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UNITED STATES OF AMERICA
NUCLEAR REGULATORY COMMISSION

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ADVISORY COMMITTEE ON NUCLEAR WASTE

+ + + + +

WEDNESDAY

SEPTEMBER 22ND, 2004

+ + + + +

The Committee met at the Suncoast Hotel,
9090 Alta Drive, Ballroom A, Las Vegas, Nevada.

Advisory Committee Members Present:

MICHAEL T. RYAN	CHAIRMAN
RUTH F. WEINER	MEMBER
ALLEN G. CROFF	MEMBER

Others Present:

KEITH ECKERMAN	Oak Ridge National Laboratory
FRED HARPER	Sandia National Laboratories
DAVID JOHNSON	ABS Consulting
BRUCE CROWE	Los Alamos National Laboratory
DR. BILL MELSON	Smithsonian National Institute
MICHAEL LEE	ACNW
JOHN LARKINS	ACNW
JAMES CLARKE	ACNW
WILLIAM HINZE	ACNW

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1 Others Present:

2 BRUCE MARSH ACNW

3 BOB BRUDNITZ LLNL -- on detail to DOE

4 LYNN ANSPAUGH University of Utah

5 B. JOHN GARRICK NWTRB

6 GEORGE HORNBERGER NWTRB

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P-R-O-C-E-E-D-I-N-G-S

8:05 a.m.

OPENING REMARKS

CHAIRMAN RYAN: Good morning. The meeting will now come to order. This is the first day of the 153rd meeting of the Advisory Committee on Nuclear Waste.

I am Michael Ryan, Chairman of the ACNW. The other members of the Committee present are Ruth Weiner and Allen Croff. Also present are ACNW consultants William Hinze and Bruce Marsh.

James Clark, another ACNW consultant will be joining us later in the meeting. He was unavoidably called away. During the next two days the Committee will conduct a working group meeting to review and discuss issues related to the evaluation of igneous activity and its consequences at a potential geologic repository Yucca Mountain, Nevada.

The Committee will gather information, analyze relevant issues and facts, and formulate proposed positions and actions as appropriate in the form of advice to the Commission.

The meeting is being conducted in accordance with the provisions of the Federal Advisory Committee Act. The rules for participation in today's

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1 meeting have been announced as part of the notice of
2 this meeting previously published in the Federal
3 register.

4 Mr. Mike Lee is the designated Federal
5 Official for these sessions. A transcript of this
6 meeting is being kept. And the transcript will be
7 made available as stated in the Federal register
8 notice.

9 It is requested that speakers first
10 identify themselves and speak with sufficient clarity
11 and volume so that they can be readily heard.

12 We have received no request for time to
13 make oral statements from members of the public
14 regarding today's sessions. Should anyone wish to
15 address the Committee, please make your wishes known
16 to one of the Committee's staff.

17 As an administrative matter, if you
18 haven't already done so, it is requested that you sign
19 in at the table in the back. We also request that, if
20 you have them, please confirm that your cell phones
21 are turned off or alternatively have been rendered
22 into silent ringing mode.

23 Lastly, for those of you who wish to do
24 so, there are comment feedback sheets available at the
25 sign-in desk. Items of interest, before starting the

1 first session, I would like to cover some brief items
2 of current interest.

3 On August 16th, 2004 President Bush
4 announced his intention to appoint ACNW members Dr.
5 John Garrick and Dr. George Hornberger to the Nuclear
6 Waste Technical Review Board.

7 Dr. Garrick was designated as the Board's
8 new Chairman. We regret their resignations from the
9 Committee and wish them well in this new endeavor.
10 Congratulations to you both in every success.

11 The Committee and I, as the previous
12 Committee Vice-Chair, have assumed the Chairmanship of
13 the ACNW. Volumes one and two of the Nureg 1710
14 series on the history of water development in the
15 Amargosa desert were recently approved for publication
16 by the ACNW's Executive Director.

17 These Nuregs were co-authored by Mike Lee
18 and Neil Coleman of the ACNW technical staff and Tom
19 Nicholson of the NRC's Office of Nuclear Regulatory
20 Research.

21 In addition to service to this Committee,
22 the ACNW has encouraged the support of the Staff's
23 efforts to publish technical reports and papers -- the
24 Agency's overall mission.

25 Lastly, Mr. Marvin Sikes, a Senior Staff

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1 Engineer with the Advisory Committee on Reactor
2 Safety, the ACNW sister Committee, has been selected
3 to fill a branch D position in NRC's region one
4 division of reactor safety.

5 He will depart for his new position in
6 mid-November, and the Committee wishes him well. The
7 ACNW has been tracking developments related to the
8 modeling of a disruptive igneous event at Yucca
9 Mountain for several years.

10 Earlier Committee views on the pertinent
11 issues can be found in five letter reports. Copies of
12 these letter reports can be found in the Committee's
13 internet web, as well as in Nureg 1423, the
14 compilation series for ACNW letters.

15 Most recently, in June 2002, the ACNW
16 conducted a workshop group meeting to learn more about
17 the issues which resulted in the letter report for the
18 Commission dated August 1st, 2002.

19 WORKING GROUP PURPOSES

20 The overall focus of the working group
21 meeting is to better understand what knowledge base is
22 available for decision making, areas of specific ACNW
23 interest, including understanding the realism of
24 existing approaches and calculations and identifying
25 areas in those approaches and calculations that may

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1 require additional work.

2 Consistent with the published agenda,
3 three technical sessions, consisting of about 15
4 presentations are planned over two days to focus on
5 the treatment of probability, consequence, and dose in
6 igneous activity performance assessment analysis.

7 To help the Committee explore the issues
8 and interrogate the invited speakers, and maybe just
9 have a conversation with the invited speakers, rather
10 than interrogate, a panel of invited experts has been
11 assembled.

12 They include Dr. Robert Budnitz from the
13 Lawrence Livermore National Laboratory, Dr. Dave
14 Johnson from ABS consulting of Irvine, California, Dr.
15 William Hinze, Professor of geology and geophysics at
16 Perdue University, and Dr. Bruce Marsh, professor of
17 igneous petrology at Johns Hopkins University, and
18 finally Dr. William Melson, Senior Scientist of
19 volcanology at the Smithsonian Institute in
20 Washington, D.C.

21 Welcome all, thank you very much for your
22 time and participation in this working group meeting.
23 At the conclusion of tomorrow's meeting, Dr. Johnson
24 will provide summary remarks concerning the issues
25 discussed in the context of the application of the

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1 risk triplet, the risk triplet being three questions.

2 What can go wrong? How likely is it? And
3 what are the consequences? So, we will be thinking
4 along those lines. The first session planned today is
5 on probability.

6 Areas of specific ACNW interest here
7 include understanding the types and kinds of geologic
8 information needed for generating probability
9 estimates, the uncertainty in that information, and
10 identifying which analytical approaches yield
11 defensible estimates.

12 And, to address these issues, three
13 presentations have been scheduled for the first
14 session. The first presentation will be by Dr. John
15 Trapp of the NRC staff, and will feature a discussion
16 of the geologic features of the Yucca Mountain region
17 considered to be important in the estimation of
18 igneous event probabilities.

19 Dr. Bruce Crowe, of Los Alamos National
20 Laboratory, former principal investigator of igneous
21 activity in DOE's Yucca Mountain programs, and a
22 subject matter expert in the 1996 probabilistic
23 volcanic hazards analysis, will share his perspectives
24 on the type of geologic information that is important
25 to decision making at the time the expert elicitation

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1 is conducted.

2 Other perspectives on the interpretation
3 of the local geologic record, and how it affects
4 probability estimates will be made in a presentation
5 by Mr. Neil Coleman of the ACNW staff.

6 He will present a paper that he co-
7 authored with Dr. Lee Abrams of NRC's Office of
8 Research and Bruce Marsh that was recently submitted
9 to geophysical research letters.

10 This paper relies on statistical methods
11 to evaluate the probability of the issue. I'll talk
12 about the second session when we begin that session.
13 So, without further ado, let me turn to our first
14 speaker, Dr. John Trapp.

15 **NRC PERSPECTIVE ON VOLCANISM MODELING ISSUES**

16 MR. TRAPP: Okay, Good morning. Like I
17 was saying, a few comments. The actual discussion on
18 probability comments will be given by Dr. Britt Hill.

19 I'm going to be presenting just a brief
20 overview of-- our program, talking about really the
21 main assumptions. That was the second one we were
22 talking about.

23 And then, in addition, talking about what
24 we feel like the risk significant items that we need
25 to understand. That's basically coming out of -- I

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1 partly should first off say some things that we will
2 not be talking about.

3 We will not be discussing any of the work
4 that is presently in progress. Everything that we are
5 talking about today from the NRC perspective will
6 material that is readily available to the public.

7 In addition, we will not be making
8 comments about DOE's licensing case. A, we really
9 don't know it, and B, it's inappropriate at this time
10 to discuss this type of things by the NRC staff.

11 Next slide please. So, what am I going to
12 be doing? I'm basically going to, like I said, be
13 providing a basic assumption, the NRC's and the RPA --
14 evaluating these.

15 Based on results that we have -- are not
16 specific. Next slide please. For those of you who
17 have not been to the area of Yucca Mountain, this was
18 just kind of a slide overview.

19 The center of the slide is Yucca Mountain.
20 If you take a look off to the west, you will see Bare
21 Mountain. And, in between Yucca Mountain and Bare
22 Mountain, there are a series of electrons down there.

23 As you come to the southeast, in the
24 Crater Flat area, what you don't see is a series of
25 other basalts, which are basically 3.7, approximately,

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1 a million years old.

2 Farther down, at the very tip of the
3 mountain, you will see the youngest igneous feature,
4 which is present in the area, Lathrop Wells. The
5 Amargosa Desert area is an area which has quite a few
6 varied igneous features and quite a few anomalies,
7 which may or may not be igneous features.

8 This is an area where DOE has run a recent
9 aeromagnetic electromagnetic survey, the results of
10 which are just starting to become available.
11 Yesterday the preliminary results from DOE -- this
12 hopefully will shed a lot of light on information
13 about the distribution of igneous bodies in the area,
14 and help us work to determine the probability.

15 Jackass Flats, which is on the west side,
16 or the east side of Yucca Mountain, has feature
17 covered mountains -- the Fortymile Wash basin, which
18 is going to be quite important in the whole
19 discussion.

20 Let's take a look at these. That's of the
21 wells that was drilled by -- there was a basalt of, I
22 believe, nine and a half, a million years,
23 approximately from that well.

24 More important, for the sake of some of
25 the discussions that will be going on negative today,

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1 if you notice, the drainage coming out of Fortymile
2 Wash, you will see just along the highway, running
3 east and west through that area 23 feet that the
4 characteristic of the drainage system changes
5 tremendously.

6 You're going from a marginal system into
7 a depositional system. This also happens to be
8 approximately in the area where the reasonably
9 maximally exposed individual, the person that we have
10 to use to characterize doses to the public too high.

11 Next slide please. So, what are some of
12 the basic assumptions? Well, if you took a look at
13 that slide, you will see that a small volume of
14 basaltic cones have occurred in the general area of
15 Yucca Mountain in the past.

16 And there is some potential that there
17 will be future basaltic igneous events that could
18 possible occur. We modeled it, the DOE has modeled
19 it, the State has modeled.

20 So far, all the models -- and there is
21 quite a bit of arguments back and forth on what the
22 probability is -- but, all the probability models come
23 out a value that's larger than regulatory requirements
24 considered in our performance assessment.

25 There is large uncertainty with this, like

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1 I said DOE has finished the aeromagnetic survey. They
2 have been doing some drilling with some of the
3 anomalies to determine which ones really are under the
4 basalt.

5 They are going to be digging those
6 basalts. All of this will hopefully produce the
7 uncertainty on this probability. Next slide. If you
8 take a look at the volcanoes that you've got in the
9 aerial, you'll see that these all produce not only
10 lava flows, but their results, the deposits, show
11 periods of sustained eruption columns with buoyant
12 tephra plumes.

13 If you take a look at the historically
14 active analog, what you will see is these type of
15 volcanoes have the capability of hitting buoyant
16 plumes and transporting them 10 to 100 kilometers
17 downwind.

18 If you take a look at some of the recent
19 results that have been published in the literature,
20 what you will find is, contrary to some of the earlier
21 modeling and some of the assumptions, these basalts
22 are actually quite wet.

23 They have got -- the best estimate would
24 be something like about four percent water. One of
25 many of the original modeling studies on these were

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1 done with much lower water percents.

2 And what we are talking about is, with
3 this high water, you definitely have potential
4 fragmenting and getting these connected dispersive
5 plumes.

6 Next slide please. One of the questions
7 that has been asked quite a few times in the past is,
8 why didn't we put any other -- a risk likely to in
9 this?

10 Well, one of the reasons is, there really
11 isn't a good way to measure how big the volcano is.
12 Here is one example. If you take a look along the
13 top, you will see that, really what it is talking
14 about is two factors.

15 How much tephra is produced in the ash?
16 And how high do these columns get? It doesn't talk
17 about the total volume of magma produced. It concerns
18 some of the other type eruption sequences.

19 If you go on down, you will look at the
20 volcanic explosive that makes number two. And,
21 basically, this is -- of all the studies we have done,
22 approximately a majority of the events -- they may
23 sneak down to a one.

24 They may sneak up to a three. But,
25 really, we're talking about a single class for all

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1 practical purposes, is ten to the minus seven cubic
2 meters tephra and columns on the order of two to five,
3 maybe seven, possibly even size ten, but I doubt that
4 high.

5 Another important point is, if you go on
6 down a line, you will see that these do not get into
7 the stratospheric level. So, this is very important
8 in talking about some of the potential health effects
9 and other considerations for that.

10 Next slide. Another assumption that we
11 have is that the waste package is intercepted by magma
12 and be subject to very high thermal stress, and very
13 large mechanic stress.

14 This is a likely caused failure of the
15 canister. And, therefore, many radioactive waste is
16 exposed to magma. We have given this problem to our
17 waste package engineers and talked about the
18 conditions that we got in this type of situation.

19 And, with the days to weeks that this
20 package would be subject to these types of thermal
21 stresses, mechanical stresses, the conclusion that we
22 come to is that this package -- well, basically, can
23 be breached.

24 Our assumptions in the -- what we have
25 done, is assumed the waste package offers no

1 protection whatsoever. This assumption has been used
2 in previous DOE analysis.

3 It may change, etcetera, but this is the
4 present assumption that we are using. Next slide.
5 Okay. We've got the package breached. So what
6 happens?

7 Well, we've got the waste sitting there.
8 And this is now assumed to be available to put in the
9 tephra column. We don't really model, like I said,
10 lava flows for a very simple fact.

11 If you take a look at all the data that
12 you've got, a lava flow by itself really doesn't pick
13 up too much. We do not assume that this waste melts
14 in the basalt, because, really, we do not have the
15 type of material that would dissolve in magma.

16 What we're following is what you see in a
17 normal eruption, the fragmentation of the wall rock,
18 the fragmentation of the material. This gets broken
19 down in small sizes and traded with the material, and
20 put up, and then transformed back.

21 Next slide. Okay, you've got stuff up in
22 the air. You've got a transporter downwind. It falls
23 to the ground. Well, when it hits the ground, we're
24 basically assuming that, yes, you can suspend the
25 stuff into the air, from which people can breathe it

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1 and get a dose.

2 This turns out to be the igneous scenario,
3 the main method by which the dose gets in the -- we're
4 using a bunch of very simplified assumptions going all
5 the way through this.

6 It really boils down to two primary
7 factors. What mass loading factor do we use, and how
8 long does the deposit last? Next slide. The next
9 assumption we talked about, we took a look at where
10 the site was.

11 Okay, and take a look at where the remedy
12 is, and try to do what I talked about, modeling
13 assumptions. The majority of the time, based on our
14 knowledge of the winds and the altitude, the tephra
15 column will not go directly to the RMEI.

16 It would sometimes. But, most of the
17 times, it would be blown somewhere east and deposited
18 at Jackass Flats. So, I'm going to get to the RMEI.

19 I'm going to get to the RMEI by two means.
20 It can be brought down by strain erosion. And, if you
21 took a look at the Fortymile Wash, what you will see
22 at the Fortymile Wash, like I said, as you go right at
23 the RMEI location, right before the erosional
24 sequence, that position of sequence.

25 It can also be brought by wind erosion,

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1 etcetera. This, we believe, is a very important
2 factor which needs to be taken care of. And we're
3 working on that with the Staff.

4 Next slide. So, what's important? Well,
5 according to our code, what is the probability of an
6 event. This is apparently straight-forward. The rest
7 is directly proportional to the probability.

8 Dr. Britt Hill will be presenting
9 information on that. Another significant thing, well,
10 the waste package is intersected by volcanic events.

11 And we're talking about the risk being
12 proportional to the amount of waste that can be
13 exposed. So far, packages in a larger area, the large
14 area was.

15 The volume of ash produced during an
16 eruption was important. And this is actually the
17 inversely proportional, because, what you end up with
18 here is a delusional package.

19 Larger volume eruptions tend to dilute the
20 amount the material that is there. Smaller volume
21 eruptions encounter larger concentrations. With these
22 two factors, especially number two that we will
23 discuss to certain extend this afternoon when we get
24 to that session.

25 Next slide. As I mentioned,

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1 remobilization of the process is important, because
2 this will keep the majority of the ash to the
3 location. Dr. Don Hooper will be discussing that, I
4 believe, in tomorrow's session.

5 He will talk about the modeling in this
6 area. And, like I mentioned, inhalation is the major
7 factor by which you get the dose to the humans.

8 So, we will have a discussion talking
9 about this fact, this subject matter, and how it is
10 handled. These are the important things that we see
11 in the load.

12 They can all be discussed in more detail
13 later.

14 CHAIRMAN RYAN: Thank you, Dr. Trapp. And
15 thank you for competing with the music next door.
16 Maybe we can get somebody to see about turning that
17 down just a tweak.

18 Thank you. Are there any openings? I
19 think John set the stage for the following
20 presentations and their own opening for John. Or
21 shall we reserve out thoughts for the more detailed
22 presentation? Yes, Bill Hinze?

23 MR. HINZE: Well, let me ask you, John,
24 you did an excellent job going through all of the
25 assumptions at various stages.

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1 CHAIRMAN RYAN: You have to flip the
2 microphone on.

3 MR. HINZE: I would like to ask you about
4 this. We all understand that there are uncertainties
5 with modeling because you use various assumptions.

6 Some of these uncertainties remain.
7 Others, we would like to -- and Britt will expand upon
8 this. Which of these has the greatest chance in the
9 next few years of decreasing the uncertainty with
10 better models, with better data?

11 MR. TRAPP: I think we can reduce the
12 uncertainty quite a bit by taking a look at the
13 remobilization. I think that is an extremely
14 important factor.

15 Again, you are correct, you have large
16 uncertainties. And we're not going to get rid of them
17 by -- coming out in the DOE program to reduce the
18 uncertainties in the probability model.

19 Again, we will not eliminate them. But we
20 will reduce them. There is work that is going on in
21 the understanding of magma flow, some of which you've
22 got some preliminary. And there is quite a bit more,
23 which we cannot discuss at this time.

24 And, yes, I think there will be some
25 reduction in uncertainty in that area, but not as much

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1 as we could probably expect over the areas of the
2 remobilization period. Britt, would you want to
3 comment on that?

4 MR. HILL: That was fine.

5 CHAIRMAN RYAN: Any other opening
6 questions or comments?

7 MR. HINZE: If Britt doesn't have a
8 comment, I would like to ask you about this dilution
9 that you mentioned. And perhaps Don will expand upon
10 this in his presentation.

11 I understand he's making a presentation on
12 this re-distribution of distribution. Yes, you
13 mentioned that you are really interested in having
14 more tephra because that leads to dilution.

15 But, according to your slide six, as we
16 have larger amounts of tephra, our column height also
17 increases.

18 MR. TRAPP: Right.

19 MR. HINZE: And that means that -- to me --
20 - you have greater dispersion. And so, does this
21 necessarily mean that, as you go from violent to
22 whatever, that you really are leading to dilution?

23 MR. TRAPP: If you could have those type
24 of eruption, yes you would be getting a tremendous
25 amount of more dilution. But, seeing no evidence that

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1 we would have eruptions or volcanic activity, it would
2 be anything more than approximately -- two.

3 So, we're really talking about a very
4 limited subset of that. You would not have something
5 like a PDI 4 or like a Mount Saint Helens.

6 MR. HINZE: It just seems to me that, if
7 you have more, you don't dilute because you're
8 throwing it up higher and spreading it out more.

9 MR. TRAPP: That's true.

10 MR. HINZE: Okay.

11 CHAIRMAN RYAN: John, just a quick follow-
12 up as just kind of a question for maybe some of the
13 other presenters as well. We kind of end up at the
14 end of the day with a question of what is in the air
15 that's inhaled by the RMEI or some theorized person?

16 You've touched on a lot of very complex
17 processes that get us to what is an irrespirable size
18 range in the fraction for that exposure scenario.

19 That's very complicated. And Bill has
20 touched on one aspect of that. So, to the extent you
21 and the other speakers can talk a little bit about,
22 you know, what part of the mobilization process in an
23 event leads us to that endpoint of irrespirable
24 particles. That would be real helpful.

25 MR. TRAPP: Part of what Britt will be

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1 talking about will cover that. Don will definitely
2 cover that. Keith Compton will go into effects.

3 CHAIRMAN RYAN: Thanks. To me, that's
4 kind of the focal point. Because, at the end of the
5 day, having uncertainty on that is really where you
6 can kind of begin, you know, be satisfied or
7 unsatisfied with the uncertainty question.

8 MR. TRAPP: Like I said, Don will be
9 discussing that.

10 CHAIRMAN RYAN: Okay, great. Thanks.

11 MR. HILL: This is Brittain Hill at the
12 CNWRA. I just wanted to clarify a little bit for Dr.
13 Hinze in response to his comment. In our performance
14 assessment and calculates, we allow the total volume
15 of tephra to be ten to the sixth to ten to the eight
16 cubic meters.

17 But, the column height is though of not
18 only of the volume, but the rate that it would come
19 out. So, we also vary the duration of the event
20 between essentially one day to like a week.

21 It's about five days, is our approximate
22 sort of mass blow. So, the column height, while it is
23 partially a function of volume, is also a function of
24 duration.

25 So, when we run a large number of

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1 realizations in our performance assessment, we can
2 have small volume events that happen over a very short
3 period of time give us a high volume.

4 We can also have larger volume events that
5 would happen over a long period of time, that would
6 give us the lower volume. It's not quite as
7 straightforward as simply larger volume, more distal
8 dispersion.

9 And, also, the source is about to vary
10 between one and ten waste packages per event. So, we
11 are getting that full sample in the variability. And
12 no one particular size is truly driving the risk
13 analysis.

14 MR. HINZE: I think we'd all like to hear
15 about that in more detail as the presentations are
16 made. I guess one of my concerns is that this is a
17 useful chart, but it is very simplistic. And that's,
18 I think, what you are saying.

19 MR. HILL: Yes.

20 MR. HINZE: Yes, don't hang your hat on
21 that.

22 MR. HILL: No, this figure was just meant
23 to be an example of the full range of volumes that
24 volcanoes can produce. And, relative to that full
25 range, here is the area of interest for a particular

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1 hazard related to potential --

2 MR. HINZE: There are a lot of problems
3 with Richter magnitude, but at least it's --

4 MR. HILL: Right.

5 CHAIRMAN RYAN: Any other questions?
6 You've accomplished the goal of the first speaker,
7 which is to get everybody's attention and stimulate
8 their interest. So, off we go.

9 NRC OVERVIEW OF IGNEOUS ACTIVITY AT THE YUCCA

10 MOUNTAIN REGION

11 MR. HILL: Good morning. It's nice to see
12 we have such a taste in laptop computers. That's the
13 correct one. I'm Brittain Hill. I'm the principal
14 investigator for igneous activity at the Center for
15 Nuclear Waste Regulatory Analysis.

16 And, this first talk this morning, I would
17 like to talk to you about some of the Staff's
18 positions and tools that we have developed for
19 assessing the effects of uncertainty on probability
20 estimates for potential volcanic eruptions at the
21 potential repository site at Yucca Mountain.

22 Next slide, please. After a brief
23 introduction, it includes a little bit of regulatory
24 basis. I would like to talk about some of the
25 uncertainties that we have in very basic probability

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1 estimates and also make sure that we all have a common
2 framework or common definition for the remainder of
3 the talk.

4 I will then focus on some of our current
5 views on the spatial and temporal uncertainties that
6 affect probability models in the Yucca Mountain region
7 and see how those uncertainties can affect the NRC
8 probability estimate, and of course, wrap it up with
9 the conclusions.

10 Next slide, please. I guess that's my
11 soundtrack. That's fine. What we are going to call
12 upon to evaluate the probability models and licensing,
13 you have to keep in mind that these probability models
14 -- performance assessment.

15 And so, requirements for review under 10
16 CFR 63.114 are going to apply. In particular, the
17 models for probability need to include actual
18 geological and engineering data, account for data
19 variabilities and uncertainties, consider the effects
20 of alternative conceptual models, evaluate events with
21 likelihoods greater than one in ten thousand in ten
22 thousand years, include events that significantly
23 affect risk calculations, and also be supported by
24 objective comparisons.

25 So, we have to keep that in mind when we

1 start looking at the currently available information
2 around Yucca Mountain and how that information affects
3 the probability model.

4 And right now, some of the questions that
5 we're asking are how many past events have there been
6 in the Yucca Mountain region? What are these igneous
7 event locations?

8 And what are the event agents. So, I
9 don't want to call this a probability triplet, but
10 there is some parallelism on number, age, and location
11 of past igneous events.

12 And, to cut to the chase, our conclusion
13 is that, from the available information, you can have
14 multiple interpretations and large uncertainties from
15 what we currently have available for assessing
16 probability in the Yucca Mountain region.

17 Next slide, please. One of the basic
18 uncertainties that we have to address, and to begin a
19 definition for any presentation, is what makes up an
20 igneous event?

21 And, taking a figure from the Department
22 of -- the NRC's technical basis document 13 on igneous
23 activity, to illustrate what the uncertainty is in
24 finding an igneous event.

25 This figure is a geologic map showing the

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1 4 million year old basalt outside Southeast Crater
2 Flat. There's a series of numbers out there. Number
3 one, two, and three mark locations that I think, as a
4 general agreement, represent the volcanic center.

5 This is a place where we have a hole in
6 the ground where molten rock came up and material was
7 dispersed in the accessible environment. But we also
8 have points four, five and six, that may represent
9 vent locations.

10 There's just a little less certainty about
11 whether these were large vents, small vents, or vents
12 that could start the beginning phase of an eruption
13 only.

14 So, how many vents were erupting at the
15 same time? How many vents may have erupted in
16 sequence, may have represented gaps in time to be
17 counted as separate volcanic episodes?

18 There's multiple interpretations that you
19 can place just on these six features. For the
20 purposes of this talk, I'm going to keep the simplest
21 definition possible.

22 An igneous event is a volcano that has a
23 hole in the ground. And we're just going to count up
24 holes in the ground or cinder cones and call those our
25 igneous event with this presentation.

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1 We also know, in igneous events, have to
2 worry about the subsurface conditions. What's going
3 on beneath the volcanoes as well? And one of the
4 things we see out here in Crater Flat, is the
5 subsurface features, which are called intrusions,
6 extend for 50 years plus laterally away from our
7 vents, and for some unknown distance longitudinally to
8 the north and south in these vents as well.

9 So, in characterizing igneous events, we
10 not only have to find out the surface expressions, but
11 the sub-surface expression as well. And one other
12 point, when you talk about igneous event, is relevant
13 to this.

14 Do you notice how these lava flows have
15 been folded and partially eroded through time? Now,
16 if you continue the deposition process out here and
17 bury these lavas between tens or even 100 meters worth
18 of alluvium, how would you interpret igneous events
19 from this disruptive feature if all you had to go on
20 was a pattern of colors in the geomagnetic map?

21 Keep that in mind when we start looking at
22 pattern analysis in the later part of this talk. We
23 may not be seeing in the subsurface impact features.
24 We have to consider the possibility that these
25 features, like this one at Crater Flat, have been

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1 disrupted, faulted, eroded, and then buried.

2 Next slide, please. One other very
3 fundamental uncertainty or assumption in probability
4 models is, what's the extend of the igneous system
5 that we're trying to model.

6 This figure is showing in red basalts
7 that's younger than about 11 million years old. And
8 all of these parts of basalt at one time or another
9 have been used to bring various definitions of what
10 makes up the Yucca Mountain igneous system.

11 These definitions have been based on
12 associations in age, location, and chemistry. And,
13 you can't quite see it, but, the potential repository
14 site is right here on the boundary of the NTS.

15 Now, there's not correct definition of
16 what makes up Yucca Mountain igneous system. The
17 point that we have to make, though, is that a basis
18 for selecting some subset of these basalt features
19 needs to have a clear, consistent basis.

20 And that basis has to be used consistently
21 throughout the probability estimate and any resulting
22 consequence analysis based on that probability
23 estimate.

24 Next slide. So, that being said, I'm
25 going to say what we think the relevant Yucca Mountain

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1 igneous system is for the purposes of these.

2 What I'm showing in this figure is the
3 regional gravity survey that's done by the U.S.
4 Geologic Survey a few years ago. What it shows is, in
5 the hot pink colors and orange colors, are areas of
6 fairly dense crustal rock.

7 The cooler colors down in the greens,
8 yellows, and blues, represent low density crustal
9 rock. The reason we are using gravity, is this is a
10 real good regional indicator of structure.

11 What we see is this long feature through
12 here with the low density rock represents an
13 extensional basin where the crust has been pulled
14 apart and in field with low density alluvium and
15 tuffaceous rock.

16 The other rocks in high density here and
17 here haven't been as disruptive in recent time, and
18 consist of older, more crisp rock, like around Bare
19 Mountain.

20 For convenience, we're just going to refer
21 to this feature as the Amargosa Trough structural
22 basin. Now, a little bit on the west, by Bare
23 Mountain by this gravity anomaly, and by the east, by
24 what's commonly referred to as the gravity fault, and
25 extending some unknown distance up towards the old

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1 caldera complexes into the mountain.

2 And, again, within the Amargosa Trough is
3 what we think the basaltic features are that are
4 relevant to our probability estimate. And, we're
5 defining igneous events in the following analyses as
6 individual volcanoes that occur within this Amargosa
7 Trough.

8 Based on that definition, we have a
9 starting point of 24 past events in the Yucca Mountain
10 region to use in the following sensitivity analysis.

11 Okay, do not adjust the dials. This is
12 actually what the data is supposed to look like, these
13 wild colors. This is the 2000 or 1999 U.S. Geological
14 Survey -- aeromagnetic survey -- for the entire Death
15 Valley/Yucca Mountain region.

16 This is the old survey. It's not the new
17 data that the Department of Energy collected this
18 summer. These data represent the magnetic
19 characteristics of the region and of the rocks that
20 are buried and exposed at the surface in this region.

21 We've gone ahead and done a little
22 filtering on these data to enhance the basalt features
23 in the region. The important point here is, we have
24 known features, known igneous events, and surface --
25 such as Red Cone, Black Cone, Lathrop Wells, that

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1 create obvious anomalies, when you know where to look.

2 The anomalies are just these patterns in
3 the magnetic data that have characteristics
4 representative of buried basalt or strongly magnetized
5 rock.

6 But, the U.S. Geological Survey and
7 ourselves have also identified other areas that are
8 representing sub-surface, buried rock that may
9 represent a very igneous event.

10 And these interpretations are shown on the
11 figure on the right, graded by competence level. The
12 red features are ones that we have high confidence in
13 representing buried basalt.

14 The green features, for example, this --
15 from L, M, N, O and two -- we have moderate confidence
16 that these anomalies represent buried basalt.

17 And, in blue, we have low confidence but
18 can't eliminate the possibility that these anomalies
19 could represent buried basalt. So, one of the primary
20 uncertainties that we're having to evaluate right now
21 is, given these anomalies, what if they represent
22 buried basalt?

23 How would the addition of these buried
24 potential features affect our probability estimate?
25 Next slide, please. We have the aeromagnetic survey

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1 that shows us features that we can detect,

2 But, along with that data, we can also see
3 that there are some features that we know exist. But
4 we haven't found them yet. They don't create obvious
5 magnetic anomalies.

6 So, we have to consider the potential to
7 have additional features located in this region buried
8 in the subsurface that the exploration techniques have
9 been unable to detect.

10 One of the ways, and not the only way, you
11 can do it -- but, one of the ways that you can try to
12 get an estimate for potentially buried features is,
13 look at the spatial density of the volcanic fields and
14 compare it to other volcanic fields and say, well,
15 there's a long list of low, and a long list of high.

16 How could you add additional events and
17 change such spatial vents? What we see is, within
18 this Amargosa Trough volcanic system, just with our 24
19 known events, we have a spatial density of one volcano
20 every 29 square kilometers.

21 For comparison, when we look at other
22 volcanic fields in the western great basin, like the
23 Cima volcanic field in California, they have the
24 density of one volcano every four square kilometers.

25 Lunar Crater up in Nevada has one volcano

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1 every six square kilometers. The older Pancake range
2 volcanoes are about one volcano every eight square
3 kilometers.

4 And finally, the Big Pine field in the
5 valley of California has a lower density of about one
6 volcano every 16 square kilometers. From the water
7 well drilling out there, we're pretty sure there is
8 some additional hidden events out in the Big Pine
9 field as well.

10 So, we can see that the spatial density of
11 volcanic features in the Amargosa Trough are very
12 pretty low compared to other similar volcanic fields
13 in the western Great Basin.

14 The exploration technique, the
15 aeromagnetic technique that has been used, we have
16 fairly high confidence that the survey has been able
17 to technique buried igneous features in the southern
18 half of the Amargosa Trough.

19 The reason for that is the basement in
20 this area is magnetically very quiet. So, strongly
21 magnetized rock like basalt, will really stand out on
22 aeromagnetic surveys.

23 So, we're not concerned about undetected
24 significant features at this stage in the Amargosa
25 Trough in the southern part. But we have these two

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1 areas throughout Jackass Flat and Crater Flat where
2 the magnetic basement is very noisy.

3 And that noise may be masking additional
4 features in the subsurface. Right now -- we have a
5 volcanic density of one volcano every 13 square
6 kilometers.

7 Just for comparison, if you wanted to get
8 that spatial density up to something comparable to
9 Cima -- the most dense volcanic field in this analysis
10 -- you're going to have 26 buried events in order to
11 get that high of a stageable density.

12 That's just a major comparison -- not that
13 we think you have to have any volcanoes out there.
14 Also, at Jackass Flat, you've got one volcano every
15 160 square kilometers.

16 Now, it is entirely possible that that is
17 the actual spatial density within Jackass Flat and
18 that there are no buried, undetected features in
19 Jackass Flat or Crater Flat.

20 But, right now we can't eliminate that
21 hypothesis. And we have to factor in our uncertainty
22 analysis the potential for undetected events, as well
23 as the events that have been detected by current
24 exploration techniques.

25 MR. HINZE: Mike, is it possible to ask a

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1 question?

2 MR. LEE: Sure, I guess so.

3 MR. HILL: Sure Bill.

4 MR. HINZE: The one volcano per 29 square
5 kilometers seems to be key to this discussion. And,
6 it seems to me that your region of the Amargosa Trough
7 does not correspond with the complete region of the
8 Amargosa Trough that you outline in the previous
9 gravity slide.

10 Am I wrong, or right? Or what's wrong
11 here?

12 MR. HILL: It's the extent of volcanic in
13 the Amargosa Trough. Now, the Amargosa Trough, as a
14 crustal structure, extends down all the way into Death
15 Valley, and all the way up into the lunar crater area.

16 MR. HINZE: And it extends considerably
17 south. So, if the Amargosa Trough is controlling
18 this, shouldn't we be concerned with the number of
19 volcanoes per square kilometer or the volcanoes per
20 kilometer, considering the Amargosa Trough problem?

21 MR. HILL: No, I don't believe so, because
22 the Trough is a structural control on ascending magma.
23 Not everywhere in the mantle, though, we believe this
24 for the production of basalt.

25 We have many areas that are extended and

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1 lack any appreciable volcanism, not only in the
2 Amargosa Trough, but in other parts of the basin and
3 range as well.

4 So, you have to have an intersection of
5 the whole extended crust and further mantel in order
6 to get volcanism.

7 MR. HINZE: Okay, so you are arbitrarily
8 selecting the north and south --

9 MR. HILL: Not arbitrarily. I'm selecting
10 the north and south boundary that, within the last
11 billion years, defines the extent of volcanism within
12 the Amargosa Trough.

13 Until you get down to Death Valley, many
14 tens of kilometers to the south, you're not seeing
15 more volcanism. In the same way, this is butting up
16 against the Caldera Mountain -- a little south of
17 Caldera Mountain.

18 But, it's the northern extent of Solitario
19 dike complex. We're coming up very close to the
20 Caldera mountain. And I think that's defining a
21 tectonal magnetic regime that we're calling the
22 Amargosa Trough.

23 CHAIRMAN RYAN: We have one follow-up
24 question from Bruce Crowe.

25 MR. HILL: Yes, Bruce?

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1 MR. CROWE: I'm Bruce Crowe of Los Alamos
2 Lab. One question I have -- I messed around with
3 doing cone densities as well. And what I tried to do,
4 though, was divide them in age increments, because you
5 really need to look at how densities have changed
6 through time.

7 And, if you look at the forming, you know,
8 record, versus say the -- you're going to get somewhat
9 different cone densities, both in Crater Flat Amargosa
10 Trough, and in Lunar and Cima. Have you tried doing
11 that?

12 MR. HILL: To an extent. One of the
13 problems is, while we have good dating in the Yucca
14 Mountain region, these other analogs we have very
15 loose dating.

16 So, I tried to give a representation of
17 the -- and Pliocene fields. But I don't think any of
18 these fields have Pliocene database that we can go
19 into.

20 As you're well aware, we have some
21 disagreements about the relevant of the Miocene. And
22 I think that's a fair interpretation. And we believe
23 the Miocene from 11 million years, and then -- to the
24 third.

25 In other words, the past 11 million years

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1 is relevant to the probability estimate. Whereas,
2 others are saying that only the past five million
3 years of volcanic history is relevant in their
4 probability estimate.

5 So, for this talk, I'm trying to be
6 consistent with our published positions using the
7 Miocene record, and in standard volcanism -- the
8 Amargosa Trough.

9 Again, these are not the only potential
10 analogs. They are the most analogous of the Western
11 Great Basin. And they are the limits of the available
12 data for age clustering.

13 Given the uncertainty in the potentially
14 varied events where we don't know the ages of them,
15 we're trying to do more refined approach at this
16 stage, really just pushing it forward.

17 But, to get to, is that -- to emphasize
18 the main points here for the spatial uncertainty. We
19 may have no undetected events. But we can eliminate
20 the potential for undetected events.

21 We have to come up with some way to
22 quantify in a traceable methodology a way to say how
23 many could there be in this area? And, by looking at
24 a general sense of spatial density, we say that, given
25 an uncertainty of one to ten present undetected

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1 events, seeing a reasonable starting point in a
2 sensitivity analysis for evaluating whether this kind
3 of an uncertainty in undetected events is significant
4 or insignificant to the probability estimate.

5 And, just as a point of comparison, if we
6 were to have ten additional events in the Amargosa
7 Trough, we would increase the spatial density within
8 this region from 29 -- or one volcano for 29 square
9 kilometers, down to one volcano for 23 square
10 kilometers.

11 So, it's not an absurd over-estimate of
12 spatial densities for the Yucca Mountain region.

13 MR. CROWE: Just a follow-up, if I may.
14 If I understand your record, you start with the
15 premise what is there at one to ten -- present but
16 undetected events.

17 I ask the question, what's the probability
18 of it being one to ten undetected events?

19 MR. HILL: Based on the currently
20 available data, we think that -- let me back up for a
21 minute. When we had a meeting about a year ago with
22 the U.S. Geological Survey, Department of Energy and
23 others to evaluation the aeromagnetic data, we all
24 agreed that there were a number of known surface
25 features that were difficult to resolve in the

1 aeromagnetic data.

2 So, they were giving a sort of anomaly
3 patterns. They didn't know that basalt was at the
4 surface. You'd have a difficult time convincing
5 yourself that that anomaly represented surface basalt.

6 So that's the first point of why I think
7 there may be contentions there. The second is, out at
8 this location, at Jackass Flat, according to early
9 warning wells, we encountered basalt at about 1,300
10 feet below the surface.

11 That basalt is in an area that has no
12 obvious magnetic anomaly. And that depth that
13 encountered basalt is likely deep enough to attenuate
14 any magnetic character of a buried well.

15 So, we have known features that don't give
16 us a clear anomaly in both the surface expression and
17 in the sub-surface expression. So I believe it is
18 reasonable to assume that there could be additional
19 undetected features here, based on the limits of the
20 current exploration technique to detect known igneous
21 features in the region.

22 I cannot give you a probability estimate
23 though. I think that's so speculative on top of a
24 speculation, on top of a hypothesis, that we really
25 can't gain much knowledge that way.

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1 CHAIRMAN RYAN: And therein is my problem
2 in that, you know, we are in a way scuba-diving around
3 in open. We really don't know where the bubbles are
4 and where uncertainties are at this point.

5 Perhaps the new aeromagnetic data helps
6 us, you know, resolve some of that. But that, to me,
7 is kind of a critical issue, because, without knowing
8 where you are in the probabilities space of all those
9 potentials, you can run into not really knowing how to
10 interpret what the hypotheses are.

11 MR. HILL: We don't need a probability to
12 evaluation the significance of alternative conceptual
13 models.

14 CHAIRMAN RYAN: But you do need the
15 probability to know which one is real.

16 MR. HILL: Conversely, you can start with
17 a reasonable range of uncertainty, let's say one to
18 ten undetected events. Let's analyze that in the
19 models and see whether it is significant.

20 And it may a lot easier to gain a
21 reviewable consensus that says, we think if there are
22 undetected volcanoes, there's less than ten of them,
23 or less than five of them in the region.

24 When we can all agree to that to develop
25 a basis for it, rather than trying to come up with

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1 probability distribution function that is going to be
2 -- by it.

3 CHAIRMAN RYAN: I appreciate the fact that
4 you are trying to -- that's different than being risk
5 informed. So, I just want to clarify these two
6 different lines of thought.

7 MR. MARSH: I think this is one way to
8 kind of justify adding events in. But, it might be
9 actually a more illustrative calculation. You just
10 started with the probability basis itself and just
11 kept adding until we became alarmed.

12 In other words, we may have to add 5,000
13 to actually make it. So, we're basically wasting each
14 other's time down at this level. And that also
15 answers Mike's question a little bit, in that it puts
16 uncertainty on this in terms of saying how much
17 seriousness do we have to put into actually adding and
18 comparing to these other fields up there that are
19 basically very homogenous in age fields that we can
20 interpret very simply and whether this field here, as
21 Dr. Hill mentioned, is.

22 We're looking at stat data over time, and
23 so forth. It might even -- to the chase. Just look
24 at the numbers, add them in directly, justify them
25 later, worry about it after.

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1 MR. HILL: The talk is -- we're around the
2 first corner, and we're cutting to the chase. I think
3 I'll address your comment in a few slides here.

4 Let's go on. Okay, and this is just a
5 very quick summary of our view of current spatial
6 uncertainties in probability. The addition of the 11
7 anomalies that are well recognized by ourselves and
8 the U.S. Geological Survey would increase --

9 CHAIRMAN RYAN: Shut the microphone off,
10 I think we can hear you.

11 MR. HILL: Well, will that affect the
12 recording.

13 COURT REPORTER: I have a back-up here.

14 MR. HILL: Okay. Can everybody hear me
15 now, without the feedback? Excellent. Okay, we're
16 looking at, with the addition of the magnetic
17 anomalies that we have high to moderate confidence in
18 that increases the spatial recurrence rates for about
19 one volcano for 40 square kilometers, to one volcano
20 for 29 square kilometers.

21 Again, a comparison with the volcanic
22 fields, the point that we had made before about the
23 limited resolution of known features, the accounting
24 of basalt in 23E is the basis for suggesting that
25 there could be additional undetected events.

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1 And, finally, for spatial uncertainty, if
2 we add one to ten additional events, the spatial
3 recurrence rate would increase from 29 -- excuse me,
4 one volcano for 29 and one volcano for 23 square
5 kilometers, a very modest increase in spatial density.

6 So, let's move on from spatial, and go on
7 to the temporal uncertainty. Now, again, there's no
8 correct definition of an igneous event. And I know
9 there are people in the audience that have alternative
10 definitions of what constitutes an igneous event.

11 But, at the outset I said, where is an
12 igneous event definition that's each individual vent
13 is an igneous event, a cinder cone event, very simple.

14 What we've done is plotted out the number
15 of cinder cones, and cinder cone remnants that we have
16 in the region against their ages. The points that are
17 in gray are the ones that -- just to be honest -- are
18 altitude interpretation that sometimes lump them all
19 together as a single event.

20 But, again, to be consistent, these are
21 the 24 individual events that we are using for the
22 purposes of this talk. And these are the basic data
23 that we have for when have past igneous events
24 occurred in the Yucca Mountain region for the past 11
25 to 11.4 million years.

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1 Next slide, please. So, our base case, if
2 you take a longer 11 million year average, we have the
3 24 events for 11 million years, to give you a temporal
4 recurrence rate of two volcanoes every million years.

5 But now, we somehow have to address what's
6 the age of these magnetic anomalies? We have higher
7 confidence in the buried basalt. We don't have any
8 dates on these anomalies.

9 So, we have to look at alternative
10 hypotheses on what these dates could be, based on our
11 interpretations of past patterns of activity in the
12 Yucca Mountain region.

13 So, let's just say in the first hypothesis
14 that these anomalies represent basalt that have ages
15 that are randomly distributed between two million
16 years and 11 million years.

17 You don't think, by the way, that any of
18 these anomalies are younger than two million years
19 old. They are too far below the subsurface to be two
20 million year old or younger basaltic features.

21 But, if we just say that they represent
22 randomly aged events, we would add in up to 35
23 volcanoes, 11 million years, temporal recurrence rate
24 goes from two volcanoes per million years, up to three
25 volcanoes per million years, not a really large

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1 increase in temporal recurrence rate.

2 Next slide, please. We also have to
3 consider that maybe these are related to a younger
4 episode of volcanism, something that's no younger than
5 five million years old and has nothing to do with the
6 past five to 11 million years.

7 So, if we just look at the available data,
8 we have 19 events in the past 5 million years of
9 temporal recurrence of four volcanoes per million
10 years.

11 Add in the 11 anomalies, and again, assume
12 that they are randomly distributed ages between two
13 million and five million years and you end up with a
14 recurrence rate of six volcanoes per million years
15 that you could use in a sensitivity analysis.

16 One of the things that you may have
17 noticed in the basic data is that the past events are
18 not uniformly distributed in time. They tend to form
19 temporal clusters.

20 Some of these clusters aren't very
21 intense, maybe three events in a couple of million
22 years. But, some of these clusters a little bit more
23 intense than that.

24 Next slide, please. And here are what
25 we're seeing, is that we have this one temporal

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1 cluster of about four million years ago where we know
2 by these definitions we have 13 volcanoes in about
3 point six million years.

4 That would give us a recurrence rate in
5 that period of time on an order of 20 volcanoes per
6 million years. Obviously that recurrence rate didn't
7 occur for a long period of time.

8 But for a geologically short interval of
9 time, roughly a half million years, there was an
10 elevated volcanic occurrence rate in that interval.
11 So, these anomalies also could represent part of that
12 pulse of past activity, four million years.

13 If they were related to that period of
14 activity, we would see the recurrence rate for a small
15 interval -- say half million years in time -- come up
16 to a rate of about 40 volcanoes per million years for
17 a short duration.

18 Next slide, please. So, there's three
19 altitude hypotheses you can use to evaluate the
20 temporal uncertainty represented by these magnetic
21 anomalies.

22 Now, depending on the time interval used,
23 these hypotheses of the age uncertainties, you have
24 about one and a half of the factor two increases in
25 temporal recurrence rate.

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1 So, we have to consider the possibility
2 that our temporal recurrence may have range from two
3 or three volcanoes per million years if you make a
4 long-term, uniform recurrence rate assumption.

5 Also, clusters of activity -- that they
6 have been as high as on the order of 40 volcanoes per
7 million years. Now, with that, we can use analog
8 volcanic fields to gain a sense of perspective for
9 what those recurrence rates mean to volcanic fields in
10 the western U.S..

11 And you can see, for Quaternary fields,
12 and again, I'm restricted to the last two million
13 years of data, because those are the only intervals
14 that have good dating in these analog volcanic fields.

15 But, with the available information in
16 Cima, your recurrence rates are 26 volcanoes per
17 million years per a period of a billion, billion and
18 a half years.

19 That would be 22 volcanoes per million
20 years. And, up at Lunar Crater, it can get as high as
21 50 volcanoes per million years. So, you can see the
22 upper bound on the range of recurrence rates that we
23 would consider in sensitivity analysis for Yucca
24 Mountain.

25 That upper bound doesn't exceed known

1 recurrence rates in the Western Great Basin. And, the
2 lower bound also would be representative of a much
3 longer lived volcanic field at the time.

4 The question really that we have to answer
5 is what is the appropriate recurrence rate for the
6 next 10,000, 100,000 or million years, not what is the
7 absolute recurrence rate to some arbitrary period of
8 time in the geologic past.

9 You have to forecast the future. And we
10 believe that we have to evaluate multiple hypotheses
11 in that evaluation of probability and not focus on a
12 single interval of time in the past.

13 Next slide, please. How we are doing the
14 sensitivity analyses. This is a familiar figure for
15 many people. This is the published NRC probability
16 model that uses clustered event locations and uniform
17 temporal recurrence rates to calculate the probability
18 estimates.

19 What we're seeing in this figure is the
20 spatial recurrence rate based on the clustering
21 algorithms that we used, normalized to the gravity
22 outline of the Amargosa Trough.

23 And, again, for our probability estimate,
24 we believe that the controlling structure that
25 localizes magma in the region is that crustal

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1 extension zone in the Amargosa Trough.

2 So, in our models it would make no sense
3 to say that there's a probability of volcanism in Bare
4 Mountain, even though the statistical clustering might
5 say that there is some likelihood in that area.

6 We believe we should pool that geologic
7 information and normalize that controlling geologic
8 structure, the Amargosa Trough, rather than allow for
9 volcanism to occur -- to the incredible places.

10 We agree that the structural weighting
11 that we use is subjective. But, it does account for
12 the available data and does provide a transparent
13 basis for review of that analysis.

14 The other good thing about the models
15 we're using is we can accommodate the spatial and most
16 of the temporal uncertainties that we're seeing in the
17 currently available information.

18 We can evaluate the significance of those
19 uncertainties using the probability analysis. Next
20 slide, please. What we're using to evaluate the
21 uncertainty is a tool called PVHA_YM, which is a
22 series of JAVA applets that -- on anybody's web
23 browser.

24 This is readily available from the Nuclear
25 Regulatory Commission on all basic data sets. This is

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1 a screen snapshot. We put in the area of extent
2 assumptions about recurrence rate and clustering
3 functions.

4 The figures that I'm going to show on the
5 next couple of slides come from screen snapshots from
6 PVHA_YM. Next slide, please. Here is our basic
7 example for the purposes of this talk.

8 Again, I'm just trying to give you a sense
9 about how we can go about uncertainty analysis based
10 on the current uncertainties in the age, location, and
11 number of features in the Yucca Mountain region.

12 This isn't mean to be our position on what
13 probability is or is not at Yucca Mountain. So, for
14 this base example, I'm taking the 24 events that we
15 previously defined, given a long-term average
16 occurrence rate of two volcanoes per million years,
17 and a simple Epanechnikov kernel that uses gravity
18 weighting at a 90th percentile.

19 So, we're re-normalizing gravity by 90
20 percent, allowing a little bit of slop around the
21 margins of the gravity anomalies and a simple
22 clustering algorithm is the plain English way of
23 looking at that.

24 So you can see, by those basic
25 assumptions, we have come up with a probability of a

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1 cone or a volcanic event intersecting the current
2 compository footprint of essentially one times ten to
3 the minus eight per year.

4 So, this is our starting point for this
5 talk. Next slide, please. What do we do if we add in
6 the 11 high to medium confidence magnetic anomalies,
7 which are shown as additional black dots?

8 And so, I can explain that the black dots
9 represent vent locations in the Yucca Mountain region.
10 So, we add in the 11 high to medium confidence
11 magnetic anomalies.

12 And let's just look at the mid-point of
13 the uncertainty, when we are going between two and 40
14 volcanoes per million years. For illustration
15 purposes, let's say the recurrence rate with those
16 anomalies is not 20 volcanoes per million years.

17 You can see our base probability would
18 increase from ten to the minus eighth, to one times
19 ten to the minus seventh per year for those
20 assumptions.

21 We can also use PVHA_YM to calculate the
22 probability of a subsurface intrusion intersecting the
23 potential repository. Given these assumptions for
24 guidelines that vary between one and ten kilometers
25 long, that probability of subsurface intersection

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1 would be on the order of seven times ten to the minus
2 seventh per year.

3 If dikes were shorter, it would be down
4 around four times ten to the minus seventh per year,
5 given these 11 magnetic anomalies represented here.

6 Next slide, please. Now, we have some
7 questions about present undetected volcanoes and how
8 significant could that be. What I've done in this
9 example -- and believe me, there are many examples you
10 can run with this -- I have added five randomly
11 located volcanoes in Jackass Flat.

12 Hit the spacebar please. There should be
13 a pop-up. There we go. Five anomalies in Jackass
14 Flat. This is randomly located to try to look at
15 sensitivity for undetected events east of the
16 potential repository site.

17 And you can see that, if we have the same
18 recurrence rate -- 20 volcanoes per million years --
19 our probability only increases from one times ten to
20 the minus seventh, to two times ten to the minus
21 seventh by adding these five locations into the
22 dataset.

23 And, again, a similar increase would occur
24 by saying that these are -- we also would have igneous
25 dikes and subsurface diversions. We go from seven

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1 times ten to the minus seventh, to eight times ten to
2 the minus seventh.

3 So, here is one of many possible examples
4 that show that adding five events that have been
5 undetected -- adding those undetected events into the
6 probability dataset, it doesn't have a very large
7 effect on the probability estimate.

8 And, again, if you want to do some
9 additional analyses, you can use your own locations,
10 own number events, and see how these models are
11 sensitive or insensitive to the addition of
12 potentially undetected events.

13 Next slide, please. So, what did we learn
14 from all of this? First, kind of interestingly, the
15 addition of the anomalies into the dataset doesn't
16 really change our spatial recurrence patterns very
17 much.

18 In other words, the anomaly locations are
19 following the known event locations, and not having a
20 profound re-alignment of our spatial patterns in the
21 Yucca Mountain region.

22 More volcanoes are located toward the
23 existing locust of activity than they are distributed
24 in areas away from that known locust around
25 Southwestern Crater Flat.

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1 We also see, by running a number of these
2 simulations, that clusters of more than five
3 undetected volcanoes appear to be meeting with a
4 change our spatial recurrence rates have is
5 significant.

6 We already thought patterns that are
7 pretty well established by the existing data,
8 including the magnetic anomalies. So, to perturb
9 those patterns in a way that would grossly affect the
10 probability of potential repository site, you have to
11 create a pretty intense cluster of undetected events
12 on the east side of the potential repository site.

13 That cluster would have to have more than
14 about five volcanoes located within a couple of
15 kilometers of one another in order to create that
16 spatial recurrence based on our models.

17 Also, we are seeing that the uncertainties
18 in the temporal recurrence rate for short periods of
19 time -- and by short, I mean 10,000 to 100,000 year
20 periods -- those variations are not really captured by
21 the existing uncertainties that we have in long-term
22 recurrence rates.

23 In other words, the million year average,
24 the variations that we see in the million year
25 averages really aren't capturing the potential

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1 variations for shorter intervals of time when we could
2 have higher recurrence rates than the long-term
3 average.

4 Finally, the cluster of past events gives
5 short-term recurrence rates that are comparable to
6 other Western Great Basin volcanic fields. Again,
7 those recurrence rates don't exist continuously
8 through time.

9 But we're not looking for 11 million years
10 in the future. We're looking for some shorter
11 interval of time in the future, time to forecast
12 what's the likelihood in that future time of volcanic
13 eruption.

14 And, finally, evaluate the large
15 uncertainty anomaly ages and anomaly locations by
16 testing alternative conceptual models and looking at
17 the sensitivity of those models to the resulting
18 probability estimate.

19 So, to wrap it up, next slide, please. In
20 looking at the current uncertainties in the number,
21 age, and location of past events in the Yucca Mountain
22 region, we have concluded that our conceptual basis
23 for the probability estimate has not been affected by
24 those uncertainties.

25 We're not seeing anomalies outside of

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1 areas that previously we had defined as the structural
2 basis for probability or clustering effects that we
3 can't account for in the current probability model.

4 We can evaluate the effects of the
5 existing spatial and temporal uncertainties on the NRC
6 probability estimate. And we had questions before
7 about what are reducible uncertainties.

8 One of the key areas for reducible
9 uncertainty is the potential for undetected events, I
10 believe is a very reducible uncertainty. And I'm
11 optimistic that the new data that are being collected
12 by the Department and the high resolution magnetic
13 survey will help to resolve that uncertainty more than
14 the current data can do.

15 Our best estimate of the effect of these
16 current uncertainties it can get a factor ten increase
17 in the NRC probability estimate relative to these base
18 models.

19 That kind of a factor on the probability
20 estimate gives us a high significance to performance
21 calculations. So, we are going to need to have a good
22 basis to review those uncertainties and a traceable
23 basis to document those uncertainties during our
24 potential license application review.

25 Finally, we also can conclude from doing

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1 this work, that the potential effects of current
2 uncertainties on the number, age, and location of past
3 events really can affect some of the assumptions in
4 the conceptual basis used in many probability models,
5 the key interpretations of past spatial and temporal
6 patterns.

7 And finally, these uncertainties can also
8 directly affect parameter ranges used in any
9 probability model for the Yucca Mountain region.
10 Thank you for your attention.

11 CHAIRMAN RYAN: Thank you. I guess we'll
12 start with any questions from the members. Allen?

13 MEMBER CROFF: In going back into this, I
14 look at your slide 15, which shows, I think, your
15 basic probability contours. I think the high being to
16 use all the exponents above 18 or 20.

17 And the Yucca Mountain site being -- I'll
18 call it eight roughly. But then, when I go back and
19 look at the diagram say, on page seven, which shows
20 the magnetic anomalies, it shows, to me, sort of a
21 clustering of these anomalies in certain areas.

22 And in other areas, such as the bedrock,
23 where the Yucca Mountain site are, and other areas of
24 bedrock, essentially zero recurrences over all time.

25 Whereas, the probability model you end up

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1 with has about a factor two probability difference.
2 And that intuitively doesn't seem right to me. Am I
3 missing something.

4 MR. HILL: There are a few points that I
5 can clarify for you. First, there is an event located
6 about 200 meters from the northwestern edge of the
7 repository site.

8 That's our roughly ten million year old
9 basaltic canyon dike and eroded vent complex. It's a
10 very small feature, but a very significant feature.

11 So, given these past events, like the
12 models have consistently said, the highest likelihood
13 for the next event would be in that southern part of
14 Crater Flat, not in that potential repository site.

15 But, through time, there has been an event
16 coming very close to that location. And, that would
17 scale as about the order of magnitude reduction in
18 recurrence rate given the number of events that we
19 have -- 20 events, 30 events, one out of 30, as
20 opposed to the two orders of magnitude or continues.

21 Second, the probability map isn't really
22 a probability map. The contour lines are spatial
23 recurrence rate. And then you have to multiply
24 spatial recurrence rate by the chemical recurrence
25 rate by the area of intersection, which is about five

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1 square kilometers for the current.

2 So, to translate in figure 15, those
3 contour lines in the probability, you have to define
4 probability in what area. We're using, of course, the
5 five square -- in this case, the seven square
6 kilometer repository footprint.

7 So, to calculate the probability, you have
8 to average some spatial recurrence over that interval
9 times the temporal recurrence, times the area.

10 So, these contours are volcanoes per
11 square kilometer using that specific kernel function.

12 MEMBER CROFF: Okay, so what is
13 approximately the difference in the probability of a
14 volcanic event in your base case, between the peak in
15 the middle of the valley, and the Yucca Mountain site?

16 MR. HILL: It would be about -- if we were
17 saying ten to the minus seventh at the potential
18 repository site, it would be approximately ten to the
19 minus sixth at the center of the locust of activity in
20 Crater Flat.

21 And it would be about ten to the minus
22 eight when you get to the edge of the Amargosa Trough
23 out there just at the western edge of Jackass Flat.

24 MEMBER CROFF: Okay, thank you.

25 CHAIRMAN RYAN: And, again, that's average

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1 per year?

2 MR. HILL: Yes, probability per year.

3 CHAIRMAN RYAN: If I could follow-up just
4 quickly, you talked about the spatial aspects. I'm
5 real interested in the temporal aspects. When I look
6 at the temporal distributions as a math problem in
7 trying to predict, you know, recurrence or a look at
8 recurrence interval, can you think of any strategies
9 to address?

10 The aeromagnetic survey updates will do
11 the spatial work. But, how do you attack the
12 uncertainties in the temporal distribution?

13 MR. HILL: Well, again, it's do we
14 evaluate this as a homogenous or non-homogenous
15 process? And, in the absence of data, you just have -
16 - you hypothesize.

17 So, we can take a rigorous statistical
18 approach to evaluate what is unconstrainable in terms
19 of the age uncertainty. What we need are the data,
20 which would be the proposed drilling program that will
21 look at some of these anomalies, drill down and sample
22 whatever is causing those anomalies.

23 It may be a welded tuff that's been
24 faulted. It may be basalt. If it's basalt, we need
25 to get those data. I think that's a very

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1 straightforward process.

2 CHAIRMAN RYAN: So really, the drilling is
3 how you get at the age distribution and prove your
4 temporal --

5 MR. HILL: Right. And, if we have that
6 age information, we can factor that into the
7 uncertainty estimate. Again, this is not our view of
8 how things will be.

9 But, it's an attempt to present to the
10 committee how we can evaluate the currently available
11 uncertainties with currently available information.

12 And then, of course, as new information
13 comes in, you can use these methods to evaluate that
14 new information for the licensing process.

15 CHAIRMAN RYAN: That's coming through
16 well. And I appreciate you clarifying that again.
17 Ruth, a question?

18 MEMBER WEINER: Is there a microphone?

19 CHAIRMAN RYAN: Oh, sorry.

20 MEMBER WEINER: I think John's -- first of
21 all, I'd like to congratulate you on making the PVHA
22 model available. I did play with that, and it works
23 very nicely.

24 And I think you all ought to be commended
25 for that.

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1 MR. HILL: Our consultants Laura Connor
2 and Chuck Connor were the real --

3 MEMBER WEINER: Well, convey to them my
4 congratulations.

5 MR. HILL: I will.

6 MEMBER WEINER: I have what's probably a
7 very simplistic question about the spatial density.
8 And that is, you outlined very carefully the area that
9 you were looking at for the Crater Flat, Jackass Flat
10 volcanoes.

11 How does that area compare with the
12 comparisons where y have volcanic fields that have a
13 higher density of events, higher spatial density of
14 events?

15 MR. HILL: I think we're looking at fairly
16 comparable. I would want to check on that. But,
17 we're not comparing huge fields or microscopic fields
18 compared to the area that we're dealing with for
19 Crater Flat, Jackass Flat.

20 The entire basin, the Amargosa Trough,
21 that contains the volcano is bigger than the
22 Quaternary part of a number of these fields. But I
23 think it is comparable to area for Lunar Crater field,
24 which is a bit more extensive. Is that addressing --

25 MEMBER WEINER: It does address it. The

1 thing that is of concern that I picked up on is, if
2 you define the area differently, how differently do
3 you need to define the area to make a significant
4 difference in the spatial density?

5 MR. HILL: On all of these definitions,
6 the area is defined by the extent of mass -- not
7 connected on the margin, but pretty close to the
8 margin of that, and also accommodating the very
9 obvious structure.

10 Like Bare Mountain, we wouldn't include
11 that potential area. And the same thing in a place
12 like Lunar Crater. You're not going to expand the
13 area out into the alluvial basins just to get a bigger
14 area.

15 You define it right around where the
16 mapped volcanoes are. And so, in the scale on order
17 of magnitude, these are comparable. In detail there
18 is going to be some variation. But we're not taking
19 a comparison with a huge volcanic field to come up
20 with spatial densities.

21 MEMBER WEINER: Okay, thanks.

22 CHAIRMAN RYAN: Let me open it up for
23 questions from our panelists and participants and
24 consultants. Bruce?

25 MR. MARSH: Yes, it's a very interesting

1 presentation. One of the things I've always been
2 amazed over in the Western United States and in
3 volcanic terrains themselves is that, if you actually
4 look at the solid rock areas, where we know the
5 geology the best, you don't see much signs of
6 volcanism compared to what we see in valleys, for
7 example.

8 Our discussion today, for example, is all
9 mostly concerned about things that perhaps we don't
10 know what's going -- buried in these valleys. And, it
11 would be interesting, I think, in some ways, to adopt
12 a different view, in other words, build a probability
13 model that didn't use anything in the valleys, but
14 only used solid rock data information.

15 The repository, for example, the mountain
16 ranges are all solid rock. We know the geology there
17 well. We can see what happened there. And, if we
18 built up, for some reason, for example, there aren't
19 a lot of cinder cones up in the mountains on the solid
20 rock areas where you see the geology very well.

21 It would be very interesting, as an
22 alternative to build a probability model using only
23 the areas of solid rock in the mountain and say, okay,
24 we know we can see the dike, maybe a cinder cone, and
25 build up a model like that, and then use that for the

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1 whole region.

2 In effect, now we are doing the reverse.
3 We are actually taking all the stuff in the valleys,
4 the alluvial fill and things we don't know, and we're
5 putting a model forward that we're pervasively using
6 in the regions where we have the best geologic
7 control.

8 And it's odd that, in many ways, you know,
9 volcanoes just don't seem to appear ever in some
10 areas, regardless of what's going on nearby. And, so,
11 have you thought of this in trying to build a model
12 like this?

13 MR. HILL: We thought about this a lot.
14 And, while maybe true in some areas, we see in other
15 areas the fact that volcanoes do erupt, which are
16 characterizing as solid rock.

17 It depends very much on what are the
18 controlling structures in the region, and what are the
19 areas of local extension, versus local compression, to
20 put it very simply.

21 In places like the Big Pine field, you see
22 them coming up the range of the Sierra. Some of them
23 are in the valley, and some are buried in the valley.

24 But other volcanoes come up and are
25 essentially sitting there in the foothills of the

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1 Sierra Nevada. In the Yucca Mountain region, we see
2 not only Solitario Canyon Dike, but also up around
3 Thirsty Mountain we see the hidden cone sitting all on
4 a bedrock bottom.

5 You know, there are plenty of alluvial
6 basins sitting around there. The reason it's all that
7 high is structural control, not anything to do with
8 whether the bedrock is above surface or below the
9 surface of baseline alluvial.

10 One of the reasons -- well, I'll back up
11 for a minute. The existing pattern of volcanism
12 already reflects that control. We have no basis to
13 say that Yucca Mountain is somehow a zone that magma
14 physically cannot get into.

15 The current patterns show that, while it's
16 less likely for it to go there, it still can go there.

17 MR. MARSH: Well --

18 MR. HILL: The greatest likelihood is down
19 where we are seeing the most volcanoes. But, at a
20 process level, the controlling structure is not
21 whether a couple hundred meters of bedrock sticks up
22 above the alluvial or is below the subsurface.

23 It depends on those structural elements
24 that are important for mobilizing the magma and
25 allowing breakout at certain points.

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1 MR. MARSH: Well, I mean, that sounds
2 interesting. But, in fact, it is the numbers and the
3 effects in the model that we really need to put in.

4 For example, we know that the regional
5 stress fields direct the localization or the local
6 dispersal of magma. So, when a cinder cone is
7 erupting, for example, these -- what we see --
8 reinforced yesterday, for example, is that there's an
9 extreme north-south predilection for the magma being
10 dispersed.

11 So, one of the things that missing, I find
12 in this probability model, is the detailed local
13 characteristics of the structure that you're
14 mentioning.

15 Structural integrity is expressed on a
16 local basis, let's say on an area that involves, let's
17 say, you know, 10,000 square kilometers, 5,000 square
18 kilometer area.

19 That detail, that granularity in the model
20 where you need to put those details in this regional
21 stress field and how that influences it, is extremely
22 important.

23 Instead of having a very dispersed line
24 sampling kernel like this, it spreads as an umbrella
25 over the whole area. It doesn't have any granularity

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1 in it for the integrity.

2 For example, in a big earthquake we know
3 we built buildings on areas that are alluvial areas
4 that may undergo -- basically quicksand. How do you
5 stabilize a building?

6 You build a big sub-structure on it. You
7 put basically a boat in the earth's crust there. And
8 this building will sit there and sway back and forth
9 and be perfectly fine.

10 If you don't know anything about that
11 granularity and detail of structure, you would predict
12 that everything would just collapse into the earth
13 when, in fact, it actually has this integrity built
14 into it to make it survive.

15 I'm worried that we're looking at detailed
16 numbers. And these numbers are so uniformly spread as
17 kind of a wide umbrella here that we're missing very
18 important granularity in this.

19 And, as you're mentioning, there are areas
20 where we cinder cone things spread up on sides of --
21 in the Sierra's, for example. We see it in
22 Antarctica.

23 We see other places. But, we don't see it
24 here. And that's something that's special to this
25 area. And I'd like to see that somehow evaluated or

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1 built into the model, because we do have a lot of
2 variations that's due to the stress fields, for
3 example.

4 But, that is a particular characteristic
5 here that doesn't seem to be in the model.

6 MR. HILL: Do you think that we have the
7 science that would allow us to make that sort of
8 deterministic approach that certain areas are
9 structurally or mechanically facilitating?

10 MR. MARSH: Absolutely.

11 MR. HILL: What do you think those would
12 be.

13 MR. MARSH: I mean, we worry about it all
14 the time. We can see things even using the models
15 that -- developed. For example, for years and years,
16 looking at stress fields around volcanoes and knowing
17 where the dispersal is going to be.

18 MR. HILL: But you're talking about around
19 a volcano, you know, gross perturbations in the local
20 and regional stress field. Here, at Yucca Mountain,
21 we're talking about first characterizing the stress
22 field in the alluvial subsurface, which would be a
23 very challenging thing to do with the available
24 information.

25 Second, the pattern -- as best we can tell

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1 -- really doesn't change from in the alluvial part of
2 the basin, out into Yucca Mountain. Those variations
3 are continual.

4 We're trying to, in the first order,
5 represent a continual variation in deviatoric stress.
6 So, what I'm really getting at is, in some areas I
7 agree, that there are profound changes in the local
8 stress field that could be used.

9 In this particular condition, though,
10 we're not dealing with huge or large variations of
11 deviatoric stress. They are very subtle.

12 MR. MARSH: Well, let me get down to
13 actual some detail here. For example, this area is
14 heavily fractured in the north-south direction. So,
15 if a magma is coming out, since there are so much
16 availability to run in north-south direction, the
17 probability, for example, if we were to look at the
18 propagation of dikes, the probability is very large
19 that it would go in a north-south direction, rather
20 than in east-west Director, for example.

21 MR. HILL: Yes.

22 MR. MARSH: So, that should be built in in
23 great detail. In other words, these cones, for
24 example, in terms of setting off a dike that would be
25 off east-west in any of these would be a very low

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1 probability event relative to a north-south event.

2 MR. HILL: When we talk about the dikes in
3 a probability estimate, we have a variation from zero
4 to 20 degrees that reflects both the regional
5 structure as well as the deviatoric stress in the
6 region.

7 We're not considering the probability of
8 east-west dike, because it just wouldn't occur in
9 this.

10 MR. MARSH: Right. I know. But those
11 ripples do not appear here. For example, if we have an
12 eruption at one of these centers, when we look at in
13 detail with exactly the same space, we should be able
14 to predict in great detail, in terms of the volumes
15 involved.

16 And that would also give us some limit on
17 the dikes, but also where the dikes are going to go.
18 We should then have a much different basic umbrella
19 probabilities than we see here.

20 MR. HILL: Well, this is the spatial
21 recurrence pattern --

22 MR. MARSH: Right.

23 MR. HILL: -- for the volcanic event.
24 It's treated as a point source, not a line source. So
25 this model was not intended to try to represent the

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1 distribution function of a linear event.

2 It's a point. That's the whole thing with
3 this probability. I can't really extemporize on how
4 you go about making a linear event probability model
5 in that sense and then look at the variations in three
6 dimensions of the regional stress field and pull that
7 in to a normal function.

8 But, my personal review is that we don't
9 have the data or ability to resolve this on scale or
10 kilometer at the Yucca Mountain region to say that, if
11 we move over one kilometer, that we can grossly
12 characterize this as favorable for magnetism or
13 unfavorable for magnetism.

14 All we're seeing is what's right there at
15 the surface. And the magma isn't coming up from a
16 very shallow inter-volcanic magma field. It's coming
17 up from depth that is controlled by the regional
18 structure as well as the local structures in the near
19 surface.

20 CHAIRMAN RYAN: John, you had a comment?

21 MR. TRAPP: Yes, this is John Trapp. Is
22 this on. This is John Trapp of the NRC. I would just
23 like to make a couple comments. Number one, it is not
24 the NRC's job in licensing to provide the probability
25 model.

1 It is our job to evaluate the probability
2 models that are presented by the Department of Energy.
3 Second point, if you want a north-south model, there
4 already is one.

5 Smith, from the State of Nevada, had
6 published models in which they have basically taken a
7 north-south structure and used this and compared the
8 results of their models with the models in another
9 direction.

10 You will find a tremendous -- up to a
11 couple orders of magnitude in the results of the
12 probability. Third, if you want to go into the detail
13 that you are talking to, again, this would be
14 something that should be directed at the Department of
15 Energy as far as characterization studies that should
16 be taken.

17 There are a lot of things that would be
18 nice to put in a deterministic patter. But we just
19 don't have them.

20 CHAIRMAN RYAN: Other questions. John
21 Garrick had a question.

22 MR. GARRICK: I wanted to talk a little
23 bit about the probability calculation itself. One of
24 the great difficulties is getting our arms around the
25 issues of uncertainty and the issues of igneous event

1 scenarios and thresholds of concern.

2 On the uncertainty issue, Rick, you
3 articulated very well a number of scenarios that
4 resulted in different volcanic frequencies. As one
5 way of getting some additional insight into the
6 uncertainties involved, have you, in getting to your
7 bottom line probability numbers, have you embedded
8 those frequencies in probability distributions to
9 reveal how the uncertainty varies with respect to the
10 different scenarios that you presented?

11 In other words, it lends itself very
12 nicely to doing that. And, you developed some
13 probability frequency curves that would really give
14 some illumination and insight as to the uncertainties
15 for the different categories of events that you
16 described.

17 MR. HILL: Right. I agree we could. We
18 have not done that. We could do that to look at a
19 distribution given these parameter ranges, parameter
20 uncertainties, what would be the resulting effect.

21 Essentially, the same as I presented a
22 very deterministic sense here, you can do a more
23 stochastic analysis that would give the full range.
24 And it could then be integrated into another
25 distribution. We have not done that.

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1 MR. GARRICK: I just see it as another
2 opportunity to get additional insight into where the
3 uncertainties are as a function of the igneous event
4 scenarios that you should be worrying about.

5 MR. HILL: I want to make sure I
6 understand. When you talk about the igneous event
7 scenarios, are you talking about a different
8 consequence scenarios or evaluating different
9 probabilities for different --

10 MR. GARRICK: Well, yes. I have trouble
11 separating the probability calculation from the
12 consequence calculation. And, when I think scenario,
13 I think from initial condition to the consequence.

14 And, with the underlying assumptions
15 associated with the consequences because there's some
16 certainty associated with them as being part of the
17 makeup of the probability.

18 So, my -- is different than the way it has
19 been presented.

20 MR. HILL: Right. But, in a very simple
21 sense, what we're looking at is the igneous event is
22 the initiating event in the event sequence.

23 MR. GARRICK: Yes.

24 MR. HILL: And then we have two branches
25 in the event, one for volcanic disruption, one for

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1 intrusive disruption. But the probability of both of
2 those branches still comes back to a singular issue.

3 We don't have a discrete probability for
4 those the way we are treating the performance
5 assessment. And, within the sub-branches of volcanic
6 or intrusive, we don't have the data to begin to say
7 that we have a probability distribution for this class
8 of initiating event gives us this sub-class on a
9 volcanic event consequence.

10 MR. GARRICK: But I suspect you have some
11 sort of evidence that would allow a certain level of
12 discrimination between your supporting evidence for
13 these different frequencies.

14 And that might turn out to be very
15 important to characterizing the overall uncertainties
16 of the probability. That's just a thought.

17 MR. HILL: Yes. It's certainly something
18 that we've thought about from day one of the program,
19 because this does appear different from how you would
20 do a seismic hazard analysis where you have a large
21 variation in the magnitude of the initiating event.

22 A large volume data to characterize the
23 frequency -- have this large range of initiating
24 events. And a hazard that is directly related to the
25 magnitude of the initiating event.

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1 But this comes back to the point that John
2 Trapp was talking about earlier. We're not looking at
3 a range of initiating events like you're doing in
4 seismic where you go off the magnitude seven and half
5 down to maybe magnitude three.

6 Our hazard and earthquake space would
7 pretty much be about a magnitude four. So, we're not
8 sampling that entire magnitude range using this
9 analogy.

10 Our initiating event is restricted to a
11 kind of earthquake analog that would be only about a
12 magnitude of four. So, we don't have to consider
13 large changes in the hazard because the initiating
14 event has a small range in consequential hazard,
15 unlike the earthquake scenario.

16 But, it's again something that we continue
17 to look at. We hold in a lot of the variability
18 within that narrow initiating event. We still have
19 variations in eruption size, eruption duration,
20 etcetera that reflect a lot of the uncertainty in the
21 event.

22 But, we're not using a strict probability
23 linkage between the larger volume range having one
24 probability to the smaller volume range having another
25 probability, for example. The data just don't support

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1 that.

2 MR. GARRICK: Thank you.

3 CHAIRMAN RYAN: We'll go to George
4 Hornberger and Bill Hinze.

5 MR. HORNBERGER: I'd like to perhaps
6 approach Bruce Marsh's first question from a slightly
7 different approach angle. So, Rick, I think if I can
8 loosely summarize here, you said that you're base case
9 temporal recurrence rate at Yucca Mountain would be
10 something like ten to the minus seventh or eighth.

11 And then if you add in the potentially
12 hidden features, it goes up to ten to the minus
13 seventh. And then your sensitivity study said it
14 could increase another order of magnitude.

15 Okay, now, my question is, with any of
16 those estimates of temporal recurrence, and now if we
17 restrict our knowledge to the hard-rock geologic
18 features where we have the best information, are any
19 or all of those temporal recurrence rates consistent
20 with the observed features within the hard-rock
21 portion of the area that you're considering?

22 MR. HILL: By hard-rock you mean the
23 surface exposures at Yucca Mountain? I'm not quite
24 sure what you mean?

25 MR. HORNBERGER: Again, Bruce was

1 suggesting -- his question, why not build a model just
2 based on features not in the valley. Now, my question
3 is, turn it around. You've estimated frequencies at
4 Yucca Mountain.

5 And, presumably, we can take Yucca
6 Mountain to be not in the valley. And so, we have
7 observation throughout the region of dikes, like you
8 said, the Solitario Canyon being the closest to Yucca
9 Mountain.

10 We can count up the number of observations
11 we have that are not in the valley. Do the number of
12 observations we have over 11 million years, are they
13 consistent with your estimate that is ten to the minus
14 seven or with your estimate that is ten to the minus
15 eight, or your estimate that's ten to the minus six?

16 MR. HILL: I think, if I understand --
17 first, I would just want to go on the record as saying
18 that I don't believe that there is a controlling
19 difference between a couple hundred meters of bedrock
20 versus bedrock being a couple hundred meters below
21 alluvium that changes -- head just isn't significant
22 in the sense of magma.

23 But, to answer your question very
24 directly, if I was to outline the bedrock exposure of
25 Yucca Mountain and say, how many events have occurred

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1 there in the past 11 million years?

2 Again, I want absolute certainty of if you
3 include Lathrop Wells in that dataset or not. But,
4 let's ignore Lathrop Wells. We just have one in the
5 past 11 million years.

6 That would be the Solitario Canyon Dike.
7 If you believe that they are discrete probability
8 issues, which I do not believe, between Yucca Mountain
9 and the adjacent part of Crater Flat and Jackass Flat
10 valleys.

11 MR. HORNBERGER: I just want to restrict
12 it to just the footprint of Yucca Mountain because we
13 have this whole area.

14 MR. HILL: Okay.

15 MR. HORNBERGER: And you have bedrock
16 exposure across the whole area. So, don't restrict to
17 Yucca Mountain. How many events do we count in your
18 database that are in the bedrock exposure.

19 MR. MARSH: Let me interject one thing
20 here. I think you do believe that exactly, Britt,
21 because you put a emphasis on the Amargosa Trough
22 region.

23 The way that's drawn and the basis of that
24 is extremely important. You're using that as a guide
25 to bring magma. If we actually exclude the mountain

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1 ranges to the east, including Yucca Mountain, of that
2 -- in other words, we did detailed gravity, and maybe
3 did the isostatic correction a little differently, the
4 Amargosa Trough would be defined in such detail that
5 the regions that we're talking about are outside of
6 it.

7 So, you actually do believe this, without
8 realizing it, because the Amargosa Trough you're
9 saying is the preferred area to go. And things that
10 are happening that -- that's a heat transfer zone.

11 So you're actually believing it without
12 realizing it.

13 MR. HILL: No, I don't believe we do. We
14 have seen many other people's interpretation,
15 including many of the U.S. Geological Survey. We have
16 looked high and low to find what it is.

17 Is there a change in structural domain
18 between the dirt in Crater Flat and the Rock in Yucca
19 Mountain. So, we don't see a crustal structure there.

20 This is just part of a continuous
21 extensional basin that's in -- with maximum extension
22 on the west and -- extension on the east. And Yucca
23 Mountain is part of that continuum of extension.

24 So, I'm not going to agree that somehow
25 there is a large or significant or controlling

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1 difference between the structural domain at Yucca
2 Mountain versus what occurs to the west, except that
3 this is an extensional basin, and you have a base
4 level alluvium that is covering part of that basin.

5 But, in terms of what controls the ascent
6 of magma, it's not the upper couple hundred meters of
7 dirt. It's that large scale structure.

8 CHAIRMAN RYAN: Bill, you had a question.

9 MR. HINZE: Well, --

10 MR. HILL: Let me go back to George.
11 Within this structural basin, we have 24 events. One
12 of those events has been within a couple of hundred
13 meters of Yucca Mountain, on that bedrock exposure.
14 So, one out of 24 in 11 million years.

15 CHAIRMAN RYAN: Thank you. Bill?

16 MR. HINZE: Why don't we have this 15th
17 illustration? If we set ourselves back 80,000 years
18 ago, I assume that this spatial temporal clustering is
19 impacted by the presence of Lathrop Wells, which you
20 pointed out there.

21 But, I suspect much greater. Have you
22 tried it out?

23 MR. HILL: Yes, we have.

24 MR. HINZE: And, how was the probability
25 changed -- the contour changed between Yucca Mountain

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1 and Lathrop Wells? In other words, I read eight times
2 ten to the minus eight at the repository and eight
3 times ten to the minus eight at Lathrop Wells.

4 If I was back there 80,000 years ago, I
5 would expect that to be eight times ten to the minus
6 eight.

7 MR. HILL: Well, certainly this is
8 something that anybody can do using the PVHA tool --
9 go in, edit the volcano dataset, go out, take Lathrop
10 Wells out of the dataset, run the model, use your
11 preferred assumptions and see.

12 This is a general guide. The addition or
13 subtraction of one event doesn't change the spatial
14 patterns significantly. So, you would see that you'd
15 have the same basic spatial pattern about, like you
16 were saying, eight per square kilometer or -- I forget
17 the exact unit -- volcano per eight square kilometer.

18 I think that was the spatial recurrence of
19 that particular point per square kilometer. And that
20 would be about the same recurrence rate -- eight of
21 ten -- that you would have before Lathrop Wells
22 existed.

23 So, in the end result, it is really
24 comparable to Lathrop Wells. It is important, Lathrop
25 Wells did form in the most intense part of the field.

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It was out there out around the eight to ten, not 18 or 15 in terms of simple recurrence, similar to what we see at the potential repository site.

CHAIRMAN RYAN: We're getting close to the end of the session, so let's --

MR. HINZE: Another brief question, or I want to make sure that we're together on nomenclature here. And I'm stepping into Bruce's space here.

The PVHA expert elicitation had this hidden event factor of 1.1 to 1.5, something like that. These are your undetected events? That's a question.

MR. HILL: I'm afraid we're getting into an area that I really can't speak to in this meeting in commenting on the Department of Energy's --

MR. HINZE: Is your definition of undetected event and hidden event factor in the PVHA the same thing?

MR. HILL: No, they are not. What we mean by undetected events is events that slight characterization has not detected in terms of volcanic features, not dike eruptions or dikes that haven't gotten to the surface.

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1 MR. HINZE: Okay, so it's not fair to
2 compare the hidden event factor effect upon the PVHA?

3 MR. HILL: Not always. Sometimes it is,
4 sometimes it isn't. It depends on whose definition.

5 MR. MARSH: I just have something brief.

6 CHAIRMAN RYAN: Yes.

7 MR. HILL: Yes please.

8 MR. MARSH: Getting back to the issue,
9 your reason on how you count events and what is event.
10 I think that's a salient issue to be worried about.

11 And, one of the things is that, an event,
12 for example, if you look at those flows out there at
13 Lathrop Wells, for example, and you've been around
14 volcanoes that are erupting, you can see that these
15 things have these big -- you know, there are small
16 lobes and there are tractor tread type things.

17 And they're kind of pushing towels ahead
18 of them. And they're moving along maybe at meters per
19 days some places, maybe meters per hour other places.

20 But if you live there -- let's say you
21 have a little hut nearby -- you'd be worried about
22 hour-to-hour. An event would be a boulder falling off
23 and rolling over your house.

24 That would be an event. So you would call
25 that an event. But, if you're actually concerned

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1 about -- you live, you know, five miles away, and
2 you're concerned about a dike coming out and hitting
3 your house or hitting your farm, that's a different
4 kind of thing to think about for an event.

5 For example, because when, as you know,
6 centers establish themselves, most of the destruction,
7 most of the dispersal of the dike's warmth is early.

8 It concentrates down to something more.
9 So, maybe we actually should think about counting
10 events in several different ways. For example, the
11 outpouring itself would be one event.

12 We would think about each one of these
13 things as a just event, no matter how many small
14 effusive cones it had near by. That would be one
15 extreme.

16 And that would be for, let's say, a
17 disruptive event -- up through Yucca Mountain. On the
18 other hand, we could have another one that sent out
19 dikes.

20 And we worry about then the radial
21 component of sampling kernel I was talking about
22 before in the stress field. And that would be a
23 different kind of event we'd talk about.

24 So, these are different ways to calculate,
25 instead of lumping them all in and saying, you know,

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1 there are going to be hidden things that are -- you
2 know, 24, 40, whatever.

3 We would actually classify these kind of
4 in a category that John was mentioning earlier, in a
5 hierarchical structure and based somewhat on outcome
6 in their potential destructiveness in terms of what
7 their potential capabilities are.

8 So, have you thought about this or tried
9 to build this up.

10 MR. HILL: I'm not sure I really
11 understand the comment, the event. First, we don't
12 have a minimum threshold event below which igneous
13 activity would not create a potential hazard if it
14 intersected.

15 Second, these are all for direct
16 disruption. In other words, the dike or the volcano
17 would have to penetrate the footprint of the
18 repository.

19 And that's the only conditional
20 probability to worry about because that's really the
21 only hazard. For the volcanic event, we're not
22 worried about the lava flow, because encapsulation of
23 a lava flow is not going to create a potential hazard
24 at a location 20 kilometers down range.

25 It's only that part of the eruption that

1 produces the dispersed tephra that truly caused the
2 hazard for the RMEI who isn't living at the volcano.
3 We're not worried about rolling rocks on the RMEI.

4 We're worrying about penetration of
5 potential repository site. So, I think a number of
6 these assumptions are already built into the basic
7 probability model.

8 The thing with event that's, I think, a
9 bit more important to you is you can also define
10 events as like the Crater Flat center. That could be
11 little cones, Black Cone, Red Cone, and Northern Cone
12 as a single event, depending on how long you want the
13 event to last, the same way Sunset Crater has multiple
14 vents and discrete hiatuses and activity.

15 The reason I chose this particular
16 definition is not because it is the correct
17 definition, but it is the simplest definition. Here
18 is a cinder cone, here is an event.

19 Here is an anomaly, here is an event. I
20 don't have to make assumptions about the nearest
21 neighbor is a part of that event in defining
22 distribution of event sizes or event areas.

23 Because, once you start say an event is a
24 series of points, then the point has an area term that
25 has to be tracked as well. Here, because the point --

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1 the vent -- is small relative to the area of interest,
2 the footprint -- five square kilometer -- it is
3 treated as a simple point-source and not worry about
4 the area itself.

5 MR. MARSH: My point is that this is a
6 pretty serious issue in that getting at one question
7 earlier -- people were saying what's your uncertainty
8 in your definition of these?

9 This is a way to get at those things. And
10 it's worth, I think, taking the time to actually look
11 at them.

12 CHAIRMAN RYAN: That's probably --

13 MR. HILL: -- ponder a paper about the
14 sensitivity of the probability estimate in event
15 definitions.

16 CHAIRMAN RYAN: Let me ask that we
17 continue the discussion after we take a break and hear
18 from the other speakers. I'm sure we'll move into the
19 details of this as the next two days go on.

20 Britt, thank you for a wonderful
21 presentation and answering all the questions. I think
22 the dialogue is wonderful. So, thank you very much.

23 We'll take a break now. It's ten o'clock.
24 We'll reconvene sharply at 10:15.

25 (Whereupon, the above-entitled matter went

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1 off the record at 9:59 a.m. and went back on the
2 record at 10:20 a.m.)

3 CHAIRMAN RYAN: We had one request for a
4 brief question from Mike Sharpton, the University of
5 Buffalo. We'll go ahead and catch that question now.

6 PARTICIPANT: Okay, thanks a lot for
7 allowing this. This question is for Britt. In your
8 reply to some of the questions from the panel, you use
9 this as an example of a volcanic event in bedrock as
10 the Solitario Canyon dike.

11 And that's ten million years old. Now, in
12 the analysis of the PVHA panel and the probabilities
13 that we've been using, we only considered volcanism
14 from four million years to the present.

15 What is the reasoning for using these
16 older rocks, because the tectonic regime was probably
17 different at ten million years from one million years.

18 MR. HILL: The simplest answer is that we
19 don't believe the tectonic regime was that much
20 different ten million years ago as to a comparable
21 current tectonic regime.

22 One of the papers by some -- reports on
23 paleomagnetic direction data for this data. It shows
24 that most of the extended and rotation that accompany
25 the end stages of -- have been accomplished by the

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1 time and in the place of a basalt, for example --
2 Canyon -- in the southern part of Crater Flat.

3 So, the tectonic regime had been set by
4 that, which is comparable to the tectonic regime that
5 we see in the present. It doesn't mean it's
6 identical.

7 But, it's near an episode of tectonic.
8 Second, the petrogenesis of those lavas -- fit the
9 Dikes are preserved, Solitario, the Miocene rocks that
10 are in the drill, the southern Crater Flat basalts.

11 The petrogenesis in the variations that
12 you see in the basalts is the same petrogenesis
13 variations that you see in the Pliocene and -- rocks
14 in Amargosa Trough.

15 They have a common petrogenesis. In
16 contrast, if you go out to places like Skull Mountain
17 and look at basalts there, you will see a very
18 different characteristic.

19 The vapors -- there's a lot of this
20 equilibrium. There's a lot of quartz, zircon -- and
21 the white elements are floating around. These are
22 giving all the signals of magma that sat in the
23 salicic crust in response to that larger tectonic
24 regime associated with the calderas.

25 So, we believe that the Miocene within the

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1 Amargosa Trough is relevant to understanding past
2 patterns of igneous activity, because the petrogenesis
3 of that basalt and the tectonics within that basin has
4 been a continuum of a similar process for the past 11
5 million years.

6 In contrast, outside the basin represents
7 a separate sort of event that doesn't give us any real
8 insight on what's happening in recent.

9 PARTICIPANT: Thank you, I appreciate it.

10 CHAIRMAN RYAN: Moving to our next
11 speaker, Dr. Bruce Crowe is here. He's going to talk
12 about the 1996 probabilistic volcanic hazard analysis,
13 one subject matter expert's perspective. Dr. Crowe,
14 welcome. Thank you.

15 MR. CROWE: I like to stand by my slides
16 and walk around with it. So, if people can hear me,
17 I would prefer talking from there.

18 CHAIRMAN RYAN: Okay, we have a pointer.

19 **1996 PROBABILISTIC VOLCANIC ANALYSIS: ONE SUBJECT**

20 **MATTER EXPERT'S PERSPECTIVE**

21 MR. CROWE: Okay, the reason I call this
22 an out-of-touch look is I left the program in '96, so
23 I want to just make clear that I have time. I reached
24 the point where I told Frank just not to talk to me.

25 So, this is defining a cobweb. So here's

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1 what I'm going to try to do. I'm going to focus on
2 how the logic and the assumptions and particularly the
3 framework geology I used to construct my PVHA model.

4 And it was fun to be in the PVHA because
5 we were allowed to say, okay, what is your best guess
6 as an expert at how you think is the best way to do
7 these calculations.

8 Mike Sheridan was involved. He can keep
9 my honest when I deviate. But, it was fun to do it
10 where we actually could inject some personal opinion
11 and some personal biases into the program.

12 So, I present that. I also put that
13 together for a book chapter that I wrote that was
14 supposed to come out two years ago. I never know when
15 it's going to come out.

16 I put together an influence diagram that
17 I tried to assemble the logic of how you do these
18 probability calculations. I'm going to step through
19 that and kind of use that as a framework for my
20 presentation.

21 And then I kept some new perspectives.
22 I've been doing a lot of probabilistic PA modeling for
23 the Environmental Management Program. And I've been
24 working with Bayesian statisticians who have educated
25 me a lot.

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1 I've really learned that the best way to
2 do is -- a geologist can do the prior, but let the
3 Bayesians do the posterior and handle all of those
4 messy curve fittings and those sort of problems.

5 And then, I have to interject some biases.
6 I'm going to talk where I think that you can put some
7 fairly logical arguments together whether there might
8 be some bounds on the probability limits for these
9 calculations.

10 And then we may be approaching the limit.
11 We're getting down. And I think it is time to move
12 on. But, again, it's a distant perspective. They
13 gave me a whole bunch of handouts.

14 And I looked at them and stole a few
15 slides from them. But, I don't profess to understand
16 everything that was in all those handouts.

17 Way back in 1978-1979 when we started on
18 this probably, when they were kind of focusing in on
19 Yucca Mountain, after they looked a number of sites at
20 the test site, I made the mistake agreeing.

21 I was told by the USGS to go look at these
22 basalt volcanoes. It would just take you a couple of
23 months and then you can move on to something more
24 interesting.

25 Here we are in 2004. But, anyway, what we

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1 always pointed out was that, rather than use risk --
2 I probably should be saying hazard -- you have a
3 fairly low hazard of disruption.

4 And ten to the minus seven and ten to the
5 minus eight numbers are low numbers. But, what we've
6 always pointed out is because you have a small number
7 of volcanoes, you're always going to have a lot of
8 uncertainty.

9 There's just no way of getting around that
10 uncertainty. You have a lot of irreducible
11 uncertainty by virtue of a limited geologic record. If
12 you had a lot more volcanoes, you would have the
13 luxury of having less uncertainty.

14 But you would have much higher risk. And
15 so, clearly you want the trade-off. But it means that
16 there are some limits to how well you can define this
17 probability.

18 And so, what we always argue at that
19 point, and I think it carries on today, is that you're
20 going to have multiple permissive models. And, in my
21 opinion, you don't have the dataset to resolve those
22 models.

23 So you really shouldn't get too caught up
24 into what is the correct model. But instead, you
25 should look at what are the impacts of a whole

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1 spectrum of models and then use that to guide your
2 intuition on the significance of the problem.

3 Okay, so here's starting out. This is
4 basic probability that I first worked out in the late
5 70's. It still kind of holds. And, basically, just
6 as the probability that for a disruption to occur --
7 the repository -- has to be an event somewhere in the
8 region or in a volcanic zone.

9 And that event has to intersect or hit
10 near the repository to be an issue. When we first put
11 this probability together, we argued that these were
12 independent events.

13 But, there are some couplings in these
14 that I'll be talking about that I think are important,
15 that affect how you assemble the probably
16 calculations.

17 So, this represents an influence diagram.
18 And I did the program -- and each box is set to show
19 the different types of variables that go into this
20 equation.

21 The square boxes represent either decision
22 uncertainty or decision assumptions that you have to
23 make in order to do the calculations. They are like
24 boundary assumptions or modeled assumptions.

25 And you really can't treat those

1 stochastically. They're basic fundamental assumptions
2 that you have made. Or, in the case of the
3 repository, this is a decision variable that the DOE
4 controls.

5 It is changed dramatically every year. I
6 always have to go look up what the new repository
7 footprint up. But, it has no uncertainty whenever the
8 DOE finally firms up what that repository area will
9 be.

10 These ovals that are here represent things
11 that you can treat as stochastic variables, or you can
12 treat them as a PDF and calculate them as stochastics.

13 And then you actually couple those
14 together to calculate the recurrence rate. And then
15 that feeds into the repository intersection. I'm
16 going to stay out of this area.

17 I don't want to go there at all. So, I'll
18 just be talking about these two things, E1 the
19 recurrence rate, and E2, the probability of repository
20 disruption.

21 Okay, so what we have to start out with
22 with the experts was they said, given the conceptual
23 model of why you think volcanoes are out there. And
24 so, I kind of stepped back.

25 This is a diagram I borrowed from one of

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1 Frank's papers. And I said, what's interesting, if
2 you look at the base -- the Great Basin, the Southern
3 Basin range, the Colorado Plateau, most of the
4 activity is volcanic activity, is concentrated on the
5 active margin.

6 But there's kind of an interesting
7 tendency that you get small bits of volcanism in the
8 interior parts, both the Great Basin, the Mohave, and
9 the Southern Basin.

10 Basalts in this probably seem to like to
11 pop out occasionally in places where you have to
12 wonder why they are popping up there. Certainly their
13 rates are much lower than these very active provinces.

14 And so, our challenge is to try to
15 understand why these basalts are occurring where they
16 do. I've given up tracing the petrogenesis models
17 because they've changed so much in my 30 years or so
18 of looking at them that I give up.

19 I think they are permissive and they don't
20 tell you a lot. And people go back and forth on what
21 they think is driving these things. But what you see,
22 as Britt described, is fields like Lunar Crater, Cima,
23 which are kind of big, high density volcanic fields.

24 But you also see phenomena where you have
25 down to just individual separate cones, like the

1 Crater, or in Death Valley, where just one single cone
2 can occur.

3 Crater Flat is interesting I think because
4 you have to call it a volcanic field. There has been
5 enough recurrence of events there. But, it's toward
6 the low end of the spectrum of volcanic fields that
7 you see in this whole province.

8 So that's fundamentally the conceptual
9 model. I don't think anybody can say we understand
10 why magma is either generated or comes up exactly
11 where it does.

12 So, I'm going to focus a little bit more
13 on what I think is a critical part of this part of the
14 Great Basin that's unappreciated. And it's Basin
15 Range.

16 But, toward the southwestern edge of the
17 Basin Range there's a very strong overprint of what's
18 been called the Walker Lane structure zone. And that
19 overprint is an overprint of stripes of faulting.

20 And what you see with -- all the basins
21 when we've looked at them in more detail, they show a
22 component of stripes of faulting associated with
23 extensions open the basins.

24 Crater flat has been proposed to be the
25 stripes that -- we have new data at Frenchman Flat

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1 that suggest it's left step pulpar associated with
2 this left step movement zone through here.

3 So, when you look at the structural
4 controls of volcanism, you really should factor in
5 this Walker Lane structural overprint that's basically
6 overprinted on top of the basin range in the caldera
7 models and the caldera cycles.

8 So, let's see. This doesn't show up very
9 well, does it? What I did is I just borrowed this
10 slide from one that I found in an NRC paper. I just
11 wanted to show that what you're faced with, if you
12 take a big zone is, how do you choose a record that is
13 representative for doing your probability calculation?

14 And we wrestled with this for decades.
15 Everybody has a slightly different opinion. And it's
16 kind of fun to read to the PVHA because you see how
17 each expert assembled them in a somewhat different
18 way.

19 And I think the most important thing is
20 not which one is right, but what's the range of
21 answers that you get out of a sampling. I wanted to
22 point out one thing right down in here that I think is
23 important.

24 That's the formation and basalts at the
25 green water range, because, at the end, I want to say

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1 a little bit about it. I think there's a big four
2 million year event that you see associated with the
3 opening of Death Valley or the one of the phases of
4 the opening of Death Valley.

5 And it may have been responsible for this
6 big event here. And I'm kind of wondering whether the
7 Amargosa Valley record that we see is responding to
8 that event.

9 There's a different tectonic event than
10 what's going on in Yucca Mountain. Next slide. Okay,
11 so here's how I put together the record that I think
12 is relevant to the problem.

13 And it is a bit different from the NRC's
14 approach. I basically -- if you look at this, there's
15 a major phase of basaltic volcanism associated in the
16 stage of the Timber Mountain and Oasis Valley caldera.

17 What you see is bi-model basalt roudades
18 with a large volume of basalts. And then you also see
19 another pulse of larger volume basalts when you look
20 at the origin of each of these basins.

21 They opened up a fairly extensive -- as
22 best we can time the extension. We can't time as well
23 as we like. But it does appear that, associated with
24 the opening of the basins, there were large volume
25 basalts.

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1 And these tend to be in the range of say
2 nine to ten, maybe eleven or twelve million years in
3 very -- basin to basin. What you see is they show up
4 mostly in the subsurface, like at Frenchman Flat,
5 Yucca Flat, Crater Flat.

6 And we now -- we drilled some holes in 9.5
7 caldera. What the typical thing that you see with
8 these is these are big volume basalts. They tend to
9 be one to ten cubic kilometers in volume.

10 And I think we're associated with this
11 pulse of tectonics. What I think we now know is that
12 that tectonism is weighing. We certainly know that
13 extension rates are much lower.

14 Although, we're still debating those, but,
15 what you see is, with the later stage basalts, is a
16 switch-over to what I call small volume, post-caldera,
17 post-extension basalts.

18 And they tend to have volumes in the order
19 of about a tenth to a cubic kilometer. And this is
20 the episode that I think is the most important thing
21 to look at for Yucca Mountain.

22 It's the most current -- what I think is
23 a current tectonic regime. Okay, next one. So, how
24 would I assemble what I think is important? Again,
25 here's a familiar map.

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1 My argument that I use in my PVHA was,
2 take a look at the volcanoes in Crater Flat. We also
3 looked at the hidden cone, Thirsty basin units and the
4 aeromagnetic anomalies in Crater Flat and Amargosa
5 Valley.

6 But you have to be careful that some of
7 those are probably associated with this older phased
8 extension. We know from the dates that some of them
9 are in the nine plus age range.

10 And, again, I mentioned that I don't think
11 that the origin of the basalts are well known.
12 There's been a constant debate over cause and effect
13 between structure and the basalts themselves.

14 I think it's absolutely clear that local
15 structure plays a role in the basalts. And whether
16 that's simply that it's the guiding pathway for the
17 last few kilometers or somehow that these waning
18 tectonic systems can trigger episodes is a big debate
19 that I don't think is going to be resolvable in the
20 time of Yucca Mountain.

21 Okay, next. So, getting back to here,
22 here's how I went to assembling this. What we found,
23 one of the interesting things in PVHA was Kevin
24 Coppersmith was the person who led the elicitation.

25 He's a seismologist. And he kind of

1 guided us to think of a new way of starting. And
2 seismologist, when they go to a problem, they come up
3 with a seismic zone.

4 And they look for typical seismic
5 characteristics of that zone and then apply recurrence
6 rates for seismicity events to those recurrence zones.

7 And that having convinced us that that's
8 probably the way we should be starting. Before we
9 always did event counts. And then we looked at zones.

10 And then we tried to combine them. But,
11 what we found out is, when you start with the zones,
12 it does constrain you on how you use your recurrence
13 event, because, depending on the structural definition
14 of your zone, you may include or exclude some events.

15 And so, it's not fair to have a maximum
16 recurrence rate but then apply it to a zone that isn't
17 relevant to those recurrence rates. And I think Bruce
18 is getting to the question that you are asking, that
19 you want to bring as much geologic record and
20 structural intuition into this problem that you can.

21 So, what I did was, I said, okay, let's
22 start with zones and look at different ways to define
23 zones. And then you also have to make some decisions,
24 which are modeled assumptions about the distribution
25 of events within those zones.

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1 I think my next slide starts into that.
2 Yes, what I did was, I said, okay, I'm going to take
3 two approaches. One is I'm just going to say, let the
4 geologic record be your guide and then see what the
5 geologic record tells you.

6 And then the other was I said, I'm going
7 to try to look at what I think are structural
8 controls. So this is what I have that I call spatial
9 models, which this represents in structural models.

10 So, what I started off with in the spatial
11 models, I just said, okay, take the events and then
12 draw areas around those events and see how those
13 evolve through time.

14 And so, what you see if you just look at
15 the record that I think is critical -- which is the
16 last five million years, as Mike pointed out -- what
17 you start off with is the oldest event is Thirsty Mesa
18 at about 4.7.

19 And then you jump down. We did date this
20 one anomaly in Amargosa Valley in 3.8. And you have
21 a 3.7. So we see a northwest trending zone that you
22 can then draw around these events.

23 It's going to change a little bit now when
24 we see some more anomalies in there. But, basically,
25 I would call this event on spatial zone. The one

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1 interesting next step is that, up toward Mesa up here,
2 at three million years, it jumps up to the -- it's in
3 the interior.

4 It's in the red tractor zone in the tephra
5 mountain caldera. So, I would just draw the zone in
6 that. And then you added the 1.1 million year event
7 and then the Sleeping Butte that I think erupted
8 around 300,000.

9 And then, finally, I can't forget Lathrop
10 Wells. Lathrop Wells at 80,000 is then down here. So
11 what you see that I think is kind of interesting is,
12 if the space defined by the first couple of events
13 kind of stays in there and doesn't get modified with
14 the exception of one event out here.

15 So, what I did is I just said, okay, I'm
16 going to use these spatial zones and I'm going to
17 define my recurrence rates based on simply the
18 spatial.

19 I'm making no structural interpretations.
20 I'm just using the geologic record. Next one. The
21 second step that I did I said, okay, I'm going to look
22 at what I think are structural models.

23 And I had a range of structural models.
24 I'm influenced by the Walker Lane that I first pointed
25 out. You have that strong overprint in the Walker

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1 Lane.

2 And you see, when you look at the patterns
3 -- just looking at spatial patterns -- you see a
4 northwest trend to the distribution of volcanoes, not
5 the local trend, but the broader trend.

6 The local trend is following local
7 structure. So, I came up with like a Walker Lane
8 structure. And I had several different definitions,
9 depending on whether I looked at the District Attorney
10 record or the -- Attorney record.

11 And then I had a Crater Flat pull-apart
12 model, which was both Pliocene and Quaternary. And
13 then I included in different components of that the
14 Walker -- I'm sorry, the Amargosa Valley.

15 So that changes there. I think I had
16 seven or eight. And then I included Jean Smith's
17 northeast trending zone. But notice, when you draw
18 these zones, you are including an excluding events.

19 And so, again, you have to be careful to
20 make sure that you sum your recurrence rates based on
21 how you do your zones. Okay, next one. So what was
22 really interesting with PVHA was, you know, I had done
23 this for years.

24 In sitting down with a panel I was amazed
25 with how many different ways people came up with

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1 different models. Mike can testify to that. We all -
2 - almost every expert had a different model that he
3 liked.

4 And each one seemed to have -- there was
5 a spectrum of similar models. But each expert had one
6 model that he would basically beat on the table and
7 say, this is the right model.

8 All the other ones are wrong. And this
9 comes back to -- well, what I wrote here. Many models
10 are possible. There is limited data, so none can be
11 disproved.

12 And nearly every expert had preferred
13 models. So, why get into a debate over which model is
14 right? Look at what the impacts of the alternative
15 models are.

16 So here's just a diagram out of the PVHA
17 which shows all kinds of different ways. Basically,
18 here's Yucca Mountain. These boundaries represent
19 different ways the experts drew their zones and then
20 applied their spatial models to those zones. A wide
21 range, it was impressive.

22 MR. HINZE: Could I ask a questions about
23 that?

24 MR. CROWE: Sure.

25 MR. HINZE: What effect did topography

1 have on the -- we've been talking about here. What
2 effect does topography have upon the decisions here?

3 MR. CROWE: That's a good question. I
4 could tell you, in my model, topography had a major
5 effect. I mean, basically, I agree with Bruce Marsh's
6 assumption.

7 Our observation that you look at the set
8 things in general. A few places they lap into
9 bedrock. But, most of the places, the concentration
10 is -- particularly Quaternary cone, let's say.

11 The Old Great Basin region tends to be an
12 alluvial valley. And, if you go talk to the
13 structural people, they say alluvial valleys is where
14 the extension is occurring.

15 That's where the basalts are going to
16 occur.

17 MR. HINZE: That's where the action is.

18 MR. CROWE: Right. And that's how would
19 I use in my model. Now, we had different ways of
20 doing that. What we ended up -- go back to that just
21 one more time.

22 What we ended up with is what became
23 really important was we had kind of a boundary. I
24 think they drew this in the PVHA. And there was this
25 raging debate over could things go in there.

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1 The way I tried to resolve it, as I will
2 point out later, was I said, okay, I'm going to locate
3 my centers within my zone, but allow the dikes to
4 extend out of there.

5 So, the dike lengths, the dike orientation
6 will dictate whether or not they can result in a
7 structural eruption. Okay, next one. So now here's
8 just some interesting things I wanted to point out
9 that I think the record tells you.

10 Again, I'm focusing on the younger ages
11 here. And this is what I call the small volume, the
12 point one to one cubic kilometer. And what I did
13 here, because I fought with geochronologists for so
14 many years, I hate to see histograms of ages where
15 they are based on the number of ages.

16 What I did is I tied them to an age and an
17 event. So, every place that I had an event and I knew
18 the age. In some cases I had to guess the age. I
19 called that one count.

20 And then I histogrammed this out. What
21 you see is some interesting patterns there. There was
22 a cluster of events in the seven to ten. These are
23 the small volume events.

24 There was a hiatus here and then another
25 cluster of three to five. And I think this represents

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1 this Amargosa to Death Valley event. And then another
2 cluster just happens to have been the two really is a
3 cluster of one and -- right here.

4 So, I think the record is showing you that
5 there may have been three discrete pulses of activity
6 possibly associated with pulse of extension.

7 And the question is, what's relevant to
8 the future hazard. All the experts debated it. The
9 majority of them used a five million years and
10 younger.

11 Some only used one million years and
12 younger. Some also included everything. But not many
13 did it. Now, here's the second thing. What I did
14 here was I just plotted the locations of these.

15 They are color coded in red as the younger
16 group and blue is the older group. Then I just
17 plotted an ellipsoid and the centroid of the
18 distribution.

19 And what you see is there are two
20 different spatial distributions. All the older ones
21 occurred mostly toward the northeastern parts of the
22 Nevada Test Site.

23 And then you see this centroid here
24 located, not surprising, down in Crater Flat. And
25 here's the anomalies of Yucca Mountain. So, to me,

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1 the record is telling you that there are some clear
2 patterns here and you should incorporate those
3 patterns in your probability models.

4 Next one. So here's a real interesting
5 thing I did just before I left. And it has been
6 buried in the paper I wrote, I doubt if anybody has
7 even read.

8 What I did was I said, okay, let's look at
9 an interesting exercise. Let's just go through the
10 geologic record and let's say, where did each volcano
11 occur and what's the sequence?

12 So, basically, these lines that I've drawn
13 is I've covered both of these. I went to -- I started
14 with the oldest events were up here. And then I just
15 drew a line where the next event was.

16 And then I continued through that. Then
17 I jumped down to here and repeated the process here.
18 And what you see is this remarkable oscillation. And
19 it tends to like a few spots.

20 But it oscillates back and forth. And, in
21 fact, if you look at Crater Flat, the first event at
22 is Thirsty Mesa, as I talked about. Then it jumps
23 down to here.

24 And then it jumps up here. Then it comes
25 back down here and goes up here and comes back down

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1 there. To my mind, there's a lack of predictability
2 there.

3 Well, let me rephrase that. There's a
4 tendency for things to cluster in groups, here and
5 here. But, when you look at it in detail, the last
6 event is a poor predictor of the next event.

7 It looks like magma will come up wherever
8 it feels like coming up. You have to be very
9 cautious. That's why I went to a random model. I
10 just felt like we just don't have enough information
11 to really say, why is it coming up where it is?

12 And so, what I did from my zones, I'd see
13 this is a random distribution of event. But let me
14 point out that there are two scales of clusters. And
15 I worked with some spatial experts to look at this.

16 There are the clusters where, when you
17 have an individual event -- let's take the 1.1 million
18 year -- that clusters as a group of four things.

19 But that's clustering like one event that
20 forms in probably a fairly narrow period of time. As
21 best we can tell, it's largely synchronism. I gave up
22 arguing with the geochronologist of whether there's
23 any differences there.

24 But, the best we can date, we don't see
25 any difference. Now, that's what I call an event

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1 cluster. And then there's a spatial cluster, which I
2 think is just where you see patterns through time.

3 So, I distinguish those too. Okay. So
4 now let's come back here. So what I did is I did my
5 two types of zones and then I used a random
6 distribution of events.

7 So now we come down to event counts. And
8 anybody who had been around Yucca Mountain knows that
9 this got debated for so many years and there have been
10 so many different models that I got tired of even
11 talking about them.

12 But, here's the parameter. You have to
13 come up with an event definition. And Britt gave you
14 one event definition. And that's basically -- it's a
15 model assumption of how you chose your events.

16 And what I think is important is to make
17 sure each expert defines that, because you can end up
18 kind of muddying the waters using different event
19 definitions and come up with recurrence rates that are
20 variable and are confused because you haven't
21 clarified your event definition.

22 You have to choose a time interval. I
23 covered that. I'll talk a little about time
24 distribution. But we argued that one -- as I
25 understand it, I think both the DOE and NRC agree on

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1 a steady state of vent rate that they have been using
2 for their probability models.

3 And then this undetected events I'll talk
4 about in a little bit. But, those are the parameters
5 you have to come up with when you do your event
6 counts.

7 And here's one. I borrowed this one out
8 of one of the things that the ACNW sent out because I
9 thought it was really neat. The vent areas show up in
10 red in this particular spectrum.

11 And what you can see is, if you take like
12 phases of volcanism, they have a discrete event
13 geometry to them. And it ranges. Britt described in
14 some detail the 3.7, what you see.

15 When I originally mapped it I thought
16 there was about fiver or six centers so that I could
17 reconstruct. So, we had a cluster of five or six
18 centers.

19 Now, are there five or six events there?
20 Or is that one event that has an event geometry that's
21 spread over an interval here. What's important really
22 is that you can look at it almost both ways.

23 But, they have different consequences.
24 So, if you're going to assign the maximum
25 consequences, which would be a large event, you have

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1 to go back to the recurrence rate and treat it as a
2 single event.

3 So, you can't over count it and then come
4 back and weight it. So, what would happen is the
5 consequences go up but the recurrence probability goes
6 down.

7 So, if you look at the record, what you
8 see is that there was about five here. We think
9 there's about four here. We've got long debates about
10 what counts as one or two.

11 And I don't think it's worth arguing over.
12 Thirsty Mesa up there, I think I mapped three distinct
13 event. Sleeping Butte has two way up here.

14 Lathrop Wells and the Mesa up here, we're
15 just thinking they had one single event. So what you
16 see is you have -- the record is telling you there's
17 a spectrum of behaviors.

18 And I think you should just treat it
19 probabilistically as probably as a uniform from one to
20 six. And that's a nice way you can treat how you do
21 your events.

22 But you have to be very careful to make
23 sure that how you do your events is tied to the
24 consequences. And then I mentioned that you have to
25 do your event counts specific to the zones.

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1 And, I mean, I made the mistake in my
2 first calculations. I treated them independently.
3 And you end up coming up with combinations that have
4 no possibility, because they don't exist in the
5 geology realty space.

6 Undetected events, to the best of my
7 memory, Bill, to answer your question, was, the way we
8 handled it in PVHA was most of the experts thought
9 that if magma is going to ascend all the way up to
10 repository depths of about 300 meters, it's going to
11 make it to the surface.

12 So they felt like it's going to be
13 unlikely to have an event that comes up into the
14 shallow crust and just stops, that you're in the depth
15 range.

16 We're starting to -- volatile. It should
17 be the driving force that's going to push it to an
18 eruption. But they felt that there could be an event
19 geometry of more undetected events with that.

20 So say at Lathrop Wells there might have
21 been some intrusions to the southeast of it. So they
22 were adding -- that's what they call undetected
23 events.

24 And that's different from having an event
25 that came up and never reached the surface and created

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1 something in the volcanic record. So, you know, I
2 completely agree with how Britt was describing that
3 there.

4 You have to be careful on how the
5 different -- PVHA looked at undetected event
6 associated with known surface volcanoes. And there
7 was a fundamental dispute over whether or not you
8 could have an intrusion pausing in the very shallow
9 crust.

10 Okay, now here's the -- these diagram I
11 hate putting up because I get in trouble every time I
12 talked about them. But, let me start with a simple
13 one first.

14 If you look at -- this is just cumulative
15 volume versus the time. What you see is a four
16 million years event where larger volume has inversed
17 slope.

18 And then the younger events have a
19 different slope. And I think these are probably
20 telling you that fundamentally they are different
21 parts of the record, that they're probably responding
22 to, I think, different tectonic regimes.

23 And, you ought to make some choices about
24 which one you think is the most relevant to the
25 future. The second one was -- this is kind of an

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1 exotic plot that I labeled.

2 What I calculated was, I took the time
3 from the previous event, which I called the reposed
4 interval -- so, in other words, like this is between
5 Thirsty Mesa and I think the anomaly in Crater Flat.

6 That would represent this reposed
7 interval. And then I plotted that versus time. First
8 I fit a nice little linear regression. So what you
9 see is a slight tendency to a decrease in that reposed
10 rating.

11 Again, your dataset is pretty limited.
12 Just for fun, I did a distance weighted square fit
13 which shows an oscillation. When I put this in a
14 paper, a reviewer said, oh, he just predicted the next
15 eruption is going to happen any time now.

16 And, I don't know if I'd go that far.
17 But, we used to have negative ages on Lathrop Wells.
18 And we used to argue, there's your next event. Okay,
19 next slide.

20 So, okay, coming back to -- then we summed
21 all these event counts in different ways. And I did
22 it for spatial and structural models. That then feeds
23 into the recurrence rate.

24 And then that recurrence rate goes into
25 the probability of repository intersection. And let

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1 me show you the way that I've ended up kind of liking
2 to do it.

3 But there's a whole bunch of different
4 ways to do this. Let me point out the variables that
5 go with this repository area, the dike ledge, which
6 you can treat as stochastic, and dike orientation, and
7 then the probability of an eruption are an intrusion.

8 So, the next one. Okay, this again is
9 just a reminder. This is how I assigned these to my
10 individual structural zones. So what I said is I
11 allowed these to have a random distribution of events
12 within each of these zones.

13 Then go to the next one. Then I worked
14 with Goulder and we used the code. And we run
15 simulations where we assign the dike height, a dike
16 length, and a dike orientation.

17 We just did simulations of the repository
18 block that's buried down under here. And we just let
19 them run. This one happens to be for the Yucca
20 Mountain region.

21 And, because we find that the outer domain
22 of our models, we put a lot of dikes in this one to
23 extend past the model domain. But, in the other ones,
24 we just basically gave a dike dimension, randomly
25 located them within that.

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1 And then we summed up three things -- two
2 things, the number of intersections in the volume
3 intersection, and then calculate that as our
4 disruption probability.

5 It's very comparable to the way most
6 experts did it in the PVHA where a geomatrix helped
7 them use kind of a geometry of intersection. They
8 treated dike length as a stochastic.

9 They treated dike orientation as
10 stochastic. And they ended up -- you have a
11 trajectory of only certain areas will actually project
12 into a disruption.

13 So they brought that geometry of dike
14 directions into an intersection. So, if you go back -
15 - so, if you go to different centers, some of them are
16 capable of a repository intersection, some are not.

17 So, disruption ratio becomes -- we
18 basically are very influenced by the modern stress
19 field, which says that dikes should be entering in a
20 north-northeast direction, basically.

21 And the stochastic was centered about
22 that. So, when you locate your events, you assign
23 that to it. So, some events are going to occur in
24 that zone.

25 But they have a virtually zero probability

1 of intersecting the repository because either the
2 orientation of the dike length takes them out of the
3 ability to intersect.

4 So, okay, that's what I did. I want to
5 say just a little bit about -- I did this right at the
6 last minute before I left Yucca Mountain. I think
7 it's also been buried and nobody has read it.

8 It's something that I didn't do until
9 after I finished my probability calculation. I came
10 up with a simple logic that says, I think there's some
11 somewhat firm bounds you can put on this probability
12 of repository disruption.

13 Here's the argument I went through. In
14 the basin and range there is a background recurrence
15 rate. Basalts tend to keep coming up. And so, I
16 said, well, if you located a repository away from a
17 defined volcanic zone or in this background, you
18 should calculate the probability of it being in a
19 background setting.

20 And, that's what I did. My particular --
21 well, I'll get to that in a second. And then I said,
22 the other -- so that would define your minimum value
23 for your probability.

24 So, in other words, the distribution
25 shouldn't get less than background, or maybe you have

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1 to go back and question your assumptions. And then,
2 on the other end what I said is, let's just take the
3 repository and put it right in the middle of one of
4 these zones that I defined.

5 And that should give you the maximum at
6 the other end. It says, we think Yucca Mountain --
7 this is open to great debate, of course -- sits
8 outside of the volcanic field, but close to it.

9 So, logically, Yucca Mountain -- the
10 probability of disruption should be greater than
11 before but less than putting it right in an active
12 volcanic zone.

13 And this becomes the big debate. How far
14 away from a volcanic zone is Yucca Mountain. And I
15 don't think that's resolvable. So, okay, let's see
16 what happens if you make those assumptions, what you
17 come up with.

18 I use the Southern Great Basin. And I use
19 this thing that was very popular during the PVHA
20 called the Amargosa Valley Isotopic Province, or AVIP.

21 It's an area where there's a unique
22 isotopic composition to most of the basalts. I'll let
23 Frank talk about that. I'd like to stay out of that
24 area.

25 But, basically, the AVIP defined this area

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1 of unique isotopic compositions of basalt. And so,
2 what I did is I said, okay, let's take a 4.5 kilometer
3 repository footprint, put it into these two provinces.

4 And I used event counts from the
5 combination of expert judged in my own regional field
6 studies where there wasn't any data. And I used that
7 -- I used the recurrence rate and then the disruption
8 ratio was simply the ratio of the area to the
9 repository area.

10 And here's what you come up with numbers
11 of what I would call background. Somewhere down in
12 the low one to three times ten to the minus nine. So,
13 what I would say is, anybody that calculates a number
14 less than that, you should question how you assembled
15 your probability calculations.

16 So, let's go to the next one. So then
17 here's what I did if I plugged into my zones. And the
18 numbers range from almost two times ten to the minus
19 seven.

20 Two is low for this at the Jean Smith
21 Northeast structural zone. And it's interesting for
22 this one because there is a restricted number of
23 events that that encircles.

24 The recurrence rate goes down. And so,
25 the probability of disruption is lower. So, actually,

1 this is the zone that includes Yucca Mountain.

2 And yet it has the lowest of the
3 calculated. So, somewhere in this range, you would
4 argue -- and I put the number up around one to one
5 point two times ten to the minus seven would be a
6 maximum bound.

7 So, what I would argue is, if you're
8 getting much higher than that, you basically aren't
9 paying attention to the geologic record and you should
10 look at your probability calculations.

11 So, let's go to the next one. So, here's
12 what I did. I love this phrase that basically you
13 have to cut off the maximum, which is the uniform
14 distribution between your min and the max.

15 Basically I'd like an uninformed prior is
16 they way I like to look at it. So, my uninformed prior
17 was the min and the max I calculated. So, I used on
18 times ten the minus nine and one times ten to the
19 minus seven.

20 That gives you a mean value of about five
21 times ten to the minus eight. And, interestingly
22 enough, our numbers -- everybody's numbers comes
23 around pretty close to that.

24 I mean, in my opinion, some of the fights
25 I've been in and I think are still occurring are

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1 you're just modeling noise around some numbers. It's
2 probably unresolvable.

3 So, why not just kind of look at it that
4 way. So, I went back and I looked at the PVHA
5 disruption, which is here. So, this is ten to the
6 minus ten, ten to the minus seven.

7 What I argue is they have a fair amount of
8 detail that goes down below the ten to the minus nine
9 range. So I'd argue that we probably should have
10 truncated that and said that those are just a little
11 bit too low.

12 And so, what you do is you reduce some of
13 this huing on this distribution. You probably shift
14 the mean a little bit over here. And then, going to
15 the NRC model, they've been talking about a ten to the
16 minus seven, ten to the minus eight for most of the
17 data they interpret.

18 And I would just argue that, instead of
19 using ten to the minus seven value -- which they do in
20 their PA calculation -- treat that as a uniform and
21 sample that distribution.

22 If you do that, the difference between
23 this uniform and about there is not enough to get
24 excited over. And I would argue it's getting time to
25 move on to consequences, where all the uncertainty is.

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1 So, next one. So, the final overview
2 comments, I just want to comment a little bit about
3 where I was when I thought I left with the
4 aeromagnetic anomalies.

5 I haven't looked at the new data. So, it
6 would be very interesting to look at it. As I
7 mentioned, we did drill the one anomaly. And the fact
8 that -- as Britt pointed out -- these anomalies are
9 buried.

10 If there was surface basalt at the
11 centers, they have to be fairly old. He used two. I
12 would argue that I bet they are going to come out
13 around four, because that's the one that we drilled,
14 at about four.

15 And it also matches a regional Death
16 Valley event that I think you see as an overprint in
17 this region. So, if these things are about four, and
18 they're mostly located down in the Amargosa Valley,
19 the dike lanes and the dike orientations are not going
20 to lead them any intersections.

21 So, you don't want to just look at the
22 recurrence rate. You want to look at both the
23 recurrence rate and the likely hood of an intersection
24 with these new events.

25 I don't think that it's going to change

1 the relationships as much as people have been saying
2 in a new era. There's going to be a range of change.

3 But, when you take into effect the
4 recurrence rate and the likelihood of disruption, I
5 don't think the numbers are going to change that much.

6 Here's the only thing -- in fact, before
7 I left -- go to my current program with the DOE to do
8 this. This is the anomaly near little cone. It has
9 a normal polarity, which doesn't match anything we see
10 in the record.

11 Everything else is reversed out there. We
12 need to find out what that is. Because, if it is
13 something in the record that we don't know of, then we
14 really need that data.

15 And I had also argued that let's explore
16 some of the anomalies in Crater Flat that are close to
17 Yucca Mountain that might have a higher potential
18 intersection.

19 And that should influence -- I mean, those
20 are just so important. And my opinion is it's
21 probably so important that you really should gather
22 data on those.

23 We have the potential, so let's just go
24 gather it. But, I would argue that, for Amargosa
25 Valley, drill one or two of them. But, if they all

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1 come out at about the four million range, I think I'd
2 walk away and feel pretty confident that you know what
3 you're doing.

4 MR. HINZE: Before you leave that, if I
5 might, the limited impact, is that based upon an
6 assumption about where these aeromagnetic anomalies
7 will be found? Could you expand on that a little bit?

8 MR. CROWE: Yes. It was based mostly on
9 what I saw in '96 with the aeromag data, which was
10 mostly Amargosa Valley. They have some new data in
11 Crater Flat that I'd want to look at.

12 So, I should cavy out that. That's a '96
13 profile that I'm presenting. But, if you looked at
14 what Britt was presenting, most of the anomalies are
15 down in Amargosa Valley.

16 I'm guessing that a lot of the ones in
17 Crater Flat are probably very tough, since it is so
18 magnetic. You can fault it and get a pretty good
19 signal.

20 MR. HINZE: How about in Jackass Flats?

21 MR. CROWE: I'm biased. But, I looked at
22 Jackass and I was doing some work. There actually is
23 a drill hole that penetrated the south in Jackass.

24 Way back in the nuclear rocket program in
25 the 60's they drilled three holes, J11, J12, and J13.

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1 And one of them hit a basalt at, I think about 1,100
2 feet.

3 I looked at the cuttings from it, and I
4 think it matches what we call a basalt of EMAD, which
5 you see at the surface, which we date at about 11
6 million years.

7 I walked the sections and looked at -- we
8 have dates through all the basalts surrounding that
9 valley. And they're all in the nine to 11 million
10 years.

11 I think it's unlikely you're going to see
12 a shallow anomaly there. But I want to see the high
13 resolution data to see if anything shows up. But, I
14 don't think I would get really excited about it.

15 The record seems to show that not much has
16 been happening in Jackass Flat. Let's see, where was
17 I? Okay. Here's on last thing I wanted to point out.

18 I really think that the Crater Flat pull-
19 apart is where the active extension is. And the
20 record is telling you that that's where the basalts
21 are coming up.

22 And that's the major part of the record we
23 should be looking at. And I think it's the critical
24 thing to calculating future probability. I think
25 people have neglected this.

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1 I worked with Will Carr here originally.
2 And he always pointed out what he called the Minot
3 Spotted range system, which is a series of -- left
4 slip faults.

5 And we now know that one of those control
6 the extension of Frenchman Flat. What's really
7 interesting is, in this basin -- I guess I would argue
8 by inference the Jackass Flat Basin -- you probably
9 have strike -- components to them in this left slip.

10 And most of the basalts that you see
11 occurred primarily at the time of extension, as best
12 we can tell. And what you see is fairly large volume
13 basalts.

14 We know -- we have penetrated basalts in
15 Frenchman Flat. In multiple cases, the -- testing
16 dated down -- the maximum plug buried up in the
17 bedrock to the west, dated five.

18 And, they are voluminous enough that they
19 look like they probably are marking the major
20 extensions. Similar arguments could be made for
21 what's in Yucca Flat.

22 In fact, I now think going -- this bedrock
23 that we dated 86 here is probably part of this
24 extension of that basin. What's kind of interesting
25 is most of the basins except Crater Flat and Frenchman

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1 Flat have one major phase of basaltic volcanism
2 associated with extension.

3 But, in Frenchman, there was the later
4 stage of about 7.2 million years later. What's
5 interesting is that this has been a persistent site of
6 volcanism.

7 I think there's a little bit of anomalous
8 for all the other basins here in that, not only was
9 the -- older stuff that floors the basin -- we
10 penetrated 11.5 million year basalt at 1,100 feet
11 below the surface here.

12 We see it in the south exposed to the
13 surface. But then there are these multiple pulses of
14 younger. And that's where Crater Flat is a little bit
15 unusual.

16 I personally think it may be a combination
17 of Amargosa and Crater Flat -- is in the intersection
18 of this spotted range Minot Mountain system.

19 And it has been influenced by a part here,
20 and possibly might be influenced by the proximity to
21 Death Valley. But, that's very speculative. And I'm
22 just going there because I can get away with it
23 because I don't go to the program.

24 CHAIRMAN RYAN: Okay, we have time for a
25 few questions. Any questions? Yes?

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1 MR. HORNBERGER: Bruce, you mentioned that
2 when you did this you had a bound, and you said that
3 less than ten to the minus nine is not credible. And
4 you didn't think that higher than ten to the minus
5 seven was credible.

6 MR. CROWE: Yes. I would go into maybe
7 three times ten to the minus seven range, but
8 somewhere in there. You might have to -- you might
9 take all the expert judgment and assemble them to see
10 how you are bound to compare. I just did my set of
11 models.

12 MR. HORNBERGER: Right. I realize that.
13 Can you think of any way consistent with your
14 knowledge of the geologic system that you could say
15 get to five times ten to the minus six?

16 MR. CROWE: No, I can't. I mean, you'd
17 have to have some preferential mechanism for focusing
18 events at Yucca Mountain. I think the geologic record
19 says.

20 Since you can go back ten million years,
21 there is that one Solitario Canyon event. But, I
22 think that's associated with the maximum extension. If
23 you go back and look at the ash record of Yucca
24 Mountain, it was a basin when the eruptions occurred
25 that formed most of the mass of Yucca Mountain.

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1 And it was elevated between the two sheets
2 of Timber Mountain. You can see this huge -- forming
3 in the geologic record there. So, most of the
4 tectonism that elevated that mountain, occurred about
5 11 million years ago.

6 And Yucca Mountain de-coupled from Crater
7 Flat in my opinion at that point. If it didn't, it
8 would still be a basin. But, it's a high standing
9 range.

10 And, again, I believe the model that
11 extension in the record all over the Great Basin shows
12 that, with seismicity, that where the extension is
13 occurring the valleys.

14 And that's where the basalts tend to
15 occur. But they can spread a little bit. That's not
16 to say it excludes penetration. But I would say our
17 best guess from the record is in the valleys of where
18 all the action is.

19 MR. HINZE: Even including the 10 million
20 year old events, you still fall within the ten to the
21 minus seven, ten to the minus eight?

22 MR. CROWE: You do, exactly right. Yes,
23 I mean, I really have -- the only plea I would like to
24 make is get on to consequences. I mean, that's where
25 your uncertainty is.

1 And you're going to just be fine-tuning
2 here. I mean, I really think you should drill these
3 anomalies. I mean, I would like to see the new
4 dataset.

5 But, the expectations are that it's not
6 going to change it too much. And, if you look at your
7 bucket of uncertainty, the consequences are so much
8 more significant.

9 CHAIRMAN RYAN: Thank you very much Bruce.
10 That was an interesting talk. Any last questions?

11 (No response.)

12 CHAIRMAN RYAN: All right, we'll press
13 onto our next speaker. Mr. Neil Coleman of the ACNW
14 staff will be talking about alternative views on the
15 likelihood of an igneous event in the Yucca Mountain
16 region.

17 And, while Neil is getting ready, let me
18 recognize Dr. Charles is in the audience, a member of
19 the ACNW. Thank you for your participation, for being
20 with us.

21 **ALTERNATIVE VIEWS ON THE LIKELIHOOD OF AN IGNEOUS**
22 **EVENT IN THE YUCCA MOUNTAIN REGION**

23 MR. COLEMAN: This talk represents
24 background research in support of the ACNW's review of
25 volcanism. I thank my co-authors, Bruce Marsh of

1 Johns Hopkins University in Baltimore, and Lee
2 Abramson of NRC's office of Research.

3 I thank them for their contributions.
4 Thanks to John Trapp of the Staff for providing NRC's
5 PVHA code and -- Center for Nuclear Waste Regulatory
6 Analyses.

7 PVHA stands for Probabilistic Volcanic
8 Hazard Assessment. I should add at this point, this
9 talk represents our views, the author's views, but
10 does not necessarily represent vies of the Commission,
11 NRC Staff, or the ACNW.

12 We suggest that our work be considered in
13 evaluations of volcanism at Yucca Mountain. I will
14 briefly describe the technical issues for volcanism
15 and provide a brief summary of volcanism in the
16 region.

17 Previous estimates of the probability of
18 volcanism will be discussed. And I will show the
19 results of our statistical and PVHA analyses. And we
20 will compare Yucca Mountain to other volcanic fields.

21 Finally, I will present conclusions and
22 recommendations. Next slide, please. A special topic
23 in the earth sciences is using geologic data to
24 evaluate very low probability events such as volcanic
25 eruptions and earthquakes, and evaluating how these

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1 could potentially have significant consequences.

2 Now, of course, the technical issue here
3 is the potential for inter-igneous activity very much
4 like the repository. Here we are looking south from
5 Yucca Mountain.

6 And, in fact, some of us from the Staff
7 were on the crest of Yucca Mountain and had that exact
8 view just yesterday. You can see the 80,000 year old
9 Lathrop Wells cone in the distance.

10 Geologically, this is the youngest known
11 volcanic event in the Yucca Mountain region. Next
12 slide. On the left is a pan view of the underground
13 repository.

14 On the right is a close-up of the waste
15 placement drift showing the potential horizontal
16 storage of alloy 22 waste packages. If the probability
17 of an igneous dike intersecting the repository is less
18 than one times ten to the minus eight per year, it may
19 not be considered in licensing.

20 However, regional studies do suggest that
21 the probability is just high enough that the
22 Department of Energy must evaluate the consequences of
23 dike intrusion.

24 Potential consequences will be discussed
25 in the next session of this working group. Next

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1 slide. I want to take a moment to just put geologic
2 time in perspective.

3 We toss these terms around, Quaternary,
4 Pliocene, Miocene. Here's a timeline that compares
5 volcanism in the Yucca Mountain region to other
6 events.

7 This figure shows the last two million
8 years. The tuffs that form the surface of the
9 mountain are quite a bit older. They erupted between
10 ten and 13 million years ago.

11 So they are off the left end of this
12 chart. Not all the basaltic events in the region are
13 shown. Here are some examples. The X axis here is in
14 millions of years before present.

15 The last 1.8 million years represents the
16 Quaternary. You can see the -- if I can find the
17 button here -- the time frame on the bottom. 1.8
18 million years is the break between Pliocene and
19 Quaternary.

20 And there's a Miocene-Pliocene boundary of
21 5.3 million. Older events are Miocene in age.
22 Approximately 11 ice ages appear since the late
23 Pliocene time.

24 Only once volcanic event at Lathrop Wells
25 cone has erupted since the advent of modern humans on

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1 earth, that's the *Homo sapiens sapiens*. That was
2 about 120,000 years ago.

3 The million year old cones in Crater Flat
4 pre-date all the pieces of the *Homo sapiens*, including
5 the Neanderthal. The famous hominid fossil Lucy,
6 right here, (*Australopithecus aphaeresis*), dates back
7 to the Pliocene time around the time that those guys
8 were occurring in Crater Flat, the Pliocene.

9 At the far left is the Solitario Canyon
10 dike that was mentioned, around 10 to 12 million
11 years. There are two dates for that one. The key
12 thing to point out at the top of the figure is that
13 the uncertainty in the actual number of volcanic
14 events greatly increases as you go back in time,
15 because you had more time to erode basaltic events
16 that occurred then.

17 Also, you had more time to cover them up
18 with younger volcanic, like sevens. Next slide,
19 please. The large surface exposures in the region
20 outside of the basin data is a tuff produced between
21 nine and 13 million years ago, the huge caldera formed
22 eruptions, pyroplastic eruptions, some of it.

23 The largest pyroplastic eruptions that we
24 know of anywhere. You see a series of these
25 overlapping calderas north of the blue star, Yucca

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1 Mountain.

2 Calderas are large areas of collapsed
3 terrain that form during and after large volcanic
4 eruptions. There are extensive Miocene and Pliocene
5 basalts that erupted in and near these calderas, which
6 represent kind of a unique structural.

7 Next slide. Dr. Crowe showed this slide.
8 I'll just mention the repository shown in blue here.
9 The black areas here are Pleistocene basalts. There
10 are eight of them, including two up in the upper left
11 hand corner, that Black Mountain vicinity.

12 Of course on the sort of black pattern
13 sort of classing basalts, and the grades in the
14 Miocene basalts, which occurred all over this area.
15 After Miocene time, volcanism clustered to the west
16 and south of Yucca Mountain.

17 There are no known Pleistocene or Pliocene
18 basalts on Yucca Mountain or to the east in Jackass
19 Flats. Next slide. Here is a satellite image. I
20 think John Trapp showed this one also.

21 The Yucca Mountain site is, again, in the
22 blue star location. The DOE has conducted
23 aeromagnetic surveys. And we saw some initial results
24 from that in an appendix.

25 They have plans to drill and date a number

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1 of suspected buried basalts. The latest drilling
2 results show that the basalt penetrated at Nye 23P,
3 they didn't encounter basalt.

4 What was sort of interesting, the
5 impression that the DOE contractors had is that this
6 may not be basalt, it may be a boulder zone that was
7 penetrated.

8 But, I would suspect that if these were
9 large boulders, they wouldn't come from very far. So
10 that probably does represent an insidious basalt
11 somewhere here nearby.

12 But the key is that this is not
13 particularly surprising to find this. There is no
14 magnetic anomaly associated with it. It is very deep,
15 400 feet deep in alluvium.

16 And the age that has been determined, the
17 Miocene age is consistent with the ages of other
18 basalts in Jackass Flats. Next slide. There have
19 been approximately four known pulses of basaltic
20 volcanism in the area.

21 And this is a different way of showing
22 what Dr. Crowe showed with changes in the estimated
23 magma volume over time. What you're seeing is volumes
24 of magma erupted in cubic kilometers that are on the
25 bottom scale.

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1 The X axis is a million years before the
2 present. The vertical axis is volume and cubic
3 kilometers. The large bar, A, represents the Miocene
4 eruptions, B, the Pliocene events, C and D the
5 Pleistocene events.

6 The tiny bar under D is the Lathrop Wells
7 cone. This figure shows the volume of volcanism was
8 basaltic and the -- were constant. Support in an
9 uncertainty increases a lot as we go back in time
10 right to about the big bar A.

11 It is most certainly too small, because
12 those Miocene results were probably buried by younger
13 basalts and alluvium in Crater Flat. Likewise, the
14 Pliocene events in B may similarly be too small.

15 The magnetic data that we saw yesterday
16 gave a preliminary look -- shows that is indeed the
17 case. The Pleistocene volumes shown by C and D are
18 much more reliable because little time was available
19 to erode or conceal those deposits.

20 Next slide. Here are estimates for
21 volcanic disruption of a repository, some of which
22 claim the probability could be much higher than
23 previously thought -- ten to the minus six per year or
24 higher.

25 That is on average one penetration of the

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1 repository per million years. I am aware of these
2 reports and believe that each represents an earnest
3 effort for the authors to come up with reasonable
4 estimates.

5 Now, the rest of this talk considers some
6 simple tests of whether the highest probabilities are
7 realistic. We look at past volcanic -- four time
8 scales, the 13 million span, the total length of time
9 that the surface rocks have existed at Yucca Mountain.

10 One million years is the to the last four
11 million years. 100,00 years, and then some inferences
12 about present day conditions. We'll look at present
13 day.

14 One impetus for a higher probability would
15 be unusual crustal activity. In 1998 Brian Wernicke,
16 et al reported in the Journal of Science that Yucca
17 Mountain has tried to pull apart.

18 This claim is countered by Savage, et al
19 in 1999 and in 2001 papers in the Journal for
20 Geophysical Research. They used a larger GPS network
21 to show that the extension rate is not anomalously
22 high for this region.

23 And, therefore, present day strain rates
24 do not indicate conditions favorable for the infinite
25 triggering of volcanism. Next slide. The rocks that

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1 make up Yucca Mountain record an integrated tectonic
2 volcanic history since the 13 million year old tuff in
3 the repository.

4 Yucca Mountain is one of the most
5 intensively studied places on earth. Over 20 years of
6 studies have included detailed surface and sub-surface
7 mapping, geophysical surveys and construction more
8 than ten parameters of tunnels.

9 DOE drilled more than 450 surfaced bore
10 holes -- depths. It seems unlikely that multiple
11 dikes could exist in the repository footprint and
12 escape detection.

13 We examined whether dike penetration rate
14 was greater than two times ten to the minus seven per
15 year are realistic given that no dikes have been found
16 in or above the 13 million years old repository block.

17 Now, it was mentioned earlier that there
18 is one event, a dike 10 to 12 million years old, that
19 was a near miss. And you can see it. There we go,
20 just to the west of the site and located within the
21 Solitario Canyon.

22 You can see the expression of fault in the
23 topography in this area. And here is a north-west
24 extension of it as well. Although it is close, it is
25 a near miss.

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1 As far as we can tell, it did not
2 penetrate the repository block. And, DOE does use
3 certain criteria for set-back from faults for tectonic
4 reasons, for earthquake reasons.

5 I should mention that, because of the age
6 of this unit, we know that in true to these upper
7 faults, during the ancient period of caldera
8 formation, when basaltic volcanism was very
9 widespread.

10 And, caldera formation had not ceased at
11 the time that this dike was in place. There was still
12 activity to the Northwest, Thirsty Canyon tuffs are
13 younger than this unit.

14 The image on the right shows this Miocene
15 dike is very close to the site. Exposures are small.
16 The whole thing is maybe about 10 to 15 in length.
17 It's about a meter across, less than one meter thick.

18 And it is highly eroded. What you see is
19 most of what is there. It is possible that other
20 features like this exist but have been undetected on
21 the mountain.

22 Geophysical methods would be poor tools
23 for finding dikes like this. And, in fact, that was
24 presented in the appendix 7 yesterday, that low
25 altitude magnetometer passes over this dike did not

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1 detect it.

2 This is the last high resolution work that
3 was done. Although, that was a very preliminary
4 result, and there were numerous other passes. But,
5 yet, it is an extremely -- dike.

6 It was found in the geologic method. And
7 the geologic method is the best tool. And the
8 emphasis that was placed on mapping out fault traces
9 maximize the possibility of finding this kind of
10 feature.

11 Of course, the best way to locate any kind
12 of dikes in the mountain, is in the underground
13 tunnels. They have been mapped in great detail. No
14 dikes have been found in them more than 10 kilometers
15 of tunnels.

16 I should also mention that, on this trip
17 when the photograph was taken, an NRC hydrologist was
18 the one who located this, an individual with almost no
19 mapping experience.

20 So, next slide. We could use the apparent
21 absence of basaltic dikes to detect -- in the
22 intrusion probability. Assuming a constant recurrence
23 probability rate, the number of penetrating dikes in
24 time, T, has a Poisson distribution with a mean of $\bar{e}T$.

25 The probability of no penetrations is the

1 exponential of minus eT for two times ten to the minus
2 seven per year. The expected number of penetrating
3 dikes is 2.6.

4 The probability of at least one
5 penetration is .93. For a recurrence probability of
6 one times ten to the minus six per year, that is a
7 very high intersection probability claim, the expected
8 number of dikes would be 13 and the probability of at
9 least one penetration, as you can see, 0.999998.

10 These results are not consistent with the
11 exploration evidence because no dikes have been found
12 in the footprint. Claims of high intrusion
13 probability failed as test over the 13 million years
14 time scale.

15 Next slide. Let's look at some younger
16 basalts. On the left is a vent complex in Pliocene
17 H in Crater Flat. At right is Black Cone, which is a
18 Pleistocene volcano dated around one million years.

19 And this series of cones -- Northern Cone,
20 Black Cone, Red Black, Blue Cones, these are all dated
21 around one million years. No features like these
22 exist on Yucca Mountain.

23 An important point to make is that
24 preservation of exposed basalts in southern Nevada
25 depends on their age and topographic setting. Miocene

1 and Pliocene basalts have been found in local basins
2 buried by alluvial basins.

3 Partial burial has been reported for
4 Pleistocene basalts. But they are too young to be
5 completely buried, even in basins. Next slide. Now,
6 you've also seen this slide before.

7 To further analyze the probability of
8 volcanism intersection we require NRC's PVHA code,
9 version two. And we analyzed the ten datasets that
10 have been published with that code.

11 Here's an example graphic from Connor et
12 al., 2000 in the Journal of Geophysical Research.
13 This slide shows the spatial recurrence rate contoured
14 for the Yucca Mountain region.

15 It's based on event cluster modeled that
16 uses a kernel function. It has built in either the
17 use of Gaussian or Epanechnikov code kernel function
18 that produce similar results.

19 It's also based on locations of Quaternary
20 volcanism for this particular case and information
21 about the density of the earth's upper crust. To
22 learn more about the code, I would refer you to that
23 JGR paper in 2000, also to a report by CNWRA by Laura
24 Connor et al., 2002.

25 Next slide. This slide summarizes our

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1 results using all ten datasets. And they are
2 described briefly in the left-hand column. These
3 datasets represent various patterns and ages of
4 volcanism.

5 The top file, all 64 events, you can see
6 that in the back, covers a region that includes parts
7 of Death Valley. It also includes some magnetic
8 anomalies that are assumed to be buried volcanoes.

9 The bottom dataset includes just the eight
10 known Pleistocene events. Eight of the datasets will
11 include five to 15 magnetic anomalies that are
12 assumed, generally without proof, to be volcanoes.

13 This makes for a robust analysis. This
14 incorporates a lot of uncertainty about the
15 possibility for buried volcanoes. For each dataset we
16 evaluated the recurrence rates in the Yucca Mountain
17 region that were required to produce repository
18 intersection rates of ten to the minus eight, ten to
19 the minus seven, ten to the minus six per year.

20 As shown in the far right column, a mean
21 rate of ten to the minus six per year prevailed in the
22 last billion years, 42 to 96 volcanoes would have
23 erupted in the Yucca Mountain region.

24 In reality, only eight events occurred
25 during all the Pleistocene, which is 1.0 million years

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1 long. That's a recurrence rate of just 4.4 per
2 million years.

3 Now, if we divide these numbers by ten to
4 reduce a time scale to the last 100,000 years, at ten
5 to the minus six per years, the expected number of
6 volcanoes is four to nine.

7 But there's only one, Lathrop Wells,
8 event. We can see that claims of high probability of
9 intersection, fatal tests of volcanic recurrence, and
10 time scales with million years and 100,000 years.

11 Now, something more should be said about
12 this because PVHA results shown here are based on a
13 Gaussian model modified to include crustal density
14 effects.

15 And you heard the discussion about that.
16 This approximately doubles the dike intrusion
17 probability at Yucca Mountain. However, gravity
18 weighting isn't limited.

19 The number shown here would double, which
20 is an extraordinary number of volcanoes. And, in
21 fact, we do recommend not using this weighting factor
22 of several reasons.

23 It is highly subjective. No basis has
24 been demonstrated for including it. Also, the kernel
25 estimator has already quantified the degree of

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1 clustering of the volcanoes.

2 The crustal density information simply
3 provides a partial geologic explanation as to why the
4 clustering occurs where it does. Finally, it should
5 be said that the seismic tectonic regime represented
6 by the density map, based on gravity, probably
7 reflects the much higher extension and volcanism
8 rates, both Miocene time and into part of the Pliocene
9 time.

10 Present day extension rates are
11 significantly lower. In other words, the primary
12 effects of the lower crustal density probably
13 manifested themselves long ago when the density
14 contrast was created.

15 The decline in volcanism over time
16 supports this interpretation. Next slide. The very
17 large recurrence intervals in the previous slides in
18 40 to 96 volcanoes per million years or four to nine
19 in 100,000 years or 80 to 192 in the last million
20 years without gravity weighting.

21 And perhaps the answer lays somewhere
22 between the sets of numbers. Let's look at other
23 volcanic codes that have this level of activity. And
24 the source of the slide is Chuck Connor, University of
25 South Florida.

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1 Also, one of the developers of the NRC
2 PVHA code. If ten to the minus six per year were true
3 at Yucca Mountain, then Crater Flat would be as active
4 as many of the volcanic fields in this table,
5 approaching one times ten to the minus four events per
6 year.

7 The volcanic field at Cima, California,
8 falls in this branch. Next slide. Cima is located
9 south of Las Vegas. Here's Las Vegas Valley. Here's
10 the location of the Cima field, to the south.

11 This volcanic range has more than 50
12 events and approximately 65 or more blows, covering an
13 area above 150 square miles. Next slide. Here we
14 have three panoramic views.

15 Crater Flat is at the top. And then there
16 are two views of the Cima field. About 30 of the
17 cones at Cima are Pleistocene in age, which means less
18 than 1.8 million years old.

19 Yucca Mountain and Crater Flat have not
20 experienced anything like this level of activity. If
21 they had, probably the best view for you to look at to
22 compare Cima is the bottom view, the widest panoramic
23 view.

24 You can see that the horizon is covered
25 with cones. You simply to not see -- your eye tells

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1 you this level of activity has not existed in
2 Quaternary time.

3 But, if it had, what should Yucca Mountain
4 look like? Next slide. This is a projection taking
5 roughly 35 to 40 events and placing them approximately
6 where they would arise.

7 A very high rate of volcanism had
8 occurred. You will see that there was, in this case,
9 a hypothetical impact at Yucca Mountain just once per
10 million years.

11 But what, in fact, do we actually see?
12 Next slide, please. Back to this figure. There were
13 eight events in Quaternary time. Only six of which
14 are here.

15 If you flip back and forth between those
16 two, just for a second, it is a dramatic difference.
17 I would say, where are all the volcanoes that should
18 exist if these very high rates had prevailed through
19 the last million years?

20 In arid to semi-arid climate of southern
21 Nevada is very hard to obliterate the evidence of
22 these very young volcanoes in Quaternary time. Next
23 slide.

24 What agree with comments made in the paper
25 Connor, et al. in JGR. Rates of basaltic volcanism

1 comparable to those in Cima or also seen the Colorado
2 Plateau volcanic fields, approximately 30 volcanoes
3 per million years have not occurred in the Pliocene
4 and Quaternary in the Yucca Mountain region.

5 And it is reasonable that the probability
6 estimates we calculate for the volcanic eruptions be
7 substantially less than those estimated for the
8 larger, more active fields.

9 Next slide, a recommendation. We would
10 recommend using the Quaternary recurrence rate to
11 estimate the frequency of repository intersections.
12 This has three advantages.

13 We are, of course, still in the Quaternary
14 period now. Compared to Pliocene time and certainly
15 compared to the Miocene time, the Quaternary best
16 represents the present day seismo-tectonic regime.

17 Also, the Quaternary fully captures the
18 most recent volcanism cluster of one million years.
19 This cluster represents five events or less. But we
20 consider also the maximum number that is somewhat
21 conservative.

22 The biggest advantage, it is a more
23 reliable recurrence rate. The uncertainty about the
24 number of Quaternary events is greatly diminished
25 compared to Pliocene events, certainly compared to

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1 Miocene.

2 There has been insufficient time to erode
3 or bury Pleistocene basalts. Next slide. Our
4 estimate from this review, we use the PVHA code and
5 the dataset of eight Quaternary events.

6 A Pleistocene recurrence rate -- that's
7 4.4 events per million years -- and zero gravity
8 rating. We estimate the intersection frequency at
9 five point four times ten to the minus eight per year.

10 Since the result is based on eight events,
11 you can get upper confidence bound, in this case 95
12 percent, using the Poisson distribution. Upper bound
13 is approximately one times ten to the minus seven per
14 year.

15 Next slide. Conclusions -- our analysis
16 raises doubts that a potential repository could be
17 penetrated by a dike once every million years. We
18 evaluated four time scales, as discussed.

19 And, at the 13 million year scale, non-
20 detection of basalts suggests an upper-bound
21 penetration rate of two times ten to the minus seven
22 per year, on an average over 13 million years.

23 At the one million year time scale, using
24 the PVHA code, it suggests 40 to 96 events to have
25 erupted in the region in the last million years.

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1 Without gravity weighting, that number goes up to 80
2 to 192.

3 But, only 80 events are known of all the
4 Pleistocene. Next slide. The results are especially
5 interesting for the 100,000 year time scale.

6 We contest a hypothesis that was discussed
7 by the expert elicitation in 1996. Is it possible
8 that the 80,000 year old Lathrop Wells cone was the
9 start of a new pulse of volcanism.

10 For a dike penetration rate of ten to the
11 minus six per year, the PVHA results indicate four to
12 nine events would have been expected in the last
13 100,000 years.

14 Without gravity weighting, we do dispute
15 the degree to which gravity -- you would expect eight
16 to 18 events. Only one is known. Our best estimate
17 for dike intrusion is more than ten times smaller than
18 the highest probability claims.

19 The future volcanism follows the
20 Pleistocene pattern. The probability of intersection
21 is 5.4 times ten to the minus eight per year using the
22 PVHA code.

23 Claims of greatly increased probability --
24 failed the simple test of reasonableness of four times
25 scales. Spatial temporal models predicting intrusion

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1 probabilities greater than two times ten to the minus
2 seven per year in the potential repository footprint
3 are overly conservative.

4 Along with ongoing work by the Department
5 of Energy, ongoing site investigation, our realistic
6 models will be developed by considering non-detection
7 of basaltic dikes in the potential footprint, and also
8 known patterns of Quaternary volcanism.

9 I have one item that's probably best taken
10 up in the discussion panel session. Listening to the
11 presentations earlier today, I see some evidence that
12 the NRC Staff approach to volcanism is not risk
13 informed.

14 In the presentations by Tim McCartin over
15 the years on performance assessment and the risk
16 informed evolution of that work, you have seen what
17 that can accomplish in other areas of the program.

18 The volcanism work that was done is not
19 part of the overall performance assessment. Numbers
20 were fed into performance assessment from that group.

21 And, particularly slide 13 Dr. Bill's, is
22 one that we may want to discuss in more detail. That
23 concludes our talk. Thanks for your attention.

24 CHAIRMAN RYAN: Thank you very much, Neil.
25 Any questions? Ruth, you had a question.

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1 MEMBER WEINER: Any chance that Neil can
2 partly answer the question that I have. And that is,
3 how does your estimate compare with what the
4 presentation of -- can you?

5 MR. COLEMAN: I believe it was mentioned
6 that the Staff currently have an assessment of
7 probability around ten to the minus eight to ten to
8 the minus seven.

9 But, with consideration of varied events,
10 they feel the probability could be as much as an order
11 of magnitude beyond that. However, that -- I would
12 just add -- that is not consistent with the record of
13 the last 100,000 years.

14 The simple tests may be enough to reject
15 this extreme tail of the probability distribution.
16 But it does not seem to be any evidence for events for
17 probabilities of intersection greater than two times
18 ten to the minus seven per year.

19 MEMBER WEINER: So, just to simplistically
20 repeat what you just said so I understand it, what
21 you're saying is the tail of their distribution is not
22 supported by the record.

23 MR. COLEMAN: Right, we do not see that
24 extreme. We do not see any evidence to support that
25 extreme end.

1 MEMBER WEINER: Thank you.

2 CHAIRMAN RYAN: Just one quick question.
3 And I guess I'm actually asking this out of ignorance.
4 Why did you pick a Poisson distribution over any other
5 to use as your model.

6 MR. COLEMAN: There were other
7 distributions that could be used. That one has long
8 been used in earth sciences for evaluating events,
9 including clustered events of low probability.

10 It has been used in earthquake analysis,
11 as well as volcanism.

12 CHAIRMAN RYAN: It's used in radioactive
13 too. But, I mean, is it a standardized model of how
14 to model these geologic events, is that what you're
15 saying?

16 MR. COLEMAN: Yes, it is commonly used.

17 CHAIRMAN RYAN: Thank you. Okay,
18 questions from the panel members or other
19 participants?

20 (No response.)

21 CHAIRMAN RYAN: Other questions from
22 Staff, or the audience? Yes?

23 MR. HINZE: A quick question. If I
24 understand you correctly, you are suggesting that, in
25 the Connor and Hill paper 2000, that the idea of an

1 Amargosa low, gravity low, is due to decompression
2 that is because of lower pressures involved.

3 It is not a viable hypothesis for the
4 concentration of volcanic activity. Is that what
5 you're saying?

6 MR. COLEMAN: I don't think that's quite
7 what I said, but --

8 MR. HINZE: But, you were suggesting that
9 the use of the gravity weighting was inappropriate.

10 MR. COLEMAN: That's absolutely right.

11 MR. HINZE: And, the reason that they use
12 the gravity weight was because they had to had did
13 hypothesis -- if I'm understanding it correctly --
14 that it speed compression effects that are localized
15 in that area.

16 So you're -- being complacent -- too
17 insufficient to cause volcanic activity.

18 MR. COLEMAN: No, I would not suggest that
19 at all.

20 MR. HINZE: What do you suggest.

21 MR. COLEMAN: I essentially agree with the
22 decompression modeling. There are a lot of discussion
23 and debates about relative depth, the rise of the
24 magmas in this region.

25 But, the idea is that the density map that

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1 we see today -- most of it was created long ago, in
2 Miocene time and into part of Pliocene time, and
3 represents these crustal deformation effects produced
4 by the very high rates of extension.

5 Then, to use it to modify the distribution
6 for trying to project future volcanism, in the current
7 Quaternary period, makes no sense. What does make
8 sense is it's a partial explanation for why the
9 continued volcanism is there at much lower rates.
10 Have I answered that for you?

11 MR. HINZE: I understand where you're
12 coming from now. Let me ask you another thing about
13 your comments about using only -- focusing on the
14 Pleistocene events to achieve a more robust analysis.

15 One of the reasons why I very much like to
16 see us extend the area of volcanism that is involved
17 is because, in this extrapolation, you need a large
18 number of events, and, if you're going to have a
19 robust analysis.

20 And, by including the Pliocene, what
21 you're doing is you're increasing the robustness of
22 the determinations. Is that not correct?

23 MR. COLEMAN: I believe that is correct.
24 But there are reasons why we would suggest using the
25 Quaternary grid. From a regulatory point of view --

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1 something for the Staff, the Committee, and others to
2 consider -- there is great power in approaches that
3 can dramatically reduce some of the uncertainties.

4 And the whole question about buried events
5 and their effects on the probability essentially
6 vanishes. If the events that we're looking at in
7 Quaternary -- we have a very high confidence in what
8 that recurrence rate is.

9 But this other point, which I actually
10 read about in the reports from the CNWRA folks, that
11 the Quaternary events actually yield a somewhat higher
12 probability because, as a group, they are somewhat
13 closer to Yucca Mountain.

14 So, I would submit that it is the robust
15 analysis in that sense. And, in a way, it partly
16 responds to the model that has been submitted by Jean
17 Smith, a model that talks about the pollution in
18 volcanic fields and where new events might occur in
19 the periphery of others.

20 This actually allows for somewhat of a
21 migration slightly closer to the sight. And that is
22 the reason that you see slightly higher probability,
23 but still very low and far below the extreme tail that
24 was presented earlier.

25 MR. COLEMAN: Thank you.

1 CHAIRMAN RYAN: Any last questions? Yes,
2 please.

3 MR. MELSON: Yes, Bill Melson, can you
4 make any comments about how low the probability is?
5 You sound pretty clearly talking the tail off the high
6 end. What do you do at the other side?

7 MR. COLEMAN: I don't have the figure here
8 with me. But, when we take our central result and use
9 the same Poisson -- the test for determining
10 confidence intervals -- I will get that for you.

11 I suspect that the number will be slightly
12 below ten to the minus eight per year. But, the
13 results from -- the results shown on my slide fifteen
14 would suggest that ten to the minus eight per year is
15 too low, that we had more events than that in the last
16 million years.

17 So, regardless of -- I think the best way
18 to answer your question is, I still see evidence that
19 the probability is somewhat higher than ten to the
20 minus eight per year.

21 So, therefore, the consequences of low
22 would need to be considered, as they will be in the
23 next session.

24 CHAIRMAN RYAN: We'll convene if there are
25 no other comments or questions. Oh, yes, Tim

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1 McCartin.

2 MR. McCARTIN: Yes, Tim McCartin, NRC
3 Staff. I would like to clarify something for the
4 Committee, the performance assessment effort, as well
5 as the development of risk insights and risk informing
6 the NRC process.

7 It has been a team effort. And, in my
8 opinion, the igneous activity is not a separate
9 activity that was done offline.

10 CHAIRMAN RYAN: Thanks. Any other
11 questions or comments? We'll reconvene our afternoon
12 session promptly at one o'clock. Thank you very much.

13 (Off the record for a lunch break.)

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1 A-F-T-E-R-N-O-O-N S-E-S-S-I-O-N

2 SESSION ONE WORKING GROUP ROUNDTABLE DISCUSSION

3 CHAIRMAN RYAN: Okay. The first thing
4 this afternoon is a panel discussion with five
5 individuals, Dr. William Melson, Dr. Bruce Marsh, Dr.
6 William Hinze, Dr. David Johnson, and Dr. Robert
7 Budnitz.

8 Let me take them in reverse order of
9 what's on my agenda. We'll start with perhaps Dr.
10 William Melson. Can we have your comments, your
11 thoughts?

12 What have you heard? What should we
13 listen to?

14 MR. MELSON: Well, as Michael said, I'm
15 Bill Melson. I'm a curator at the Smithsonian. I've
16 worked with the TRV since about 1889, the volcanic --

17 CHAIRMAN RYAN: Since 1889?

18 MR. MELSON: I'm sorry, 1989.

19 (Laughter.)

20 MR. MELSON: My comments on the morning
21 session are generally that I thought it went very
22 well. Quite frankly, it's not adding a lot to what we
23 already know.

24 But, I think we've come along further.
25 And yet, I do wonder about what we can learn by

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1 examining some of the PVHA issues again. I think
2 Bruce's comments about you have a small dataset that's
3 been looked at in many different ways and, if it is
4 change, will it really see the uncertainty limits on
5 what we've already done.

6 So, for better or worse, I would think we
7 ought to look at that very carefully if it's not too
8 late about going ahead with it, and be sure there's a
9 very strong feeling and rationale as to why it needs
10 to be redone.

11 I was very gratified by the comment or
12 presentation by Neil Coleman and actually using the
13 repository as an experimental body to look for to use
14 it to look at the big frequencies or likelihood of
15 dike injection repository.

16 That was new. And, it certainly is
17 consistent with staying fairly low probabilities of
18 intersection. It doesn't, to me, raise any new flags
19 that we need to be concerned about.

20 I think Bruce Crowe's comment son drilling
21 and sort of finishing up some of the work of the
22 anomalies very near the site would be well worthwhile.

23 I think, for now, that's about all I have
24 to say.

25 CHAIRMAN RYAN: Okay. Well thanks, that's

1 a great start. We appreciate your comments. Dr.
2 Bruce Marsh, sir?

3 MR. MARSH: Yes. I was very pleased with
4 the morning's presentations. And one of the things I
5 was particularly struck by also was the fact that it's
6 always a big problem in geology.

7 We look at layers upon layers and layers
8 of things that have happened in sorting through those.
9 But that's really not our firmament. That's really
10 what we have in the record, the historical record, the
11 geology when we look at it.

12 So, one of the things that I don't think
13 has been emphasized enough -- it came up in Bruce
14 Crowe's comments -- is that the tectonic development
15 in the area, the history of that, can be read pretty
16 carefully because we have ash loads and we have
17 erosional surfaces, and we have fault histories and
18 things, and questions, for example, of whether or not
19 this block that Yucca Mountain's on, and that whole
20 area, is still structurally attached to what's going
21 on in the basin to the west of it.

22 It is a very important issue. And there's
23 a lot of cogent things that can be said about that. A
24 lot of the style of what caused the tectonic style
25 that basically encouraged the volcanism and gave rise

1 to what see today was set up in the Miocene, 15
2 million years ago or something.

3 And yet, when we look today at these
4 things, it's like looking at the heat flow. The heat
5 flow of the surfaces is reflecting the thermal
6 conditions in the crust ten million years ago.

7 We can become confused a bit by that in
8 thinking that, you know, we're in the middle of an
9 onslaught of something new. So, it's very nice to
10 carefully sort out that and realize what kind of
11 environment we are in today and to look at that in
12 terms of the last one million years, two million
13 years.

14 And so, it's very important to put the
15 geology into the models carefully -- topography,
16 what's in the basins, what structural units are
17 talking to each other, and which ones aren't.

18 The deeper we go -- there's an interesting
19 phrase by -- I believe it was Francis Birch -- who
20 said that it's interesting, the deeper we go in the
21 earth, we know less and less, but our description of
22 it becomes more and more exact.

23 And, this is what happens a lot. We
24 actually go down in the mantle and say, well, we're
25 melting the mantle, and we have fertile mantle,

1 depleted mantle.

2 We have a thermal pulse here. We have a
3 small plume. We have thermal convection. Really,
4 objectively speaking, I mean, we have a hard enough
5 time understanding how a volcano that's about to erupt
6 is going to erupt.

7 And we have no chance of actually using
8 any of that deeper information. So, in other words,
9 in putting -- we tend to use that Poisson distribution
10 for time and for spatial events.

11 If we go down deeper in the crust we know
12 that basically we have an exponential decay of what we
13 understand. In only using the geology of things, we
14 understand very little to be used in a predictive
15 model as we go deeper into the earth.

16 And so, there's a cut-off. We should use
17 that. We should put stuff, and model it, we really
18 know something about, and ignore stuff that's pretty
19 below the horizon in terms of being able to
20 scientifically say cogent things about it.

21 So, the other thing that's an interesting
22 thing is that, I think, you know, at least we're all
23 in the same room, in terms of we don't have -- there
24 are disparities.

25 But I think they can be brought into line.

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1 And I think the interesting thing about it is that
2 there is sort of a common sort to Socratic element
3 here that can be used to adjust each other's points of
4 view or to convince one another whether or not we
5 should take certain bounds seriously or not.

6 The idea of looking at things between ten
7 to the minus eight and ten to them minus seven, and
8 just agreeing that that window -- not worrying about
9 so much where we are in that window, would be a very
10 interesting way to approach these problems. Thank
11 you.

12 CHAIRMAN RYAN: Thank you. Dr. Hinze?

13 MR. HINZE: I enjoyed this morning
14 because, one of the reasons I think is that, as a
15 result of this morning, your job is going to be less
16 difficult than perhaps it could have been.

17 There's a certain amount of unanimity in
18 the conversations that we heard, that we don't have
19 all the answers. Bruce added on that very well. We
20 don't have all the answers.

21 We're not going to have all of the
22 answers. And, Bruce Crowe and Britt Hill both
23 commented on the fact that we're not going to mover
24 people very far from their models.

25 But the point is that the models aren't

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1 making -- the difference in the models are not making
2 that much difference. Now, I think, in terms of the
3 probability, ten to the minus seven, ten to the minus
4 eight, and 1.1 times ten to the minus seven or two --
5 I don't think we are smart enough to worry about those
6 type of things.

7 I think we have to keep a pretty broad
8 swab here. Let me say a few things about the
9 recurrence rate. What we're dealing with here is a
10 science where we're dealing with a situation where we
11 don't have precursors that are in the right timeframe.

12 We have only the very basic knowledge of
13 the -- Bruce we only have a very fundamental
14 knowledge, basic knowledge of the physics or the
15 geological control.

16 I'm sure that was said several times here
17 today. And so, what we have to do is we have to
18 extrapolate from what we do see. And extrapolation
19 means that we're going to need to deal with
20 probabilities, which we are.

21 We have uncertainties -- and certainly we
22 do have those. How do we cut down on those
23 uncertainties? Where are the points where we can go
24 in and cut down on those uncertainties?

25 I don't think there are many points that

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1 we can go to to cut down on the uncertainties, the
2 definition of igneous event or the length of the dike,
3 etcetera.

4 And it turns out, as Bruce and others have
5 suggested, that this is not making much of a
6 difference in the results. But one thing that can
7 make difference is the number of igneous events that
8 have occurred within the last timeframe.

9 I personally would like to see a timeframe
10 that extends to four or five million years. And I
11 think that's backed up by the ten independent
12 scientists that worked in the PVHA.

13 And, we have been saddened with inadequate
14 way to look at these past events. The 1999 survey of
15 the USGS solved the purpose of not accounting, and
16 perhaps some others.

17 But it didn't solve at all the problem of
18 the events that may be hidden in the -- beneath the
19 alluvium in particular. And so, the DOE, I think very
20 appropriately, has set -- embarked upon this new
21 magnetic survey, which we've just seen the first light
22 of.

23 You have to realize that there are -- I
24 hope I'm not duplicating what Britt said this morning.
25 But, there are basically three types of magnetic

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1 anomalies that we are observing in the area.

2 First we are dealing with long wavelength
3 anomalies that are derived from rather massive
4 structures within the rocks and pre-cambrian rocks.

5 And those are long wavelength and should
6 be able to be discerned pretty well. But, they are
7 going to overlap in spectrum with the magnetic
8 anomalies that are buried within the rather deep
9 alluvium.

10 The second type of anomaly is the anomaly
11 due to the permanent and susceptibility, magnetic
12 susceptibility, permanent magnetization and magnetic
13 susceptibility of the tuffs.

14 These will produce anomalies that --
15 particularly where they are faulted or whether it's
16 been structural disruption or some variation. And,
17 finally, we have the basaltic rocks that we are
18 interested in.

19 The problem is -- one of the problems is
20 that the latter two types of anomalies may give
21 somewhat the same signatures. And so, we have to be
22 smart enough in analysis.

23 And we must have the right data in order
24 to differentiate that. Ideally, the specifications of
25 the magnetic survey were such that we could make great

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1 strides.

2 Due to an unfortunate set of
3 circumstances, some of the data is going to be
4 degraded -- is degraded -- from what the DOE wished to
5 have.

6 And that's going to have a serious impact
7 on the results. The DOE has very correctly attempted
8 to -- or is going to attempt to differentiate between
9 these two types of magnetic anomalies that have
10 somewhat the same spectra.

11 That is the tuffs and the basalts. By
12 feeding them against the electromagnetic response, and
13 in this way attempt to identify the higher
14 susceptibility basaltic rocks.

15 I'm going through this because I want to
16 make it clear -- at least in my mind -- that the
17 results of this new survey are going to have an
18 impact, could have an impact.

19 But it isn't guaranteed at all that it's
20 going to have an impact. There are many problems in
21 interpreting these data. And one of them is
22 especially the above mean terrain clearance, which has
23 been degraded some bit, especially in the rich areas.

24 But, also, there is an overlap in the
25 susceptibilities between the tuffs and the basalts,

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1 which may make the EM impossible to differentiate. In
2 addition to that, there are some conductive zones,
3 alteration zones, and fault zones where there's
4 alteration as well, that are going to complicate the
5 interpretation by trying to differentiate basalt from
6 the top using the EM data.

7 My own very quick review of the data is
8 that there's nothing that comes out and bangs you in
9 the face and says this is obviously going to change
10 the probability, the recurrence rate, in a quick look
11 at the data.

12 But there are a lot of very interesting
13 anomalies. And there are a lot of interesting
14 anomalies, particularly to me in Jackass Flats that I
15 think could have an impact upon the PVHA if the PVH
16 goes in, as I understand it is.

17 But, it's going to take time, and it's
18 going to take some effort. I think that prejudging
19 the aeromagnetic results based upon the quick look
20 that we had yesterday morning and yesterday afternoon
21 is very -- it doesn't give credit to the DOE, nor
22 their efforts to come to resolution on this.

23 So, the recurrence rate, which is the
24 major way we can get an uncertainty in that
25 probability factor, was going to be able to decrease

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1 with the new set of -- the interpreters will do that.

2 But, it's going to be a difficult process.

3 And it's going to take some time, some effort, some
4 resources. I guess I'll leave it at that.

5 CHAIRMAN RYAN: Okay, thanks Bill. Dr.
6 Johnson?

7 MR. JOHNSON: Thank you. I think first I
8 should say a few words about my background. I'm not
9 a geophysicist. My field is in developing
10 probabilistic formats and methods to support the
11 decision making.

12 So, from that point of view, I hear the
13 presentations about whether or not the frequency of
14 intrusion is ten to the minus eight or ten to the
15 minus seven, or even six.

16 From my background, my experience, and
17 again not knowing much about the chronological issues,
18 those tend to be in violent agreement in my mind.

19 That said, I think it is important. I
20 think it was said earlier that it would be a useful
21 exercise to have the experts go and try to present
22 their findings, if you will, in a format of a
23 probability of frequency format so we understand what
24 their key assumptions are and how they affect their
25 results.

1 This is obviously -- would be useful for
2 more fundamental understanding of what's going on, but
3 also as new information is derived in the future and
4 issues pop up.

5 It might provide a pretty sound basis for
6 quickly reacting to those sorts of events. I am kind
7 of waiting for the so what to all of this. I am
8 anxious to see what the scenarios look like from an
9 initial condition to the final end states of the
10 analysis.

11 I think once we have that in hand we can
12 then go back and make judgments from judgments on
13 whether or not our understanding of the frequency of
14 volcanic intrusion is something we need to focus more
15 on.

16 I do think that there is some -- for
17 investigations of some of the near field anomalies
18 that would make a lot of sense to resolve. I think
19 I'm waiting to see what the big picture looks like
20 before I go on any further. Thank you.

21 CHAIRMAN RYAN: Thank you very much. Dr.
22 Budnitz?

23 MR. BUDNITZ: I'm Bob, Budnitz. I'm
24 Lawrence Livermore Laboratory. But I'm on detail to
25 the Yucca Mountain project DOE. So I'm here with a

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1 DOE hat on.

2 But, I need to give you some background.
3 I'm not an earth scientist at all, never mind a
4 volcanologist. But, you know, it comes in by osmosis,
5 so I now know about ten to the minus four about these
6 other stuff, which is a hell of a lot more than ten to
7 the minus eight.

8 I want to tell you something about the
9 status and explain why DOE is in here. The Department
10 has written the license application. We're sending it
11 in December.

12 I imagine you will have it by then. It's
13 only three months away. And right now it is an
14 intense review, everything, not just the igneous
15 piece, everything.

16 It is an intense review for consistency
17 and to make sure that we do the validation and the
18 quality assurance checks, and make sure everything
19 that we're going to send in in December hangs together
20 into a coherent application.

21 I'm sure you understand that. And,
22 because that process is right now in its final stages,
23 we found ourselves not in a position of being ale to
24 talk too much about the details because it's just now
25 coming together into something that's final.

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1 But, I insist on December 15th it's going
2 to go in, as we say. And we're going to put it in.
3 And, when that's true, it's going to be a public
4 document.

5 And anybody in the public can read and
6 review it. Certainly if you were submitting it -- the
7 regulatory Commission, the Staff and ultimately the
8 Commission for a review for -- to get a construction
9 authorization.

10 But, it will be in the public domain.
11 And, at that time, anybody who wishes to review it
12 will be able to do so. I have two or three things to
13 say about the license application that are relevant to
14 what we heard about this morning.

15 First off, everything we've done in the
16 license application is risk informed and, in parallel,
17 responsive to the Yucca Mountain review plan, which is
18 the NRC's -- you know, the Staff review plan.

19 We know they are going to review it again
20 some time. And so, which means because we have to be
21 responsive with the Yucca Mountain review plan, some
22 of the stuff that is in the license application isn't
23 risk informed because the review plan isn't
24 necessarily risk informed, although our criteria in
25 the end is.

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1 You know, the part 63, individual dose
2 based criteria is, of course, dose informed. It's not
3 risk informed, it's risk based. So, because of that,
4 the license application analysis is intrinsically
5 probabilistic through the -- because our regulation is
6 probabilistic.

7 So, what you're going to see is a
8 probabilistic analysis intrinsically -- that's just
9 the way it is -- modified by the fact that, of course,
10 we do have to respond to the Yucca Mountain review
11 plan, which means a whole lot of stuff is in there
12 that is supportive -- or in some cases we review other
13 things that aren't.

14 You can imagine what that means. And, of
15 course, we have to be attentive to the technical
16 issue, you know, the agreements that we made. And, I
17 guess, in that sense, we just look forward to
18 submitting it.

19 And, somewhat later, the ACNW along with
20 the Staff will have a meeting like this in which we
21 can discuss what we've done, which we've just --
22 position to talk about here.

23 A couple of other things that are very
24 important to say. And that is, although the work to
25 support license application is done by definition --

1 we have not stopped work in the igneous area.

2 People know that. The aeromagnetic work
3 that was done from March until June is just now being
4 analyzed, and will be available perhaps the next six
5 to eight weeks for public review.

6 We'll let the NRC review it at that time.
7 And, after that, there's a plan which was discussed at
8 yesterday's NRC meeting, to do some drilling of
9 several of the sites.

10 Exactly what drilling will be done hasn't
11 been decided yet. We're going to have to sort out
12 exactly which targets and we don't have enough money
13 to drill a thousand of these things.

14 We're just going to drill a few of them.
15 And, how to select those, is a difficult choice
16 between different agendas. Secondly, and I suppose
17 many of you know, but I should tell the rest, we are
18 beginning a new PVHA, probabilistic volcanic hazard
19 analysis.

20 The first meeting to kick that off is in
21 the second week in October. It's the data needs
22 workshop in which the data needs for the PVHA are
23 going to be discussed amongst the experts.

24 And that will kick that off. The PVHA,
25 the new or revised, is due to be completed in the

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1 first half of the fiscal year '06, a year and a half
2 away.

3 And, both of those -- and, of course, PVHA
4 is supposed to integrate a whole lot, both of those
5 are confirmatory in nature when we do. That is, we
6 believe the license application is strong enough as it
7 is, but we're doing confirmatory work because, as we
8 must, we are going to continue to do that work over
9 the years.

10 You never know whether you find something
11 that doesn't confirm with the expected. And we're
12 going to proceed on that basis, and challenge the data
13 and assumptions and so on.

14 And, as other work may emerge that needs
15 to be done over the future years, we will consider
16 doing that too. We just don't know what that would
17 be.

18 So, I'm just here to tell you that we're
19 very close to having something that everybody will be
20 able to look at and review. It will be a public
21 document with the license application, with all the
22 supporting data and everything else that supports it.

23 We are proceeding with more technical work
24 now. And, whether more than that is going to be
25 needed, we just don't know yet. We're going to let

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1 those chips fall where they may as time goes on.

2 CHAIRMAN RYAN: Thanks very much. Any
3 comments from the Committee or the panel? Ruth?

4 MEMBER WEINER: Just a brief comment on
5 Dr. Budnitz's last comment. If you're beginning new
6 PVHA, and I assume that means the new expert
7 elicitation --

8 MR. BUDNITZ: Yes.

9 MEMBER WEINER: What kind of differences
10 do you expect to happen?

11 MR. BUDNITZ: We have no idea until it is
12 done. We just don't know. The nature of this is it's
13 a scientific investigation, like they all are. And,
14 how it comes out will depend on how it comes out.

15 I'm not ducking that question. I
16 literally couldn't say, because we have an open mind
17 as to what the data will -- how it will be understood
18 and what models will be used.

19 And who's going to argue with who about
20 what?

21 MEMBER WEINER: What was the primary
22 driver for this? I mean, I'm just curious, because
23 it's late in the day.

24 MR. BUDNITZ: The last one was seven years
25 ago. And a lot more is understood now than then. And

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1 so, we believe that we will be in a better position
2 than we would otherwise be by doing it.

3 Otherwise, we would have nothing but that
4 old thing and some patches here and there. Instead,
5 we're going to have a -- how should I say it -- a
6 coherent, consistent PVHA that is intended to
7 integrate all of this into a framework that we believe
8 the community will participate in and endorse.

9 CHAIRMAN RYAN: Thank you. Yes, sir?

10 MR. MELSON: Bill Melson. Bob, will the
11 DOE's volcano assessment be close to what you've put
12 out in, I think, January 9th of this year? We have
13 gone through that and it seems like a pretty strong
14 document.

15 So, is that, to your knowledge, what's
16 going to go ahead?

17 MR. BUDNITZ: You're asking me to part
18 with something that I'm not willing to do?

19 MR. MELSON: Well, I'm just wondering,
20 because that gives us a preview, I suspect.

21 MR. BUDNITZ: You can peak if you want.
22 You'll know on December 15th. I'm not ducking. It's
23 just that it's hard to respond.

24 MR. MELSON: Yes.

25 MR. BUDNITZ: Whether something is close

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1 to, of course not to be completely different, how
2 close it is? You know what I mean?

3 CHAIRMAN RYAN: Questions from the Staff,
4 other comments? I guess I want to try and summarize
5 a bit if I can. I know it's a daunting task,
6 particularly for a non-geologist.

7 My geology mentor on the ACNW, Dr.
8 Hornberger told me geology is easy. They always want
9 to dig one more hole. That seems to be the case
10 today.

11 I guess that's one of two important
12 elements. The aeromagnetic data seems to be a
13 critical issue. I think, Dr. Crowe, you suggested
14 some drilling and some value that could be acquired
15 through that drilling.

16 I've heard three or four folks endorse
17 that idea, that that might actually help reduce some
18 uncertainties. And then I think the theme that we
19 really haven't touched on, and I would like the other
20 panel members to talk about, is -- except for David
21 who mentioned kind of a more formal probabilistic
22 assessment here.

23 I think Dr. Garrick mentioned that earlier
24 in the morning, that a more rigorous treatment of
25 probability analysis or a probabilistic approach,

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1 maybe a Bayesian approach to the event side of things,
2 maybe in line with some of the things that Neil
3 Coleman presented, might be crucial.

4 It's kind of the third leg of the three
5 major components I heard this morning. It might be
6 different or enhancements, or improvements to what we
7 know now.

8 Is there any reaction to that? I mean,
9 could any of you talk a little bit more about the
10 probabilistic approach?

11 MR. GARRICK: Well, I want to follow-up
12 with what you said, because it might make a
13 difference. I was curious about the new PVHA. And
14 Bob said that we've learned a lot more now, and we'll
15 want to incorporate that.

16 And, I had a couple questions. One, is
17 the same team that did the PVHA one going to do PVHA
18 two?

19 MR. BUDNITZ: Eric Smithstead from the DOE
20 Staff in Las Vegas, I think, can an answer that
21 question.

22 CHAIRMAN RYAN: I'm sorry, John, maybe you
23 can repeat your question so everybody can --

24 MR. GARRICK: Yes, I was very curious
25 about the second time around. Bob indicated that

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1 we've learned quite a bit. And that's one
2 justification for the second time around.

3 I wondered if the same team that did the
4 PVHA one is going to do PVHA two.

5 MR. SMITHSTEAD: Right. Eric Smithstead
6 at DOE, that's what we're looking at right now, is
7 trying to reassemble the same team. We won't get
8 everybody back. But we'll have the majority of them.

9 MR. GARRICK: Thank you. While I have the
10 microphone, I wanted to ask Bill Hinze a question.
11 Bill, are you awake?

12 MR. HINZE: With you talking, John, how
13 can I help it?

14 (Laughter.)

15 MR. GARRICK: You mentioned a couple of
16 categories, some things that you thought ought to be
17 done, but probably wouldn't make much of a difference
18 with respect to the probably and some things that you
19 think ought to be done that will make a difference.

20 Can you elaborate on that a little bit as
21 to why we want to do the things that aren't going to
22 make a difference? How would you prioritize what we
23 should do?

24 MR. HINZE: Well, we wanted to decrease
25 uncertainty. And I think that's one of our functions.

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1 And, the major uncertainty in the PVHA was recurrence
2 rate.

3 And that very well typifies the fact that
4 in reducing uncertainty, that we want to look at the
5 number of volcanic events in the last million, two
6 million, five million years, six and seven.

7 And so, that has to be done or -- the
8 process of being done. I also pointed out that -- and
9 this is not my original thought -- that there are some
10 particular areas on that map that are interesting in
11 the surrounding area.

12 And one of those is Jackass Flat. And the
13 reason for that is that if we had the Quaternary
14 volcanism jump across the ridge on Crater Flat to the
15 other side of the repository, we would be -- I think
16 this would cause contemplation on the part of any
17 analyzer of the data.

18 I am reminded of Mike Sheridan's comment.
19 At the last appendix seven meeting on the aeromagnetic
20 that -- July of last year, as I recall -- Mike was the
21 only one at that meeting that was part of the PVHA.

22 And Mike stood up and -- paraphrasing him,
23 he can speak for himself usually -- if we knew
24 volcanic sediments were found to the east of the
25 repository and extension to the south, that he would

1 need to reanalyze his PVHA position.

2 MR. GARRICK: I always find that --

3 MR. HINZE: Did I do it correct Mike?

4 MR. SHERIDAN: Correct.

5 MR. HINZE: Did I read it right?

6 MR. GARRICK: I was trying to get at what
7 you would consider to be the biggest action that would
8 give us the biggest bang for the buck.

9 MR. HINZE: Exactly.

10 MR. GARRICK: Yes. And I want to take the
11 opportunity to indicate that, in the category where
12 you said it wouldn't change the probability much, but
13 it would change the uncertainty, it certainly changes
14 the risk.

15 And we want to make that distinction. So,
16 both categories have substantial impact on risk.

17 MR. HINZE: Yes.

18 CHAIRMAN RYAN: Any other comments from
19 participants? Yes, Dr. Marsh?

20 MR. MARSH: One that that's a little bit,
21 I think should be of some concern is that DOE submits
22 its application, the way that they've assessed, or
23 estimated, or come to grip with the probability
24 hazards for volcanism, is basically using panel
25 experts.

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1 And they have built into it their
2 knowledge, geology background, exhibit with volcanism,
3 etcetera, a series of series of estimates.

4 And that's really a substantial amount of
5 experience. But, it's not a computer program. On the
6 other hand, what we've heard this morning, the Center
7 and the NRC has a program, Connor et al.

8 And that has built into it several things
9 on various premises. And, that's almost a different
10 language than the other. So, we're going to use one
11 set of principles to evaluate another set of results.

12 It's almost passing in the dark. It could
13 be. In other words, you could be speaking different
14 languages. And so, I think the in-between land --
15 worrying about where all the geological influences
16 that the various experts would use to modulate their
17 results, where do those exactly fit into a computer
18 program or a program that someone would have?

19 Where are the analogs? Where do these
20 things go in, you know, in layers that we can put in
21 and take out? And, I think, to do an effective
22 evaluation, you really need to have that expertise
23 built in or you have to have that flexibility in the
24 evaluation.

25 For example, this focus on putting

1 topography in and out, but, the mantle in and out --
2 slide these filters in. Otherwise, I really wonder,
3 you know, how will anybody do an effective evaluation
4 in the DOE program.

5 CHAIRMAN RYAN: Dr. Budnitz?

6 MR. BUDNITZ: I just want to be sure to
7 point out that the PVHA, the structure, if it is
8 executed properly, will do just that, as it is
9 intended to be, by structure, a form for such
10 explorations among the experts who bounce some things
11 off of each other, and considering literature that
12 isn't in the room.

13 And they arrive at a common understanding
14 of all the underlying data and all the different
15 models that explains those data, in order to deduce
16 what is sort of the best you can do.

17 I don't know of any better structure than
18 that to do that. In the end, there are -- that is, to
19 structure such a way to pull out what the community's
20 knowledge is and the different approaches to it.

21 And, if it is successful, why there won't
22 be any stone left unturned. Although, of course, the
23 experts themselves are the ones that have to sort out
24 which are the important and which are the less
25 important issues, which models may -- while they fit

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1 the data -- don't make sense for another reason,
2 whatever comes to mind.

3 I know something about this because the
4 methodology used in the PVHA is called the SSHAC
5 methodology. SSHAC stands for the Senior Seismic
6 Hazard Analysis Committee that developed methodology
7 the probabilistic seismic testing analysis.

8 And I plead guilty to having chaired the
9 SSHAC committee for three years. So I think I
10 understand how all that works. And, if it is
11 successful, it will, in fact, not only allow, but
12 require the consideration of all the different models
13 and data we've got there.

14 CHAIRMAN RYAN: Thanks. I guess I'd like
15 to maybe turn a question to the NRC presenters from
16 this morning. And, you know, I took note on, Dr.
17 Trapp, your comment that you're on the vote of
18 reviewing an application.

19 So, you don't have the burden to come up
20 with the answers to all these wonderful questions we
21 thought up today. But, I wonder if maybe you could
22 talk about the following things, or Dr. Hill, either
23 one.

24 You know, we've heard a lot about
25 deterministic values, about bounding analysis, and now

1 a little bit of discussion about probabilistic
2 analysis and so forth.

3 Could you talk a little bit more about how
4 all that fits together in your mind from your
5 reviewing of potential application? Or is that too
6 broad of a question to dive in on?

7 MR. TRAPP: If you take a look at part 63,
8 it does require a risk informed analysis. It does
9 require that you go through the whole probabilistic
10 analysis to get to the end.

11 I'm not really sure how to answer your
12 question.

13 CHAIRMAN RYAN: Well, I guess I'm reacting
14 to a couple of comments that Britt made where you had
15 deterministic kinds of thinking in the structure of
16 your presentation.

17 How does that fit when you're trying to
18 assess a probabilistic assessment?

19 MR. TRAPP: That normally is used to get
20 some kind of value, etcetera. And a lot of times it
21 is used when you do not have a good handle or can't
22 resolve some of the underlying scientific basis.

23 CHAIRMAN RYAN: But isn't that a risk that
24 you'll either include or miss something when you just
25 decide on the deterministic value for a key parameter?

1 MR. TRAPP: I'm sure Dr. Garrick would say
2 yes.

3 MR. HILL: This is Brittain Hill from the
4 Center. The hope is to -- the licensing interaction,
5 is to come up with -- especially for the conceptual
6 models.

7 Now, at this stage, I don't think we're
8 gaining a lot of information by trying to an
9 artificial distribution on the limited range of
10 alternative conceptual models because, ultimately,
11 we're trying not to find the simple tendency and
12 cluster of models, but look at, given the current
13 uncertainties, and given the current testable
14 hypothesis, what is the potential significance and the
15 risk calculation from these alternative hypothesis.

16 So, this really is more of a testing
17 methodology than trying to arrive at the mean value
18 that we use to make a regulatory decision. So, that's
19 why we haven't gone through the exercise.

20 We're trying to come up with a
21 distribution function using alternative conceptual
22 models for both the probability itself, and some of
23 the probability parameters. Does that help?

24 CHAIRMAN RYAN: It helps.

25 MR. GARRICK: It helps. Well, before we

1 leave this probability discussion, I wanted to get a
2 couple of licks in. I think that one of the things
3 that the regulators are faced with always is how to
4 make the analyses we've performed as transparent, as
5 understandable as possible.

6 We talked a little bit this morning about
7 how to, and at the same time how to reveal what's
8 really going on, how to reveal the truth. And we
9 talked about these igneous event scenarios, these
10 categories, and the volcanic frequencies that you had
11 associated with these categories, and how
12 characterizing those and embedding those frequencies,
13 because there's uncertainty in those frequencies and
14 uncertainty in probability distributions to kind of
15 convey with time and with conditions with aging and so
16 forth, how the uncertainty changes depending on those
17 conditions.

18 It can be very illuminating. Another
19 thing that I think would be very illuminating, to pick
20 up on Dave Johnson's comment and Mike Ryan's earlier
21 comment, would be to more deliberately manifest the
22 value added of new evidence systematically.

23 And, of course, Bayesian application are
24 perfect for that kind of thing. And there's very
25 little of that that's been done in the past in any

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1 truly systematic and constructive way.

2 And I think that, if we could somehow
3 create a map of what value is added in terms of our
4 knowledge about the risk as a function of pieces of
5 evidence, I think that would be enormously beneficial
6 in aiding the whole process of what the probabilistic
7 analysis is telling us.

8 I would hope the second time around
9 advantage would be taken not only of what we have
10 learned about the earth and about Yucca Mountain and
11 its geology and the rock, but also what we've learned
12 in practice with respect to how to characterize risk
13 in our analyses.

14 And much has changed in the last few years
15 about that. I hope that we take full advantage of
16 that, especially with respect to the transparency
17 issue.

18 CHAIRMAN RYAN: Thanks, any other
19 comments? Yes?

20 MR. JOHNSON: Just to add something, what
21 I meant by saying that the model builder talked to
22 embrace the uncertainty in the models as much as they
23 can and then try to articulate it.

24 For example, if we're saying that there's
25 a zero chance of these relatively recent volcanoes to

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1 occur in Jackass Flat, your belief is that's
2 absolutely zero.

3 But, we ought to go look to see if -- any
4 better. So, what we're really saying is we're not 100
5 percent certain of the fact that -- volcanoes from
6 that particular source.

7 So, if that model were to express that
8 level of uncertainty, it might be very certain, but
9 not 100 percent. Then, as new evidence arrives, then
10 the model can accommodate that, or we can look at the
11 model.

12 It can tell us how important that new
13 evidence can be. I think it's just a more robust
14 explanation of our experts.

15 CHAIRMAN RYAN: Any other comments in the
16 audience? Yes?

17 MR. MELSON: Yes, Bill Melson. I just
18 want to comment that I've heard a little bit of the
19 rumors going on about the appointment of the PVHA and
20 who is going to be on it.

21 And I think Bruce's concern can be
22 lessened somewhat. And that will include someone who
23 -- the core of a lot of the NRC's contract work. So,
24 I think they will not pay us in the night, these
25 things will be.

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1 That's the intention, I think, of some of
2 the planning.

3 CHAIRMAN RYAN: Yes, in the audience.

4 MR. REITER: I'm Leon Reiter with the
5 technical review board staff. I have a question about
6 the PVHA. I guess the first is to NRC. Weren't there
7 any methodological concerns with the other PVHA that
8 DOE had to take into account?

9 And, if there were, did DOE take that into
10 account?

11 MR. TRAPP: The PVHA actually was started
12 a little bit before the Nureg -- PVHA or this time of
13 elicitations. Two areas that really were of concern
14 with the original PVHA panel was the criteria,
15 documentation of the criteria, selection of it, and
16 then, basically, the total documentation of the
17 analysis itself.

18 These are areas that we thought could be
19 improved and were areas that, in this panel would be
20 better.

21 CHAIRMAN RYAN: Yes, a question.

22 MS. KEEFER: Susan Keefer, University of
23 Illinois. I'm an incoming member of the NWTRB, but
24 I'm sitting here until I master my acronyms. My
25 question, Lathrop Wells, my reading of the literature

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1 was that there was water interaction indicated both at
2 the beginning and the end of that sequence.

3 Am I misunderstanding it, or has that been
4 considered?

5 CHAIRMAN RYAN: Do we have someone who can
6 answer that question.

7 MR. CROWE: There is some controversy over
8 the height of volcanic features. But they probably
9 occur about midsection in Lathrop Wells. And then
10 there always is -- the cone itself is an unusually
11 large -- the ration of pyroplastics to lava is unusual
12 for a typical cinder cone fields.

13 But, I think there's strong evidence that
14 it's hydro-volcanic. But, I mean, Leon was on many
15 field trips. We had maybe, eight or ten people at an
16 outcrop that I thought was unequivocally surged.

17 And we had two or three people who swore
18 up and down that it wasn't. So, there is some
19 uncertainty in identifying those deposits. I think
20 the majority of people feel that there was a hydro
21 volcanic component, probably predating the main final
22 cone that we see out there.

23 MS. KEEFER: I'm a consultant to the
24 NWTRB.

25 CHAIRMAN RYAN: Any other questions or

1 comments in the audience. I want to thank all of the
2 -- I'm sorry, is there a question.

3 MR. KESSLER: John, Kessler, EPRI. Two
4 comments, both from a performance assessment
5 perspective, surprise. One is a specific comment, and
6 then the another a more general comment.

7 The specific one was on the discussion
8 this morning and Britt Hill's talk about the temporal
9 variability in the number of volcanic events that
10 occurred and how one might deal with that performance
11 assessment space.

12 And, Britt talked about, well, you could
13 have maybe as many as something like 40 some events
14 per million years if you look at the right million
15 years, and then, maybe look at the mid-point between
16 that and the long-term average.

17 Well, I would argue that, if you're going
18 to look at the maximum, the mid-point isn't with the
19 long-term average, but it's with the other end. It
20 might be something like zero events in a million
21 years.

22 And, from a performance assessment
23 standpoint, you could say, sure, I'll show you
24 everything that I see. I'll show you, for any million
25 year interval, I'll show you a table of zeros, maybe.

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1 And then I'll show you one where you have
2 something when you get a higher dose risk for that
3 particular million year interval. So, it comes back
4 to what is it that's the compliance number you can
5 use.

6 And I'd still say come back to whatever
7 the scientists agree is the right number of years to
8 average over. But you're still going to take the
9 average, unless you have anything that suggests you
10 know what's going to happen in the next million years
11 or in the next time period, or whatever.

12 If you don't have any way of
13 distinguishing that, to me, the long-term average,
14 whatever the long-term you choose to use, is what
15 seems to be the right course of action to take.

16 You can go ahead, of course, and add
17 sensitivities on any particular variability around
18 that, from zero up to forty some. But, in the end, I
19 would think that, from the compliance standpoint, that
20 would probably be what you would want to do.

21 Now for my general comment. It really
22 falls right along the lines about what John Garrick
23 was talking about, which is, I felt that a lot of the
24 discussion this morning, where we're talking about
25 maybe changing probabilities by factors of five, maybe

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1 ten, is interesting and all.

2 But the uncertainties in just the volcanic
3 dose risk assessment, when you look at the risk
4 triplet, what can happen, you know, what are those
5 probabilities, and then what are the consequences, it
6 seems that, you know, we should be focusing on a lot
7 of other aspects, given the kind of changes we are
8 talking about.

9 So, why are we talking about this? I
10 mean, a lot of money, and a lot of time have been
11 spent. Well, why is that? That's because assumptions
12 have already been made about what the consequences
13 are.

14 For example, in John Trapp's talk, he had
15 one line where he said, well, we've had some analysis
16 that said the containers are going to fail, which is
17 assuming they go poof and that's it.

18 Okay, that's why the numbers come out the
19 way they do. Those kinds of assessments were done, I
20 don't know how much thinking went into that part of
21 the assessment, but it's just as important in terms of
22 coming up with the dose risk numbers, as getting this
23 probability that we're talking about that may change
24 by a factor of two, ten, something like that.

25 So, my point is that one can do a lot of

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1 expert elicitations on all kinds of aspects of the
2 system. I just don't know, if I was king for the day,
3 I'd spend my money on redoing the PVHA, versus some
4 other aspect of this problem or something else in
5 terms of getting at reducing uncertainties.

6 The uncertainties that are being looked at
7 here are ones that, it could increase the uncertainty
8 because we've already made all these other
9 conservative assumptions.

10 That's why this one is being looked at.
11 I think if we looked at what can happen and shouldn't
12 just look at what can go wrong, but what can go right.

13 If we're looking at a best estimate
14 assessment, if there are possibilities that we can
15 replace the model that's conservative with maybe
16 something else that's less so that could dramatically
17 affect the dose risk numbers, perhaps that should be
18 looked at just as strongly -- things towards what
19 could go wrong.

20 CHAIRMAN RYAN: Thanks John. Any other
21 last comments before we finish up the session. Yes,
22 sir?

23 PARTICIPANT: My name is John. I'm a
24 consultant. I have point of clarification and comment
25 of concern, if I may. My point of clarification is

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1 with respect to our early warning drilling program,
2 23P, where we first talk about intercepting the
3 basalts there.

4 Our geologists have taken a look at it.
5 And they do not believe that they basalt below. They
6 believe that they basalt boulder. Their reason for
7 that is because, in looking at the cuttings, they saw
8 roundings on the upper portions of it.

9 And they saw a rounding of the cuttings.
10 The second thing is, they saw no alteration or
11 difference in the sediment of any of the above and
12 below.

13 So, we're not sure that we had basalt
14 flow. I'll tell you, in terms of a comment, I don't
15 think that enough time is being spent in looking at
16 the structural relationships between these existing
17 volcanic centers and flows, and what we're trying to
18 get at here.

19 One thing that really jumps out to me in
20 one of the presentations, Dr. Crowe put at the basin
21 range and showed how the big activity was on the
22 margins of it.

23 When we look at the magnetic math from
24 yesterday and the previous work that -- commissioned
25 back in 1999, we saw major magnetic east-west

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1 liniments being transected with north-south ones.

2 And, when you look at where these recent
3 volcanics are, they are at the intersection of these.
4 We think there needs to be more work done like this
5 gentleman talked about in looking at the structural
6 aspects of it.

7 And then, finally, on the concern, and
8 this came out yesterday, the new magnetic work is
9 exquisite. We can see some really good things in
10 there.

11 We can see our EWDP wells, where we put in
12 deep steel cases. Those show up on that survey quite
13 clearly. My concern from yesterday was, when we're
14 laying it out at the workshop, I pointed to one
15 anomaly.

16 And I said, well, what's this one? And
17 the response was, oh, we hadn't noticed that one. And
18 that's a concern to us. We're not volcanologist.

19 We're not probabilistic people. We have
20 to rely on other people to do these things. But, when
21 we look at something after ten minutes and say, what
22 about this one?

23 Then we think there's a concern there,
24 that maybe there's a method that needs to be looked at
25 to make sure that all the anomalies are indeed

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1 identified.

2 And the County doesn't have to come along
3 after the fact and say --

4 CHAIRMAN RYAN: Thanks. I'll just add
5 that I think many of us here, me in particular,
6 weren't at yesterday's meeting. So, we're a little
7 bit blind.

8 And we didn't have an appreciation for
9 what you have in those. But, thanks for sharing it
10 with us at this point.

11 PARTICIPANT: Again, they're doing a
12 terrific job. It is a repressed timescale. People
13 concentrate on what they want.

14 CHAIRMAN RYAN: Thanks very much. Any
15 other questions or comments? Yes?

16 MR. MELSON: Mike, I'm wondering if an
17 integration of changing the PVHA at this time, and it
18 does examine as well a probability for canister
19 dysfunction, and that the membership be changed -- the
20 attitude, so there are experts of the type that we'll
21 be hearing from some of the work that's been done this
22 afternoon to really try to go after some of these
23 volcanological issues quantatate.

24 PVHA has updated. It remains what it was.
25 But it also --

1 CHAIRMAN RYAN: I think that is an
2 interesting suggestion, Bill. If we can maybe hold
3 that in our minds as we hear the second two pieces of
4 the parts that go into the risk triplet, maybe we'll
5 come to a better appreciation of how to answer your
6 question.

7 We can sure think about it, because that's
8 a great segway to our next segment, which is the
9 repository interaction with the magma, and then on to
10 the dose consequence aspect of it a little later on
11 tomorrow.

12 We are scheduled at the moment for a
13 public comment period. Do we have any other comments
14 that folks would like to make? Yes?

15 MR. McCARTIN: This is Tim McCartin, the
16 NRC Staff. In relationship to the previous comment,
17 for the record, I would like to state the appendix
18 seven meeting between the NRC and DOE is a way for NRC
19 to get information from the Department as quickly as
20 it is available.

21 And, not that I have to defend the
22 Department, it's not my role, but, in the spirit of an
23 appendix seven, we got this information.

24 And DOE has barely looked at it. They
25 have not spent a lot of time analyzing it. And so,

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1 the fact that they may not have seen everything there
2 is, we did press them to get this information as
3 quickly as possible.

4 And so, there's a lot of analysis time
5 with the information that we have that is yet to come.
6 But, this meeting was held before they had done this
7 analysis.

8 CHAIRMAN RYAN: So, a further study of
9 their's and NRC's will be underway. Is that a fair
10 comment?

11 MR. MCCARTIN: Yes.

12 CHAIRMAN RYAN: Thank you. I appreciate
13 the clarification. Any other comments or questions?

14 (No response.)

15 CHAIRMAN RYAN: That being said, and it
16 being just slightly after two o'clock, we are
17 scheduled for a 30 minute break. I'm going to propose
18 that we limit our break to perhaps 20 minutes.

19 Let's come back at, say, 25 minutes after
20 two. And that way we'll have continuity with our
21 presentations for the rest of the afternoon on
22 discussions thereof.

23 (Whereupon, the above-entitled matter went
24 off the record at 2:03 p.m. and went back on the
25 record at 2:25 p.m.)

1 CHAIRMAN RYAN: We'll begin our afternoon
2 session with a presentation again by Dr. Britt Hill
3 entitled NRC Review Capabilities for Evaluation of
4 Potential Magma Repository Interaction Processes.

5 I might make a note at this point that the
6 next talk that's on the agenda will not be held. The
7 speaker and the other members of that panel were not
8 available to be here today.

9 They all had other prior commitments. But
10 the consequence review panel is documented. It has
11 been presented previously to the ACNW and to the
12 NWTRV.

13 And that is a matter of public record.
14 Those records will be available. Mike Lee will help
15 anybody that wants to find those references. He will
16 help them get the identity of those records.

17 So, that presentation will not be held.
18 And then, after our first presentation by Dr. Hill,
19 Dr. Kozak, with introductory remarks by John Kessler,
20 will follow-up with the EPRI alternate views on the
21 modeling of magma repository interactions.

22 After that, we will have a closing working
23 group roundtable sessions with panel members to talk
24 about that second part of the igneous event. Thank
25 you, without further ado, Dr. Hill, welcome back.

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1 NRC STAFF PERSPECTIVE ON THE MODELING OF
2 MAGMA/REPOSITORY INTERACTIONS

3 MR. HILL: First I want to make sure that
4 we recognize a number of people who have made the key
5 contributions in this work on developing our review
6 capabilities for potential magma repository
7 interactions.

8 They are essentially the lead authors on
9 a lot of this work, O. Bokhove, Anne Marie Lejeune,
10 Steve Sparks, and Andrew Woods. Notice they are all
11 from the Netherlands.

12 Unfortunately they couldn't be here today
13 to help with these presentations. Next slide, please.
14 What I'd like to do this afternoon is give a brief
15 overview of why these potential magma/repository
16 interactions can be significant to do performance
17 calculations.

18 I'll set the stage a little bit for why we
19 are doing some of this work. I'll talk for a few
20 minutes on some recent developments that have come
21 about in understanding the water contents of Yucca
22 Mountain basalts.

23 This has been done from direct --
24 experiments that I think provides real interesting
25 insights on some of the important physical conditions

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1 that we have to consider in any sort of numerical
2 model for igneous processes in the Yucca Mountain
3 region.

4 I will give a fairly brief overview of
5 some of our previously focused models for initial
6 magma interaction processes, but not spend too much
7 time on that, and then discuss some of the newer work
8 we've done in the past year or so on modeling magma
9 flow in conduits that have elastic wall-rock problems,
10 sort of looking at some of the relationships between
11 fluid pressure and wall-rock response, but building up
12 the initial models we presented several years ago.

13 And finally, I'll give a pretty quick
14 summary of where this current information is leading
15 us. Next slide, please. Why does this matter? Why
16 do we care about potential magma repository
17 interaction processes?

18 The simplest answer is these processes
19 control the source-term for igneous intrusive and
20 extrusive events that we're modeling in the
21 performance assessment.

22 Very simply, you can think of this in
23 terms of three very basic conceptual models. The
24 first one would be that the rising magma would come up
25 along a dike.

1 And we're looking kind of down the plane
2 of the dike. It would vertically intersect the
3 repository with the probabilities that we've been
4 talking about this morning, and then continue to rise
5 to the surface, and, during the course of an eruption,
6 form a vertical conduit.

7 And, the expansion of that conduit through
8 time is the one that could -- waste packages in the
9 conduit footprint. Now, alternatively, sort of a
10 model that was first developed in the Woods et al
11 paper in 2001, we are looking for the potential for
12 developing a breakout at some place distant from the
13 point of initial interception of the dike.

14 This could be for a variety of reasons.
15 But, basically, we're looking at a zone of weakness
16 that was easier for magma to propagate along a
17 secondary plane than along the initial plane of
18 intersection.

19 And, again, we have no real good
20 historical analogs or any geologic analogs for how
21 rising magma would come up to 300 meters below the
22 surface, interact with a void that extend for hundreds
23 of meters laterally and are five meters in diameter.

24 So, we have to take a conceptual model
25 approach that looks at information that we can use for

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1 historically active volcanoes, experimental analogs,
2 and numerical models.

3 And, one of the insights that we can gain
4 from the historical basaltic volcanism, is this third
5 option, or third thing that we have to consider.

6 During the course of an eruption, you get
7 very simply the breakouts that can occur from the main
8 magma system and rise at some distance away from the
9 central volcanic system.

10 I'll talk in a little more detail about
11 that later in the presentation. Basically, before we
12 start to talk about any numerical modeling or
13 experimental modeling of these potential conceptual
14 models, we have to constrain some of the physical
15 characteristics of the magma system that we're trying
16 to simulate.

17 One of the most important model
18 uncertainties, really just about any model that we do
19 for all of the buried basalts, is understanding what
20 are our magnetic water contents.

21 Those of you that aren't really familiar
22 with the geology of this area, you can dissolve
23 certain amounts of water into the molten rock. This
24 dissolution of water isn't from groundwater.

25 It occurs very deep in the mantle where

1 the basalt originates from. There are minerals in the
2 mantle and in the crust that contain water in them.

3 And, when these minerals melt, they
4 release water into the melt. That water stays
5 dissolved until you depressurize the metal at some
6 point.

7 Then the water and other volatiles, like
8 carbon dioxide, come out of the solution and form a
9 gas base. Is this expansion of the gas base as the
10 magma rises up to shallow depths, say on the order of
11 a kilometer or so, that governs a lot of the eruption
12 characteristics that we're trying to understand.

13 And these, in effect, are mass flow
14 characteristics as well. So, we have to kind of give
15 that language between volatile contents and the mass
16 flow characteristics.

17 And, in addition, some of the
18 uncertainties that we're going to have to address
19 include the crustal properties, things on the order of
20 the elastic properties of the rock as it goes down
21 deeper into the crust, as well as the distribution of
22 stress within the crust, not just at the surface, but
23 with increase.

24 Next slide, please. Okay, the basaltic
25 magmas that we see out here in the Yucca Mountain

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1 region are characterized by about three to five
2 percent of larger crystals of olivine to small amounts
3 of the mineral amphibole.

4 There's a little bit of the mineral called
5 plagioclase and pyroxene. We're not going to worry
6 about that at the moment. Amphibole is a hydra-
7 silicate mineral.

8 It is unusual in basalts, but does occur
9 in the basalts in the Yucca Mountain region. The
10 reason we are concerned about amphibole is it's a
11 mineral that has water in the crystal lattice.

12 So, obviously, if you have a hydrated
13 mineral in this basalt, you have to have some activity
14 of water in order to format the mineral in the first
15 place.

16 Previously we had looked at some
17 experiments that were done in basalts that weren't
18 directly related to Yucca Mountain, but had some
19 analog characteristics to the basalts that we see in
20 Yucca Mountain.

21 Based on those analog experiments we would
22 say that you would need probably greater than two
23 weight percent water in order to crystallize this
24 amphibole mineral you can see in the picture in the
25 basalts in the Yucca Mountain region.

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1 But, it was a very loose correlation, and
2 we didn't have a good sense for two percent, three
3 percent, etcetera, for these specific basalts. About
4 two years ago, Jim Luhr of the Smithsonian Institution
5 and Tom Housh at UT Austin had done some measurements
6 with a very high powered microscope, as well as some
7 work where they take a look at inclusions in the
8 mineral.

9 When these minerals form, they trap glass
10 in part of the -- when the basalt cools, turns into a
11 glass. That trapped glass traps the amount of magma
12 and dissolved volatiles in it at the time of
13 crystallization.

14 So, we're capturing processes that are
15 fairly even in the crust. When Luhr and Housh
16 analyzed these glass inclusions, what they found was,
17 for the ones that hadn't leaked, there was anywhere
18 from three and a half to four and a half weight
19 percent water in these cracked melt inclusions in the
20 minerals in places like Little Cone in the Yucca
21 Mountain region.

22 They also were able to measure anywhere
23 from 600 and 900 parts per million of a dissolved
24 carbon dioxide in these math conclusions as well. So,
25 this is the first really direct measurements that we

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1 had that looked at what did the volatile contents be
2 in Yucca Mountain type results.

3 And they're a bit higher than the two
4 weight percent estimate that we had from the
5 experimental analogs. Next slide, please. We were
6 really lucky that this year there some experiments
7 that were published in the Journal of Geology by
8 Nicholis and Rutherford, where they had gone out and
9 evaluated some direct experiments, all basalts
10 collected at Lathrop Wells volcano and Little Cone
11 volcano.

12 Essentially they are the same basalt
13 composition. Now, what we see is, at Little Cones,
14 they have done these experiments. Let me just take a
15 moment to explain this diagram.

16 What we see on the lower axis is
17 temperature increasing from left to right, and
18 pressure increasing from the base up to the top. Now,
19 this pressure is the total pressure in the
20 experimental system.

21 They've also added water into this
22 experimental system. So, the total amount of water,
23 the pressure of that water is equal to the total
24 pressure in the system.

25 For a comparison, these pressures here

1 from anywhere from 60 to 220 megapascal, corresponds
2 to dents in the crust about five kilometers deep down
3 here at 100 megapascal, up to roughly ten kilometers
4 deep at 220 megapascal.

5 So, we're still dealing in the brittle
6 crust, down there around ten kilometers or so beneath
7 the Yucca Mountain region. What these experiments are
8 showing is that, in order to form the mineral
9 amphibole, which is occurring in this shaded area
10 right here, and olivine, you need to have water
11 contents of about four weight percent.

12 In other words, you have to have pressures
13 with total water pressure of about eight kilometers
14 depth in order to stabilize that amphibole at about
15 980 degrees centigrade, because this is the mineral
16 assemblage that we see at little cone.

17 It is olivine with a bit of amphibole.
18 Now, if we only have three weight percent water in the
19 melt, that saturation would be here at about five
20 kilometers depth.

21 Three weight percent would correspond to
22 about 100 megapascal of pressure. If we have three
23 weight percent water, what we would see as that magma
24 cools, is first olivine, then plagioclase, and then
25 this mineral pyroxene would come in.

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1 And, only have you crystallized a large
2 amount of plagioclase and pyroxene in addition to the
3 olivine, would you being to see the amphibole. The
4 fact that we consistently see olivine plus or minus
5 amphibole and don't see any plagioclase in pyroxene,
6 is telling us that the water pressures had to be high,
7 and on that order of four weight percent water to get
8 the observed mineral assemblies.

9 So, we have all the lines of evidence from
10 these direct experiments on Yucca Mountain basalt, as
11 well as the glass inclusions in Yucca Mountain basalt.

12 They are saying, in general, we are
13 looking at four weight percent dissolved water when
14 the magmas were down at a depth of about 10
15 kilometers.

16 This is a really important number when we
17 talk about numerical modeling of the eruption
18 processes, because we don't have four weight percent
19 water in these magmas. Gas bubbles begin form at
20 about eight kilometers depth.

21 When you start rising that magma up,
22 whether it's simple solubility models, simple bubble
23 models, you begin to form bubbles when that magma is
24 about eight kilometers below the ground.

25 By the time you get to within one

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1 kilometer depth, under an equilibrium sort of descent,
2 when you're at one kilometer with this amount of water
3 in the melt, you've got a very large volume of gas,
4 something on the order of 70 percent of the total
5 volume of material that's rising up, is going to be
6 the gas phase.

7 You have so much gas, that you're
8 beginning to break apart the magma in a very simple
9 way of looking at things. In contrast, if you had
10 only a couple of weight percent water, you wouldn't
11 have a lot of bubbles, maybe 40 percent bubbles, by
12 the time you got below one kilometer.

13 So, this difference in volatile content or
14 water content makes a real important distinction in
15 how you model a gas magma mixture at depths that
16 correspond to the potential repository.

17 Next slide. Okay, I've been going on and
18 one about volatiles. I need to shift gears now on how
19 we're going to evaluate the potential magma-repository
20 interactions.

21 I think the easiest way to do this is
22 start with the simplest models. When we evaluate
23 experimental analogs in fluids that lack volatiles, --
24 sorry about that, but that's just the easier way of
25 explaining it.

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1 What we're using at first is an
2 experimental system that looks at potential
3 interactions between a volatile-absent magma -- in
4 this case a volatile-absent golden syrup fluid --
5 interacting with a subsurface drift.

6 What we're using is a golden syrup,
7 essentially a sucrose syrup that has viscosities on
8 the order of one to 100 pascals. A real plain English
9 way of looking at that is it's kind of like a stiff
10 syrup all the way down to a very weak syrup.

11 So, these are very sticky kinds of syrups. What
12 we're doing is we're setting up an experiment lab
13 where we have a reservoir just off the plain of the
14 pitcher.

15 That reservoir is under pressure, and we
16 have a pressure in this horizontal tube, and a little
17 gate down here. So, we have the fluid come up into
18 this vertical cell, which simulates a dike, a gate,
19 and then a different pressure in this drift.

20 We open up the gate and watch the fluid
21 response across different pressure gradients between
22 the reservoir system and the tunnel system, and also
23 for different viscosities.

24 And we could go into a lot of these
25 experiments. They are documented in a report that's

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1 available from the Center by Lejeune et al, 2002.
2 But, basically, the goal of all this work is to
3 develop a numerical model on how volatile-absent
4 fluids could flow into a potentially intersected drift
5 and scale those experiments to repository conditions,
6 and try to get a first-order feel for the flow rates
7 and kinds of flow that you would see for volatile-
8 absent flow.

9 This lower slide just shows a very
10 simplified simulation where you have a different
11 pressure to measure and measure the flow front through
12 time.

13 The shape of the flow front is going to
14 change depending on viscosity and the pressure
15 gradient, and whether or not you develop a
16 gravitational front on this so it looks more like a
17 lava flow, versus more of a pressure driven flow, that
18 doesn't have a gravitation front to it.

19 This could affect our understanding of how
20 this would interact with structures in the potential
21 drift. Next slide, please. One of the things we see
22 from these experiments is that the viscous drag in the
23 dike system is a much more important process than any
24 sort of drag effects in the tunnel.

25 So, the dissipation of pressure in a magma

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1 system is really going to be controlled by friction in
2 the dike, not really tunnel interaction effects. The
3 controlling process that we have to worry about is the
4 pressure gradient, which we need the pressurized fluid
5 that's rising in this simulate the dike system, and,
6 of course, the drift system pressure.

7 What we're looking at here are a series of
8 open tunnel experiments where we're allowing the end
9 of the experimental apparatus to be opened at the
10 sphere.

11 So, we're not getting any air compression
12 effects as the syrup flows in the tube system. And
13 this could be similar to a permeable drift wall
14 scenario where you have a lot of gas escape as magma
15 would flow into the open system.

16 What we're seeing is that the flow
17 velocities that you would expect in these simulations
18 really is controlled by the overpressure in the magma
19 system.

20 And here we're looking at pressures of
21 five megapascal pressure drops, with the drift and the
22 dike, up to ten megapascal pressure drops, and seeing
23 of change in flow velocity on that order of eight to
24 twelve meters per second.

25 A real simple way of thinking of these

1 experiments is, if we have volatile-absent model
2 coming up into this potential drift system, and we
3 have a pressure differential that corresponds to
4 lithostatic to maybe five megapascal over hydrostatic
5 or lithostatic in this system.

6 We should expect to see magma flowing into
7 the drifts on that order of around ten meters per
8 second, which isn't a huge flow rate. But, again, we
9 have to remember that this is for a volatile-absent
10 flow.

11 Next slide, please. We also have known
12 some numerical models on volatile-rich interactions in
13 each trip. And I know there's been a lot of
14 discussion of these models.

15 Some of them are documented in the Woods
16 et al, 2002 reports. But, basically, we looked at a
17 very simplified model for initial magma repository
18 interactions.

19 Some of the simplifications you get with
20 this first order model, assuming we had no gas loss in
21 the magma system, that the dike would instantly open
22 into the drift, there were no elastic effects, there
23 was no feedback between magma pressure and wall
24 opening, and also that we have a closed drift, no gas
25 loss, and the drift was completely smooth.

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1 The important point of these experiments
2 or these numerical models is that, if you optimize the
3 system for rapid decompression, you would get a flow
4 acceleration on the order of about 100 meters per
5 second and potentially generate some sort of a shock
6 as that accelerated flow reacted with the end of the
7 drift wall and started to bounce back.

8 The important point is not whether or not
9 a shock would develop, but, by using these optimized
10 modeling conditions, the magnitude of that shock is
11 still very low compared to the strength of a waste
12 package.

13 So, we are not worried about transient
14 overpressures from potential initial interactions
15 between bubble bearing magma and an open drift. One
16 of the concerns that arose from these sort of models,
17 though, was the potential to generate fractures along
18 the drift, especially at the end of the drift where we
19 would have reflection phenomena, because, at this
20 depth of about 300 meters below ground, it only takes
21 about five megapascal to hydrofracture the rock at
22 this depth.

23 So, the question was, even though we
24 weren't disrupting the waste package, couldn't we be
25 creating fractures during this initial stage of

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1 interaction that would then be exploited at some time
2 by rising magma during the course of eruption.

3 Next slide, please. So, again, even
4 though we're not looking at this happening initially,
5 of the magma shooting out through a fracture, we still
6 have to consider that, although we have more likely
7 scenario where the rising magma continues along a
8 plane of vertical ascent, there still appears to be a
9 possibility of exploitation, all these secondary
10 fractures at some time during an eruption.

11 Now, the reason we're concerned about
12 this, and why we still have residual uncertainty about
13 this alternative flow path, is that we know, from the
14 numerical modeling that we've done, as well as the
15 modeling of others, in shallow magma systems, you can
16 have variations in the overpressure and underpressure
17 that occurs within the conduit system.

18 This is, again, just a very simple
19 numerical model in the Woods et al, 2001 paper that
20 shows more steady state conduit assumptions. And, by
21 the way, this assumes a rigid conduit wall.

22 There are no elastic effects built in on
23 this conduit. You can see that, in these assumptions,
24 the conduit would be under pressure until we got to
25 around 300, 200 meters below ground surface.

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1 And then, because of choking effects in
2 the event, we would have some overpressure in the
3 system on the order of several megapascals. Again,
4 this is just a snapshot representation of a moment in
5 time of an eruption.

6 During the eruption, because of variations
7 in mass flow, and variations in conduit response these
8 over and under pressures can be much more dynamic than
9 you get from a standing model.

10 And, again, we're trying to look at, not
11 just in any one instant of time, but what could
12 happen, what needs to be considered in our risk
13 assessments for the duration of an igneous event?

14 And one of the reasons we're still looking
15 at this as a potential concern is that, in some cinder
16 cone volcanoes, we see these secondary breakouts
17 occurring at various stages during an eruption.

18 Next slide, please. Here's two examples
19 of these secondary breakouts. We volcanologists
20 commonly call them boccas. One of these that we're
21 looking down on Paricutin Volcano, this is a big --
22 from Luhr and Simkin's book on Paricutin.

23 We're looking down from the air on
24 Paricutin Volcano. And what we're seeing are these
25 series of fractures that went from the southwest to

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1 the northeast, that represent some of the initial
2 fracturing at Paricutin.

3 The main volcano localized, the main
4 cinder cone formed at Paricutin right through here.
5 But, during the course of the eruption, we have a
6 secondary breakout occur, one of these boccas occur
7 along the plain of initial intersection.

8 And this vent was active for only part of
9 the eruption, but still effused lava and had an effect
10 on the eruption characteristics for some time.

11 These are the most common kinds of
12 breakouts that you see at these volcanoes, ones that
13 are along the initial plane of intersection. They are
14 not ubiquitous.

15 But, certainly, it's very common to find,
16 along the direction of dike propagation, these sort of
17 secondary breakouts. Another example is shown on the
18 right, a very simplified geologic map of the cinder
19 cone eruption in Russia in 1975.

20 We're seeing three main cinder cones that
21 formed along a generally north, north-east trending
22 fissure system. And there's a number of these boccas,
23 the diffused lavas, that are localized pretty much on
24 trend with this main fissure system.

25 These are not what we're concerned about.

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1 We're concerned about these sort of secondary
2 breakouts that occurred over a kilometer and a half
3 from the plane of initial intersection.

4 I don't think any of us really understand
5 the details of how the magma system is forming these
6 breakouts from the main magma system. But, it does
7 appear that, some time in some eruptions, you get
8 these lateral breakouts from the system, away from the
9 plane of initial intersection.

10 So we need, in our conceptual models, in
11 our uncertainty analyses, to consider the likelihood
12 of this kind of a condition occurring at the potential
13 repository site.

14 Okay, next slide. What I'd like to do now
15 is just provide a real quick summary of some of the
16 newer work that we've been doing on modeling magma
17 flow in elastic wall conduits.

18 Now, the results of these analyses are in
19 the woods et al 2004 report that I believe was
20 distributed before the meeting. Again we're starting
21 off with building on previous models in which we've
22 used rigid conduit walls.

23 And one of the criticisms and concerns
24 that we had received on those models is that there
25 really was no linkage between pressure in the system

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1 and wall rock response.

2 Because we can model some of this wall
3 rock response in the simple elastic process, we know
4 there's going to be deformation on the wall rock as
5 the magma rises through time, and also as the volcanic
6 conduit responds to pressure variations through time.

7 So we need to develop some sort of a basis
8 to evaluate how this feedback between elastic
9 properties of the rock and fluid pressure in the
10 system can affect the eruption characteristics.

11 In here, we're using the base assumption
12 I think is very generally accepted that fractures
13 dilate from magma pressures that are going to be
14 greater than the minimum principle horizontal struts.

15 So this model is going to be sensitive to
16 the assumptions -- its stress ratio, in other words
17 minimum stress to maximum stress in the horizontal
18 domain.

19 What we're doing is assuming a two
20 dimensional elastic conduit wall, and allowing that
21 pressure to have a feedback in the elastic response of
22 the conduit with the total pressure in the system.

23 These are essentially basic flow dynamic
24 models that use mass in the land of conservation to
25 evaluate this response with some different assumptions

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1 regarding magma pressure, horizontal structure ratio,
2 and also taking a look at variations in volume
3 content.

4 Next slide. For Yucca Mountain type
5 ascent rates and magma viscosities, and dike
6 geometries, we're ending up with buoyancy driven flows
7 that have fairly low Reynold's numbers.

8 These are not very turbulent forms. We
9 see Reynold's numbers on that order of two to 20. So
10 we can make a number of simplifying assumptions in
11 numerical models.

12 What we see in the model, and again these
13 are models in the Woods et al 2004 paper, where the
14 conduit width is going to be controlled by the
15 difference between the pressure in the magma and what
16 we're calling simply the pressure in the rock.

17 Where this rock pressure integrates the
18 density variations, the stress ratios, and the elastic
19 properties in the rock as well, we're seeing that the
20 viscous drag in the system is much more important than
21 the turbulent drag.

22 But again, this is only for the model with
23 poor magma conditions. And what we're doing in this
24 example is looking at a stress ratio of .7, in other
25 words the minimum horizontal stress is 70 percent of

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1 the maximum horizontal stress.

2 And a magma viscosity on the order of 100
3 pascal seconds, which we think is pretty
4 representative of the bubble-absent, crystal-poor salt
5 in Yucca Mountain region.

6 And for these different mass flow rates,
7 and the units here are meters squared per second,
8 that's flow rate per unit length of a dike, anywhere
9 from .3 to about 10 meters per second along a meter of
10 dike.

11 We're seeing an opening in that dike from
12 that depth around two and a half, three meters wide,
13 narrowing up to on order of one, one and -- or one to
14 a half meter wide as we approach the surface.

15 As the pressure in the system is
16 dissipated by frictional losses against the dike as
17 well as the variation in the duction in magmatic
18 pressure as we increase distance away from the source
19 region.

20 Now we're assuming in these calculations
21 that the viscosity remains constant. We know as we
22 bring volatiles out of solution and have a little bit
23 of cooling to the magma, we can be varying the
24 viscosity.

25 And also, as we have bubbles come and

1 appear in the magma, the viscosity is much more
2 dynamic than the assumptions that were made. Again,
3 we're not trying to realistically model with this
4 stage the full range of magma set processes, but gain
5 some understanding of how feedback between the wall
6 rock and the magma system can affect flow processes in
7 the shallow sub-surface.

8 Next slide please. And the reason this
9 model becomes more important is when we starting
10 adding bubbles into the volume. We have the volatiles
11 in there. We model the exsolution of gas and start to
12 account for the compressibility affects of the magma.

13 When you don't have bubbles in the magma,
14 you don't have the volatile phase, that magma is very
15 difficult to compress. It's essentially a big
16 compressible fluid.

17 So its density doesn't change in
18 compression. But once you start adding bubbles in to
19 the system, it's much more sensitive to the density
20 with pressure variations because a porous gas can be
21 compressed much more easily than a fluid.

22 Here, when we add this volatile phase into
23 the magma, again we're assuming the constant
24 viscosity, we have to account for laminar, as well as
25 the turbulent drag effects by having this volcanic

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1 fluid.

2 But we're still modeling this as a single
3 phase flow. We're not trying to get into modeling
4 different velocities between the gas phase and the
5 magma phase.

6 So again, it is a simplified model, but
7 getting into two dimensional, or two phase flow is a
8 much more complicated step. Within those limits
9 though, what we're seeing is that when this magma
10 system is allowed to reach the surface and vent, we
11 impose a condition of choke flow.

12 In other words, the flow velocities can't
13 exceed the local velocity of the speed of sound at the
14 vent. It's a common modeling assumption in volcanic
15 processes.

16 Because of these choke conditions; we end
17 up very simply getting an overpressure in the system
18 that inhibits gas exsolution. So because of
19 accounting for these compressibility and pressure
20 effects, we end up suppressing a lot of the bubble
21 growth and gas exsolution in the shallow subsurface.

22 That gives us a very different
23 understanding of how bubbly these mixtures would be
24 under a steady choke flow condition. Then if we made
25 just a simple equilibrium ascent like I was talking

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1 earlier, here in this model again from Woods et al,
2 we're seeing depending on the assumptions you make in
3 a set velocity you'd have anywhere from about ten to
4 20 percent bubble fracture, in this case void fracture
5 in the melt, once you fully realize the choke
6 conditions and elastic effects in the conduit model.

7 For comparison if you didn't account for
8 those effects, at 300 meters depth you'd have over 70
9 percent void fractions for a simple lithostatic
10 pressure model, again using about two weight percent
11 volatiles in the melt.

12 So this gives us a very different
13 understanding of the kind of decompression effects
14 that we may have to consider if we talk about
15 decompression at 300 meters into an atmospherically
16 pressure filled drift.

17 Next slide please. Now we haven't just
18 done these simple flow models. A terrible effort's
19 been gone -- has been undertaken by O. Bokhove and his
20 colleagues at the university.

21 He used some non-linear invective
22 diffusion equations. At the stage that I can record
23 we've only been using non-compressible flows for these
24 models, but the advantage is by using these non-linear
25 equations, we can evaluate non-steady flow.

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1 In real simple language terms, is that we
2 can look at the variations in these flow processes
3 through time, rather than previous models which are
4 pretty much a snapshot in time.

5 We're looking at a single equilibrium sort
6 of ascent, or equilibrium ascent and flow in the Woods
7 et al models. And here we're using a full dynamic
8 realization of the evolution of the system through
9 time.

10 One of the reasons -- one of the ways we
11 can gain some confidence in this approach is take our
12 non-linear equations and set it to a steady flow
13 condition.

14 In other words, make it kind of similar to
15 the model I was just showing. When we make a steady
16 flow assumption, we end up getting the same sort of
17 variations in dike width with depth that we saw in the
18 Woods et al 2004 model.

19 So, very simply, this alternative approach
20 using the invective diffusive equations when we set it
21 to the same flow conditions of steady flow gets you
22 the same basic dike width depth relationships that you
23 get from an alternative approach using the maximum
24 interval versions to the Woods et al.

25 Now the value isn't in duplicating this

1 figure, but in the next slide. On the left hand side
2 we're looking at a time history of how dike width
3 would evolve from a reservoir located simply three
4 kilometers away from the surface.

5 Each one of these lines represents a time
6 step that's about three minutes in the simulation.
7 And so depth is increasing from these meters below a
8 surface to up to the surface here.

9 And dike width is going from one meter
10 wide down to about a tenth of a meter wide. So, in
11 the first step in the simulation at the reservoir, the
12 dike would be around .7 meters wide, with decreasing
13 depth, the dike gets narrower and narrower.

14 As the simulation progresses through time,
15 in other words if the magma is rising, the dike width
16 increases near the reservoir but also increases with
17 decreasing depth.

18 Until finally you get an equilibrium
19 condition right before it vents to the surface, to
20 where you get a small variation from .9 to .8 meters
21 width as the dike goes from three kilometers up to the
22 surface.

23 Now, again, we're not using any eruption
24 at the surface in this first simulation. And it's a
25 non-compressible, non-volatile bearing basaltic magma.

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1
2 The minimum stress assumption that we're
3 using is that the minimum stress is again 70 percent
4 of the maximum horizontal stress for the elastic
5 response in lava rock.

6 On the right-hand figure, we're taking the
7 same basic model in black and then allowing it --
8 excuse me, the same basic model in the red dash lines,
9 and then allowing the dike to vent to the surface, and
10 allow conditions of flow.

11 And what we see from -- faintly see the
12 red lines that correspond to this figure here, but
13 once break out occurs into the surface, there's this
14 drop in dike width, and then dike width reappear --
15 re-widens until it becomes a steady condition of flow
16 out to about a meter to about a meter to a meter and
17 a half wide as flow to the surface is established.

18 So we're seeing a nice dynamic realization
19 here for the non-volatile barium melt that shows that
20 once we have eruption at the surface we get a pressure
21 drop that transmits through the surface, the dike
22 closes and then reopens, but forms a conduit that's
23 going to be wider than a condition for no flow.

24 Next slide. So what did we learn from all
25 this modeling? A lot of things, but putting at the

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1 top level points, we're seeing that models are pretty
2 sensitive to our assumptions on magma viscosity,
3 volatile contents, and our assumptions of minimum to
4 maximum horizontal stress.

5 Even though the model's apparently
6 sensitive to the assumptions, I think we have a pretty
7 good technical basic to evaluate the range of
8 uncertainty in these different parameters, and
9 evaluate the sensitivity of those parameters and
10 models.

11 One of the things is that flow choking at
12 the vent can cause pressure variations in the magma
13 system that really affect the characteristics of the
14 gas phase.

15 So, when we look at the course of an
16 eruption and try to simulate potential interaction
17 processes for the duration of the event, we need to
18 consider feedback between choke conditions at the vent
19 and flow conditions in the sub-surface, and how those
20 conditions of flow in the vent can influence bubble
21 evolution and gas evolution back through the
22 continuous magma system, not just at the surface but
23 down into the sub-surface as well.

24 What we're seeing from the models is that
25 the model eruption rates for a one kilometer long

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1 dike, which are on that order of 100 to 3,000 cubic
2 meters per second, those eruption rates correlate
3 pretty well with the eruption rates that you'd measure
4 historical basaltic cinder cone eruptions, which again
5 are on that order of 100 to 1,000 cubic meters per
6 second.

7 So the mass flow relationships that we're
8 using in the simplified model -- excuse me, we're
9 deriving in the simplified models -- to the first
10 order correspond to the kind of mass flow
11 relationships that we see at typical basaltic scoria
12 cone eruptions.

13 And finally what we're beginning to
14 understand is that the model dike or conduit system
15 can respond to changes in pressure on the order of
16 minutes to hours.

17 So when we start talking about dynamic
18 pressure variations in the course of an eruption, we
19 can be looking at pressure variations that are
20 transmitted through the magma system with velocities
21 on the order of -- excuse me, on time periods of order
22 of minutes to hours of response throughout the system,
23 to these overpressure and underpressure conditions
24 related to potential interaction processes.

25 So to conclude, as a summary of our

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1 current information, we see that the available
2 information is supporting that water contents on the
3 order of four weight percent appear characteristic of
4 the one million year old and younger basaltic magmas
5 in the Yucca Mountain region.

6 If the rising basaltic magma also
7 intersects non-backfilled drifts in the potential
8 repository, that magma may flow into the drifts on
9 that order of a hundred to ten meters per second.

10 Again these aren't fast enough velocities
11 to report any chemical damage, we think, to the waste
12 package, but represents a fairly rapid infilling of
13 flow into the drift system on the scale of minutes to
14 hours.

15 Continued vertical ascent following
16 potential interaction appears to be the more likely
17 scenario following intersection with the drifts. So
18 we think that the available information would favor
19 continued ascent along the plane of initial
20 intersection during the early stages of the eruption.

21 However, we're still concerned about the
22 development of these additional breakouts which
23 sometimes are referred to as doglegs that may occur
24 for short periods during an eruption the same way that
25 we see these breakouts occur in some basaltic scoria

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1 cone eruptions that have been observed historically.

2 And finally I don't want to leave you with
3 the impression that these are completed models. There
4 is still a lot of work that is ongoing to evaluate
5 specific conditions and uncertainties that are
6 appropriate for the potential repository site.

7 We have that work or unfortunately have
8 not gotten it to the stage of reports that are in
9 public domain just yet. So with that I'd like to
10 thank you again for listening to a fairly raspy voice,
11 and open it up for discussion.

12 CHAIRMAN RYAN: Thank you Britt. Any
13 questions from members? Allen? Okay.

14 MEMBER CROFF: I'm going to take a try at
15 a question and I'm not sure I can articulate it very
16 well, but given the statement earlier this morning
17 that the NRC was assuming that there would be waste
18 package failures from you know when magma interacts
19 with the package, I'm sort of struggling to relate
20 what you've said you know sort of this relatively
21 abstract modeling of magma flow to the conclusion that
22 the packages will fail, and that radionuclides will
23 presumably be released to the magma.

24 Can you or anybody sort of elaborate a
25 little more on the underpinnings of this assumption

1 which seems to be fairly critical?

2 MR. HILL: Okay. This isn't really an
3 assumption. There is a lot of the work that underlies
4 that basis has been presented in like the 1999 issue
5 resolution status report.

6 It considers the -- well let me back up
7 for a minute. We had two scenarios that we have to
8 make sure we're talking about. One is for the waste
9 packages that would remain in a potentially
10 intersected drift. And we have waste packages that
11 would be entrained in the erupting conduit.

12 Part of the situation that we have to
13 consider is that we're not instantly developing the
14 conduit and throwing a waste package into the erupting
15 volcano.

16 These conduits that we've interpreted from
17 a lot of geologic information, some of this work again
18 is documented in a publication by -- conduits open
19 through time gradually.

20 So we have a scenario where the rising
21 magma intersects a drift, it goes up to the surface,
22 and then in the course of days to potentially weeks
23 the drift -- the dike opens from a one meter wide
24 conduit to essentially a cylindrical conduit that
25 could be on the order of meters in diameter to tens of

1 meters in diameter, that widens gradually through
2 time.

3 But while it is widening you still have
4 waste packages in the intersected drift that are
5 exposed to the molten magma, or the thermal effects
6 form that magma.

7 So first we have to talk about
8 incorporation into an eruption conduit of a waste
9 package that is at essentially magmatic temperatures.

10 One of the things that we've been
11 frustrated by is that there have been bits and pieces
12 of mechanical analysis but we haven't really been able
13 to develop the full mechanical analysis of waste
14 package response to the range of conditions that would
15 occur in a potential igneous event.

16 For example, we know that there's going to
17 be gas overpressure in a system as in the waste
18 package as the temperature rises up to and over 1,000
19 degrees C.

20 We also know that the C22 -- out of the
21 way, and the stainless steel innerpack have different
22 thermal expansivity. This innerpack is 30 percent
23 more expansive than the outerpack.

24 We have done some sculpting analysis that
25 show that for the small gap -- and don't quote me but

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1 I think it's on order of half a millimeter between the
2 innerpack and the outerpack, that's not enough to
3 accommodate different thermal expansion between the
4 stainless steel and C22.

5 So, as it comes up the temperature, you've
6 got differential expansion and significant radial and
7 hoop stress on the waste package itself. So, by the
8 time we talk about a waste package potentially getting
9 to see the vertically erupting conduit, it is already
10 at temperature and has a significant material stress
11 from internal pressurization from both gas expansion
12 as well as differential expansion in the alloys.

13 We're then taking that material and
14 putting it into a very complex pressure regime, where
15 material is flowing by anywhere on the order of tens
16 of meters per second to 100 meters per second under
17 this dynamic pressure variation.

18 Our waste package people -- and I'm not
19 going to speak for anybody but our waste people at
20 this stage -- have concluded that given the kinds of
21 mass that we have in an eruption volcano, and the
22 condition of the waste package at the time of
23 incorporation, that we would not see resiliency of the
24 waste package when it is thrown into the conduit of an
25 active erupting volcano.

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1 That's been documented in, for example,
2 our issue resolution status report in 1999.

3 MEMBER CROFF: So I should take from your
4 talk that that part of it is assumed, and you're
5 trying to better understand the rate at which the
6 conduit forms, widens and intersects waste packages?

7 MR. HILL: We're trying to understand,
8 yes. Some of the -- ultimately we want to understand
9 better the mechanics of magma flow in a conduit when
10 we have this drift system in the subsurface.

11 There is a number of effects that we have
12 to consider that we never really have considered
13 before in volcanology, because we're having this
14 horizontal tube full of a volume of magma that is a
15 bubbly magma.

16 Where are the bubbles going to go in this
17 tube? We're going to have segregation of the gas
18 phase and the liquid phase, and the possibility of
19 return parameters.

20 In addition, we have this unusual
21 geometry, and unusual stress distribution, that you
22 normally wouldn't see around a volcanic conduit where
23 you've got this drift system sitting here.

24 We need to understand at a better level
25 the uncertainties in the mechanical response in rock,

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1 and of the conduit given this perturbation in the
2 system, because we can't use a simple analog which
3 doesn't exist.

4 So the first part of your question is this
5 has been a long standing series of analyses and
6 information that's been in the issue resolution status
7 report, and in many of the performance assessments.

8 It's not a new assumption in that sense.
9 And second, the reason that we're looking at these
10 sort of flow models is really to understand this
11 perturbation in the system that arises from the
12 presence of the engineered system in the subsurface,
13 not to better understand the volcanoes themselves.

14 MEMBER CROFF: Does the experience from a
15 reactor accident analysis and the experiments that
16 offer any insights for the two phase flow in this
17 case?

18 MR. HILL: I don't really think it's
19 analogous for the two phase flow dynamics. There is
20 a good body of literature and a lot of work that's
21 gone on to understanding volcanic eruptions.

22 One of the reasons we're working with the
23 consultants that we are --Anne Woods and Steve Sparks
24 in particular - is they're some of the worlds leader
25 in understand mechanics of eruption dynamics, fluid

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1 dynamics of molten rock systems.

2 It's more a matter of trying -- a lot of
3 the uncertainty really is trying to understand what's
4 going on in the subsurface from very indirect
5 evidence.

6 You know, we can't really observe physical
7 conditions in the eruption volcano. The eruption
8 products that we see have been highly modified by the
9 time we see them from their condition in the
10 subsurface.

11 So some of the work, some of the modeling
12 that was initially developed to understand volcanic
13 flow processes, did arrive from basic fuel cooling and
14 basic thermal fluid dynamic relationships.

15 But we're not trying to derive from first
16 principles these models. They're already from a
17 fairly established volume of literature.

18 CHAIRMAN RYAN: I appreciate the modeling
19 effort you have underway to improve the modeling of
20 magma, but I'm sitting here thinking of the question
21 how did any of these variations of your modeling
22 impact your basic assumption, which was the packages
23 entrained, and if I understood John Trapp earlier
24 correctly, that the package offers no confinement or
25 containment so all the radioactivity is in the magma?

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1 MR. HILL: There are two -- three ways
2 that this affects the source for the igneous
3 scenarios. I can't demonstrate this yet because we
4 haven't finished the modeling.

5 But the potential here is to understand
6 conduit widening processes in this disturbed geologic
7 settling -- excuse me, disturbed geologic setting.

8 Right now we're making an assumption that
9 the diameter of the volcanic conduit is completely
10 unaffected by the presence of repository grips. We
11 want to understand the stress distribution and flow
12 response in this disturbed regime to say whether or
13 not that assumption is supportable, or should have a
14 larger variation of uncertainty because of these
15 conduit drift interaction processes.

16 In other words could the conduit be larger
17 or elongated, or have a larger source-term for
18 volcanic eruption because of these flow interaction
19 processes that we're currently assuming?

20 CHAIRMAN RYAN: Let me ask you to follow
21 up on that point. If all the radioactive material in
22 a package or a number of packages is entrained, isn't
23 the source-term constant?

24 The concentration will vary based on how
25 much magma you have but the amount of -- that's

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1 involved is cut down constant.

2 MR. HILL: For one waste package.

3 CHAIRMAN RYAN: Well for any -- pick a
4 number I mean one, ten or 50. But what he's saying is
5 that one package or ten. Okay. I think I'm
6 understanding a little better.

7 MR. HILL: All right. I'm sorry I didn't
8 make that clearer.

9 CHAIRMAN RYAN: Okay.

10 MR. HILL: That as the conduit widens, the
11 number of waste packages intersected would also
12 increase.

13 CHAIRMAN RYAN: Okay. And then of course
14 there's the complicating feature of is it entrained or
15 is it sequestered in the end of some tunnel or
16 something. That's what I'd imagine.

17 MR. HILL: That was the second part of
18 this story, of the three part story.

19 CHAIRMAN RYAN: Okay.

20 MR. HILL: Is we have -- we're talking
21 about the direct volcanic release but we also have
22 what was called the indirect release scenario. Where
23 all the waste packages that remain in the drift.

24 Now we want to understand a better
25 mechanical approach to how these igneous events, for

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1 the duration of event and afterwards, can affect the
2 waste package performance.

3 We believe there is sufficient information
4 to show that as the magma was in place, and cooled,
5 and the stresses involved, would cause breaching of
6 the waste package for these waste packages that are
7 left within the drift.

8 By understanding the variations in
9 pressure through time and temperature through time, we
10 have a much better mechanistic basis to evaluate
11 potential waste packages response to the physical
12 conditions of magnetism, and evaluate the conservatism
13 or non-conservatism of the assumption of damage extent
14 in an intersected drift, as well as waste form
15 behavior following a potential intrusive event.

16 So even though we're not getting those
17 damaged waste packages out during the igneous event,
18 we're still following the even at the resumption of
19 normal hydrologic flow and transport.

20 And so the scenario here, the risk
21 significance, is that the conditions don't cause waste
22 package failure. Then we don't have a large number of
23 waste packages fail following an igneous event, and
24 there's no increase in the hydrologic source-term.

25 Conversely, if we're intersecting a number

1 of drifts, and the contact of magma with the waste
2 package is sufficient to cause failure of the waste
3 packages in those drifts, we have a large source-term
4 that has to be considered in performance assessment
5 given the condition of the igneous event.

6 CHAIRMAN RYAN: I guess one friendly
7 amendment I'd ask you to think about is that it's
8 really not a source-term yet. It's an available
9 inventory.

10 MR. HILL: Okay. I'm using that --

11 CHAIRMAN RYAN: Are we on the same page
12 here?

13 MR. HILL: -- loosely. There are many
14 steps to go between disruption --

15 CHAIRMAN RYAN: Okay.

16 MR. BRITT HILL: -- or potential
17 disruption of a package, and the release mechanism.

18 CHAIRMAN RYAN: I just wanted to make that
19 point that's fine. And the third one?

20 MR. BRITT HILL: The third one was these
21 horizontal doglegs and breakouts. We have it as an
22 alternative hypothesis in the Woods et al 2002 paper.

23 I think it's fair to say that we are less
24 concerned about that condition occurring during the
25 initial stage of the event, but still need to think

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1 out, and get a good basis for evaluating, developing
2 these breakouts at any time during an eruption.

3 And, of course, if we had a horizontal
4 flow path away from our existing conduit, that could
5 also entrain and potentially eject more waste packages
6 than we're currently assuming in the performance
7 assessment calculations.

8 So there's the three risk significant
9 impacts for the work that we're doing in this area.

10 MR. GARRICK: I just wanted to follow up
11 what Allen Croff said about the -- and the accident
12 aggression analysis --

13 CHAIRMAN RYAN: Turn that up please.

14 MR. GARRICK: That'd be a good idea. I
15 wanted to follow up Allen Croff's question as to
16 whether or not the technologies that have developed in
17 accident compression analysis of reactors have any
18 impact on the source-term development, particularly
19 with respect to entrainment, because there have been
20 enormous amount of work gotten done in this area, and
21 in some cases some major surprises as to the
22 confinement capability of the debris.

23 And it seems to me that this could have an
24 impact on the form and of the material that eventually
25 is in the cloud. And I just was curious, I wanted to

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1 press on that point a little bit, has this technology
2 been examined at all?

3 MR. HILL: That's one of the areas of
4 ongoing investigation. I'm hoping we're going to get
5 some insights today or tomorrow from some of the other
6 presentations, but I'm afraid that it's not at the
7 stage that I can really comment or report on.

8 We haven't had any major breakthroughs in
9 that area, and one of the major limitations is trying
10 to relate the physical conditions of those scenarios
11 with the physical conditions of an igneous event.

12 They're not very comparable, but still
13 trying to look at the -- how difference in physical
14 conditions may or may not affect our understanding.
15 It's not very straight forward.

16 And that's why we haven't been able to
17 make a lot of rapid process in that area. But
18 certainly this was something that Dr. Weiner had
19 mentioned earlier in the year.

20 We have been following up on that area all
21 -- a lot of other areas. I'm just afraid at this
22 stage it hasn't come to fruition.

23 CHAIRMAN RYAN: John, do you want to make
24 a comment?

25 MR. TRAPP: Just one comment. I did state

1 this morning that yes, we are making this assumption
2 on a waste package. Now, in order to understand how
3 a waste package is going to respond to this type of
4 environment, you have to better understand the
5 environment.

6 This is what this work is doing. It's
7 getting us a better handle of the mechanical thermal
8 environment than we have to put the waste package in.

9 And therefore there may be a possibility
10 that there could be some say to use Dr. Garrick's
11 terminology, a more realistic model that maybe does
12 say that some of a package can survive.

13 But unless you understand what is really
14 happening from a mechanical and thermal response in a
15 volcano you're just -- and there's more.

16 CHAIRMAN RYAN: And I have been looking at
17 the volcanology and flow of magma in it's purest form
18 to understand that. I appreciate that, but at the end
19 of the day it is important only in the context of a
20 waste package and its interaction with it, and then
21 what happens down the line.

22 And again I apologize for jumping ahead a
23 bit, but I'm trying to keep a whole range of parts and
24 pieces of this question in my head at the same time.
25 I appreciate your comment. Ruth?

1 MEMBER WEINER: Well I have -- should I
2 use the mic? I have a couple of possibly unrelated
3 questions. First one is, your talk is titled NRC
4 Review Capabilities for Evaluation Potential Magma
5 Repository Interactions.

6 Could you expand a little bit on how
7 you're going to use this model to review what DOE has
8 done? Are you going to say -- well I'll let you
9 respond to that question.

10 MR. HILL: All of us are faced with an
11 extraordinary challenge in trying to evaluate this
12 process. It's unprecedented in trying to look at the
13 interactions between a volcanic system and an
14 integrated system.

15 We have no good natural analogs, we have
16 no objective basis of comparison to use the part 63
17 terminology. So really about the only way that we can
18 try to look at eventually the risk significance of
19 this is by doing some of this actual work.

20 It's very typical to review something
21 that's state of the art if you're not actually doing
22 some things that are kind of state of the art. We
23 don't use this modeling as the correct way.

24 This is one insight of many possible ways
25 to approach this problem, but it does give us a

1 knowledge base to kind of understand in doing this
2 model, what's important, what's not important to
3 process level, and how this would affect our
4 understanding of the downstream risk impacts for this
5 challenging problem.

6 But again I want to emphasize, we're not
7 viewing this as setting the baseline for a comparison
8 of whether we're right or wrong or any other group is
9 right or not. Is that answering --

10 MEMBER WEINER: It does.

11 MR. HILL: Indirectly?

12 MEMBER WEINER: That's a partial answer,
13 and one of my questions was, well suppose DOE comes up
14 with an entirely different model, are you going to say
15 well ours is right and yours is not or ours is not and
16 yours is?

17 MR. BRITT HILL: Well --

18 MEMBER WEINER: But --

19 MEMBER WEINER: -- I would just very
20 speculatively say if -- there's the two conditions.
21 DOE comes up with a model for magma flow, and it
22 completely disagrees with our model, and we have a
23 completely different risk insight from it.

24 Were going to have a challenge in
25 evaluating who is right in that sense, or how we're

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1 going go. And it may be we just have to let the
2 licensing process take care of that.

3 But, conversely, we have alternative
4 modeling approaches, and they could be very close in
5 approach. We could be getting the same basic
6 insights.

7 And I think that would be a useful thing.
8 To -- we have no proof. We have no basis to say it's
9 right or wrong, unlike some other things. So we've
10 done an independent effort. The department has done
11 an independent effort.

12 And all these answers appear to be
13 conversing in about the same risk difference. And we
14 can evaluate the differences, and they could be
15 insignificant.

16 I think we've all done the best job we can
17 in that case. So I'm going to hope that we're going
18 to be successful in this. Unfortunately I can't
19 comment on ongoing reviews.

20 But I'm optimistic that this approach that
21 we're that using is not completely out in left field.
22 And let me leave it at that.

23 MEMBER WEINER: Thank you. That's a very
24 comprehensive answer. My totally unrelated question
25 is, your slide six has a rather elegant experiment

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1 analog, elegant in its simplicity.

2 And I commend you for that. Is there any
3 way to take your experimental setup and add analog
4 waste packages to it? You would face some difficulty
5 in scaling the scale up.

6 MR. HILL: Yes.

7 MEMBER WEINER: But is there some way that
8 you can do that because it seems to me that would give
9 you some insight at least into the pressures and gas
10 bubble interactions.

11 MR. HILL: This apparatus is designed to
12 look at the initial fold. I think the state of fluid
13 modeling is sufficient to show that, given these
14 conditions of flow for simple geometry, you know we
15 couldn't make an analog that was anything more than a
16 simple waste package anyway.

17 Whether or not these conditions are
18 sufficient to entrain or bump things around, and even
19 -- I'll fall back on the Woods 2002 paper -- even
20 under the conditions of optimized accelerated flow,
21 there may be slight movement but nothing that was high
22 enough velocity given this very simple geometry to
23 really pick things up and move them around.

24 So I think we have the insights we need
25 for the risk assessments that, with this apparatus and

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1 those conditions, we're not really concerned about low
2 impacts on a waste package.

3 We do have ongoing experiments that are
4 looking at sustained flow and circulation. And some
5 of those will be considering experimental analogs for
6 engineered systems, again, trying to gain insight and
7 model verification for the simplified calculations on
8 how these circulation effects may or may not affect
9 material located in various parts of the analog
10 system.

11 MEMBER WEINER: Is there a fluid dynamics
12 model that could help you model entrainment better?

13 MR. HILL: Yes. There's a number of them.

14 MEMBER WEINER: Yes.

15 MR. BRITT HILL: And again we are doing a
16 lot of work. This is why we're doing the basaltic
17 work at the University of Bristol crew, who have Steve
18 Sparks.

19 One of the people who's currently working
20 with us is Dr. Jerry Philips. He's not in the current
21 -- or presentation but he's another one of the leading
22 experimentalists fluid dynamics people for magma
23 repository interactions, also a parallel reference
24 going at Cambridge University, with Professor Andrew
25 Woods, and some of his colleagues, to really come up

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1 with solid fluid dynamic basis to evaluate these
2 processes to the best of our ability, and without
3 uncertainties.

4 MEMBER WEINER: Thank you.

5 CHAIRMAN RYAN: Panel members, any other
6 questions, comments? Bill?

7 MR. HINZE: A couple of quickies. The
8 ICPR, the Igneous Consequence Peer Review Panel
9 certainly made it clear that they felt -- I think the
10 words are most unlikely that there is a dogleg.

11 And I gather that your work suggests that
12 that's not the case, that a dogleg this possible. Or
13 are you talking -- is the difference here related to -
14 - of preexisting zonal weakness?

15 How do you rationalize what you're saying
16 with what the ICPR has come up with?

17 MR. HILL: I'm afraid I can't comment very
18 much on the DOE's peer review comments. What I can say
19 is that there have been -- we have at a subjective
20 level, and like I said in this presentation, we don't
21 believe that this is a likely scenario.

22 However, all the assessments have been
23 very qualitative and subjective about -- well it seems
24 less likely, but we're not -- have not received a
25 input that says it cannot happen.

1 And we have to be careful in
2 distinguishing between what may or may not occur in
3 the initial stage of potential interactions, versus
4 what may or may not occur for the duration of an
5 event.

6 And I think it's fair to say that none of
7 us have presented a model or an analysis that truly
8 looks at the evolution of the system for the duration
9 of an event.

10 Most of the work is focused on the initial
11 stage of interaction. So I can't go too much farther
12 with that.

13 MR. HINZE: Do we have a -- an explanation
14 for the current boccas in nature? Is this a chocking,
15 rocking of the main conduit? What causes this?

16 And how do you get at the likelihood that
17 this may happen as you say in, in -- to carry it up?

18 MR. HILL: I certainly -- my personal
19 opinion is we don't have a great understanding. But
20 what we can see from observation is especially -- I
21 used -- as the best example right now, just because
22 they're isolated cinder cones.

23 Mass flow rates are very comparable to
24 what we would expect in Yucca Mountain. And there
25 were a number of boccas that formed during the course

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1 of this event.

2 There wasn't really a hiatus in activity
3 in the central cone, and a shift to the bulk. What
4 you saw was simultaneous eruption of lava and tephra.

5 Sometimes that eruption is occurring from
6 the central cone itself. Other times the diffusion
7 rate of lava may decrease slightly from the central
8 vent, but increase at a bocca.

9 But it's not really a straight forward
10 thing where you shut down the main conduit and have
11 everything coming out of the bocca. It's a much more
12 complex plumbing system than that.

13 So I'm drawing a cartoon in a cartoonish
14 view that's a great simplification. But I'm in that
15 intermediate position of here's what happens in
16 nature.

17 This could potentially affect our risk
18 understanding, but I know there's a lot of
19 complexities. And it's not a one for one for one
20 scaling relationship either. So we're working on
21 that.

22 MR. HINZE: Speaking of the natural
23 conditions, I remember forty years ago going on a
24 field trip to Kings Fall, and the Snake River plains,
25 and seeing the basaltic dikes which obviously had

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1 backwash.

2 MR. HILL: Correct.

3 MR. HINZE: And where are you in modeling
4 or could you be in modeling that considers backflow
5 and how do you approach it, and so forth?

6 MR. BRITT HILL: What I can say is this is
7 part of our ongoing work at the University of Bristol.
8 We are looking at circulation and flow effects for
9 steady and non-steady flow conditions in a conduit
10 system.

11 Now, again, it's very hard to get people
12 to climb up to the center cone to look down when
13 there's a hiatus in the eruption. I had a colleague
14 in Nicaragua that did this, and he was very lucky, but
15 he didn't see too much because you get a lot of rubble
16 coming in.

17 The cinder cone conduits, it's very hard
18 to say whether you're going to get a hiatus in the
19 eruption, that would cause -- to rain to hundreds of
20 meters below the surface.

21 It seems kind of unlikely, but I can't
22 eliminate the possibility. But it is a very different
23 kind of conduit system than what we would see at these
24 rift dominated systems like Craters of the Moon, which
25 are very Hawaiian.

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1 And you've seen drain back and cessations
2 of eruption as a very typical feature in those kind of
3 systems.

4 MR. MELSON: Yes, but I think certainly we
5 it should be happening here. It's one of those models
6 that I think should be considered, at least just
7 proven that it can happen.

8 MR. BRITT HILL: And I think with the
9 ongoing work, we will be able to evaluate whether that
10 kind of drain back phenomena would have an effect or
11 no effect on the engineered system.

12 MR. HINZE: Let me ask a last question.
13 The last bullet of your summary of current information
14 continues to refine these models. Where are you?

15 How much do you have left? Do you have a
16 feel for this at all? What are your plans?

17 MR. HILL: Well, what I think it's pretty
18 obvious that we haven't presented anything that looks
19 at the specific geometry. So, obviously we're going
20 to be applying that through a complex geometry that
21 would represent potential dike drift systems.

22 That's why the first step if, of course,
23 makes your numerical, gain confidence in the model.
24 Start simple, build upon that, and then apply it once
25 we have an understanding that we're not just modeling,

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1 but modeling something in a reasonable way, and then
2 apply that to the complex system as the final stage.
3 That is ongoing work.

4 MR. HINZE: I understand how research
5 works and how we can keep solving more problems, but
6 if we leave this problem out and -- have a duration.
7 Are you writing this proposal for two years, four
8 years?

9 Where are we now from zero to ten, in this
10 whole understanding of the intersection of the magma
11 with the repository? Where are we, from zero to ten?

12 Give me a number and I'll give you what
13 percent we need.

14 CHAIRMAN RYAN: That's a so odd question
15 Bill?

16 MR. HILL: The goal is to have the work
17 completed and of course written by December 15th.

18 MR. HINZE: Okay.

19 CHAIRMAN RYAN: Bruce, you had a question?

20 MR. MELSON: I sat through as a number of
21 people did here, the Igneous Consequence Peer Review
22 Panel, where much of it dealt with the Woods paper,
23 which I assume was contracted by your group and what
24 not.

25 And a lot of interchange went on there of

1 a very substantive nature. I don't know if you've
2 ever gone through the documents that came out of there
3 or not but I'm just -- I mean it's so construct in the
4 sense of allowing you to look at your work, and you
5 said you couldn't talk about that.

6 And so I was just wondering -- because you
7 never did mention it, as if it never even happened.
8 But I assure you it did happen. And I'm wondering, is
9 it a license thing or something?

10 Is there some legal reason you can't deal
11 with people who are feeding what I hope is
12 constructive criticism into your work.

13 MR. HILL: I thoroughly appreciate the
14 desire to be able to have open communication and open
15 interchange on topics all over the map. Unfortunately
16 we are approaching a very complex legal arena.

17 And we have to be very sensitive on what
18 we communicate in terms of publicly available
19 information and the format in which we communicate
20 that information.

21 It is unfortunate that we have not
22 documented the results of our review and our thoughts
23 of the DOE's Peer Review Panel. We have not done
24 that.

25 And so, in this format, I unfortunately

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1 can't give extemporaneous feedback on our thoughts
2 that are not in the -- and, again, this is solely
3 because of the approaching license application
4 deadline, and a concern about Staff's independence,
5 and the various roles that different groups are going
6 to need to maintain during this complex legal
7 proceeding.

8 So that is why there are very obvious gaps
9 in this presentation. I am not commenting on our
10 views of the current DOE models, or anybody's models.
11 I can't, not at this stage.

12 CHAIRMAN RYAN: John?

13 MR. TRAPP: Just -- one thing I can say is
14 there is a document that is currently in review at the
15 Nuclear Regulatory Commission. It's called the
16 Integrated Issues Resolution Status Report, revision
17 two, or one.

18 It's the revised issue resolution status
19 report. I think will be issued in the next couple of
20 months.

21 PARTICIPANT: Sooner than that.

22 MR. HILL: Sooner than that. Well, we're
23 hoping to get it out as soon as possible. That will
24 have the staff's current view of many aspects of the
25 DOE program, including some of the comments, and major

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1 process level concerns that came about from the peer
2 review meeting.

3 So, I guess a delicate way of putting it,
4 a lot of those comments from the DOE Peer Review have
5 been incorporated into the latest DOE documents as
6 well.

7 So, while we may not be commenting in the
8 integrated Issues Resolution Status Report directly on
9 the peer review conference, we will be commenting on
10 the Department of Energy analysis and lava reports
11 that have already incorporated those comments from the
12 review.

13 So we're not just leaving these comments
14 out of the vacuum. They will be addressed. But
15 unfortunately I can't do that today.

16 CHAIRMAN RYAN: Britt, thanks for your
17 detailed explanation of that. I think we can get a
18 clear picture of where you are and why, so --

19 MR. BRITT HILL: Sorry to be --

20 CHAIRMAN RYAN: It's all right. John, you
21 want to make a comment?

22 MR. TRAPP: I just want to carry on with
23 what Britt was saying, because when we got into this
24 whole thing and talked to our lawyers, what did you
25 see?

1 CHAIRMAN RYAN: You know what, I'll tell
2 you. We're really here to discuss technical issues,
3 and the position that you're in with regard to all of
4 that we kind of understand that --

5 MR. TRAPP: I just want to say that we
6 cannot discuss a lot of these things that we'd like
7 to. And I'm sorry, but that's --

8 CHAIRMAN RYAN: I think we all appreciate
9 that very clearly. And you have both done a nice job
10 of explaining that to us so I want you to realize that
11 it's, from at least the Committee's viewpoint, not
12 negative.

13 We understand that. So we appreciate the
14 position you're in. Neal you had a comment?

15 MR. COLEMAN: Yes. Neal Coleman, ACNW
16 staff. At the spring AGU in Montreal, I had the
17 pleasure of hearing a lecture by Michael Menga, who's
18 a leading authority on behavior of magmatic conduits.

19 I noted from your slide eight, under the
20 assumptions in your simplified model, no gas losses.
21 He stressed in that talk that it is very critical how
22 much gas escapes.

23 It has a very key role in determining the
24 power -- the energy of an eruption. So you -- this
25 isn't a question it's a suggestion. You want to

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1 include that kind of information, and also recognize
2 that the tuffs at Yucca Mountain are highly permeable.

3 And I've given you something to start a
4 range of say 10 to the minus 10, to ten to the minus
5 12 meters squared.

6 MR. HILL: What's the diameter of the
7 drift? About 20 meters square. So yes there is the
8 permeability, we agree. But it's not completely
9 permeable.

10 In relation to the area there isn't that
11 much permeability, if you're talking about
12 compressibility.

13 MR. COLEMAN: What I'm giving is --

14 MR. HILL: These assumptions aren't
15 clearly stated. It is not meant to be a realistic
16 representation of all things. Now to address your
17 second or first comment.

18 We've also done some very simple
19 calculations that anybody can do about bubble rise
20 speeds in magma and the magma that would be ascending
21 on the order of a tenth of a meter per second or a
22 meter per second, in the crust.

23 And bubble rise speeds are orders of
24 magnitude below the magma ascend speed. So, when you
25 want to talk about bubbles escaping from the magma

1 system, you have to come up with a mechanism to
2 segregate the bubbles beyond just simple buoyant rise.

3 While I agree that, yes, there are gas
4 escape effects, there isn't that much gas escape in
5 the rising magma until you get to very shallow levels.

6 And again when you want to talk about a
7 fully realistic realization and simulation of magmatic
8 process, that's great. We all know that these are not
9 single phase flows.

10 But to try to model them as two phase
11 flows is a heck or a challenge. It takes a lot of
12 computational resource. We're just trying to gain
13 first order insights on this, not think that we're
14 going to come off with a fully four-dimensionally
15 realistic model. That's not our role.

16 CHAIRMAN RYAN: I hate to cut off the
17 exchange but we are running long on this talk. So if
18 we can perhaps finish up with perhaps, Neil, one last
19 comment. And then I think Dr. Marsh has a question
20 and we'll finish there.

21 MR. COLEMAN: There may be a lot to learn
22 in the work that we're finishing up from a world
23 expert on this topic.

24 MR. HILL: That's why we're involved in
25 world experts.

1 MR. MARSH: This is very interesting,
2 Britt. And these kind of calculations are very
3 difficult to do in a single phase. The problem with
4 a lot of calculations is -- in the earth for example
5 is that -- especially in these type of calculations
6 and all of our sciences that initial conditions are
7 always a problem because, in general, in the earth we
8 don't know any initial conditions.

9 And yet, to solve the problem we need
10 initial conditions. And we don't know initial
11 conditions when the earth formed. We don't know the
12 initial conditions for how any single crystal is
13 grown.

14 We don't have initial conditions for how
15 a volcano starts out.

16 But -- so what do you do? Well you
17 actually impose the problem so that the solution does
18 not depend heavily, entirely on the initial
19 conditions.

20 Now, if you go back to the conduit
21 experiment there, your horizontal -- could you show --
22 Britt.

23 MR. HILL: I don't have control.

24 MR. MARSH: Okay, Bruce, could you go back
25 to the experimental level? Six. So, for example,

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1 just to use this as an example, Britt was saying
2 himself about how things start up.

3 They start off very gradually and open
4 very small. In fact you can see arrested sills and
5 arrested dikes in places and you can actually see way
6 out -- a kilometer two kilometers in front of them.

7 It's a tiny little one centimeter or even
8 just a millimeter crack that starts out and enlarges
9 maybe up to a couple hundred meters or so. So, in a
10 situation like this, you have to start the problem
11 somewhere of course.

12 And so they start the problem at the
13 nozzle, give it overpressure. And suddenly you open
14 this thing. And so if you actually tweaked it open
15 just very, very slowly, you would get a different
16 solution than if you just open it up, or if you
17 puncture a membrane or something like this.

18 So the initial conditions are very
19 important. Also the pressure drop is enormously
20 important. And you can see that the ensuing
21 velocities that you get -- and this is a flow that has
22 no volatiles in it.

23 But it has -- you know the meters per
24 second. So in other words that flow would be in Las
25 Vegas in a few hours, and -- from Yucca Mountain for

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1 example, the volumes would be enormous.

2 So how do you actually temper these
3 things? You temper these kinds of calculations by
4 looking at the geological record of what you actually
5 see in terms of how fast the lava is actually --
6 emerge, how fast the flow field gets larger and
7 larger.

8 And I think, for example, Lathrop Wells
9 and things like that you actually have some control
10 over looking at how fast these things advanced. When
11 you actually use two faces flow and tephra shooting
12 out it's a little bit different.

13 But the other factor -- another assumption
14 of course is they all start out at one meter wide --
15 dike in the calculations and you move from that. Now
16 that's an initial condition assumption also.

17 The other thing is that there's no
18 solidification whatsoever in depressurizing. The
19 volatile is coming out of solution. What is does it
20 changes the entire phase.

21 If you want to go up to the phase diagram
22 -- five. Let's go to five. So this phase diagram for
23 example, up in the left hand, the dashed or shaded
24 area there, if the magma starts in that region it
25 starts to ascend to the surface.

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1 So you have pressure on the left, so it
2 has to come to the surface. And so, what happens is
3 it can't hold its -- a magma's like a diver, it gets
4 the bends. This stuff comes out of solution, and as
5 the bubbles can't escape, Britt talked about it drives
6 the whole eruption and you get this back and forth all
7 through.

8 But never the less it has to come out in
9 solution so with the phase diagram it starts going to
10 a low water content and the whole thing shifts to the
11 right.

12 And at those temperatures, 980 degrees --
13 it actually starts out at that by the time it gets to
14 the earth's surface, it wants to be solid, entirely
15 solid more or less.

16 Something at 55 degrees -- 55 percent
17 crystals can't erupt because it's at maximum packing.
18 It's a solid and chokes and conduit, becomes
19 explosive, for example.

20 But it will actually shut down the whole
21 system. That's what volcanoes are in many ways.
22 They're like -- they're trying to shut down so they
23 build and pile on top of it until they actually shut
24 themselves down and they degas.

25 So the suffocation effects are enormous,

1 in terms of loosing -- just loosing volatiles. And
2 that's a way that the volatiles actually get trapped
3 out of the system, partitioned out.

4 And this is like many have talked about
5 somewhat. You have an enormous solidification effects
6 and they start eating out this stuff.

7 What happens in fact the volatiles are
8 richest where the solidification is the greatest of
9 course, because the amount of melt is low, and there
10 is excess melt of fluid around it?

11 So not only here is it very important and
12 this is coming out, but in a volatile free
13 environment, any time the magma rises up into conduit,
14 a major problem is it starts undergoing thermal death.

15 In other words the solidification is at
16 right angles to the flow. So heat loss is at right
17 angles to the flow field. So, no matter what the flow
18 does, it cannot actually impede solidification it just
19 starts growing and growing and growing and trying to
20 stop the conduits.

21 And the conduit either has to overpressure
22 itself, keep opening itself to offset this
23 solidification. But these effects can be enormous. So
24 the thing that I would like to see in this -- these
25 kinds of -- this kind of work -- you know an

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1 investigation of initial conditions, as related to
2 what kinds of things we see on the surface, in terms
3 of related to the volumes of output and the times, of
4 course big time duration effects.

5 Systems start out -- it's like puncturing
6 a balloon and then you pressurize -- and affects its
7 solidification of the reoccurring viscosity. These
8 are really big factors that I'm sure you agree that
9 they're very hard to model, but they -- these are
10 critical issues.

11 And it would be nice to see that kind of
12 approach in here somehow.

13 MR. HILL: Well again, the approach we've
14 been taking is trying to build in the realism if we
15 can. You know, certainly I agree that the effect of
16 volatiles is very profound in fully thermal mechanical
17 effects that is going to occur during a complex
18 depressurization.

19 But we do know, for example, with these
20 volatile contents, that we've erupted a number of
21 volcanoes in the Quaternary that have a cone phase, a
22 tephra phase, and a lava fall phase.

23 So there still is the ability of these
24 magmas for -- of descent to avoid the thermal death
25 and crystallization choking, because we see them at

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1 the surface.

2 CHAIRMAN RYAN: I think we could go on for
3 a long time. I'd like to bring this to a close.

4 MR. MARSH: The real question is that, the
5 incorporation of these effects in the modeling. We
6 certainly see magma under the surface. That's not the
7 question today.

8 But are you going to try to put these
9 kinds of serious issues in the modeling, that's --

10 MR. HILL: We're trying to build in the
11 variable viscosity. We're trying to build in thermal
12 control as well. But you're again getting into, like
13 you're saying, very complex models.

14 But you know if you want to start building
15 in some sort of experimental apparatus that has
16 variable viscosity, variable openings, and variable
17 temperatures, that's quite a challenge.

18 We're starting with the apparatus, because
19 it's an established apparatus, and it would gain us
20 the insights that we needed in the initial state. So,
21 while I appreciate the desire like we all have to make
22 this as realistic as possible in our models, we do
23 have a limitation in the knowledge, and in the ability
24 to duplicate this unusual situation in the lab as with
25 a computer.

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1 We're working on it. A number of people
2 are working on it.

3 CHAIRMAN RYAN: Thanks. Thank you for
4 your presentation and for the good discussion there
5 after. I'd like to move now to our next presenter, if
6 we could. I think Dr. Matt Kozak will be making a
7 presentation on the Alternative Views of Modeling of
8 Magma Repository Interaction at Yucca Mountain.

9 And that -- you'll be handling all
10 comments?

11 MR. KOZAK: Yes.

12 CHAIRMAN RYAN: All right. Is John going
13 to make a comment?

14 MR. KOZAK: No John just said it --

15 CHAIRMAN RYAN: Okay, Great.

16 MR. KOZAK: Now this is still --

17 CHAIRMAN RYAN: No that's a lot easier.

18 MR. KOZAK: It's a lot easier. I don't need
19 to project from the diagrams.

20 CHAIRMAN RYAN: No. You can if you want.

21 ALTERNATIVE VIEWS ON THE MODELING OF
22 MAGMA/REPOSITORY INTERACTIONS AT YUCCA MOUNTAIN

23 MR. KOZAK: Well first I'd like to say thank
24 you very much for the opportunity to come here and
25 present this work. I'd like to emphasize that this is

1 the work of the project team.

2 Next slide please. Project team
3 contributors, a number of whom are here today, myself,
4 Mr. Apted, Mr. Bursik, Shane Findlan, Randy James,
5 John Kessler, Frassier King, Mick Morrissey.

6 MR. MARSH: Excuse me, might I -- are there
7 handouts for us?

8 CHAIRMAN RYAN: You did not bring hard
9 copies for us.

10 MR. KOZAK: Yes.

11 CHAIRMAN RYAN: You did?

12 MR. KOZAK: Yes.

13 CHAIRMAN RYAN: They're on their way,
14 thanks.

15 MR. KOZAK: Before I really get into it I'd
16 like to remind you about the EPRI's role in this whole
17 process is. That is where federal agencies have -- we
18 look around us and we follow the work of the Federal
19 agencies and we try to fill in things that perhaps
20 they're not doing.

21 And so we looked at this and we looked at
22 some of the arguments related to the probability of
23 the event that we talked about earlier. And decided
24 that our time was really better spent, and we could
25 take a bigger impact by starting to look at

1 consequence side of things.

2 First and foremost was to look at the waste
3 package. We have this waste package that's one of the
4 toughest materials known to humanity. And really in
5 and really in the DOE and NRC analyses up through
6 TSPASR which is the most recent information that we
7 have available to work on.

8 The waste package plays no role. It
9 essentially disintegrates as soon as it's contacted by
10 the magma, and is blown out to the top. So we
11 thought, that by looking a little bit more in detail
12 at that, we might be able to develop a more reasonable
13 assessment.

14 Next Please. And to do that we're trying to
15 develop a reasonable expectation base. To give you a
16 heads up of where we're going with this, what we end
17 up finding out is that we can't get the material out
18 of the waste package when we start taking into account
19 the degradation processes that can occur during the
20 eruption cycle.

21 That's not to say that the igneous -- the
22 magma does not have an effect on the waste packages,
23 but during the time frame of the eruption, the waste
24 packages are not damaged sufficiently to cause
25 releasing during that period.

1 They may be sensitized so that as we heard
2 earlier that the ground water releases at later times
3 would be affected. And we're going to be looking at
4 that more in a future report.

5 Next please. Let's just go through. It's
6 sort of instructive to look at it when we start
7 looking at the waste packages themselves, to look at
8 the sequence of the events of the eruption.

9 And the first is that this dike rises sort
10 of as a sheet to intersect with the repository. And
11 the -- as that rises through, there is significant
12 amount of degassing that goes on in there.

13 The surface area at that point that could
14 even take into account surface of the amount or
15 degassing what might have taken effect is very hard
16 because it's perhaps a kilometer long dike that's
17 coming up as a sheet through the zone.

18 And so there could be significant degassing
19 that's going on during that process. So the dike
20 eventually raises to the repository level. Next one
21 please.

22 We get perhaps an intersection if it
23 intersects with a drift, you get an intersection
24 perhaps with a single waste package. Meters -- meter
25 diameter -- meter with dike coming up make it a single

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1 waste package.

2 So we consider that to be an initially
3 impacted waste package, which is what we're thinking
4 of when we're talking about the initially impacted
5 waste package in this presentation.

6 And then over a period of time, that dike
7 evolves into these cylindrical vents. Next please.
8 And, at one or more places along the dike, it may
9 evolve into these vents.

10 Now, over a period of time those vents can
11 widen, so that initially impacted waste package, which
12 may or may not be on the center line of where a vent
13 would occur.

14 If it's not, then, when it evolves to the
15 vent stage, then that waste package is not going to be
16 contributing to releases up the vent. If it does
17 happen to form a vent on the center line of a drift,
18 then that waste package and perhaps a few more to
19 either side of that waste package can contribute to
20 the release that might occur coming out of the vent as
21 it heads toward the surface.

22 Now if you look a little bit further out
23 down the drift, down here, we have the magma is coming
24 up and it's going to start flowing out down the drift.

25 And the heat losses that are going to

1 happening as the magma flows down, would sufficient
2 that fairly quickly the magma's going to solidify and
3 we're going to get a basalt plugs down the drift.

4 So, the number of waste packages that can
5 contribute is the number that are actually actively in
6 this vent at that state of the eruption site. And the
7 ones that are further down there's really no driving
8 force for radionuclides, even if those waste packages
9 were breeched, there's no driving force for that --
10 the radionuclides to come up a gradient of magma, to
11 think of it that way, to come back up and go up the
12 conduit.

13 So if you can identify the width of this
14 conduit, and the number of waste packages that could
15 conceivably be affected by that, and consider the fact
16 that there could be more than one conduit, then that
17 gives you an idea of the bounds on the number of waste
18 packages that could be affected by it that contributed
19 at this state in the eruption.

20 Next please. And then at that stage we have
21 to get this tephra plume to come out here 18
22 kilometers. One of the important things to recognize
23 is that as this initial dike comes up, the initial
24 dike intersects with the drifts at the drift level, at
25 that stage, the type of eruption that comes up when

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1 the dike hits the surface is fire -- and at lava
2 flows.

3 There's very little tephra production at
4 that stage. It's not until it evolves into this vent
5 stage that we start getting these significant tephra
6 plumes.

7 So we find that, if the waste package
8 initially were to just dissolve as soon as the magma
9 were to hit it, and the fuel were to just come up as
10 chunks -- which is in a way a conceptual model for
11 some of the DOE and NRC models, except the waste
12 package is just instantaneously gone -- that if it
13 were to come up at that point along with the dike
14 eruption, that it would be coming out in lava, and it
15 wouldn't come down and it wouldn't affect the record.

16 It's only if the waste package does not fail
17 at that initial stage, it fails at some intermediate
18 stage later on, so that whatever radionuclides are
19 released could come out during this period, then it
20 would be associated with tephra that could get down to
21 producing dose.

22 Next please. To look at this we broke the
23 system up again to initially impact the waste packages
24 and neighboring waste packages that might be in a
25 vent.

1 And what we did was go through sort of a
2 logical diagram of what type of effects we need to
3 take into account at each stage of the eruption. And
4 the main thing to recognize in this initially impacted
5 waste package is, as you come down there, if the C22
6 and the stainless steel shells were to completely be
7 destroyed upon initial impact, again, if you come down
8 here to bottom, you have decision branch to say
9 whether or not it's actually going to get down to the
10 RMEI.

11 And, for these types of eruptions, as it
12 comes out, it's not getting to the RMEI to produce
13 zero dose if that were the case. So, but we did want
14 to evaluate that.

15 We wanted to find out whether or not the
16 conditions at that stage of the eruption were severe
17 enough to cause this extreme damage. And we'll be
18 talking about that a little bit later in the
19 presentation.

20 We also have a logical diagram of how the
21 waste packages can fail and the types of failure
22 mechanisms that would come in for these waste packages
23 that are not initially in the dike.

24 CHAIRMAN RYAN: You don't want us to read
25 that?

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1 MR. KOZAK: No. The details of this are,
2 you know, we'll be going through each of the sort of
3 the phases of that as I go through the talk.

4 CHAIRMAN RYAN: Okay.

5 MR. KOZAK: One of the things that we wanted
6 to do was to look at this issue of magma down the
7 drift and the evaluation of this dogleg of magma that
8 could come down a drift and then go up someplace else.

9 We wanted to look at the waste package
10 failure mechanisms. Next please. I wanted to talk a
11 little bit about this Nicholis and Rutherford paper.

12 And I'll come back to it a little bit later
13 in a different context. But, when we look at the view
14 of the TSPASR conditions of which the initial dike
15 hits the drift, they're looking at magma temperatures
16 up to 1,200 degrees centigrade and on the order of
17 four centimeters per second for the upper end of that.

18 This paper by Nicholis and Rutherford, which
19 was referred to in the previous presentation, one of
20 the other implications of that paper is that the
21 temperatures that are consistent with the observations
22 of the basalt, would be much lower.

23 That's very important for waste package
24 performance. And it is very important for the
25 viscosity and the cooling rate of the magma as it

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1 potentially flows down the drift, and also in terms of
2 the ascent rate.

3 If you have a much more highly viscous
4 fluid, it's going to be moving at much slower rate.
5 This actually -- the conclusion that Nicholis and
6 Rutherford came to said that it needed to be greater
7 than four centimeters per second.

8 But my argument here is that, we're probably
9 at the lower end of the magma ascent rates because of
10 the higher viscosity and lower temperature.

11 That's an interpretation of their
12 information. It's not from their paper directly. And
13 we actually did not use this information as part of
14 our evaluation of the extrusive scenario.

15 So, the results that we will be showing here
16 may actually be more conservative than they would be
17 if we were to have done it, including the Nicholis
18 Rutherford evaluation.

19 The first thing that we want to look at was
20 the rise of the magma into the drift, the flow of
21 magma along the drift, and to look at whether or not
22 these shockwaves and the potential formation of a
23 dogleg, a new pathway to the surface types of
24 phenomena could occur as we show in the Woods et al
25 model.

1 The evaluation that we did with using this
2 computer code called SAGE -- and Dr. Morrissey is here
3 to fill us in on the details later if we'd like to.

4 Essentially this is a code that was
5 developed for evaluation of underground nuclear
6 testing. So, it's fully coupled -- it's a big,
7 elaborate, fully coupled heat mass -- to transport,
8 and has a good deal of acceptability in certain very
9 exotic communities of subsurface flow phenomena.

10 In the SR, the DOE -- some of the initial
11 temperatures on the order of 1,200 C and corresponding
12 viscosities on the order of 140 pascal seconds.

13 If we look at the Nicholis and Rutherford
14 information, we are actually down in the much lower
15 temperature range and the viscosities are right near
16 a break point.

17 So, as soon as the temperature starts to
18 drop, the viscosity is going to go up very quickly.
19 And that's an important point because, as the dike is
20 rising up to the repository level, if the magma begins
21 to flow down the drift, it's going to cool very
22 rapidly.

23 There's no other heat source around there.
24 And, as it starts to cool, that viscosity is going to
25 sky rocket. So, we're not going to get these huge

1 water-like flows going down the drift.

2 This is going to be more like the flow at
3 the surface -- at the ground surface where we see sort
4 of clinkers forming at very gradual lava-like flows
5 into the drift and solidifying and setting up as the
6 initial stages of basalt formation at anything offline
7 from where the drift comes up.

8 Next please. This is -- go back. Yes.
9 Could you back up one? There we go. Okay. This is
10 the first few seconds of the dike rising into the
11 drift.

12 If you squint very carefully, off to the
13 right in the drift you can see that it's a little bit
14 narrower. And that's the effect of the waste
15 packages.

16 The waste packages are in this half of the
17 drift. And there's no waste packages in that part of
18 the drift. What this is a plot of pressure as a
19 function of time just as the initial tip of the dike
20 hits the drift.

21 And, what we see is the formation of a high
22 pressure center immediately above the dike, and
23 relatively low pressure propagating out from there.
24 There is no indication of the shockwaves of pressure
25 moving up and down.

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1 By the way, this is essentially a two
2 dimensional version of the Woods model that showed the
3 shock pumps.

4 MR. MARSH: What's the real time?

5 MR. KOZAK: This is on the order of a few
6 tenths of a second, I think.

7 PARTICIPANT: What it's showing you is you
8 take the Woods et al model of gas entering into the
9 tunnel. And, if you put the vertical dike into the
10 horizontal drift, what you're going to see is this
11 pressure concentration on top of the drift where it's
12 fairly high.

13 And so, that's what we really wanted to
14 show. You're not going to get these larger pressures
15 down the drift to the dogleg scenario. You're going
16 to have the continuation of any dike up through.

17 MR. KOZAK: And this very rapidly exceeds
18 the fracture stress limit for the rocks above the
19 drift. And so, the dike just continues going on the
20 way it wants to go, rather than coming down here and
21 creating a dogleg.

22 Next, this is the temperature on the same
23 timescale, same space-scale. Here we're assuming no
24 heat transfer at the boundaries. You can see the heat
25 moving up here.

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1 The main reason for putting this up is to
2 show that we can do fairly complex heat transfer
3 behavior in this zone and look at what the effect of
4 the heat on the waste packages is, and, to demonstrate
5 the behavior.

6 Next, please. So, the implication of this
7 modeling is that the down-drift pressure is much less
8 than the pressure above the drift. The pressure above
9 the dike rapidly exceeds the fracture stress limit.

10 So, the dike will continue straight to the
11 surface. And we don't have a dogleg in that case
12 because there is not these high pressures to create
13 them.

14 And that shockwave appears to be an artifact
15 of their 1D model. And, when we put it into the
16 similar kinds of conditions into the 2D model, that
17 disappears.

18 Next please. The next thing we did was to
19 look at the effect of this initially impacted waste
20 package -- rising up, hitting it to a single waste
21 package that just happens to be located over the dike.

22 We wanted to see what happened. So, what we
23 wanted to do was look at sort of a worst case on this.
24 And, what we did was to evaluate a -- even though
25 we're looking at the dike stage where this is a

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1 plainer source of magma that was easier
2 computationally to look at something that's more like
3 a cylindrical jet coming up and hitting the waste
4 packages.

5 And so, this is a rising column of magma
6 with magma densities and treated as a -- space for the
7 purposes of calculations. The properties are
8 associated with the temperature of the magma and the
9 temperature assumed to be of the repository early on
10 in the repository history for the initial conditions of
11 the waste package.

12 And, what we did was a collision calculation
13 using this detailed finite element model with this
14 code called ABAQUS/Explicit, which is an EPRI code, I
15 believe.

16 Well, it was developed to support EPRI
17 programs for different kinds of waste package and
18 other pressure vessel failures under collision
19 scenarios.

20 So, what we do is we model it as an impact
21 with very conservative conditions. Again, we consider
22 the 100 meters per second to be an absolute maximum
23 that you can see under these circumstances, -- for my
24 diagram width of the magma column, and then we're
25 going to look at what happens.

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1 Next please. These are very early in the
2 collision. This is on the order of a 100th of a
3 second. You can see the level of detail that goes
4 into the model.

5 We have -- the waste package internals are
6 modeled and the deformation of the shell is calculated
7 as a function of time. Move on to the next one.

8 This is a couple hundredths of a second
9 later. And you can see that, because of the
10 collisions, it is showing that the internals of the
11 waste package are being disrupted.

12 They are not completely filling the waste
13 package. So, it's talking into account all of these
14 things. This analysis, even though we're only showing
15 the first couple hundredths of a second, the analysis
16 was carried all the way through for the full
17 collision.

18 And actually, because of the extreme
19 conditions that we put on it, there was a second, very
20 substantial collision with the roof of the drift.

21 So, the magma comes up, it hits the roof of
22 the drift, and then we look at what happens and how
23 much damage there is. Next please. There is a cut-
24 away view of a sequence of what happens to the fuel
25 elements and the internals to the waste package.

1 Initially we've got it in pretty good shape.
2 By the end, there is a fair amount of damage to the
3 internals. And, just to summarize this without going
4 into the bloody detail of the calculations.

5 Our estimate is that the energy applied to
6 the waste package, if you look at the direct
7 coefficient kind of calculations, is on the order of
8 100 to 10,000 times what it would actually be
9 experiencing and the type of rise rate that we think
10 are reasonable.

11 So, it's extremely conservative. Based on
12 this paper by -- that I was talking about earlier, the
13 Rutherford paper, the rise rates may be on the order
14 of centimeters per second rather than on an order of
15 tens of meters per second.

16 But, even under this very extreme condition,
17 it wouldn't break the package. We had a structural
18 dent, and possible minor tearing on the C22 shell.

19 There was damage to the internal elements.
20 But there was no rupture of the internal structural
21 shell, the stainless steel shell. And so, to
22 summarize, simply the impact, none of the other
23 failure mechanisms, but just from the impact, we don't
24 get any release.

25 PARTICIPANT: Is the weld as strong as the

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1 container?

2 MR. KOZAK: The weld, I believe -- we'll get
3 out of my area of expertise pretty quickly. I think,
4 at high temperatures, the weld beings to lose its
5 strength faster than the material.

6 But, at the temperatures that we're at here,
7 which is at the repository temperature, they are of
8 comparable strength. So, that becomes more important
9 later on.

10 Then, as the temperature really elevates,
11 then we look at a little bit longer term effects.
12 But, even there, it's not the weld that's the critical
13 failure or critical potential failure.

14 So then, if it doesn't fail because it has
15 been slammed by the dike, now we're going to look at -
16 - we still only have one waste package that is
17 affected, until the eruption cycle starts to go into
18 the conduit cycle of eruption.

19 That stage may last on the order of weeks or
20 longer. And we essentially took the information from
21 DOE from a TSPASR in terms of duration of eruption.

22 And we used that as a probability density
23 function. We used the DOE's information on that one.
24 But, over a longer timeframe, the failure mechanism
25 diagram that you guys couldn't read before -- we're

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1 starting to get into what some of the other failure
2 mechanisms are.

3 There's a concern of erosion, that there is
4 a corrosive and abrasive material flowing past the
5 waste package. And the conclusion, this was from
6 Frazier King from Canada, who's been with waste
7 management business for a very long time, went through
8 an evaluation of what the effects of erosion corrosion
9 would be.

10 And, he came up with an estimate of an
11 erosion corrosion rate and put it together with the
12 duration of the eruption. He came up with a maximum
13 on the order of two millimeters.

14 So, a failure or a stripping away of the C22
15 shell is bound to be very highly unlikely. Next
16 please. The second mechanism that we consider was
17 failure by this internal overpressure, which, in some
18 of the documents, this is the failure mechanism that
19 is used to justify the neglect of waste package from
20 the TSPA analysis.

21 And, essentially, the idea is that because
22 we've got the temperature going up in the waste
23 package, that we generate the stresses from the
24 differential between the expansivity of the steel
25 versus the C22.

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1 And we also have internal overpressure from
2 the air pressure going up. And, you can think that if
3 the pressure exceeds some threshold value and
4 threshold value will be going down as a function of
5 temperature because the yield strength of the material
6 is going to go down with the temperature, if it
7 exceeds some threshold, you could get it popping open
8 like a can of soda bursting.

9 It's generally the concept of what we're
10 talking here by internal overpressure. So, material
11 strength is decreasing at an increasing temperature.
12 But, one of the things that has not been taken into
13 consideration in these evaluations, and I think we're
14 the first ones to look at this, is that, whatever
15 internal pressures and internal stresses are built up,
16 they are offset by static pressure exerted by the
17 column of magma over it.

18 We have the waste package sitting in the
19 column of magma, very dense fluid. As it rises to the
20 surface, we have hundreds of meters of this heavy
21 fluid sitting over the waste package, exerting a
22 positive pressure on the outside of the waste package.

23 And, going through some calculations, we can
24 show that, after a very short time, the stress on the
25 waste package becomes compressive, so the net stress

1 on the waste package is not from the inside out, it's
2 from the outside in.

3 And that's an important factor because the
4 waste package is stronger than it is the opposite.
5 Okay. So we consider a range of magma conditions.
6 And our conclusion was that the waste package will not
7 fail on overpressure.

8 The next thing to consider was creep
9 failure. At the high temperatures, we don't have
10 creep data for some of these high temperatures, so we
11 have to extrapolate from lower temperature.

12 So that is kind of an uncertainty in our
13 evaluation. The contact temperatures of magma on the
14 waste package are lower than has been considered by
15 DOE or NRC.

16 Based on the Nicholis and Rutherford paper,
17 if we were to do this analysis today, we would
18 actually consider even lower still. For creep rupture
19 to occur, we have to have some way of accumulating
20 strain.

21 In other words, there has to be a
22 differential stress across the waste package. And,
23 because of the geometric constraints of the waste
24 package in the drift and the C22 next to the stainless
25 steel inner shell, there is not enough space in the

1 different constraints for it to develop enough strain
2 to fail by creep failure.

3 The final -- I think we're up to the final
4 one. We have considered several failure mechanisms
5 now. Now we're up to corrosion. There is very
6 limited available data for nickel chromium alloys in
7 magma.

8 And so, what we did was we went and we
9 looked at -- we evaluated literature data on nickel
10 chromium alloys and a variety of electrolytes did some
11 probabilistic analysis of corrosion based on those
12 literature data.

13 And we were able to come up with some
14 initial estimates for what the corrosion rate would
15 be. But, not being satisfied with that, we decided to
16 do some experiments and to find out what the corrosion
17 rates that we could experimentally measure alloy 22
18 would be if it were immersed in magma.

19 Next. So, what we did was took samples of
20 alloy 22. We got samples of basalt, melted the
21 basalt, and put the alloy 22 samples into the magma.
22 And so, this is a picture of the alloy 22 sample being
23 removed from a graphite crucible.

24 This is from a one hour exposure. Once we
25 take it out, the magma solidifies very rapidly. Here

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1 are some results from a one week test, doing some
2 micrographs of what the surface effects were.

3 This was old magma used with an inert gas
4 purge for a one week test. C22 remained intact during
5 the test and showed some degree of surface voiding.

6 There was no evidence of inter-granular
7 attack, which actually was one of the primary
8 mechanisms that the metallurgists were concerned
9 about, the potential for inter-granular attack by the
10 type of contaminated materials that are in the basalt.

11 Next, please. Going on to two weeks, very
12 similar results to the one week test with a bit of an
13 increase in void/inclusion density, and still no
14 evidence of inter-granular attack.

15 Going out to a month, the surface voiding
16 was more extensive, deeper, up to 600 microns from the
17 surface and still no evidence of inter-granular or
18 other degradation attack.

19 The net result of magma contact is the C-22
20 shell was not breached by the mechanical impact. The
21 inner shell was not breached by the mechanical impact.

22 Essentially, here is a picture of C-22
23 samples, which you can think of as being C-22 waste
24 packages if you care to, embedded in the basalts after
25 it has been sitting in molten lava.

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1 And so, there is virtually no effect on the
2 -- little bit of surface attack, but not nearly enough
3 to -- the waste package during the time frame of an
4 eruption.

5 Now, what this may do, the metal may become
6 heat sensitized and corrosion rates may go up over the
7 long term. So, once the eruption has ended and water
8 starts coming back down through, the long-term
9 corrosion rates may go up.

10 And that's something that we're evaluating
11 now. We've got some initial results on that. And, we
12 will be evaluating that as part of an intrusive
13 scenario of the evaluation.

14 We will be publishing more on that next
15 year. So, our conclusion is that we have reasonable
16 expectation, given all these failure mechanisms that
17 we've gone through.

18 And they have found no way to breach the
19 waste package under any reasonable conditions that we
20 apply to the waste package. We have reasonable
21 expectations that we will get no waste packages to
22 fail during the eruption.

23 And, again, that doesn't refer to the period
24 after the eruption when we haven't had the corrosion.
25 We may have effects that we can maybe look at then.

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1 But, during the period of eruption, we were
2 not able to get radioactivity out of the waste
3 package. Thank you. So let me transition here for
4 just a minute, because up to now we've been building
5 our case solely by looking at the evaluation of the
6 eruption as it would affect the repository and as it
7 would affect the waste packages.

8 We wanted to do more than that because we
9 wanted to look at the total system as a whole. So,
10 the remainder of the presentation is going to be an
11 assumption that, even though we think that within
12 reasonable expectation no waste packages would fail,
13 we're going to assume that the waste package fails
14 anyway.

15 So now we're going to take into account
16 effects if we haven't perhaps taken into account
17 extreme enough effects in the eruption or something
18 that we haven't taken into consideration.

19 What this allows us to do is look at the
20 other aspects of the system and see how much they
21 contribute to performance and to look at alternative
22 assumptions that we can put into other aspects of
23 TSPA.

24 Next slide, please. So, what we're going to
25 look at is the ash dispersal modeling, biosphere

1 analysis, and we're going to do a series of
2 sensitivity studies.

3 And we call these conditional results rather
4 than just sensitivity studies, because all of them are
5 conditional on the assumption that a waste package
6 fails, even though we don't think it will.

7 So, we're going to assume it fails anyway,
8 go through and come up with a credible release
9 mechanism for how radionuclides could get out of a
10 partly failed waste package.

11 Given how tough this material has proven to
12 be, we are going to take some credit into account of
13 the waste package itself. And, what we really want to
14 do is demonstrate defense in depth from each part of
15 the system.

16 So, we're looking at the multiple variables
17 now. The first part that I want to talk about is the
18 ash dispersal model that we did. We went through and
19 did an evaluation with multiple models.

20 We used the ASHPLUME code that is used by
21 NRC. There's a commercial version called TEPHRA. We
22 used that one. And we also compared that against
23 three other models that are common in the ash
24 dispersal literature.

25 Ultimately for the results I present in the

1 TSPA analysis, we focused on the results from TEPHRA.
2 And, what we found was that we had a lot of
3 realizations when we went through and did probably
4 calculations using TEPHRA to look at a variety of
5 eruption magnitudes, eruption energies, column
6 heights, things like that that were all within a
7 credible range based on our range of analog eruptions
8 that we were looking at.

9 And what we found was that a very large
10 number of them had negligible deposition of tephra at
11 the -- point of the record. In fact, depending on
12 what you want to call negligible, it was between 60
13 and 80 percent of the realization produced no
14 deposition of tephra.

15 So, even though we don't think the waste
16 package fails, if it fails, probably 80 percent of the
17 time we get negligible accumulation and negligible
18 dose down wind.

19 The other thing to point out about this, is,
20 this may be a little bit hard to see just what it is.
21 But these are our tephra contours here, tephra iso-
22 depth contours coming from the TEPHRA model, from
23 ASHPLUME model.

24 This is compared to a different model, a
25 PUFF model. And what you'll see is the receptors

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1 would be right down here. This is the point at which
2 the RMEI occurs.

3 And so you'll see, even though 80 percent of
4 the realizations gave negligible accumulation, it's
5 still a very conservative model. The PUFF model
6 actually has most of the deposition going someplace
7 else.

8 And so, it would give a lower number of
9 realizations that would -- that you would calculate
10 using the ASHPLUME model.

11 MR. HINZE: When do you get a different
12 distribution?

13 MR. KOZAK: The PUFF model takes into
14 account variability. The ASHPLUME model assumes that
15 the wind blows toward the receptor through the
16 duration of the eruption.

17 There's a number of effects. One is the
18 wind distribution, the other one is that the PUFF
19 model takes into account the thermal circulation that
20 occurs around the eruption during the period.

21 So, that's why you have a lot -- it's kind
22 of hard to see here. But, the eruption is actually
23 right down there. So you actually have a lot of
24 deposition upwind.

25 In fact, it would be upwind from the volcano

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1 because of these thermal cycles -- convective cycles
2 into the air. And so it is moving it in every
3 direction.

4 It doesn't just move nicely down the wind
5 grade. The final point to make about this is that the
6 particle sizes, when they are deposited, aren't
7 respirable.

8 They are much bigger than respirable size.
9 In fact, the respirable size end up in Canada
10 someplace for some of these calculations.

11 (Laughter.)

12 And so, we need to take that into account
13 when we do this. We also have to do biosphere
14 modeling to convert this deposition of ash into dose.

15 And, one of the things that we wanted to
16 look at, I don't know if you guys have looked in depth
17 into the way biosphere modeling is done for this
18 scenario.

19 But, essentially, probabilistically, you
20 have to sum all of the previous volcanoes that may
21 have occurred, assuming that the ash stays around for
22 a very long period of time and you add up all the
23 previous potential volcanoes at the time that you want
24 to calculate the dose.

25 So, it's not just the dose that occurs

1 during that year. It's the dose that occurred ten
2 years ago, that the ash is still sitting around that
3 you combine all these probabilistic.

4 It's a very complicated way of doing it.
5 And, so one of the things that we wanted to look at --
6 when you look at past analogs, meaning how people
7 behave after an eruption, they clean up.

8 Here's an example of Mount Pinatubo. People
9 don't just sit there under several centimeters of ash
10 because it is nasty and awful, and it clogs up
11 machinery.

12 This is a significantly different way of
13 doing the dose calculation of the biosphere. We can
14 argue whether or not that's appropriate. In fact,
15 there are some interesting philosophical arguments we
16 can have over a beer about whether or not we should
17 take into account cleanup because it's not
18 intervention.

19 You're doing a radiological assessment for
20 this practice. But a point of fact, the cleanup is
21 going on, not because of the radiological content of
22 the ash, but because of the ash itself.

23 Even if people are unaware of the
24 radiological content of the ash, they will clean it
25 up. They will clean it up to differing extents on

1 agricultural land as around their home and so forth,
2 but they will clean up.

3 CHAIRMAN RYAN: I just wanted to point out
4 the fact that there's actually radiological content in
5 the picture you showed us.

6 MR. KOZAK: That's right. That's a whole
7 different -- yes, you're absolutely right. So, the
8 implications of this in terms of a biosphere model is
9 that our dose is predominantly in the first year and,
10 to a lesser extent, comes from later years, depending
11 on what we assume about how quickly it is distributed
12 in the system.

13 As a result, we don't have to add up doses
14 over many years. I'm not sure quite how many years
15 the DOE and NRC assume that the radioactivity can
16 persist in the biosphere.

17 But, we assume that it goes over the -- so,
18 we have different important pathways in the first year
19 and in later years. And we wanted to look at this.

20 So, as one of our sensitivity studies, we
21 are going to look at this issue of consistence of
22 biosphere. But it's not part of our sort of base
23 case.

24 The ash particulates, as I said before, are
25 not respirable. And, indeed, in the first year dose,

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1 you could argue that a lot of people are going to be
2 acting like this guy right here.

3 They're not going to be going around and
4 just breathing the dust in or the ash in, because it
5 is going to be unpleasant, not because they are
6 worried about radiological content.

7 But there will be some degree of removal
8 using masks and things like that to reduce the amount
9 that they will inhale. Just normal behavior of
10 people.

11 I don't think it is reasonable to
12 incorporate human behavior that's not normal after an
13 eruption. So what we did was evaluate the biosphere
14 dose inversion factors for different particle size
15 ranges and differing deposition.

16 Of course, we want to look at that as --
17 sensitivity right there. So, our conditional
18 analyses, again, are conditional on a release from the
19 waste packages that we don't think will occur.

20 So, what we did was we assumed that there
21 was some failure. And, in this case, we assumed that
22 the failure did occur along the weldment and split the
23 waste package open to some degree.

24 And then we have diffusive releases from the
25 waste package into the magma as it goes by. There is

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1 a whole series of things that have to go on for the
2 magma to contact the internals of the waste package.

3 There's sort of a whole series of events in
4 which magma needs to flow into it, contact the fuel,
5 dissolve the fuel, diffuse back out through the waste
6 package, and get to the outside.

7 By and large we're not looking at that
8 because we just didn't have the time. It was
9 complicated. So, what we did was look at a diffusion
10 layer across an opening in the external shell of the
11 waste package.

12 So, our base case assumptions are that we
13 have this type of release from the waste package.
14 Something that I haven't talked about that we discuss
15 in the report is that the thermal field around the
16 repository changes the stress field and has the
17 potential as the dike is rising up, that it could
18 divert.

19 The peer review panel discusses this concept
20 that has a potential to divert the dike because of the
21 changes in the normal stress field. And so, as part of
22 our base case, we assumed, for lack of a better
23 assumption, that the higher the temperature was, that
24 the more likely the dike would be diverted.

25 So, if it were diverted, that lowers the

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1 probability that the event occurs at all. So, at
2 early times, when the temperature is highest, you have
3 a lower probability of the event occurring at later
4 times.

5 At later times it ramps up until the year
6 2000, in which the thermal field is essentially gone.
7 At that point it reverts back to the normal pre-waste
8 placement probability of the event happening.

9 And we're going to look at that sensitivity
10 study too. So, if you don't like that assumption,
11 we'll work around that. Another one that we're going
12 to assume is that the waste package doesn't limit
13 releases from the fuel inside.

14 So we have sort of a dissolution mechanism
15 inside and a diffusion out of it. We have no cleanup
16 of ash occurring at compliance points. We have ash
17 fall respirable particles, even though we know that
18 that's not going to happen.

19 We have to take into account the fact that
20 the ash is breaking down in size or if there is some
21 other mechanism. These are things that we have to
22 explore as part of it.

23 So, we're carrying these out in spite of
24 there being an initial assessment that we have a zero
25 release during the attack. Okay. So this is the

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1 first conditional case.

2 We have between one and nine waste package
3 failures. This is based on geometrical considerations
4 of how big the vents are. A single vent could get us
5 up to three waste package failures.

6 If we have these multiple events that the
7 multiple events in a single event, that we could -- up
8 to three vents would give us up to nine waste package
9 failures.

10 In this one we have the temperature
11 dependent dike diversion. And what we find, we have
12 the 95th percentile and a mean right here. We get
13 about nine orders of magnitude dose lower than TSPA-
14 SR.

15 Next please. One of the sensitivity cases
16 that we wanted to look at is persistence in the
17 environment. We didn't set up our analysis to do this
18 in the elaborate way that DOE does it.

19 So what we did was in an approximate manner.
20 We found, after going through the calculation, that
21 the doses in year two to about ten were about
22 constant.

23 So we just took that constant and added that
24 for as many years as we wanted to. So, our worst case
25 analysis for this particular assumption, the

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1 persistence in the environment, assuming that it stays
2 in the environment forever, and all it does is change
3 by decay, it would be this top curve up here.

4 So, it increases the doses over our sort of
5 nominal conditional case. We don't want to call it
6 nominal. But, our first conditional case that was
7 sort our base of assumptions.

8 It increased it by about two to three orders
9 of magnitude. But it's still very low doses.

10 CHAIRMAN RYAN: Let me just ask a very quick
11 question while we are --

12 MR. KOZAK: Yes.

13 CHAIRMAN RYAN: That no depletion case is
14 the case where you're assuming the radioactivity
15 that's deposited is available at that same deposition
16 forever?

17 MR. KOZAK: That is correct. That no
18 depletion in the ground, it doesn't blow away, it
19 doesn't redistribute, it just sits there.

20 CHAIRMAN RYAN: I'm going to ask the
21 audience a question. I want to peak around the corner
22 and see where that red graph flattens out. Can you
23 tell me when that happens?

24 MR. KOZAK: That's a million years?

25 CHAIRMAN RYAN: Oh, it's a million? I

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1 thought it was ten thousand. I see, sorry.

2 MR. KOZAK: Based on the way we did the
3 calculation, it will continue going up because what
4 you're doing is adding in the consequences of all the
5 previous eruptions.

6 So, essentially this is the number of years
7 that have occurred prior to it.

8 CHAIRMAN RYAN: There are more and more
9 eruptions as time increases?

10 MR. KOZAK: Yes.

11 CHAIRMAN RYAN: I got it. All right.

12 MR. KOZAK: Exactly.

13 CHAIRMAN RYAN: Thank you.

14 MR. KOZAK: The effect of waste packages,
15 now we took out the waste package. And so we assume
16 that, when the dike hits it, it will disappear. And
17 we get on the order of five orders of magnitude
18 increases over a base case.

19 We're still, because of just doing this one
20 parameter at a time, one sensitivity variable at a
21 time, we're still very low, six orders of magnitude
22 below the dose standard.

23 We looked at a whole bunch of other events,
24 some of them positive, some of the negative. The
25 effect of dike diversion, if we -- if the temperature

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1 effect is really very strong in changing the direction
2 of the dike, we could get full dike diversion in 2,000
3 years, which means we couldn't get an event for the
4 first 2,000 years.

5 On the other hand, if this effect doesn't
6 actually occur, then we actually just increased the
7 dose of early times by less than an order of
8 magnitude.

9 Going and looking at respirable particle
10 sizes versus non-respirable particles sizes, it was
11 less than an order of magnitude increase in dose.
12 Now, the reason for that was we were looking at -- the
13 dose factors that we were looking at included the
14 nasal pharyngeal contribution to inhalation.

15 The increase in lung dose for the smaller
16 particle size wasn't that much. That actually
17 surprises quite a bit that it wasn't that big of an
18 effect.

19 The other thing that we looked at that would
20 be a positive effect, that we didn't include, was the
21 conditional probability that the vent occurs in
22 between the drifts.

23 We assume that, for our basic conditional
24 analysis, the dike intersects it and we start to form
25 these vents, that they hit a waste package,

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1 essentially.

2 This one we took into account the space in
3 between the drifts, which is quite a bit of space, and
4 drops it by about a factor of six. This one is where
5 we took all of the conservative assumptions, all these
6 different orders of magnitude, put them all together
7 in one analysis.

8 The interesting thing about this is that,
9 when we look at this early time, that's up where the
10 TSPASR dose level is. So, we can recover the orders
11 of management of dose that DOE dose by adding in all
12 these conservatisms.

13 So, the point of this is, by doing what we
14 consider to be a more reasonable analysis, we can show
15 the level of conservatism that's associated with each
16 of those assumptions.

17 When we add them all back in, we still go up
18 to where the DOE dose analysis brings us, and we still
19 comply. So we can demonstrate the relevant fact that
20 each part of the analysis can provide conservatism.

21 CHAIRMAN RYAN: Is the next step of taking
22 this to a probabilistic approach, looking at each one
23 of these, considering each of these variables?

24 MR. KOZAK: I'm sorry, I wasn't clear
25 enough. This is a probabilistic calculation.

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1 CHAIRMAN RYAN: This is a probabilistic
2 calculation.

3 MR. KOZAK: This is run using at risk --

4 CHAIRMAN RYAN: Okay.

5 MR. KOZAK: -- with a thousand realizations
6 -- sample. And each of the distributions is laid out
7 in the report.

8 CHAIRMAN RYAN: Okay.

9 MR. KOZAK: They are subjective probability
10 distributions, so --

11 CHAIRMAN RYAN: I understand. That's fine.

12 MR. KOZAK: So, to summarize, our reasonable
13 expectation approach has led us to a conclusion that
14 we would get zero release during the eruption.

15 We looked at multiple different kinds of
16 failure mechanism in the waste package. We couldn't
17 really come up with a credible waste package failure
18 mechanism for any of the circumstances that we looked
19 at.

20 The key lines of evidence that lead us to
21 that, the conditions of the drift level are not as
22 extreme as has been assumed previous by the NRC and
23 DOE in our judgment, based on our evaluation of the
24 available data, based on our model.

25 The magma entering the drifts is much less

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1 violent than has been assumed previously. The
2 conclusions by Woods et al -- we have shown that the
3 shockwaves and these extreme things that go on are not
4 going to happen when we take into account two
5 dimensional flow of the magma.

6 The waste package provides a very
7 significant barrier to release. I think it is
8 extremely conservative. Our analyses show an order of
9 six orders of magnitude conservative to ignore it.

10 The magma entering the drifts is going to
11 cool and solidify pretty quickly to isolate dike and
12 event. And so, to conclude, we think that the
13 analysis that shows up in the TSPASR is extremely
14 conservative on the order of nine orders of magnitude.

15 But the thing to keep in mind is that,
16 despite all the conservatism, it still complies. So,
17 any potential changes that we could do -- you know,
18 when we start talking about pushing it in the
19 direction of being more conservative.

20 There's always the tendency when we do TSPA.
21 Everybody is always trying to think of a more
22 conservative analysis. There's a lot of these things
23 that will drive it very strongly to being less
24 conservative.

25 And so, we were able to demonstrate the

1 amount of conservatism introduced by different parts
2 of the analysis. And, of those, the waste package is
3 by far the most important.

4 CHAIRMAN RYAN: Thanks very much.

5 MR. KOZAK: Thank you.

6 **SESSION TWO WORKING GROUP ROUNDTABLE DISCUSSION**

7 CHAIRMAN RYAN: I guess I'm intrigued by
8 your modeling of the -- the detonation modeling
9 capability that you mentioned. Could you expand on
10 that?

11 MR. KOZAK: I'm going to completely differ
12 that to Megan Morrissey. That's her specialty.

13 CHAIRMAN RYAN: Could you come up here and
14 tell us who you are.

15 PARTICIPANT: Can't hear your question.

16 CHAIRMAN RYAN: I'm sorry. I asked the
17 question, if there could be a little bit of expansion
18 on the underground modeling and the use of it for this
19 magma modeling.

20 I just would like to know little bit more
21 about the model itself, how it is used, what it is
22 used for, and so forth.

23 MS. MORRISSEY: My name is Megan Morrissey.
24 I am in the Colorado School of Mines. The model was
25 developed at Los Alamos through the thermal nuclear

1 group.

2 They allow me to use it to do volcanic
3 simulations of flow-through cracks and whatever. So
4 I tied it to -- well, we first of all wanted to know
5 to interpret the Wood and other's pressure time group.

6 So, I was looking at that and said, okay,
7 let's really put it in a two dimensional, vertical and
8 horizontal. What it does is it is a compressible
9 fluid flow, multi-gas, multi-phase.

10 The walls -- there is some expansion to it.
11 But we used a rigid case in what we showed today. You
12 can set it up with any geometric configuration you
13 like.

14 So what we did was formed a dike similar
15 geometry with analysis configuration. And we used
16 steam, increased the density a little bit to account
17 for ash.

18 And we allowed it to -- you know, just let
19 it go into an empty drift. And what you saw was the
20 actual -- actually what you saw was a shockwave did
21 develop, not the shockwave that they believe had
22 occurred in their simulations, which is a whole other
23 story?

24 CHAIRMAN RYAN: They?

25 MS. MORRISSEY: The Woods et al. What

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1 happens is you set up a shockwave, an expansion fan
2 type that is large enough that it expands onto the top
3 of the drift.

4 At the same time it reflects off and you see
5 these oblique shockwaves moving down the drift. But
6 those shockwaves are within the steam, the magnetic
7 gas moving down.

8 You can see the front. And so that's just
9 moving down slowly. It's going to be at the end of
10 the drift soon and start filling up. But what we
11 wanted to try and demonstrate was the fact that that
12 expansion fan with the shockwave on top creates a
13 pressure concentration right at the top of the drift,
14 right above the intersection of the dike into the
15 drift.

16 So that was essentially what we were showing
17 there. And we can use it for a full range of
18 pressures, different temperatures, different starting
19 conditions, and keep expanding on the complexity of
20 the problem.

21 So that's what we've done. But we just
22 showed you a fairly simple scenario to show what the
23 Woods et al model would look like in a true vertical
24 horizontal situation.

25 MR. MARSH: Was the dike open to begin with?

1 MS. MORRISSEY: It was open to begin with,
2 yes.

3 MR. MARSH: So you didn't open it with the
4 fluid flow.

5 MS. MORRISSEY: No, it was a little nozzle.
6 It started out at rest, and let the pressure --

7 MR. MARSH: The fluid was there behind the
8 nozzle?

9 MS. MORRISSEY: Yes. The fluid was there
10 behind the nozzle. It was a very, very narrow nozzle,
11 and let it open, just let it go. And that's -- the
12 thing is, we're not trying to do anything realistic.

13 We're just reproducing based on the Woods
14 model.

15 MR. MARSH: If you let the dike get out, it
16 would open gradually.

17 MS. MORRISSEY: Yes, something like that.
18 And that's what the DOE model were going in that
19 direction, showing, okay, here's the opening, it's
20 going to go straight up.

21 And one model I didn't show, the little --
22 just a little pinhole above it, and a lot of the fluid
23 just goes straight up. And you do get diversion down
24 the drift.

25 But pressures are not lowered if you want to

1 use the same 10 to 20 megapascal reservoir pressure.

2 MR. MARSH: If you add solidification it
3 even --

4 MS. MORRISSEY: Yes, this was a gas. So
5 we're going to the extreme of a compressible, high
6 discharge. But, if you considered a de-gassed lava or
7 magma, it is very viscous at the 980 degree C
8 temperature.

9 It's going to start moving down. It's going
10 to have viscosity around 100 pascal seconds or higher.
11 And, as it decreases in temperature with
12 crystallization, that viscosity is going to keep going
13 up and up, really prohibit --

14 MR. MARSH: Maybe choke the drift off.

15 MS. MORRISSEY: I don't like to use the work
16 choked. But, it will slow up the flow. It will --
17 plug and let the rest go up, yes. Exactly. That's one
18 scenario.

19 CHAIRMAN RYAN: John, you had a question.

20 MR. GARRICK: Not necessarily for me. I
21 realize that your reference case here was the Woods
22 model. But, looking at it from a total system point
23 of view, of course, the consequence is very dependent
24 upon -- is it not -- the time at which the magma event
25 occurs.

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1 For example, if your latent time where the
2 waste package started their degradation process are
3 degrading substantially, you're certainly going to get
4 a different result than if the waste package is still
5 at their full integrity.

6 And the other thing -- and you can comment
7 on that -- the other thing that's true here is that
8 now you have to have a disruptive event. And, even
9 though there is zero release, you have disrupted the
10 total system performance in the sense that, number
11 one, you now have deformed and damaged waste packages.

12 And so, you've accelerated the mobilization
13 of the radionuclide process. And you may have
14 introduced and set the stage for other events, such as
15 downstream criticality or what have you. Would you
16 care to comment on those types of things?

17 MR. KOZAK: Yes, your first comment is very
18 well taken. We have not taken into account the long-
19 term degradation of the waste packages prior to the
20 event in this analysis.

21 If we were to do so, the worst case would be
22 the last one that I showed, where I assumed that the
23 waste package didn't contribute to the release. So, at
24 worst, it's going to increase it by some orders of
25 magnitude.

1 But, yes, you're right. Certainly after, on
2 the order of -- depending on who's model you believe
3 on the degradation, but, when you get out to the
4 100,000 year range, when the degradation is advanced,
5 it certainly won't have the structural strength that
6 we've assumed in this analysis.

7 That's completely correct. On the second
8 point, you're also absolutely correct. That does
9 affect the overall total system performance. And we
10 are going forward with that.

11 We have done some initial calculations on
12 the increase in degradation rate, increase of
13 corrosion rate that would be caused by the
14 sensitization of the material by the magma.

15 And we have incorporated that into some new
16 calculations. Right now all we have is sort of a very
17 conservative one where we assume that the eruption
18 occurs essentially at time zero.

19 And so then, that enhanced corrosion rate
20 applies to the rest of the duration of the facility.
21 To do it in proper ways, it would be very complicated
22 because we have to assume that there's a certain
23 degree of corrosion up to -- if we're doing it for
24 30,000, if it corroded up to 30,000.

25 And then we have to hit it with magma and do

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1 a different corrosion rate thereafter. So, it gets to
2 be a complicated analysis to do that. But, we're
3 looking into ways of doing that for next year.

4 And so, yes, you will have more rapid
5 failure of waste packages, and you will have increased
6 corrosion in that. The thing to remember is that
7 that's going to be increment over the nominal case, as
8 opposed to the relatively high doses that you get from
9 an extrusive case.

10 It's an increase in the sort of nominal dose
11 over the dose over the nominal case. But, the nominal
12 case is a probability of one. This has a probability
13 very low.

14 And so, the net effect of that is probably
15 not going to be very large. The effects, you
16 mentioned criticality, and I have not thought about
17 that.

18 I don't know where that can come from. It
19 wouldn't be any -- it wouldn't really be that much
20 different to having this kind of event in terms of
21 flow and transport processes compared to the nominal.

22 MR. GARRICK: Well, I'm only thinking that
23 any time you change the geometry of the fuel
24 assemblies. I'm not thinking during the time of the
25 event itself.

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1 I'm thinking that you just changed the
2 vulnerability of the waste package to a -- issue, such
3 as criticality. That's all I'm saying.

4 MR. KESSLER: I'd like to address that, John
5 Kessler of EPRI. We're not going to look at
6 criticality in this particular case. We think, as
7 Matt has already explained, that the amount of
8 deformation due to this magma blast of this single
9 container is probably, you know, the maximum amount of
10 deformation.

11 We already know that there are analyses out
12 there done by DOE that show criticality and how that
13 changes if you have collapse of the container
14 internals.

15 Our understanding is that the maximum
16 criticality is probably about in the -- as original
17 case, because you've got a near operable moderation.
18 So, we're feeling that it's not an issue. But, we are
19 not planning to look at it.

20 MR. GARRICK: Right, I'm not concerned about
21 criticality. I don't think it's a major issue either.
22 But, from the point of view of other people, you have
23 changed the conditions of the fuel.

24 MR. KESSLER: We certainly get it on the
25 transportation side. And that's certainly the hottest

1 issue right now for transportation, fuel
2 reconfiguration.

3 MEMBER WEINER: I have a couple of
4 questions. When you model with ASHPLUME AND TUFF, how
5 does your plume go before it starts to move down?

6 MR. KOZAK: Mike, would you want to answer
7 that? That was worked on by --

8 MR. SHERIDAN: This is Mike Sheridan from
9 the University of Buffalo. This work was done by a
10 colleague, Marcus Bursik, who is an expert in volcanic
11 ash plumes.

12 He did a lot of iterations, I think 30,000
13 or something.

14 MR. KOZAK: There was a range that was
15 considered based on the power of the eruption. And
16 the power of the eruption came back to Mike's work on
17 which eruptions were appropriate analog behavior.

18 So, there's a whole thread of logic that's
19 gone into answer that -- it's hard to answer.

20 MR. SHERIDAN: Regarding the height of the
21 plume, we can say all of the ranges of plume heights
22 in that diagram that Britt showed of say, VI4, three
23 and two, were replicated in the simulations.

24 MEMBER WEINER: Thanks. I asked because, as
25 it happened, I was in Washington State during and

1 after the Mount Saint Helens eruption. And I happened
2 to go hiking on Snoqualmie pass a year after the
3 eruption.

4 And you could still see deposited fine
5 particulate on the vegetation there. And that's a
6 good long distance. I just wondered --

7 MR. KOZAK: Well, yes, to address that,
8 Mount Saint Helens is not a good analog at all for
9 these eruptions. These are much smaller, much more
10 quiescent, and nowhere near Mount Saint Helens type of
11 behavior.

12 MEMBER WEINER: That was exactly the point
13 of the question. The other question I have is what
14 kind of particle size distribution or settling
15 velocity distribution did you have?

16 MR. KOZAK: That's in Marcus' report again.

17 MR. SHERIDAN: This has been a big concern
18 of ours concerning particle size distributions because
19 the question is, the total particle size produced by
20 volcanic eruption.

21 And that sort of data is difficult to
22 determine because, generally, we find the products
23 only at one location, which had been size sorted by
24 falling through to the atmosphere.

25 So, the compilations of total grain size

1 distributions are extremely hard to come by. But
2 there have been some for these strombolian types of
3 eruptions.

4 And this is the particle size of the
5 volcanic particles. But we're also concerned about
6 the radioactive particles from the canisters. And
7 this is something that I don't think anybody knows the
8 answer to.

9 And it's a great puzzle to me of how the
10 material came from the canisters into the plume
11 and be transported in this very fine size, because,
12 within the canisters, it's in size of centimeter
13 scale. We're talking about micron size. So, --

14 MR. KOZAK: Mike, let me --

15 MR. SHERIDAN: Okay.

16 MR. KOZAK: Let me comment on that because,
17 one of things that you'll see is we are doing this on
18 a very tight time scale. And we weren't sure which of
19 these mechanisms would become important.

20 So, when you look at the report, you will
21 find that we have models for mechanisms that
22 ultimately don't show up very much in the TSPA. And
23 one of them is fracturing along ring boundaries in the
24 fuel and being transported out of these particulates.

25 I didn't present it here. But we did ash

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1 plume calculations where the ash was represented by
2 UO_2 particles and that UO_2 density. And so, it is
3 fallout of particulates.

4 The ones that I showed were for dissolution
5 of contamination into the ash so it's being
6 transported as dissolved in ash or dissolved in magma,
7 which then evolves as ash.

8 CHAIRMAN RYAN: Matt that's really important
9 assumption. The action is where is the radioactive
10 material.

11 MR. KOZAK: Yes, that's right.

12 CHAIRMAN RYAN: So you just offer the
13 radioactivity, the radionuclide content of fuel into
14 the ash?

15 MR. KOZAK: For these calculations, yes.

16 CHAIRMAN RYAN: The volcanologic of it,
17 nothing left behind in the chunks of fuel. All the
18 radioactivity is dissolved in the ash.

19 MR. KOZAK: Not all of it gets out during
20 the eruption.

21 CHAIRMAN RYAN: That's what I want to know.
22 What fraction get to the ash?

23 MR. KOZAK: The mean case, I mean, it's a
24 distribution, because it's all probabilistic. But,
25 the mean was about ten to the minus fifth, I think, of

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1 the inventory of a waste package would get out.

2 CHAIRMAN RYAN: Hopefully we'll address that
3 kind of question a little more in detail when we hear
4 about the perspectives on issues on aerosol and
5 modeling, etcetera.

6 MR. KOZAK: But, keep in mind, that was
7 based on our model of release from a waste package.

8 CHAIRMAN RYAN: Right.

9 MR. KOZAK: Which actually itself was
10 probably pretty conservative.

11 CHAIRMAN RYAN: Right.

12 MR. KOZAK: Because, it only accounted for
13 diffusion across the boundary of that opening. I
14 mean, something that's diffusing along the line of all
15 the reactor -- the waste package internals, that's a
16 very long diffusion path.

17 It has to come through all this magma as it
18 is solidifying, and all these other considerations.
19 We just didn't have time to develop something --

20 CHAIRMAN RYAN: I don't agree or disagree,
21 but I just want to kind of establish in everybody's
22 minds the realism is not so much where the particles
23 or how to they get created, or where they go.

24 It is where is the radioactive material.
25 And, it may or may not be distributed uniformly, non-

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1 uniformly. And we need to understand how people
2 assume that.

3 I guess the models tend to put more of it
4 into the respirable particle size range for the longer
5 haul -- perhaps be a little bit aggressive in putting
6 the radioactivity in the air where it's, I think, less
7 might be in the air.

8 MR. KOZAK: Yes. See, our original thought
9 when we start putting the model together is we
10 expected that there would be at least some
11 circumstances when we were going to have blow chunks
12 of UO₂ out like a whole waste package, you know,
13 granularized and blow it out.

14 So, we did a lot of work on that and got a
15 really nice model in one of the appendixes that we
16 never ended up using. But we developed it, so we put
17 it in there, of how they react as they're going up the
18 vent.

19 You know, all these kinds of things are in
20 there. There reason they're in there is because we
21 had to develop a parallel. And then, once the
22 evidence started coming through on what was important,
23 we found out some that we didn't need.

24 CHAIRMAN RYAN: If you're rating the dose
25 perspective is unimportant -- the plutonium an

1 americium.

2 MR. KOZAK: Yes, but it would all be -- I
3 mean, the majority of it is going to behave like
4 uranium particulates, not from a dose perspective, but
5 from the particulate perspective.

6 It is all associated with fuel. Just to let
7 you know, the report number, if you're interested, you
8 can either get it from Monitor or from EPRI.

9 It's an EPRI report number 1008169. And it
10 was published June 2004.

11 MEMBER WEINER:

12 MR. BRITT HILL: I just have two more. When
13 you calculated the inhalation doses did you include
14 resuspension?

15 MR. KOZAK: Yes.

16 MEMBER WEINER: What resuspension model did
17 you use?

18 MR. KOZAK: I believe it was -- let me
19 think. I'm pretty sure it was simply a resuspension
20 factor.

21 MEMBER WEINER: Yes, but, did you use the
22 same resuspension factor that everybody has used
23 forever, which is basically the one --

24 MR. KOZAK: I'd have to go back and check
25 the specific file.

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1 MEMBER WEINER: Okay, thanks.

2 MR. KOZAK: I'm not sure. I think it was
3 probably a conservative value.

4 MEMBER WEINER: Yes, I didn't mean to put
5 you on the spot.

6 MR. KOZAK: So, even when we try to do
7 reasonable expectation, we find ourselves slipping
8 back into doing conservatism. It's just we fall into
9 it all the time.

10 MEMBER WEINER: Okay, my very last question
11 is, what has been the NRC and/or DOE response to this
12 contention that you just mentioned?

13 MR. KOZAK: This is the first time we have
14 presented it in this form. There has been no official
15 response.

16 CHAIRMAN RYAN: You are the response.

17 MEMBER WEINER: I am the response.

18 CHAIRMAN RYAN: Al?

19 MEMBER WEINER: Unless somebody from NRC
20 wants to comment.

21 MEMBER CROFF: Well, I would --

22 MEMBER WEINER: Tim is.

23 MR. McCARTIN: This is Tim McCartin, NRC. It
24 would be premature for the NRC to comment on this
25 information.

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1 PARTICIPANT: My name is Sharon. Regarding
2 the internal overpressure temperature, is your
3 contention that, because of the static pressure due to
4 the magnetic forces on the waste package, that that
5 would offset the internal pressure?

6 MR. KOZAK: Yes, in part the analysis is
7 based on that, that is correct.

8 PARTICIPANT: And I would agree with that
9 for fully embedded case. Have you looked at scenarios
10 where you have a partially embedded waste package or
11 a waste package that's not impacted at all magma, but
12 is exposed to the high temperatures?

13 MR. KOZAK: We have not looked at that as
14 yet. We expect that to be more important for the next
15 part of the analysis, which is the influence on the
16 nominal scenario, essentially the intrusive scenario
17 where you're looking at ground water releases.

18 PARTICIPANT: Okay.

19 MR. KOZAK: Because, if it happens away from
20 the vent, it's not going to contribute to what gets
21 back up in.

22 PARTICIPANT: Okay, so for intrusive
23 scenarios?

24 MR. KOZAK: Yes. We haven't looked at that
25 yet.

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1 PARTICIPANT: Okay.

2 MR. KOZAK: That's going to be a complicated
3 analysis. And we haven't looked at it yet on how
4 important it will be.

5 PARTICIPANT: Okay.

6 MR. KOZAK: But it probably will play a
7 role.

8 MR. MARSH: Along those lines, what's the
9 thermal state of the canisters at the time before the
10 magma hits them?

11 MR. KOZAK: It is assumed to be the thermal
12 state of the repository at that time, which is it is
13 elevated in temperature, but it's not --

14 MR. MARSH: They're not hot?

15 MR. KOZAK: They're not -- it's the hot
16 repository design. So, their temperature is from the
17 usual time history of the repository temperature that
18 you see.

19 MR. MARSH: For example?

20 MR. KOZAK: Well, 176 Fahrenheit sticks in
21 my head.

22 MR. MARSH: So you might have to worry about
23 magma quenching against the canisters.

24 MR. KOZAK: Yes.

25 MR. MARSH: Magma can quench against the

1 canisters.

2 MR. KOZAK: Yes.

3 MR. MARSH: Then that actually makes a
4 bigger thermal --

5 MR. KOZAK: And we are considering that.
6 The question would be how fractured that basalt would
7 be at the end of it. And we're looking at that.

8 We haven't really come to any -- but the
9 initial calculations that we did, because we didn't
10 know at that point, we assumed it was fractured.

11 MR. MARSH: Sure. But, in terms of
12 corrosion, it changes your corrosion because it
13 actually helps the whole corrosion probably because it
14 --

15 MR. KOZAK: It will keep the water away.

16 MR. MARSH: Well, you don't have molten
17 magma right next to the thing at months at a time.

18 MR. KOZAK: That's right.

19 CHAIRMAN RYAN: Allen?

20 MEMBER CROFF: A follow-up to John Garrick's
21 initial question. As the waste package degrades, the
22 radionuclide inventory is also going away.

23 So, to compensate in effect, I don't know
24 how it works out in longer times. You don't have too
25 many acronyms left. And then the question, I'm not

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1 even going to try anything in a detailed technical
2 level. Has this report been peer reviewed?

3 MR. KOZAK: Within the team, yes. We went
4 over it very heavily. Outside of the EPRI team, no.
5 In answer to your first question, those curves that I
6 showed account for all the decay and in growth for the
7 entire inventory.

8 We worked based on the SR inventory, because
9 that's -- there's a screened SR inventory, which is
10 the best information that we have. It's the best
11 information that's publicly available.

12 CHAIRMAN RYAN: I guess at this point let's
13 open it up to any and all panel members and the
14 audience, or staff members, anybody.

15 **PUBLIC COMMENTS**

16 MR. McCARTIN: This is Tim McCartin, NRC.
17 Your curves for the persistence in the environment, I
18 guess I'm somewhat puzzled by the no-depletion results
19 in getting a peak at 10,000 years, because radioactive
20 decay very significant for the main contributors to
21 the inhalation dose.

22 And I'm just not sure. When you have no-
23 depletion, what exactly are you doing? Or the mass
24 loading or the resuspension factor. Is there
25 something else going on there? Or is it just

1 persistence of the deposit?

2 CHAIRMAN RYAN: I guess -- can have that
3 back. Again, I'd offer it would be helpful to take
4 that out several more decades, the shapes of those
5 curves might, or at least a couple.

6 MR. KOZAK: Yes, which curve are you looking
7 at Tim?

8 MR. McCARTIN: Well, the title is
9 persistence in the environment.

10 MR. KOZAK: Yes.

11 MR. McCARTIN: And I'm assuming the largest
12 values are for the no-depletion.

13 MR. KOZAK: Yes.

14 MR. McCARTIN: And, it would appear that the
15 peak is at 10,000 years, at least the way I read it.

16 MR. KOZAK: No, that continues to go up.

17 CHAIRMAN RYAN: That's what I asked about.
18 He said it was going up.

19 MR. KOZAK: Yes, because you have to sum all
20 the previous years. Each year that you add adds an
21 incremental dose to it.

22 MR. McCARTIN: Right, but radioactive decay
23 continues to go on. And, generally those disappear
24 for say the first couple thousand years. And what
25 exactly are the assumptions when you say no-

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1 depletions?

2 Are you burying everything else in your
3 analysis except the deposit persist forever?

4 MR. KOZAK: Oh, I see.

5 MR. MCCARTIN: Is that the only thing you're
6 doing when you say no depletion? I guess that's my
7 question.

8 MR. KOZAK: Yes, this is an approximately
9 way of doing this. There's no question about that. We
10 haven't set up our analysis to really take into
11 account the release at this year of decays in the
12 environment out to there.

13 What we found was, if you look at this right
14 here, after a certain point, it's approximately
15 constant for the latter year doses. And so, all we
16 did was take that as a constant.

17 And we'll multiply that by the number of
18 years prior. It's just to give an indication of what
19 the doses are. And that's probably part of the reason
20 -- I don't remember why I cut it off at 10,000.

21 That's probably part of the reason why,
22 because you get more and more of the limiting
23 assumption to do that. This is definitely an
24 approximate way of doing it just to give an indication
25 of the orders of magnitude.

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1 MR. McCARTIN: Okay. I think you'd get a
2 different result if you actually try to simulate --

3 MR. KOZAK: Yes, absolutely.

4 MR. McCARTIN: -- events in multiple years.

5 MR. KOZAK: Absolutely. And that's a
6 complicated analysis to do. And we didn't have the
7 time to do it.

8 MR. McCARTIN: Okay.

9 CHAIRMAN RYAN: Other questions? Yes?

10 MR. HINZE: Just a quick question. I was
11 taken with Britt's suggestion that differential
12 thermal expansion would lead to stresses and other
13 stresses that might cause the canisters to lose their
14 integrity.

15 And yet, I heard from you that these
16 differential thermal expansions were inconsequential.
17 Is that an overstatement? Or do we have -- where are
18 we?

19 MR. KOZAK: We did an evaluation of those.
20 I wasn't the one to do it. And I know the discussion
21 of those stresses is in the report. And I don't want
22 to speculate on what the answer to that is. I don't
23 remember.

24 MR. KESSLER: John Kessler, EPRI. My
25 recollection, again, I'll have to remember what it is

1 that Frazier did exactly, but, the arguments that he
2 included were how much creep he felt the alloy 22
3 could manage at temperature given its yield strength
4 at temperature.

5 I don't remember exactly whether there's a
6 specific analysis about thermal expansion and how much
7 expansion it is compared to the amount of creep that
8 he thought alloy 22 could manage at temperature.

9 I think we've got enough to piece together
10 an answer for you that would suggest alloy 22 can
11 manage that amount of creep very easily. But, we're
12 going to have to go back to look to get you a more
13 specific answer.

14 MR. HINZE: Let me ask another similar
15 question. In the initial contact of the magma with
16 the canisters, do I understand that you did not
17 consider a collision effect whereby you might have
18 impacts on succeeding canisters, essentially --

19 MR. KOZAK: The train wreck kind of thing.

20 MR. HINZE: Where you might cause the
21 integrity of the canisters to be -- the acceleration
22 of the deterioration of the canister as a result of
23 the impact with effect. Is that taken into account?

24 MR. KOZAK: Acceleration of the
25 deterioration, if that becomes -- one package to the

1 next, no we didn't, because this initial impact one
2 was so much more severe than the secondary impacts.

3 In fact, we did a secondary impact on the
4 roof of the drift for that one waste package. But,
5 that secondary -- that secondary collision was not in
6 that analysis, the analysis of the finite element
7 analysis.

8 MR. HINZE: Right.

9 MR. KOZAK: No, we didn't because,
10 essentially, we're lifting it up and hitting it up
11 against the roof.

12 MR. HINZE: Yes.

13 MR. KOZAK: But, the increase in the
14 degradation rate may play a role in the longer term
15 release, but not during the eruption.

16 MR. HINZE: Another quick question, did you
17 take into account the geochemistry of volatiles in
18 terms of their impact upon the canisters?

19 MR. KOZAK: For corrosion?

20 MR. HINZE: Corrosive types.

21 MR. KOZAK: Yes. And, in fact, that was one
22 of the things we wanted to test by doing the
23 experiments, by looking at putting the C-22 in magma.

24 We weren't sure what some of the other
25 constituents of the magma were going to be.

1 MR. KOZAK: But, the basalt itself that you
2 melt might be quit different than the magma that
3 you're developing.

4 PARTICIPANT: Two comments and two
5 questions. Remember, this is a small diameter sheet.
6 I mean, this isn't a large -- your train wreck lava
7 coming down. We're looking at the initial micro
8 second, two tenths of a second of impact of a
9 relatively narrow sheet coming through.

10 So, we're only looking at one package. And,
11 you know, gravitationally rising, there's not a
12 possibility for a train wreck during this sort of
13 initial impact analysis being done.

14 On your chemical question, one of the
15 reasons Matt put up sort of the old basalt, we
16 purposefully looked where basalt was loaded with a
17 number of these sort of chemical factors, phosphorous
18 and so on, in which we sort of tested at very adverse
19 types of basalts.

20 CHAIRMAN RYAN: Just for the reporter.

21 MR. MARSH: I didn't hear anything about
22 drip -- where's the drip --

23 MR. KOZAK: We didn't consider it.

24 MR. REITER: Leon Reiter, NWTRB staff. A
25 couple questions, in your dip and dunk experiments,

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1 you didn't have any volatiles in the gas escape.
2 Wouldn't that -- symptoms of corrosion rate?

3 I assume the patches you get by magma
4 volatiles -- there could be gasses that could affect
5 the corrosion rate.

6 MR. KOZAK: It could. But, again, corrosion
7 over these time scales was insignificant. If you look
8 at --

9 MR. REITER: Well, but that's based on what
10 you assumed.

11 MR. KOZAK: But it was also based on
12 literature, data, and a variety of other evaluations,
13 which the experiments were consistent with.

14 There are multiple threads of evidence
15 there.

16 MR. REITER: And they also assume the
17 presence of volatiles?

18 MR. KOZAK: I would have to go back to the
19 report, I'm not sure.

20 MR. REITER: Another question. You have
21 five different mechanisms of waste package failure.
22 And you said that when you release those you increase
23 it by five orders of magnitude.

24 MR. KOZAK: Yes.

25 MR. REITER: Which of those are the most

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1 important? Which of those could cause the most
2 damage?

3 MR. KOZAK: Do you mean which --

4 MR. REITER: There was a creep failure,
5 there was magma contact with erosion by the magma,
6 there was impact -- do you have a ranking of those as
7 to which --

8 MR. KOZAK: In terms of the amount of
9 release that they gave, they're all the same because
10 we couldn't get it out of any, no matter what we did.

11 But, in terms of -- I mean, we did the
12 impact analysis by such an extreme calculation, we
13 feel that that's -- any credible behavior in the
14 mountain -- you're going to have a hard time just
15 breaking it open and exposing the waste.

16 That seems to be completely out of bounds.
17 Of the ones that are left, I don't know. It's hard to
18 say. What we ended up doing the calculation on, we
19 based it on overpressure because the corrosion rates
20 were so low and the erosion rates were so low and
21 everything.

22 So we said, of the ones that are left, we'll
23 do it on overpressure. But it still wasn't -- we
24 couldn't credibly get releases by that mechanism.
25 John, did you want to --

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1 MR. KESSLER: Leon, we didn't do an analysis
2 that way. And I would say though, the one I would
3 probably care about the most is the general corrosion
4 failure because, if we have creep failure or something
5 else, we have the small breach in part of the
6 container, we still have a container there that will
7 tend to do a lot of blocking.

8 That's what Matt's next set of analyses
9 were, that that slow, diffusive pathway, or whatever
10 pathway you've got, really reduced the amount of
11 release that can occur.

12 Of course, if the container completely
13 disappears because it is completely corroded, then you
14 would expect that you would get more release. That
15 would make me think that maybe that's the one I care
16 about the most.

17 But, in terms of ranking them, in terms of
18 which one you think would go first, we didn't look at
19 that.

20 MR. KOZAK: Yes, and I wouldn't expect that
21 to happen during the eruption. If there was something
22 that was accelerated by the constituents in the magma
23 or whatever, it would be part of an intrusive analysis
24 of the impacts on something later on.

25 None of these things -- the duration of the

1 eruption is just too short like that. General
2 corrosion is just -- certainly that would be of
3 concern if it could happen.

4 But, over a couple of weeks, you can't get
5 it.

6 MR. REITER: One last question. You
7 mentioned a lot a Nicholis and Rutherford article. It
8 seems to me that you've put a lot of importance in
9 what that's telling you about the background.

10 Is there any way that you can give us a
11 quantitative idea of what kind of impact, assuming
12 those kinds of temperatures, would have?

13 MR. KOZAK: Do you want to call it a
14 quantitative?

15 MR. REITER: You know, what kind of order of
16 magnitude, you know, if assume releases, let's say
17 with one temperature and with the -- of the
18 temperature, say with the Nicholis and Rutherford
19 temperature, or the velocities of them.

20 MR. KOZAK: Quantitatively, I don't think I
21 can do that without going through some more
22 evaluation. Qualitatively, I can certainly say the
23 alloy 22 is much stronger at lower temperatures.

24 The magma is at much higher viscosities, so
25 it's moving much slower. So, the dynamic effect are

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1 a lot less. But, there's a lot of things. It's very
2 important.

3 There's a lot of implications, particularly
4 at the lower temperatures and the higher viscosities,
5 I think, are absolutely crucial. I can't give you a
6 quantitative response to that.

7 MR. APTED: Leon, I think you have to look
8 at what we did in the report prior to that type -- so
9 we looked at tremendously adverse higher temperatures
10 and higher impact and so on.

11 Everything about that event would make our
12 calculations, which we find robustly defensive, in
13 terms of protecting the package during this impact,
14 weeks to months, to maybe a year or two years type of
15 event, absolutely even more conservative.

16 So, I mean, I think -- is more. As we're
17 doing our evaluation in an intrusive case, in terms of
18 looking at the sub-variance of how much basalt goes
19 down these drifts, the release of gas, and so on.

20 MR. REITER: Right.

21 MR. APTED: I'd like to sort of pose a
22 question maybe, that it strikes me always that sort of
23 like you juggle three balls and then somebody says,
24 can you juggle four balls?

25 And then you juggle four balls. And then

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1 you juggle five balls. So, I think EPRI is juggling
2 about five balls here. NRC, DOE is maybe juggling one.

3 (Laughter.)

4 And so, I'd like to ask maybe somebody like
5 Dave Johnson what -- you know, we talked earlier about
6 sort of the risk side adversity probability.

7 And we find it there and in consequence.
8 It's sort of after hearing these folks.

9 MR. JOHNSON: Well of course, the two talks
10 -- thank you, but the two talks were quite different.
11 I had reaction to the NRC one kind of wondering how
12 they were using risk information to direct their
13 ongoing research and kind of concluded maybe there's
14 not enough information there to really have a
15 direction.

16 So they're chasing the science there on its
17 own. It may be unfair. I'm certainly surprised and
18 a little bit overwhelmed by the EPRI presentation.

19 It does go from the initial conditions, if
20 you will, to the final end stage, which is what I and
21 others have been looking for. I guess my initial
22 reaction is, you know, how robust are the models?

23 Can we visualize reasonable scenarios not
24 included? And, again, I'll fallback on not being a
25 geologist. But, a scenario that the initial dike

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1 doesn't intersect the drift, but only does so after it
2 has grown wide enough -- can I transport a package up
3 to the surface and not have this counteractive
4 pressure?

5 It may be a silly question, but, I think
6 maybe at you're at the time where a detailed peer
7 review would really benefit that. And maybe that, you
8 know, could involve NRC and come to some sort of
9 consensus here.

10 MR. KOZAK: We actually did think about
11 whether or not the waste package could be entrained.
12 First off, the buoyancy courses aren't high enough to
13 actually bring them up.

14 But, if they do, if the waste package were
15 to get all the way out, it's still zero dose. It's
16 got to get 18 kilometers further down before you get
17 a dose.

18 MR. JOHNSON: But it changes the scenario.

19 MR. KOZAK: The whole thing comes out, it is
20 deposited on the surface, which isn't going to happen.
21 But, if it did, it's still no-dose. There is a
22 multiple barrier existing.

23 So, some of those things we have tried.
24 You're right, it is conceivable, I suppose, that
25 someone could. That's what we all do in this

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1 business, we look for the more conservative case and
2 say, well, have you thought about this and, you know,
3 try to push and prod at it. And that's got to happen.

4 MR. JOHNSON: I'm not necessarily looking
5 for the more conservative case. I'm looking for, are
6 there reasonable, at least as credible scenarios that
7 aren't in your model now?

8 We certainly can't answer that today. But,
9 it's an impressive amount of work. I can't speak to
10 the science of it, though, I'll say that also.

11 CHAIRMAN RYAN: To me there's a science in
12 the probabilistic analysis part of it. I think that,
13 at the end of the day, is what substantiates and
14 brings together all the pieces of it.

15 And, to be fair, I must say I think the NRC
16 is juggling at least more than one, if not as many as
17 everybody else. They're just presenting parts and
18 pieces of it today.

19 And I think it would be very fortuitous to
20 think about the constraints in which they gave the
21 presentations today. And, to be fair to them, I think
22 there is a broader spectrum.

23 They are addressing across the same
24 spectrum. It's just that we heard the whole piece
25 today. So, I offer you that to think about. As we

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1 conclude, I think, with any last couple of questions
2 that we might have. Yes, I'm sorry.

3 MR. MELSON: I just have a quick one about
4 your immersion experiment where you put the wire in
5 the magma. You said you put that in a -- gas
6 atmosphere. Is that correct?

7 MR. KOZAK: Yes.

8 MR. MELSON: Well, you know that's reducing?

9 MR. KOZAK: Yes.

10 MR. MELSON: And so, that would give you
11 more stability, unless corrosion and -- where you have
12 a oxide buffer.

13 MR. KOZAK: Yes. But, the conditions in the
14 mountain are expected to be reducing, which is why we
15 did that. The conditions of -- I'll let the chemist
16 answer.

17 MR. APTED: I think that the conditions --
18 the conditions in the basalt rising. The basalts are,
19 you know, nickel oxide, something like that. That is
20 very reducing.

21 The activity margin is extremely low,
22 probably -- than anything, doesn't achieve in reducing
23 the conditions enough compared to those that would
24 occur.

25 CHAIRMAN RYAN: Last question, please.

1 MR. MARSH: There is the possibility that
2 you could actually do real experiment. I mean, we
3 have canisters. And you do it just like this. And,
4 the world is producing slag everyday -- lots of it.

5 And, you can dig a pit, you can pour it on
6 top of canister, and you can see what happens. It
7 might be very interesting. And I've actually been
8 involved in trying to actually make a small vat of
9 magma.

10 So I've actually been into this in some
11 detail in talking to companies. Many of them think
12 I'm crazy. Other ones are more open to this. So, if
13 you want to, I have a foothold and can help you get
14 into the system.

15 MR. KOZAK: Well, I wish you hadn't
16 mentioned the slag, because we were hoping that it
17 would get transferred to Hilo.

18 CHAIRMAN RYAN: Well, that's a great lead-
19 in, Bruce, actually, to our perspective tomorrow on
20 aerosol modeling issues. You know, I have a slag
21 experiment.

22 But it's a very small version of it. It's
23 things like level gauges in steel mills that happen to
24 get -- melt. It would be interesting to think about
25 how much of that radioactive material stays in the

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1 ball of steel that hits the floor and how much ends up
2 in the bag house.

3 That might be an interesting series of cases
4 to take a look at. So, we'll hear a little bit about
5 those kinds of issues from Dr. Fred Harper, actual
6 experiments in the aerosol generation area.

7 Resuspension modeling issues from Dr. Lynn
8 Anspaugh, a person who has written on this over the
9 years. And, of course, Dr. Keith Eckerman is going to
10 talk about dose modeling and perhaps give us some
11 insights on conservatisms and none conservatisms,
12 particularly for our radionuclides of interest --
13 plutonium and americium.

14 So, we're kind of getting to the third leg
15 of this school on the igneous activity discussion,
16 which will be one of the aerosol generation
17 characteristics, resuspension and dose modeling
18 characteristics.

19 And we'll close out our approach to re-
20 examining this question. Let's see, our schedule is
21 that we'll convene at eight o'clock. We'll be in
22 promptly at eight o'clock and press on according to
23 the schedule.

24 (Whereupon, at 5: 44 p.m. the above-entitled
25 conference was concluded.)

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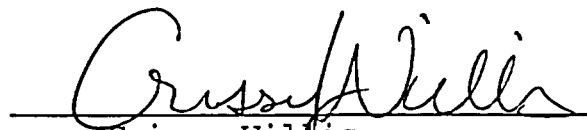
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