/W)

### **RLME Sources**

- New Madrid Fault System
- Charleston
- Charlevoix
- Wabash Valley

### Seismictectonic Sources

- Atlantic Highly Extended Crust (AHEX)
- Extended Continental crust Atlantic Margin (ECC-AM)
- Great Meteor Hotspot (GMH)
- Illinois Basin Extended Basement (IBEB)
- Midcontinent Craton (MIDC-A/B/C/D)
- Northern Appalachian (NAP)
- Paleozoic Extended Crust (PEZ-N/W)
- Reelfoot Rift (RR) and Reelfoot Rift-Rough Creek Graben (RR-RCG)
- St. Lawrence Rift (SLR)

### Select One of the Seismic Sources (Seismotectonic, MIDC Source)



# **CEUS-SSC Model Describes the following** parameters for Seismic Source MIDC:

**Maximum Magnitudes for MIDC:** 





# **Earthquake Rates**

#### 24 Alternative Rates Per Source



e.g., Case A, Realization 1















#### Case B 8 Realizations

40.74

- 35 N





100 W

50 x

10 W

80'W

70.00

10.94

2014

70 W



80'W

10.01

90'W

100'W



10'W

90 W

100 W



#### Case E 8 Realizations

# **Ground Motion Prediction Equations:**

- Defined at seven distinct ground motion frequencies
   0.5Hz, 1Hz, 2.5Hz, 5Hz, 10Hz, 25Hz, 100Hz (PGA)
- Composite model including four sets of sub-models (clusters): C1, C2, C3 and C4
  - C4 is only used when seismic hazard is primarily from large magnitude sources (e.g., RLMEs)
  - Within each cluster, three different median models capture the epistemic uncertainty

In our example of calculating seismic hazard from the MIDC seismic source at the Fermi Unit 3 site, we select 3 clusters (C1, C2, and C3), with 3 alternative models, leading to a total of 3x3=9 GMPEs

# **PSHA Calculations for a Single Source**

To calculate the mean seismic hazard for this source, we use:

- 9 GMPEs
- 24 earthquake recurrence models
- 5  $M_{max}$  values

This results in a total of 9x24x5=1080 individual seismic hazard curves for a single source and single ground motion frequency. These 1080 curves represent all plausible alternative hazard levels this source could produce. These suite of curves are later used in the fractile calculations.

<u>Important to note</u>: Each seismic hazard curve carries its total weight, calculated using the weights assigned to each logic tree branches.

# Results of PSHA: Seismic Hazard Curves Calculated for the Fermi Unit 3 Site (Single Source: MIDC)





# Mean and Fractile Calculations in PSHA

# Seismic Hazard Curves' Weights



### GMPEs

**C1** (0.3512) **C1-L** (0.185) **C1-M** (0.630) **C1-H** (0.185)

- **C2** (0.3985) **C2-L** (0.185) **C2-M** (0.630) **C2-H** (0.185) **C3** (0.2503) **C3-L** (0.185)
  - **C3-M** (0.630)

**C3-H** (0.185)

(0.125)

e.g., The seismic hazard curve calculated using Mmax=5.6, Case A/R1, C1-L would have a weight of 0.000246

# **Fractile Calculations (From Single Source Curves)**





### **Seismic Hazard Curves From a Single Source**





# From a Single Source to Multiple Sources

The mean of the 1080 hazard curves (obtained for MIDC) is added to the means calculated from all other seismic sources (that are identified to impact the site) to obtain the total mean seismic hazard curve for the site for a given ground motion frequency. Since there are seven different frequencies in the EPRI (2004, 2006) GMPEs, this process is repeated for all seven ground motion frequencies.

Fractile calculations, however, get much more complex when more than one source impacts the site, (as it is always the case in the CEUS-SSC model). Numerous combinations of seismic hazard curves need to be identified.

# A Realistic Scenario (Fermi Unit 3 PSHA)

### Mmax Sources

- Study Region
- Mesozoic Extension (W/N)
- Non- Mesozoic Extension NMESE (N/W)

# RLME Sources

- New Madrid Fault System
- Charleston
- Charlevoix
- Wabash Valley

### Seismictectonic Sources

- Atlantic Highly Extended Crust (AHEX)
- Extended Continental crust Atlantic Margin (ECC-AM)
- Great Meteor Hotspot (GMH)
- Illinois Basin Extended Basement (IBEB)
- Midcontinent Craton (MIDC-A/B/C/D)
- Northern Appalachian (NAP)
- Paleozoic Extended Crust (PEZ-N/W)
- Reelfoot Rift (RR) and Reelfoot Rift-Rough Creek Graben (RR-RCG)
- St. Lawrence Rift (SLR)

### **High-level Logic Tree of the CEUS-SSC Model**



### High-level logic tree showing in seismic source models


#### Mmax branch of the high-level logic tree





#### Seismotectonic branch of the logic tree (1/2)



#### Seismotectonic branch of the logic tree (2/2)



#### RLME sources are also added



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#### Each RLME source has its own logic Tree: e.g., NMSZ



# A Simplified Example to Estimating Alternative Seismic Hazard Curves and Their Weights

Since the full models are complex and involve many thousands of seismic hazard curves, let's examine a simpler model to demonstrate the process used to calculate the fractiles when multiple seismic sources are involved:

- 2 GMPEs (C1-M & C3-M)
- 2  $M_{max}$  sources with two  $M_{max}$  values/source (M1 & M2)
- 2 Seismotectonic sources with 2 alternatives
- 2 Earthquake rates (R1 & R2) for each source
- No RLME





Given C1-M (GMPE): 16 alternative hazard curves for the top branch of the logic tree

MESE(R1,M1) + NMESE(R1,M1) MESE(R1,M2) + NMESE(R1,M1) MESE(R2,M1) + NMESE(R1,M1) MESE(R2,M2) + NMESE(R1,M1)

MESE(R1,M1) + NMESE(R1,M2) MESE(R1,M2) + NMESE(R1,M2) MESE(R2,M1) + NMESE(R1,M2) MESE(R2,M2) + NMESE(R1,M2) MESE(R1,M1) + NMESE(R2,M1) MESE(R1,M2) + NMESE(R2,M1) MESE(R2,M1) + NMESE(R2,M1) MESE(R2,M2) + NMESE(R2,M1)

MESE(R1,M1) + NMESE(R2,M2) MESE(R1,M2) + NMESE(R2,M2) MESE(R2,M1) + NMESE(R2,M2) MESE(R2,M2) + NMESE(R2,M2)

# Part (A) hazard curve weights

No.	Total Hazard Curves	Weight
1	C1-M & MESE(1,1) & NMESE(1,1)	.3x.4(.2x.2)=0.0048
2	C1-M & MESE(1,2)& NMESE(1,1)	.3x.4(.2x.2)=0.0048
3	C1-M & MESE(2,1)& NMESE(1,1)	.3x.4(.3x.2)=0.0072
4	C1-M & MESE(2,2)& NMESE(1,1)	.3x.4(.3x.2)=0.0072
5	C1-M & MESE(1,1)& NMESE(1,2)	.3x.4(.2x.2)=0.0048
6	C1-M & MESE(1,2)& NMESE(1,2)	.3x.4(.2x.2)=0.0048
7	C1-M & MESE(2,1)& NMESE(1,2)	.3x.4(.3x.2)=0.0072
8	C1-M & MESE(2,2)& NMESE(1,2)	.3x.4(.3x.2)=0.0072
9	C1-M & MESE(1,1)& NMESE(2,1)	.3x.4(.2x.3)=0.0072
10	C1-M & MESE(1,2)& NMESE(2,1)	.3x.4(.2x.3)=0.0072
11	C1-M & MESE(2,1)& NMESE(2,1)	.3x.4(.3x.3)=0.0108
12	C1-M & MESE(2,2)& NMESE(2,1)	.3x.4(.3x.3)=0.0108
13	C1-M & MESE(1,1)& NMESE(2,2)	.3x.4(.2&.3)=0.0072
14	C1-M & MESE(1,2)& NMESE(2,2)	.3x.4(.2&.3)=0.0072
15	C1-M & MESE(2,1)& NMESE(2,2)	.3x.4(.3x.3)=0.0108
16	C1-M & MESE(2,2)& NMESE(2,2)	.3x.4(.3x.3)=0.0108 <sup>103</sup>

# Part (B) hazard curve weights

No.	Total Hazard Curves	Weight
17	C1-M & MIDC-A(1,1)&RR(1,1)	.3x.3(.2x.2)=0.0036
18	C1-M & MIDC-A(1,2)&RR(1,1)	.3x.3(.2x.2)=0.0036
19	C1-M & MIDC-A(2,1)&RR(1,1)	.3x.3(.3x.2)=0.0054
20	C1-M & MIDC-A(2,2)&RR(1,1)	.3x.3(.3x.2)=0.0054
21	C1-M & MIDC-A(1,1)&RR(1,2)	.3x.3(.2x.2)=0.0036
22	C1-M & MIDC-A(1,2)&RR(1,2)	.3x.3(.2x.2)=0.0036
23	C1-M & MIDC-A(2,1)&RR(1,2)	.3x.3(.3x.2)=0.0054
24	C1-M & MIDC-A(2,2)&RR(1,2)	.3x.3(.3x.2)=0.0054
25	C1-M & MIDC-A(1,1)&RR(2,1)	.3x.3(.2x.3)=0.0054
26	C1-M & MIDC-A(1,2)&RR(2,1)	.3x.3(.2x.3)=0.0054
27	C1-M & MIDC-A(2,1)&RR(2,1)	.3x.3(.3x.3)=0.0081
28	C1-M & MIDC-A(2,2)&RR(2,1)	.3x.3(.3x.3)=0.0081
29	C1-M & MIDC-A(1,1)&RR(2,2)	.3x.3(.2&.3)=0.0054
30	C1-M & MIDC-A(1,2)&RR(2,2)	.3x.3(.2&.3)=0.0054
31	C1-M & MIDC-A(2,1)&RR(2,2)	.3x.3(.3x.3)=0.0081
32	C1-M & MIDC-A(2,2)&RR(2,2)	.3x.3(.3x.3)=0.0081104

# Part (C) hazard curve weights

No.	Total Hazard Curves	Weight
33	C1-M & MIDC-B(1,1) & RR-RCG(1,1)	.3x.3(.2x.2)=0.0036
34	C1-M & MIDC-B(1,2) & RR-RCG(1,1)	.3x.3(.2x.2)=0.0036
35	C1-M & MIDC-B(2,1) & RR-RCG(1,1)	.3x.3(.3x.2)=0.0054
36	C1-M & MIDC-B(2,2) & RR-RCG(1,1)	.3x.3(.3x.2)=0.0054
37	C1-M & MIDC-B(1,1) & RR-RCG(1,2)	.3x.3(.2x.2)=0.0036
38	C1-M & MIDC-B(1,2) & RR-RCG(1,2)	.3x.3(.2x.2)=0.0036
39	C1-M & MIDC-B(2,1) & RR-RCG(1,2)	.3x.3(.3x.2)=0.0054
40	C1-M & MIDC-B(2,2) & RR-RCG(1,2)	.3x.3(.3x.2)=0.0054
41	C1-M & MIDC-B(1,1) & RR-RCG(2,1)	.3x.3(.2x.3)=0.0054
42	C1-M & MIDC-B(1,2) & RR-RCG(2,1)	.3x.3(.2x.3)=0.0054
43	C1-M & MIDC-B(2,1) & RR-RCG(2,1)	.3x.3(.3x.3)=0.0081
44	C1-M & MIDC-B(2,2) & RR-RCG(2,1)	.3x.3(.3x.3)=0.0081
45	C1-M & MIDC-B(1,1) & RR-RCG(2,2)	.3x.3(.2&.3)=0.0054
46	C1-M & MIDC-B(1,2) & RR-RCG(2,2)	.3x.3(.2&.3)=0.0054
47	C1-M & MIDC-B(2,1) & RR-RCG(2,2)	.3x.3(.3x.3)=0.0081
48	C1-M & MIDC-B(2,2) & RR-RCG(2,2)	.3x.3(.3x.3)=0.0081105



Additional 48 hazard curves are calculated using C3-M GMPE in a similar fashion. Since C1-M has a weight of 0.3, and C3-M has a weight of 0.7, in terms of the weights, the only difference would be that all weights in the C3-M case would have 2.33 times of the individual weights calculated for C1-M.

This produces 96 possible seismic hazard curves. The sum of all weights in these 96 curves must be equal to 1.0.

#### **Revisiting the Questions**

1. Why doesn't the uncertainty increase appreciably as you go from small spectral accelerations to very high spectral accelerations?



### **Answer to Question 1**

- The uncertainty or range in low to high fractile curves does increase with increasing spectral accelerations but does so at varying rates for each of the seven frequencies
- Lower frequency SA hazard curves (0.5,1, and 2.5 Hz) are much more sensitive to key parameters such as M<sub>max</sub> than moderate to higher frequency curves (5, 10, 25, and 100 Hz)

2. Why is the uncertainty small for very high spectral accelerations at 25 Hz when compared to the uncertainty for very high spectral accelerations at 0.5 Hz?



# Answer to Question 2 (1/2)

- Only a limited number of parameter combinations (M<sub>max</sub>, rates, GMPEs, distances) are able to produce larger 0.5 Hz SA values
  - To generate large low frequency (0.5 Hz) spectral accelerations need seismic sources capable of producing large earthquakes at close in distances
  - For Fermi RLMEs are very distant and do not produce large low frequency (0.5 Hz) spectral accelerations
  - CEUS-SSC distributed seismicity sources are capable of large magnitude earthquakes (MidC  $M_{max}$ 5.6 to 8.0) as well, but rates are very low for higher magnitude scenario earthquakes ( $\lambda_{M7}$  about 1.5E-06/yr for a typical cell)

### Answer to Question 2 (2/2)

- Nearly all parameter combinations (Mmax, rates, GMPEs, distances) produce larger 25 Hz SA values
  - Both CEUS-SSC distributed seismicity source zones and RLME sources can produce large high-frequency (25 Hz) accelerations at rock sites in the CEUS
  - Earthquakes that contribute the most to hazard at 10<sup>-4</sup> annual exceedance frequency (AEF) over the 5 to 10 Hz range for CEUS sites are typically moderate-sized earthquakes (M5.5-M6) from nearby distributed seismicity sources (10-30 km)

# **Answers to Question 3**

- Uncertainties are inherent in PSHA calculations and handled through well-established processes. Once all alternative models and variations are described and built into a PSHA input model (e.g., CEUS-SSC model), calculations are objective.
- Seismic source geometries and source parameters (such as Mmax, earthquake rates) as well as ground motion models contribute to uncertainties, each having variable levels of impacts.

# Conclusions

- Capturing the appropriate level of uncertainty is an integral part of the hazard characterization for new reactor siting
- Consistent with RG 1.208, the GMRS is developed using the mean hazard curves at each spectral frequency
  - The mean hazard curves incorporate model uncertainties
- Fractile hazard curves are key to understanding how applicants characterize the hazard for their sites
- While the fractile hazard curves are site-specific, M<sub>max</sub> and GMPEs are shown to provide larger portions of the variations in the fractile curves

# **Backup Slides**

















## Fermi Contribution to Hazard by Source



#### Contribution from Individual Sources at the

### Fermi Deaggregation of 10<sup>-4</sup> Mean Hazard



#### Fermi Deaggregation of 10<sup>-6</sup> Mean Hazard



Magnitude Interval

#### **Probability of Low Frequency Motions at Fermi**

- We can use CEUS Ground Motion Prediction Equations to determine the probability of low frequency (0.5 Hz) motions at Fermi from local and distant sources
- $\Pr(SA_{0.5 Hz} > x | M = 7.5, R = 25, 200, 320 \text{ km})$
- Probability of large 0.5 Hz spectral accelerations decreases significantly with increasing source to site distances



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#### Ground Motion Prediction Equations

- Developed for a specific ground motion frequency (f)
- General form for median ground motion- (Central and Eastern North America)
- Ln ( $Y^{Pred}$ ) = C<sub>1</sub>+C<sub>2</sub>\*M+(C<sub>3</sub>+C<sub>4</sub>\*M)\*Ln(R+exp(C<sub>5</sub>))+ C<sub>6</sub>\*(M-M<sub>1</sub>)<sup>2</sup> + (C<sub>7</sub>+C<sub>8</sub>\*M)\*R<sub>1</sub>
- Express the estimate of the median ground motion parameter of interest (Y- often PSA) in terms of explanatory variables: magnitude (M) and distance (R)

• 
$$Ln(Y^{obs}) = Ln(Y^{pred}) + \delta$$

•  $\delta$ -delta is the residual term




-2

-4

-4

-2

0

Standard Normal Quantile

2

4

Ln (Y<sup>obs</sup>) = GMPE<sup>pred</sup> f(M, R, S) +  $\delta$  = GMPE<sup>pred</sup> f(M, R, S) +  $\epsilon^* \sigma$